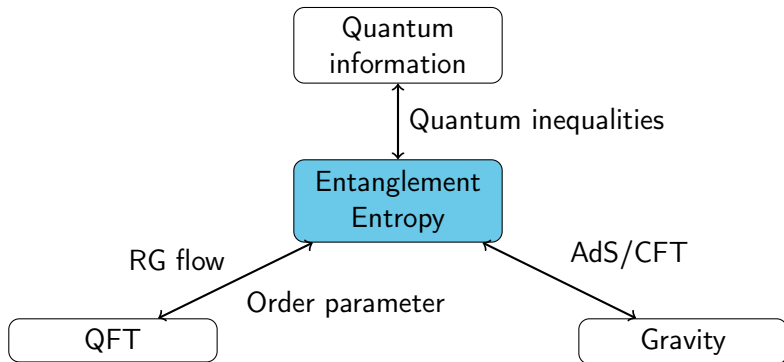


The uses of entanglement entropy in QFT and holography

Tatsuma Nishioka (University of Tokyo)

The 9th Asian Winter School, Busan, Korea

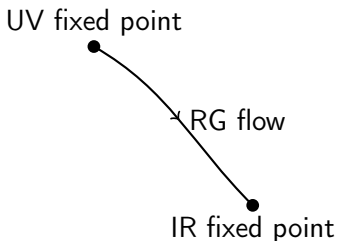
Entanglement Entropy



- Entanglement entropy as a measure of degrees of freedom
- Construct a monotonic function $c(\text{Energy})$ of the energy scale
 - Entropic c -theorem in two dimensions
 - F -theorem in three dimensions

What's the role of entanglement entropy in QFT?

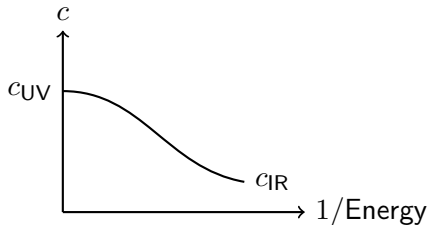
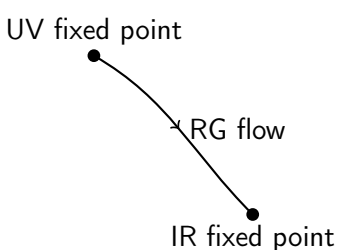
- Entanglement entropy as a **measure of degrees of freedom**



- Construct a **monotonic function $c(\text{Energy})$** of the energy scale
 - Entropic c -theorem in two dimensions
 - F -theorem in three dimensions

What's the role of entanglement entropy in QFT?

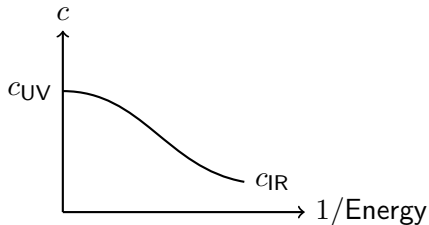
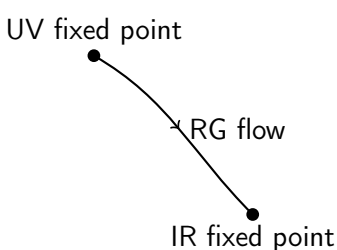
- Entanglement entropy as a **measure of degrees of freedom**



- **Construct a monotonic function $c(\text{Energy})$ of the energy scale**
 - Entropic c -theorem in two dimensions
 - F -theorem in three dimensions

What's the role of entanglement entropy in QFT?

- Entanglement entropy as a **measure of degrees of freedom**



- **Construct a monotonic function $c(\text{Energy})$ of the energy scale**
 - Entropic c -theorem in two dimensions
 - F -theorem in three dimensions

- An **order parameter** for various phase transitions
 - Confinement/deconfinement (like Polyakov loop)
 - Quantum phase transition (no symmetry breaking, no classical order parameter)
- Reconstruction of bulk geometry from entanglement
 - Similarity between MERA and AdS space
 - 1st law of entanglement and linearized Einstein equation of GR

What's the role of entanglement entropy in QFT?

- An **order parameter** for various phase transitions
 - Confinement/deconfinement (like Polyakov loop)
 - Quantum phase transition (no symmetry breaking, no classical order parameter)

- Reconstruction of bulk geometry from entanglement
 - Similarity between MERA and AdS space
 - 1st law of entanglement and linearized Einstein equation of GR

What's the role of entanglement entropy in QFT?

- An **order parameter** for various phase transitions
 - Confinement/deconfinement (like Polyakov loop)
 - Quantum phase transition (no symmetry breaking, no classical order parameter)
- **Reconstruction of bulk geometry** from entanglement
 - Similarity between MERA and AdS space
 - 1st law of entanglement and linearized Einstein equation of GR

What's the role of entanglement entropy in QFT?

- An **order parameter** for various phase transitions
 - Confinement/deconfinement (like Polyakov loop)
 - Quantum phase transition (no symmetry breaking, no classical order parameter)
- **Reconstruction of bulk geometry** from entanglement
 - Similarity between MERA and AdS space
 - 1st law of entanglement and linearized Einstein equation of GR

What's the role of entanglement entropy in QFT?

- An **order parameter** for various phase transitions
 - Confinement/deconfinement (like Polyakov loop)
 - Quantum phase transition (no symmetry breaking, no classical order parameter)
- **Reconstruction of bulk geometry** from entanglement
 - Similarity between MERA and AdS space
 - 1st law of entanglement and linearized Einstein equation of GR

Holography geometrizes the renormalization group (RG) flow

$$[R, G] = RG - GR = 0$$

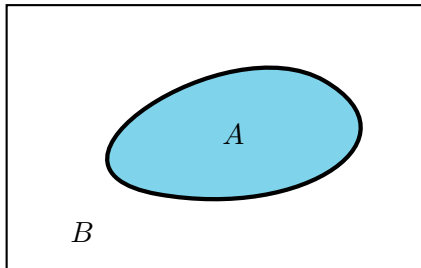
- 1 Basics of entanglement entropy
- 2 Field theoretic methods
- 3 Conformal field theory
- 4 Holographic method
- 5 Renormalization group flow
- 6 Perturbation
- 7 Summary

- Nielsen-Chuang, "Quantum Computation and Quantum Information", Cambridge university press, 2010
- Calabrese-Cardy, arXiv:0905.4013
- Casini-Huerta, arXiv:0903.5284
- Solodukhin, arXiv:1104.3712
- TN-Ryu-Takayanagi, arXiv:0905.0932
- Takayanagi, arXiv:1204.2450

- 1 Basics of entanglement entropy
- 2 Field theoretic methods
- 3 Conformal field theory
- 4 Holographic method
- 5 Renormalization group flow
- 6 Perturbation
- 7 Summary

Definition of entanglement entropy

- Divide a system to A and $B = \bar{A}$: $\mathcal{H}_{tot} = \mathcal{H}_A \otimes \mathcal{H}_B$

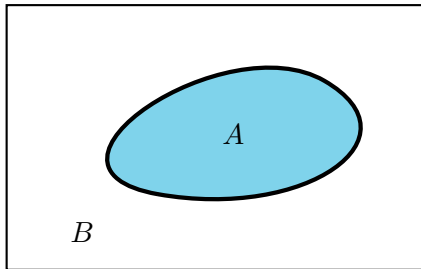


Definition

$$S_A = -\text{tr}_A \rho_A \log \rho_A$$

Definition of entanglement entropy

- Divide a system to A and $B = \bar{A}$: $\mathcal{H}_{tot} = \mathcal{H}_A \otimes \mathcal{H}_B$



Definition

$$S_A = -\text{tr}_A \rho_A \log \rho_A$$

Definition

$$S_A = -\text{tr}_A \rho_A \log \rho_A$$

- $|\Psi\rangle$: wave function of a ground state

$$\rho_{tot} = \frac{1}{\langle \Psi | \Psi \rangle} |\Psi\rangle \langle \Psi|$$

Definition

$$S_A = -\text{tr}_A \rho_A \log \rho_A$$

- $|\Psi\rangle$: wave function of a ground state

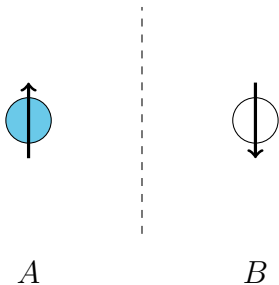
$$\rho_{tot} = \frac{1}{\langle \Psi | \Psi \rangle} |\Psi\rangle \langle \Psi|$$

- Reduced density matrix:

$$\rho_A = \text{tr}_B \rho_{tot} = \sum_i \langle \psi_B^i | \rho_{tot} | \psi_B^i \rangle$$

$\mathcal{H}_B = \{|\psi_B^1\rangle, |\psi_B^2\rangle, \dots\}$ orthonormal basis

Example: two spin system



- Hilbert spaces: $\mathcal{H}_A = \{|\uparrow\rangle_A, |\downarrow\rangle_A\}$, $\mathcal{H}_B = \{|\uparrow\rangle_B, |\downarrow\rangle_B\}$

Example: two spin system

- Given a ground state ($\langle \Psi | \Psi \rangle = 1$):

$$|\Psi\rangle = \cos \theta |\uparrow\rangle_A |\downarrow\rangle_B + \sin \theta |\downarrow\rangle_A |\uparrow\rangle_B$$

Example: two spin system

- Given a ground state ($\langle \Psi | \Psi \rangle = 1$):

$$|\Psi\rangle = \cos \theta |\uparrow\rangle_A |\downarrow\rangle_B + \sin \theta |\downarrow\rangle_A |\uparrow\rangle_B$$

- Reduce density matrix:

$$\begin{aligned} \rho_A &= {}_B \langle \downarrow | \Psi \rangle \langle \Psi | \downarrow \rangle_B + {}_B \langle \uparrow | \Psi \rangle \langle \Psi | \uparrow \rangle_B \\ &= \cos^2 \theta |\uparrow\rangle_A \langle \uparrow| + \sin^2 \theta |\downarrow\rangle_A \langle \downarrow| \end{aligned}$$

Example: two spin system

- Reduce density matrix:

$$\begin{aligned}\rho_A &= {}_B\langle\downarrow|\Psi\rangle\langle\Psi|\downarrow\rangle_B + {}_B\langle\uparrow|\Psi\rangle\langle\Psi|\uparrow\rangle_B \\ &= \cos^2\theta|\uparrow\rangle_A\langle\uparrow| + \sin^2\theta|\downarrow\rangle_A\langle\downarrow|\end{aligned}$$

- Matrix notation:

$$\rho_A = \begin{pmatrix} \cos^2\theta & 0 \\ 0 & \sin^2\theta \end{pmatrix}$$

Example: two spin system

- Matrix notation:

$$\rho_A = \begin{pmatrix} \cos^2 \theta & 0 \\ 0 & \sin^2 \theta \end{pmatrix}$$

- EE as a function of θ : $|\Psi\rangle = \cos \theta |\uparrow\rangle_A |\downarrow\rangle_B + \sin \theta |\downarrow\rangle_A |\uparrow\rangle_B$

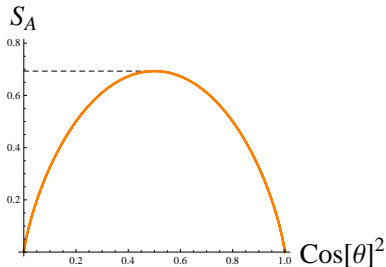
$$\begin{aligned} S_A &= -\text{tr}_A \rho_A \log \rho_A \\ &= -\cos^2 \theta \log(\cos^2 \theta) - \sin^2 \theta \log(\sin^2 \theta) \end{aligned}$$

Example: two spin system

- EE as a function of θ : $|\Psi\rangle = \cos\theta |\uparrow\rangle_A |\downarrow\rangle_B + \sin\theta |\downarrow\rangle_A |\uparrow\rangle_B$

$$\begin{aligned} S_A &= -\text{tr}_A \rho_A \log \rho_A \\ &= -\cos^2\theta \log(\cos^2\theta) - \sin^2\theta \log(\sin^2\theta) \end{aligned}$$

- $\cos^2\theta = \frac{1}{2}$: Maximally entangled, $S_A = \log 2$
- $\cos^2\theta = 0, 1$: No entanglement, $S_A = 0$



Suppose $|\Psi\rangle = \sum_{i=1}^{d_A} \sum_{j=1}^{d_B} c_{ij} |\psi_A^i\rangle |\psi_B^j\rangle$, $d_{A,B} \equiv \dim \mathcal{H}_{A,B}$

Suppose $|\Psi\rangle = \sum_{i=1}^{d_A} \sum_{j=1}^{d_B} c_{ij} |\psi_A^i\rangle |\psi_B^j\rangle$, $d_{A,B} \equiv \dim \mathcal{H}_{A,B}$

$c_{ij} = c_i^A c_j^B$: pure product state

$$|\Psi\rangle = |\Psi_A\rangle |\Psi_B\rangle, \quad |\Psi_{A,B}\rangle \equiv \sum_i c_i^{A,B} |\psi_{A,B}^i\rangle,$$

$$\rho_A = |\Psi_A\rangle \langle \Psi_A| \Rightarrow S_A = 0$$

Suppose $|\Psi\rangle = \sum_{i=1}^{d_A} \sum_{j=1}^{d_B} c_{ij} |\psi_A^i\rangle |\psi_B^j\rangle$, $d_{A,B} \equiv \dim \mathcal{H}_{A,B}$

$c_{ij} \neq c_i^A c_j^B$: entangled state

$$c_{ij} = U_{ik} \lambda_k V_{kj}, \quad U, V : \text{unitary},$$

$$|\Psi\rangle = \sum_{k=1}^{\min(d_A, d_B)} \lambda_k |\tilde{\psi}_A^k\rangle |\tilde{\psi}_B^k\rangle, \quad \lambda_k \geq 0, \quad \sum_k \lambda_k^2 = 1,$$

$$\Rightarrow S_A = - \sum_k \lambda_k^2 \log \lambda_k^2$$

Suppose $|\Psi\rangle = \sum_{i=1}^{d_A} \sum_{j=1}^{d_B} c_{ij} |\psi_A^i\rangle |\psi_B^j\rangle$, $d_{A,B} \equiv \dim \mathcal{H}_{A,B}$

$c_{ij} \neq c_i^A c_j^B$: entangled state

$$c_{ij} = U_{ik} \lambda_k V_{kj}, \quad U, V : \text{unitary},$$

$$|\Psi\rangle = \sum_{k=1}^{\min(d_A, d_B)} \lambda_k |\tilde{\psi}_A^k\rangle |\tilde{\psi}_B^k\rangle, \quad \lambda_k \geq 0, \quad \sum_k \lambda_k^2 = 1,$$

$$\Rightarrow S_A = - \sum_k \lambda_k^2 \log \lambda_k^2 = S_B$$

Quantum mechanical system

Suppose $|\Psi\rangle = \sum_{i=1}^{d_A} \sum_{j=1}^{d_B} c_{ij} |\psi_A^i\rangle |\psi_B^j\rangle$, $d_{A,B} \equiv \dim \mathcal{H}_{A,B}$

$c_{ij} \neq c_i^A c_j^B$: entangled state

$$c_{ij} = U_{ik} \lambda_k V_{kj}, \quad U, V : \text{unitary},$$

$$|\Psi\rangle = \sum_{k=1}^{\min(d_A, d_B)} \lambda_k |\tilde{\psi}_A^k\rangle |\tilde{\psi}_B^k\rangle, \quad \lambda_k \geq 0, \quad \sum_k \lambda_k^2 = 1,$$

$$\Rightarrow S_A = - \sum_k \lambda_k^2 \log \lambda_k^2 = S_B$$

Maximally entangled state

$$\text{For } \lambda_1 = \lambda_2 = \dots = 1/\sqrt{\min(d_A, d_B)},$$

$$S_A = \log \min(d_A, d_B)$$

For a pure ground state

$$S_A = S_{\bar{A}}$$

For a pure ground state

$$S_A = S_{\bar{A}}$$

Strong subadditivity

$$S_{A \cup B \cup C} + S_B \leq S_{A \cup B} + S_{B \cup C}$$

$$S_A + S_C \leq S_{A \cup B} + S_{B \cup C}$$

for any three disjoint regions A , B and C

For a pure ground state

$$S_A = S_{\bar{A}}$$

Strong subadditivity

$$\begin{aligned} S_{AUBUC} + S_B &\leq S_{AUB} + S_{BUC} \\ S_A + S_C &\leq S_{AUB} + S_{BUC} \end{aligned}$$

for any three disjoint regions A , B and C

Mutual information

$$I(A, B) \equiv S_A + S_B - S_{AUB} \geq 0$$

for any disjoint two regions A and B

n -th Rényi entropy

$$S_n(A) = \frac{1}{1-n} \log \operatorname{tr}_A \rho_A^n$$

n -th Rényi entropy

$$S_n(A) = \frac{1}{1-n} \log \operatorname{tr}_A \rho_A^n$$

It reduces to the entanglement entropy in $n \rightarrow 1$ limit

$$S_A = \lim_{n \rightarrow 1} S_n(A)$$

n -th Rényi entropy

$$S_n(A) = \frac{1}{1-n} \log \operatorname{tr}_A \rho_A^n$$

It reduces to the entanglement entropy in $n \rightarrow 1$ limit

$$S_A = \lim_{n \rightarrow 1} S_n(A)$$

Inequalities

$$\partial_n S_n \leq 0$$

$$\partial_n \left(\frac{n-1}{n} S_n \right) \geq 0$$

$$\partial_n ((n-1)S_n) \geq 0$$

$$\partial_n^2 ((n-1)S_n) \leq 0$$

For two states ρ and σ

$$S(\rho||\sigma) = \text{tr} [\rho(\log \rho - \log \sigma)]$$

It measures the **distance** between the two states

For two states ρ and σ

$$S(\rho||\sigma) = \text{tr} [\rho(\log \rho - \log \sigma)]$$

It measures the **distance** between the two states

Properties

$$S(\rho||\rho) = 0$$

$$S(\rho||\sigma) \geq \frac{1}{2} \|\rho - \sigma\|^2 \quad \text{Positivity}$$

$$S(\rho||\sigma) \geq S(\text{tr}_p \rho || \text{tr}_p \sigma) \quad \text{Monotonicity}$$

For two states ρ and σ

$$S(\rho||\sigma) = \text{tr} [\rho(\log \rho - \log \sigma)]$$

It measures the **distance** between the two states

Properties

$$S(\rho||\rho) = 0$$

$$S(\rho||\sigma) \geq \frac{1}{2} \|\rho - \sigma\|^2 \quad \text{Positivity}$$

$$S(\rho||\sigma) \geq S(\text{tr}_p \rho || \text{tr}_p \sigma) \quad \text{Monotonicity}$$

The strong subadditivity follows from the last inequality

- 1 Basics of entanglement entropy
- 2 Field theoretic methods
- 3 Conformal field theory
- 4 Holographic method
- 5 Renormalization group flow
- 6 Perturbation
- 7 Summary

- $\dim\mathcal{H} = \infty$ in QFT
- Useful trick:

$$S_A = -\partial_n \log \text{tr}_A \rho_A^n \Big|_{n=1} \quad (\text{tr}_A \rho_A = 1)$$

- Z_n : partition function on n -covering space

$$\text{tr}_A \rho_A^n = \frac{Z_n}{(Z_1)^n}$$

- $\dim \mathcal{H} = \infty$ in QFT
- Useful trick:

$$S_A = -\partial_n \log \text{tr}_A \rho_A^n \Big|_{n=1} \quad (\text{tr}_A \rho_A = 1)$$

- Z_n : partition function on n -covering space

$$\text{tr}_A \rho_A^n = \frac{Z_n}{(Z_1)^n}$$

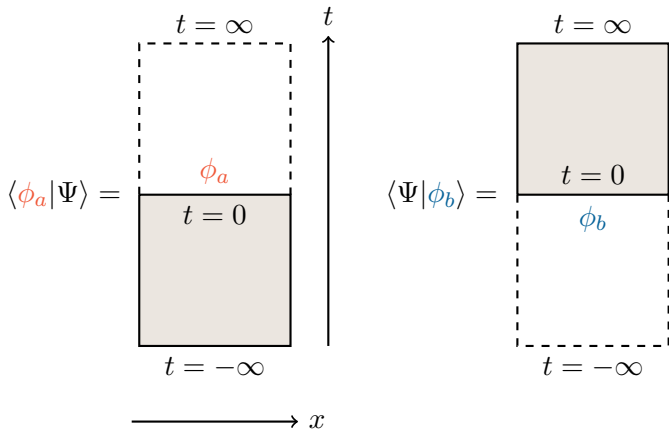
- $\dim \mathcal{H} = \infty$ in QFT
- Useful trick:

$$S_A = -\partial_n \log \text{tr}_A \rho_A^n \Big|_{n=1} \quad (\text{tr}_A \rho_A = 1)$$

- Z_n : partition function on n -covering space

$$\text{tr}_A \rho_A^n = \frac{Z_n}{(Z_1)^n}$$

Path integral representation of the wave function



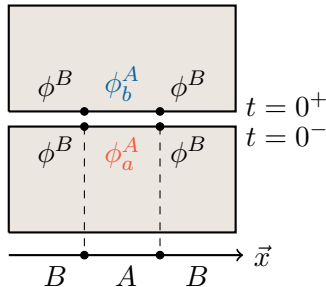
States $|\phi_{a,b}\rangle$ are the boundary conditions at $t = 0$

$$[\rho_A]_{ab} = \frac{1}{Z_1} \int [\mathcal{D}\phi^B(t=0, \vec{x} \in B)] (\langle \phi_a^A | \langle \phi^B |) |\Psi\rangle \langle \Psi| (| \phi_b^A \rangle | \phi^B \rangle) ,$$

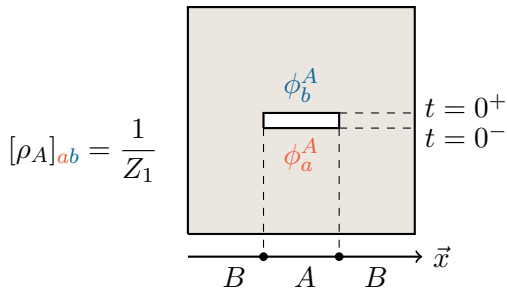
Replica trick and covering space

$$[\rho_A]_{ab} = \frac{1}{Z_1} \int [\mathcal{D}\phi^B(t=0, \vec{x} \in B)] (\langle \phi_a^A | \langle \phi^B |) |\Psi\rangle \langle \Psi| (| \phi_b^A \rangle | \phi^B \rangle) ,$$

$$= \frac{1}{Z_1} \int [\mathcal{D}\phi^B(t=0, \vec{x} \in B)]$$

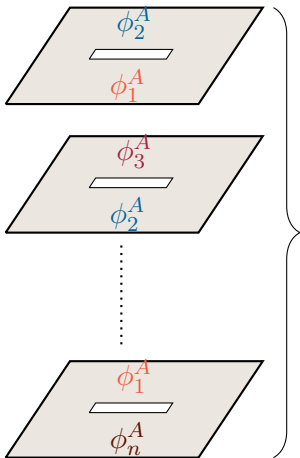


Replica trick and covering space



Replica trick and covering space

$$\mathrm{tr}_A \rho_A^n = \frac{1}{(Z_1)^n}$$



n copies $\equiv Z_n$

$$= \frac{Z_n}{(Z_1)^n}$$

Entanglement entropy

$$S_A = -(\partial_n - 1) \log Z_n \Big|_{n=1}$$

All we need to know is the partition function Z_n on the n -fold cover \mathcal{M}_n !

Entanglement entropy

$$S_A = -(\partial_n - 1) \log Z_n \Big|_{n=1}$$

All we need to know is the partition function Z_n on the n -fold cover \mathcal{M}_n !

Comment

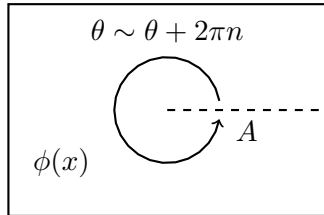
Regarding $\beta = 2\pi n$ as an inverse temperature

$$S_A = (\beta \partial_\beta - 1) (\beta F) \Big|_{\beta=2\pi}$$

where $\beta F(\beta) = -\log Z_n$

Example: A half space

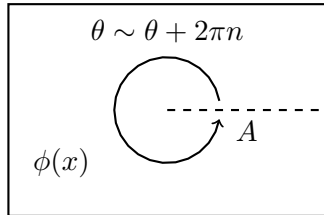
- Suppose $A = \{x > 0, t = 0\}$ on $\mathcal{M} = \mathbb{R}^2$
- $\mathcal{M}_n : ds^2 = dr^2 + r^2 d\theta^2$
with $r \geq 0, \theta \sim \theta + 2\pi n$
- $\log Z_n =$
 $-\frac{1}{2} \log \det(-\nabla^2 + m^2)|_{\mathcal{M}_n}$
- $S_A = -\frac{1}{12} \log(m^2 \epsilon^2)$
 $\epsilon \ll 1$: UV cutoff



$$I = \frac{1}{2} \int d^2x [(\partial_\mu \phi)^2 + m^2 \phi^2]$$

Example: A half space

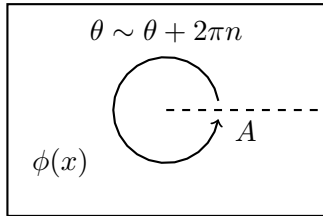
- Suppose $A = \{x > 0, t = 0\}$ on $\mathcal{M} = \mathbb{R}^2$
- $\mathcal{M}_n : ds^2 = dr^2 + r^2 d\theta^2$
with $r \geq 0, \theta \sim \theta + 2\pi n$
- $\log Z_n =$
 $-\frac{1}{2} \log \det(-\nabla^2 + m^2)|_{\mathcal{M}_n}$
- $S_A = -\frac{1}{12} \log(m^2 \epsilon^2)$
 $\epsilon \ll 1$: UV cutoff



$$I = \frac{1}{2} \int d^2x [(\partial_\mu \phi)^2 + m^2 \phi^2]$$

Example: A half space

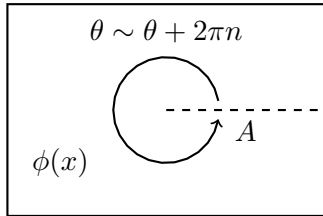
- Suppose $A = \{x > 0, t = 0\}$ on $\mathcal{M} = \mathbb{R}^2$
- $\mathcal{M}_n : ds^2 = dr^2 + r^2 d\theta^2$
with $r \geq 0, \theta \sim \theta + 2\pi n$
- $\log Z_n =$
 $-\frac{1}{2} \log \det(-\nabla^2 + m^2)|_{\mathcal{M}_n}$
- $S_A = -\frac{1}{12} \log(m^2 \epsilon^2)$
 $\epsilon \ll 1$: UV cutoff



$$I = \frac{1}{2} \int d^2x [(\partial_\mu \phi)^2 + m^2 \phi^2]$$

Example: A half space

- Suppose $A = \{x > 0, t = 0\}$ on $\mathcal{M} = \mathbb{R}^2$
- $\mathcal{M}_n : ds^2 = dr^2 + r^2 d\theta^2$
with $r \geq 0, \theta \sim \theta + 2\pi n$
- $\log Z_n =$
 $-\frac{1}{2} \log \det(-\nabla^2 + m^2)|_{\mathcal{M}_n}$
- $S_A = -\frac{1}{12} \log(m^2 \epsilon^2)$
 $\epsilon \ll 1$: UV cutoff



$$I = \frac{1}{2} \int d^2x [(\partial_\mu \phi)^2 + m^2 \phi^2]$$

- The partition function has UV divergences

$$\begin{aligned} \log Z_n[g_{\mu\nu}] = & C_d \int_{\mathcal{M}_n} d^d x \sqrt{g} \Lambda^d + C_{d-2} \int_{\mathcal{M}_n} d^d x \sqrt{g} \Lambda^{d-2} \mathcal{R} \\ & + C_{d-4} \int_{\mathcal{M}_n} d^d x \sqrt{g} \Lambda^{d-4} \mathcal{R}^2 + \dots \end{aligned}$$

where $\Lambda \gg 1$ is a UV cutoff scale, \mathcal{R} is a Ricci scalar

- The n -fold cover \mathcal{M}_n differs from $\mathcal{M} \equiv \mathcal{M}_1$ near the entangling surface $\Sigma \equiv \partial A$

$$\int_{\mathcal{M}_n} \mathcal{R}^i - n \int_{\mathcal{M}} \mathcal{R}^i \sim \int_{\Sigma} \#$$

- The partition function has UV divergences

$$\begin{aligned} \log Z_n[g_{\mu\nu}] = & C_d \int_{\mathcal{M}_n} d^d x \sqrt{g} \Lambda^d + C_{d-2} \int_{\mathcal{M}_n} d^d x \sqrt{g} \Lambda^{d-2} \mathcal{R} \\ & + C_{d-4} \int_{\mathcal{M}_n} d^d x \sqrt{g} \Lambda^{d-4} \mathcal{R}^2 + \dots \end{aligned}$$

where $\Lambda \gg 1$ is a UV cutoff scale, \mathcal{R} is a Ricci scalar

- The n -fold cover \mathcal{M}_n differs from $\mathcal{M} \equiv \mathcal{M}_1$ near the **entangling surface** $\Sigma \equiv \partial A$

$$\int_{\mathcal{M}_n} \mathcal{R}^i - n \int_{\mathcal{M}} \mathcal{R}^i \sim \int_{\Sigma} \#$$

UV structure of entanglement entropy

- The entropy has UV divergences coming from the correlation near Σ

UV structure of entanglement entropy

$$S_A = c_{d-2}\Lambda^{d-2} + c_{d-4}\Lambda^{d-4} + \dots ,$$

with coefficients schematically written as

$$c_{d-2i} = \sum_{l+m=i-1} \int_{\Sigma} \mathcal{R}^l \mathcal{K}^{2m} ,$$

\mathcal{K} : the extrinsic curvature

- It starts from the **area law divergence**

$$c_{d-2} \propto \text{Vol}(\Sigma)$$

- The entropy has UV divergences coming from the correlation near Σ

UV structure of entanglement entropy

$$S_A = c_{d-2}\Lambda^{d-2} + c_{d-4}\Lambda^{d-4} + \dots ,$$

with coefficients schematically written as

$$c_{d-2i} = \sum_{l+m=i-1} \int_{\Sigma} \mathcal{R}^l \mathcal{K}^{2m} ,$$

\mathcal{K} : the extrinsic curvature

- It starts from the **area law divergence**

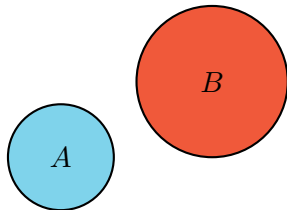
$$c_{d-2} \propto \text{Vol}(\Sigma)$$

- For two disjoint regions A and B the mutual information

$$I(A, B) = S_A + S_B - S_{A \cup B}$$

- The UV divergences cancel out!

$$\int_{\Sigma(A)} + \int_{\Sigma(B)} - \int_{\Sigma(A \cup B)} = 0$$



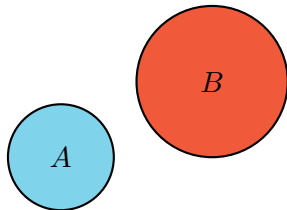
- The mutual information is scheme independent

- For two disjoint regions A and B the mutual information

$$I(A, B) = S_A + S_B - S_{A \cup B}$$

- The UV divergences cancel out!

$$\int_{\Sigma(A)} + \int_{\Sigma(B)} - \int_{\Sigma(A \cup B)} = 0$$



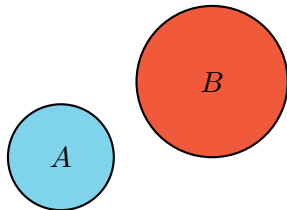
- The mutual information is scheme independent

- For two disjoint regions A and B the mutual information

$$I(A, B) = S_A + S_B - S_{A \cup B}$$

- The UV divergences cancel out!

$$\int_{\Sigma(A)} + \int_{\Sigma(B)} - \int_{\Sigma(A \cup B)} = 0$$



- The mutual information is **scheme independent**

- Consider a free massive scalar whose effective action is

$$\log Z_n = -\frac{1}{2} \log \det(-\nabla^2 + m^2)$$

- The heat kernel coefficients $a_i(\mathcal{M}_n)$ depends on the geometry \mathcal{M}_n and known for a smooth manifold

- Consider a free massive scalar whose effective action is

$$\begin{aligned}\log Z_n &= -\frac{1}{2} \log \det(-\nabla^2 + m^2) \\ &= \frac{1}{2} \int_{\epsilon^2}^{\infty} \frac{ds}{s} \operatorname{tr} K_{\mathcal{M}_n}(s) e^{-m^2 s}\end{aligned}$$

- The heat kernel coefficients $a_i(\mathcal{M}_n)$ depends on the geometry \mathcal{M}_n and known for a smooth manifold

- Consider a free massive scalar whose effective action is

$$\begin{aligned}\log Z_n &= -\frac{1}{2} \log \det(-\nabla^2 + m^2) \\ &= \frac{1}{2} \int_{\epsilon^2}^{\infty} \frac{ds}{s} \operatorname{tr} K_{\mathcal{M}_n}(s) e^{-m^2 s}\end{aligned}$$

The expansion of the heat kernel $K_{\mathcal{M}_n}(s) \equiv e^{s\nabla^2}$

$$\operatorname{tr} K_{\mathcal{M}_n}(s) = \frac{1}{(4\pi s)^{\frac{d}{2}}} \sum_{i=0}^{\infty} a_i(\mathcal{M}_n) s^i$$

- The heat kernel coefficients $a_i(\mathcal{M}_n)$ depends on the geometry \mathcal{M}_n and known for a smooth manifold

- Consider a free massive scalar whose effective action is

$$\begin{aligned}\log Z_n &= -\frac{1}{2} \log \det(-\nabla^2 + m^2) \\ &= \frac{1}{2} \int_{\epsilon^2}^{\infty} \frac{ds}{s} \operatorname{tr} K_{\mathcal{M}_n}(s) e^{-m^2 s}\end{aligned}$$

The expansion of the heat kernel $K_{\mathcal{M}_n}(s) \equiv e^{s\nabla^2}$

$$\operatorname{tr} K_{\mathcal{M}_n}(s) = \frac{1}{(4\pi s)^{\frac{d}{2}}} \sum_{i=0}^{\infty} a_i(\mathcal{M}_n) s^i$$

- The heat kernel coefficients $a_i(\mathcal{M}_n)$ depends on the geometry \mathcal{M}_n and known for a smooth manifold

- a_i decompose to bulk and surface parts in $n \rightarrow 1$ limit

$$a_i = a_i^{\text{bulk}} + (1 - n)a_i^{\Sigma} + O((1 - n)^2)$$

where the bulk part satisfies

$$a_i^{\text{bulk}}(\mathcal{M}_n) = n a_i^{\text{bulk}}(\mathcal{M}_1)$$

- The entropy is determined by only the surface part

$$S_A = \frac{1}{(4\pi)^{\frac{d}{2}}} \left[\frac{a_0^{\Sigma}}{d} \frac{1}{\epsilon^d} + \frac{a_1^{\Sigma} - m^2 a_0^{\Sigma}}{d - 2} \frac{1}{\epsilon^{d-2}} + \dots \right]$$

- a_i decompose to bulk and surface parts in $n \rightarrow 1$ limit

$$a_i = a_i^{\text{bulk}} + (1 - n)a_i^{\Sigma} + O((1 - n)^2)$$

where the bulk part satisfies

$$a_i^{\text{bulk}}(\mathcal{M}_n) = n a_i^{\text{bulk}}(\mathcal{M}_1)$$

- The entropy is determined by only the surface part

$$S_A = \frac{1}{(4\pi)^{\frac{d}{2}}} \left[\frac{a_0^{\Sigma}}{d} \frac{1}{\epsilon^d} + \frac{a_1^{\Sigma} - m^2 a_0^{\Sigma}}{d - 2} \frac{1}{\epsilon^{d-2}} + \dots \right]$$

On a smooth manifold \mathcal{M} [E.g. Vassilevich, hep-th/0306138]

$$a_0^{\text{bulk}} = \int_{\mathcal{M}} 1, \quad a_1^{\text{bulk}} = \frac{1}{6} \int_{\mathcal{M}} \mathcal{R}$$

- Apply it to a regularized geometry $\widetilde{\mathcal{M}}_n$

$$ds_{\widetilde{\mathcal{M}}_n}^2 = f_\epsilon(r) dr^2 + r^2 d\theta^2 + \dots$$

where $f_\epsilon(r)$ is a smooth function that behaves as

$$f_\epsilon(r \rightarrow 0) = n^2, \quad f_\epsilon(r > \epsilon) = 1, \quad \epsilon \ll 1$$

On a smooth manifold \mathcal{M} [E.g. Vassilevich, hep-th/0306138]

$$a_0^{\text{bulk}} = \int_{\mathcal{M}} 1, \quad a_1^{\text{bulk}} = \frac{1}{6} \int_{\mathcal{M}} \mathcal{R}$$

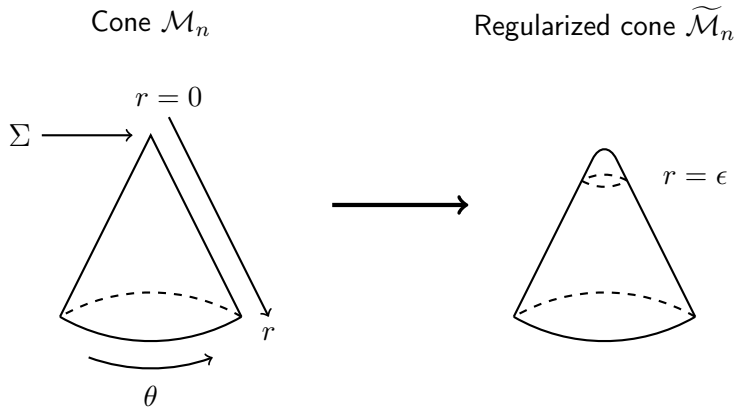
- Apply it to a **regularized geometry** $\widetilde{\mathcal{M}}_n$

$$ds^2_{\widetilde{\mathcal{M}}_n} = f_\epsilon(r) dr^2 + r^2 d\theta^2 + \dots$$

where $f_\epsilon(r)$ is a smooth function that behaves as

$$f_\epsilon(r \rightarrow 0) = n^2, \quad f_\epsilon(r > \epsilon) = 1, \quad \epsilon \ll 1$$

Heat kernel coefficients on \mathcal{M}_n



On the regularized geometry $\widetilde{\mathcal{M}}_n$ [Fursaev-Patrushev-Solodukhin 13]

$$\int_{\widetilde{\mathcal{M}}_n} 1 = n \int_{\widetilde{\mathcal{M}}_1} 1$$

$$\int_{\widetilde{\mathcal{M}}_n} \mathcal{R} = n \int_{\mathcal{M}_1} \mathcal{R} + 4\pi(1-n) \int_{\Sigma} 1 + O((1-n)^2)$$

which yields $a_0^{\Sigma} = 0$ and $a_1^{\Sigma} = 2\pi \text{Vol}(\Sigma)/3$

On the regularized geometry $\widetilde{\mathcal{M}}_n$ [Fursaev-Patrushev-Solodukhin 13]

$$\int_{\widetilde{\mathcal{M}}_n} 1 = n \int_{\widetilde{\mathcal{M}}_1} 1$$

$$\int_{\widetilde{\mathcal{M}}_n} \mathcal{R} = n \int_{\mathcal{M}_1} \mathcal{R} + 4\pi(1-n) \int_{\Sigma} 1 + O((1-n)^2)$$

which yields $a_0^\Sigma = 0$ and $a_1^\Sigma = 2\pi \text{Vol}(\Sigma)/3$

Again we obtain the **area law divergence**

$$S_A = \frac{1}{6(d-2)(4\pi)^{d/2-1}} \frac{\text{Vol}(\Sigma)}{\epsilon^{d-2}} + \dots$$

(the subleading terms are similarly obtained)

Outline

- 1 Basics of entanglement entropy
- 2 Field theoretic methods
- 3 Conformal field theory
- 4 Holographic method
- 5 Renormalization group flow
- 6 Perturbation
- 7 Summary

- Under the conformal transformation

$$\bar{g}_{\mu\nu}(x') = \Omega^2(x)g_{\mu\nu}(x) ,$$

CFT is invariant for some Δ

$$I[\bar{g}_{\mu\nu}, \bar{\phi}] = I[g_{\mu\nu}, \phi] , \quad \bar{\phi}(x) = \Omega^{-\Delta}(x)\phi(x)$$

- Example: A conformally coupled scalar field with $\Delta = d/2 - 1$ on a curved space

$$I[g_{\mu\nu}, \phi] = \frac{1}{2} \int d^d x \sqrt{g} \left[\partial_\mu \phi \partial^\mu \phi + \frac{d-2}{4(d-1)} \mathcal{R} \phi^2 \right]$$

- Under the conformal transformation

$$\bar{g}_{\mu\nu}(x') = \Omega^2(x)g_{\mu\nu}(x) ,$$

CFT is invariant for some Δ

$$I[\bar{g}_{\mu\nu}, \bar{\phi}] = I[g_{\mu\nu}, \phi] , \quad \bar{\phi}(x) = \Omega^{-\Delta}(x)\phi(x)$$

- Example: A conformally coupled scalar field with $\Delta = d/2 - 1$ on a curved space

$$I[g_{\mu\nu}, \phi] = \frac{1}{2} \int d^d x \sqrt{g} \left[\partial_\mu \phi \partial^\mu \phi + \frac{d-2}{4(d-1)} \mathcal{R} \phi^2 \right]$$

- For CFT, the variation of the action is zero for $\delta g_{\mu\nu} = 2\delta\Omega g_{\mu\nu}$

$$0 = \delta I[g_{\mu\nu}] = \int d^d x \delta g_{\mu\nu} \frac{\delta I[g_{\mu\nu}]}{\delta g_{\mu\nu}} = - \int d^d x \sqrt{g} T_{\mu}^{\mu} \delta\Omega(x) ,$$

- The trace of the stress-energy tensor should vanish classically

$$T_{\mu}^{\mu} = g^{\mu\nu} \frac{2}{\sqrt{g}} \frac{\delta I}{\delta g^{\mu\nu}} = 0$$

- Quantum mechanically, however, it does not for **even** d

- For CFT, the variation of the action is zero for $\delta g_{\mu\nu} = 2\delta\Omega g_{\mu\nu}$

$$0 = \delta I[g_{\mu\nu}] = \int d^d x \delta g_{\mu\nu} \frac{\delta I[g_{\mu\nu}]}{\delta g_{\mu\nu}} = - \int d^d x \sqrt{g} T_{\mu}^{\mu} \delta\Omega(x) ,$$

- The trace of the stress-energy tensor should vanish classically

$$T_{\mu}^{\mu} = g^{\mu\nu} \frac{2}{\sqrt{g}} \frac{\delta I}{\delta g^{\mu\nu}} = 0$$

- Quantum mechanically, however, it does not for **even** d

Conformal anomaly

- For CFT, the variation of the action is zero for $\delta g_{\mu\nu} = 2\delta\Omega g_{\mu\nu}$

$$0 = \delta I[g_{\mu\nu}] = \int d^d x \delta g_{\mu\nu} \frac{I[g_{\mu\nu}]}{\delta g_{\mu\nu}} = - \int d^d x \sqrt{g} T_{\mu}^{\mu} \delta\Omega(x) ,$$

- The trace of the stress-energy tensor should vanish classically

$$T_{\mu}^{\mu} = g^{\mu\nu} \frac{2}{\sqrt{g}} \frac{\delta I}{\delta g^{\mu\nu}} = 0$$

- Quantum mechanically, however, it does not for **even** d

Conformal anomaly

$$\langle T_{\mu}^{\mu} \rangle = \frac{(-1)^{\frac{d}{2}+1}}{2} A E_d + \sum_i B_i I_i$$

Conformal anomaly

$$\langle T_{\mu}^{\mu} \rangle = \frac{(-1)^{\frac{d}{2}+1}}{2} A E_d + \sum_i B_i I_i$$

- E_d : the Euler density ($\int_{S^d} E_d = 2$)
 I_i : the independent Weyl invariants in d dimensions
- The coefficients A and B_i are the **central charges**

Conformal anomaly

$$\langle T_{\mu}^{\mu} \rangle = \frac{(-1)^{\frac{d}{2}+1}}{2} A E_d + \sum_i B_i I_i$$

- E_d : the Euler density ($\int_{S^d} E_d = 2$)
 I_i : the independent Weyl invariants in d dimensions
- The coefficients A and B_i are the **central charges**

Conformal anomaly in entanglement entropy

- A scaling of length $l \rightarrow e^\sigma l$ is equivalent to $g_{\mu\nu} \rightarrow e^{2\sigma} g_{\mu\nu}$

$$l \frac{d}{dl} \log Z_n = \int_{\mathcal{M}_n} d^d x \sqrt{g} \langle T_\mu^\mu \rangle$$

- The entanglement entropy satisfies

$$l \frac{d}{dl} S_A = \int_{\mathcal{M}_1} d^d x \sqrt{g} \langle T_\mu^\mu \rangle - \lim_{n \rightarrow 1} \partial_n \int_{\mathcal{M}_n} d^d x \sqrt{g} \langle T_\mu^\mu \rangle \equiv c_0$$

- If the rhs does not vanish (it can happen in even dimensions), EE has a logarithmic divergence

$$S_A \supset c_0 \log(l/\epsilon)$$

- A scaling of length $l \rightarrow e^\sigma l$ is equivalent to $g_{\mu\nu} \rightarrow e^{2\sigma} g_{\mu\nu}$

$$l \frac{d}{dl} \log Z_n = \int_{\mathcal{M}_n} d^d x \sqrt{g} \langle T_\mu^\mu \rangle$$

- The entanglement entropy satisfies

$$l \frac{d}{dl} S_A = \int_{\mathcal{M}_1} d^d x \sqrt{g} \langle T_\mu^\mu \rangle - \lim_{n \rightarrow 1} \partial_n \int_{\mathcal{M}_n} d^d x \sqrt{g} \langle T_\mu^\mu \rangle \equiv c_0$$

- If the rhs does not vanish (it can happen in even dimensions), EE has a logarithmic divergence

$$S_A \supset c_0 \log(l/\epsilon)$$

- A scaling of length $l \rightarrow e^\sigma l$ is equivalent to $g_{\mu\nu} \rightarrow e^{2\sigma} g_{\mu\nu}$

$$l \frac{d}{dl} \log Z_n = \int_{\mathcal{M}_n} d^d x \sqrt{g} \langle T_\mu^\mu \rangle$$

- The entanglement entropy satisfies

$$l \frac{d}{dl} S_A = \int_{\mathcal{M}_1} d^d x \sqrt{g} \langle T_\mu^\mu \rangle - \lim_{n \rightarrow 1} \partial_n \int_{\mathcal{M}_n} d^d x \sqrt{g} \langle T_\mu^\mu \rangle \equiv c_0$$

- If the rhs does not vanish (it can happen in even dimensions),
EE has a logarithmic divergence

$$S_A \supset c_0 \log(l/\epsilon)$$

- In two dimensions, only $E_2 = \mathcal{R}/(4\pi)$ exists and choosing $A = c/3$

$$c_0 = \frac{c}{24\pi} \left[\int_{\mathcal{M}_1} \mathcal{R} - \lim_{n \rightarrow 1} \partial_n \int_{\mathcal{M}_n} \mathcal{R} \right]$$

- Applying the formula

$$\int_{\widetilde{\mathcal{M}}_n} \mathcal{R} = n \int_{\mathcal{M}_1} \mathcal{R} + 4\pi(1-n) \int_{\Sigma} 1 + O((1-n)^2)$$

- In two dimensions, only $E_2 = \mathcal{R}/(4\pi)$ exists and choosing $A = c/3$

$$c_0 = \frac{c}{24\pi} \left[\int_{\mathcal{M}_1} \mathcal{R} - \lim_{n \rightarrow 1} \partial_n \int_{\mathcal{M}_n} \mathcal{R} \right]$$

- Applying the formula

$$\int_{\widetilde{\mathcal{M}}_n} \mathcal{R} = n \int_{\mathcal{M}_1} \mathcal{R} + 4\pi(1-n) \int_{\Sigma} 1 + O((1-n)^2)$$

- In two dimensions, only $E_2 = \mathcal{R}/(4\pi)$ exists and choosing $A = c/3$

$$c_0 = \frac{c}{24\pi} \left[\int_{\mathcal{M}_1} \mathcal{R} - \lim_{n \rightarrow 1} \partial_n \int_{\mathcal{M}_n} \mathcal{R} \right]$$

- Applying the formula

$$\int_{\widetilde{\mathcal{M}}_n} \mathcal{R} = n \int_{\mathcal{M}_1} \mathcal{R} + 4\pi(1-n) \int_{\Sigma} 1 + O((1-n)^2)$$

EE of an interval of width l in CFT_2

$$S_A = \frac{c}{3} \log(l/\epsilon) + (\text{finite})$$

- There are one Euler density and one Weyl invariant

$$E_4 = \frac{1}{32\pi^2} (\mathcal{R}_{\mu\nu\rho\sigma}^2 - 4\mathcal{R}_{\mu\nu}^2 + \mathcal{R}^2)$$

$$I_4 = \frac{1}{16\pi^2} \left(\mathcal{R}_{\mu\nu\rho\sigma}^2 - 2\mathcal{R}_{\mu\nu}^2 + \frac{1}{3}\mathcal{R}^2 \right)$$

- There are general formulae for the Riemann tensors on the regularized manifold $\widetilde{\mathcal{M}}_n$

CFT₄ with central charges $A = a$, $B = c$

$$S_A = \frac{c_2}{\epsilon^2} + c_0 \log \frac{l}{\epsilon} + (\text{finite})$$

$$c_0 = -\frac{a}{2} \int_{\Sigma} E_2 + \frac{c}{6\pi} \int_{\Sigma} \left(\mathcal{R}_{\Sigma} + 2\mathcal{R}_{abab} - \mathcal{R}_{aa} + \frac{1}{2}(\mathcal{K}_{\mu}^a)^2 - 2\mathcal{K}_{\mu\nu}^a \mathcal{K}^{a\mu\nu} \right)$$

- There are one Euler density and one Weyl invariant

$$E_4 = \frac{1}{32\pi^2} (\mathcal{R}_{\mu\nu\rho\sigma}^2 - 4\mathcal{R}_{\mu\nu}^2 + \mathcal{R}^2)$$

$$I_4 = \frac{1}{16\pi^2} \left(\mathcal{R}_{\mu\nu\rho\sigma}^2 - 2\mathcal{R}_{\mu\nu}^2 + \frac{1}{3}\mathcal{R}^2 \right)$$

- There are general formulae for the Riemann tensors on the regularized manifold $\widetilde{\mathcal{M}}_n$

CFT₄ with central charges $A = a$, $B = c$

$$S_A = \frac{c_2}{\epsilon^2} + c_0 \log \frac{l}{\epsilon} + (\text{finite})$$

$$c_0 = -\frac{a}{2} \int_{\Sigma} E_2 + \frac{c}{6\pi} \int_{\Sigma} \left(\mathcal{R}_{\Sigma} + 2\mathcal{R}_{abab} - \mathcal{R}_{aa} + \frac{1}{2}(\mathcal{K}_{\mu}^a)^2 - 2\mathcal{K}_{\mu\nu}^a \mathcal{K}^{a\mu\nu} \right)$$

In general even dimensions

- There are one Euler density and several Weyl invariants
- Using a formula

$$\int_{\tilde{\mathcal{M}}_n} E_d = n \int_{\mathcal{M}_1} E_d + (1 - n) \int_{\Sigma} E_{d-2}$$

CFT_d

$$S_A = \frac{c_{d-2}}{\epsilon^{d-2}} + \frac{c_{d-4}}{\epsilon^{d-4}} + \cdots + \frac{c_2}{\epsilon^2} + c_0 \log \frac{l}{\epsilon} + (\text{finite})$$

$$c_0 = \frac{(-1)^{\frac{d}{2}+1}}{2} A \int_{\Sigma} E_{d-2} + \cdots$$

- There are one Euler density and several Weyl invariants
- Using a formula

$$\int_{\tilde{\mathcal{M}}_n} E_d = n \int_{\mathcal{M}_1} E_d + (1 - n) \int_{\Sigma} E_{d-2}$$

CFT_d

$$S_A = \frac{c_{d-2}}{\epsilon^{d-2}} + \frac{c_{d-4}}{\epsilon^{d-4}} + \cdots + \frac{c_2}{\epsilon^2} + c_0 \log \frac{l}{\epsilon} + (\text{finite})$$

$$c_0 = \frac{(-1)^{\frac{d}{2}+1}}{2} A \int_{\Sigma} E_{d-2} + \cdots$$

In even dimensions

$$S_A = \frac{c_{d-2}}{\epsilon^{d-2}} + \frac{c_{d-4}}{\epsilon^{d-4}} + \cdots + \frac{c_2}{\epsilon^2} + c_0 \log \frac{l}{\epsilon} + \cdots$$

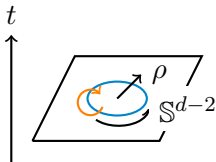
c_0 : depends on the central charges

In odd dimensions

$$S_A = \frac{c_{d-2}}{\epsilon^{d-2}} + \frac{c_{d-4}}{\epsilon^{d-4}} + \cdots + \frac{c_1}{\epsilon} + (-1)^{\frac{d-1}{2}} F$$

F : scheme independent constant

- Let A be a ball $\{\rho \leq R, t = 0\}$ in \mathbb{R}^d



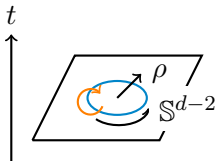
\mathbb{R}^d

$$ds^2 = dt^2 + d\rho^2 + \rho^2 d\Omega_{d-2}^2$$

Conformal map for spherical entangling surface

- Let A be a ball $\{\rho \leq R, t = 0\}$ in \mathbb{R}^d
- The coordinate transformation [Casini-Huerta-Myers 11]

$$t = R \frac{\sin \tau}{\cosh u + \cos \tau}, \quad \rho = R \frac{\sinh u}{\cosh u + \cos \tau},$$



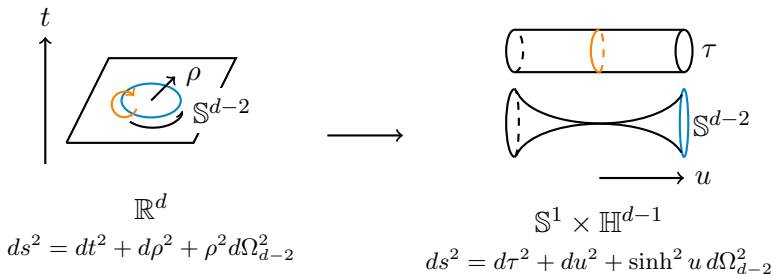
\mathbb{R}^d

$$ds^2 = dt^2 + d\rho^2 + \rho^2 d\Omega_{d-2}^2$$

Conformal map for spherical entangling surface

- Let A be a ball $\{\rho \leq R, t = 0\}$ in \mathbb{R}^d
- The coordinate transformation [Casini-Huerta-Myers 11]

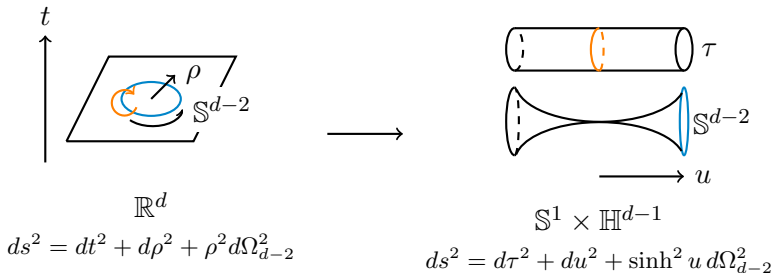
$$t = R \frac{\sin \tau}{\cosh u + \cos \tau}, \quad \rho = R \frac{\sinh u}{\cosh u + \cos \tau},$$



Conformal map for spherical entangling surface

- Let A be a ball $\{\rho \leq R, t = 0\}$ in \mathbb{R}^d
- For CFT, the partition function is invariant

$$Z_n[\mathbb{R}^d] = Z[\mathbb{S}^1 \times \mathbb{H}^{d-1}]|_{\tau \sim \tau + 2\pi n} = \text{tr}(e^{-\beta H})|_{\beta = 2\pi n}$$



- The conformal map to the hyperbolic coordinates leads to the equivalence of the EE across \mathbb{S}^{d-2} and the thermal entropy on \mathbb{H}^{d-1} at $T = 1/(2\pi)$

$$S_A = S_{\text{therm}}[\mathbb{H}^{d-1}]|_{T=1/(2\pi)}$$

- This relation will be derived from the holographic viewpoint

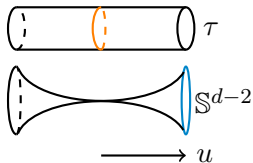
- The conformal map to the hyperbolic coordinates leads to the equivalence of the EE across \mathbb{S}^{d-2} and the thermal entropy on \mathbb{H}^{d-1} at $T = 1/(2\pi)$

$$S_A = S_{\text{therm}}[\mathbb{H}^{d-1}]|_{T=1/(2\pi)}$$

- This relation will be derived from the holographic viewpoint

Further map to sphere

- A coordinate transformation $\sinh u = \cot \theta$ yields a map to \mathbb{S}^d

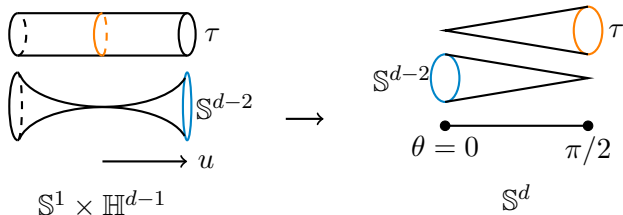


$$\mathbb{S}^1 \times \mathbb{H}^{d-1}$$

$$ds^2 = d\tau^2 + du^2 + \sinh^2 u d\Omega_{d-2}^2$$

Further map to sphere

- A coordinate transformation $\sinh u = \cot \theta$ yields a map to \mathbb{S}^d

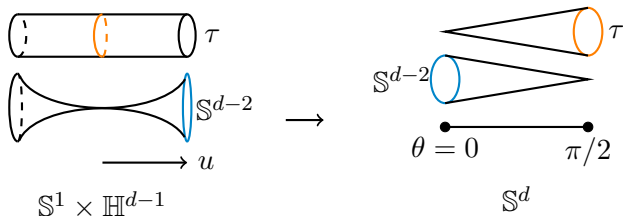


$$ds^2 = d\tau^2 + du^2 + \sinh^2 u d\Omega_{d-2}^2$$

$$ds^2 = d\theta^2 + \sin^2 \theta d\tau^2 + \cos^2 \theta d\Omega_{d-2}^2$$

Further map to sphere

- A coordinate transformation $\sinh u = \cot \theta$ yields a map to \mathbb{S}^d



$$ds^2 = d\tau^2 + du^2 + \sinh^2 u d\Omega_{d-2}^2$$

$$ds^2 = d\theta^2 + \sin^2 \theta d\tau^2 + \cos^2 \theta d\Omega_{d-2}^2$$

Replica partition function

$$Z_n[\mathbb{R}^d] = Z[\mathbb{S}^1 \times \mathbb{H}^{d-1}]|_{\tau \sim \tau + 2\pi n} = Z[\mathbb{S}^d]|_{\tau \sim \tau + 2\pi n}$$

- After the conformal transformation, the entropy is mapped to a sphere partition function

For CFT and spherical entangling surface

$$S_A = \log Z[\mathbb{S}^d]$$

- \mathbb{S}_n^d : the n -fold cover of \mathbb{S}^d
 $ds^2 = d\theta^2 + \sin^2 \theta d\tau^2 + \cos^2 \theta d\Omega_{d-2}^2$, $\tau \sim \tau + 2\pi n$
- This relation allows us to calculate EE exactly for free field and SUSY gauge theories!

- After the conformal transformation, the entropy is mapped to a sphere partition function

For CFT and spherical entangling surface

$$S_A = \log Z[\mathbb{S}^d]$$

- \mathbb{S}_n^d : the n -fold cover of \mathbb{S}^d
 $ds^2 = d\theta^2 + \sin^2 \theta d\tau^2 + \cos^2 \theta d\Omega_{d-2}^2$, $\tau \sim \tau + 2\pi n$
- This relation allows us to calculate EE exactly for free field and SUSY gauge theories!

- 1 Basics of entanglement entropy
- 2 Field theoretic methods
- 3 Conformal field theory
- 4 Holographic method**
- 5 Renormalization group flow
- 6 Perturbation
- 7 Summary

The AdS geometries

- Consider the flat $(d + 2)$ -dimensional pseudo Euclidean space defined by

$$ds^2 = -dy_{-1}^2 - dy_0^2 + dy_1^2 + \cdots + dy_d^2$$

- The AdS_{d+1} space with the radius L is defined as a submanifold satisfying

$$-y_{-1}^2 - y_0^2 + y_1^2 + \cdots + y_d^2 = -L^2$$



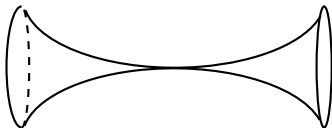
The AdS geometries

- Consider the flat $(d + 2)$ -dimensional pseudo Euclidean space defined by

$$ds^2 = -dy_{-1}^2 - dy_0^2 + dy_1^2 + \cdots + dy_d^2$$

- The AdS_{d+1} space with the radius L is defined as a submanifold satisfying

$$-y_{-1}^2 - y_0^2 + y_1^2 + \cdots + y_d^2 = -L^2$$



- The coordinate transformations

$$y_{-1} = \frac{L^2 - t^2 + z^2 + \sum_{i=1}^{d-1} x_i^2}{2z}, \quad y_d = \frac{-L^2 - t^2 + z^2 + \sum_{i=1}^{d-1} x_i^2}{2z}$$

$$y_0 = Lt/z, \quad y_i = Lx_i/z, \quad (i = 1, \dots, d-1)$$

- The metric becomes

$$ds^2 = L^2 \left[\frac{dr^2}{r^2} + r^2 \left(-dt^2 + \sum_{i=1}^{d-1} dx_i^2 \right) \right]$$

- These coordinates cover **half** of the whole AdS_{d+1} space and the Euclidean boundary at $r = \infty$ is \mathbb{R}^d

- The coordinate transformations

$$y_{-1} = \frac{L^2 - t^2 + z^2 + \sum_{i=1}^{d-1} x_i^2}{2z}, \quad y_d = \frac{-L^2 - t^2 + z^2 + \sum_{i=1}^{d-1} x_i^2}{2z}$$

$$y_0 = Lt/z, \quad y_i = Lx_i/z, \quad (i = 1, \dots, d-1)$$

- The metric becomes

$$ds^2 = L^2 \left[\frac{dr^2}{r^2} + r^2 \left(-dt^2 + \sum_{i=1}^{d-1} dx_i^2 \right) \right]$$

- These coordinates cover half of the whole AdS_{d+1} space and the Euclidean boundary at $r = \infty$ is \mathbb{R}^d

- The coordinate transformations

$$y_{-1} = \frac{L^2 - t^2 + z^2 + \sum_{i=1}^{d-1} x_i^2}{2z}, \quad y_d = \frac{-L^2 - t^2 + z^2 + \sum_{i=1}^{d-1} x_i^2}{2z}$$

$$y_0 = Lt/z, \quad y_i = Lx_i/z, \quad (i = 1, \dots, d-1)$$

- The metric becomes

$$ds^2 = L^2 \left[\frac{dr^2}{r^2} + r^2 \left(-dt^2 + \sum_{i=1}^{d-1} dx_i^2 \right) \right]$$

- These coordinates cover **half** of the whole AdS_{d+1} space and the Euclidean boundary at $r = \infty$ is \mathbb{R}^d

- Choose the coordinates as

$$\begin{aligned}y_{-1} &= L \cosh \rho \sin \tau, & y_0 &= L \cosh \rho \cos \tau \\y_i &= L \sinh \rho e^i, & (i &= 1, \dots, d)\end{aligned}$$

where e^i satisfy $\sum_{i=1}^d (e^i)^2 = 1$

- The metric becomes

$$ds^2 = L^2 \left[-\cosh^2 \rho d\tau^2 + d\rho^2 + \sinh^2 \rho d\Omega_{d-1}^2 \right]$$

- The Euclidean boundary is \mathbb{S}^d

- Choose the coordinates as

$$y_{-1} = L \cosh \rho \sin \tau, \quad y_0 = L \cosh \rho \cos \tau$$
$$y_i = L \sinh \rho e^i, \quad (i = 1, \dots, d)$$

where e^i satisfy $\sum_{i=1}^d (e^i)^2 = 1$

- The metric becomes

$$ds^2 = L^2 [-\cosh^2 \rho d\tau^2 + d\rho^2 + \sinh^2 \rho d\Omega_{d-1}^2]$$

- The Euclidean boundary is S^d

- Choose the coordinates as

$$y_{-1} = L \cosh \rho \sin \tau, \quad y_0 = L \cosh \rho \cos \tau$$
$$y_i = L \sinh \rho e^i, \quad (i = 1, \dots, d)$$

where e^i satisfy $\sum_{i=1}^d (e^i)^2 = 1$

- The metric becomes

$$ds^2 = L^2 [-\cosh^2 \rho d\tau^2 + d\rho^2 + \sinh^2 \rho d\Omega_{d-1}^2]$$

- The Euclidean boundary is \mathbb{S}^d

- The coordinate transformations:

$$y_{-1} = r \cosh u, \quad y_0 = \sqrt{r^2 - L^2} \sinh \frac{t}{L}$$
$$y_d = \sqrt{r^2 - L^2} \cosh \frac{t}{L}, \quad y_i = r \sinh u e^i$$

- The resulting metric is

$$ds^2 = - \left(\frac{r^2}{L^2} - 1 \right) dt^2 + \frac{dr^2}{\frac{r^2}{L^2} - 1} + r^2 (du^2 + \sinh^2 u d\Omega_{d-2}^2)$$

- Cover **half** of the whole AdS_{d+1} space with Euclidean boundary $\mathbb{S}^1 \times \mathbb{H}^{d-1}$
- Event horizon at $r = L$ with $\beta = 2\pi L$

- The coordinate transformations:

$$y_{-1} = r \cosh u, \quad y_0 = \sqrt{r^2 - L^2} \sinh \frac{t}{L}$$
$$y_d = \sqrt{r^2 - L^2} \cosh \frac{t}{L}, \quad y_i = r \sinh u e^i$$

- The resulting metric is

$$ds^2 = - \left(\frac{r^2}{L^2} - 1 \right) dt^2 + \frac{dr^2}{\frac{r^2}{L^2} - 1} + r^2 (du^2 + \sinh^2 u d\Omega_{d-2}^2)$$

- Cover half of the whole AdS_{d+1} space with Euclidean boundary $\mathbb{S}^1 \times \mathbb{H}^{d-1}$
- Event horizon at $r = L$ with $\beta = 2\pi L$

- The coordinate transformations:

$$y_{-1} = r \cosh u, \quad y_0 = \sqrt{r^2 - L^2} \sinh \frac{t}{L}$$
$$y_d = \sqrt{r^2 - L^2} \cosh \frac{t}{L}, \quad y_i = r \sinh u e^i$$

- The resulting metric is

$$ds^2 = - \left(\frac{r^2}{L^2} - 1 \right) dt^2 + \frac{dr^2}{\frac{r^2}{L^2} - 1} + r^2 (du^2 + \sinh^2 u d\Omega_{d-2}^2)$$

- Cover **half** of the whole AdS_{d+1} space with Euclidean boundary $\mathbb{S}^1 \times \mathbb{H}^{d-1}$

- Event horizon at $r = L$ with $\beta = 2\pi L$

- The coordinate transformations:

$$y_{-1} = r \cosh u, \quad y_0 = \sqrt{r^2 - L^2} \sinh \frac{t}{L}$$
$$y_d = \sqrt{r^2 - L^2} \cosh \frac{t}{L}, \quad y_i = r \sinh u e^i$$

- The resulting metric is

$$ds^2 = - \left(\frac{r^2}{L^2} - 1 \right) dt^2 + \frac{dr^2}{\frac{r^2}{L^2} - 1} + r^2 (du^2 + \sinh^2 u d\Omega_{d-2}^2)$$

- Cover **half** of the whole AdS_{d+1} space with Euclidean boundary $\mathbb{S}^1 \times \mathbb{H}^{d-1}$
- Event horizon at $r = L$ with $\beta = 2\pi L$

- The AdS/CFT relates the partition functions

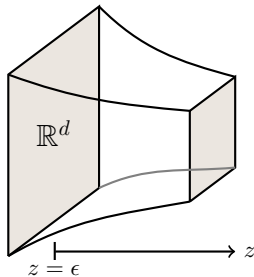
GKP-W relation

$$e^{-I_{\text{bulk}}[\mathcal{B}=\text{AdS}_{d+1}]} = Z_{\text{CFT}}[\partial\mathcal{B}]$$

- Consider the Einstein-Hilbert action

$$I_{\text{bulk}}[\mathcal{B}] = -\frac{1}{16\pi G_N} \int_{\mathcal{B}} d^{d+1}x \sqrt{g} \left(\mathcal{R} + \frac{d(d-1)}{L^2} \right)$$

AdS_{d+1} space



- The AdS/CFT relates the partition functions

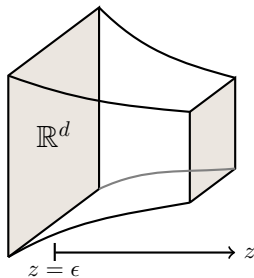
GKP-W relation

$$e^{-I_{\text{bulk}}[\mathcal{B}=\text{AdS}_{d+1}]} = Z_{\text{CFT}}[\partial\mathcal{B}]$$

- Consider the Einstein-Hilbert action

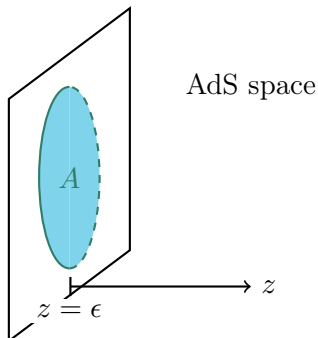
$$I_{\text{bulk}}[\mathcal{B}] = -\frac{1}{16\pi G_N} \int_{\mathcal{B}} d^{d+1}x \sqrt{g} \left(\mathcal{R} + \frac{d(d-1)}{L^2} \right)$$

AdS_{d+1} space



Following the GKP-W

$$S_A = \lim_{n \rightarrow 1} \partial_n (I_{\text{bulk}}[\mathcal{B}_n] - n I_{\text{bulk}}[\mathcal{B}])$$

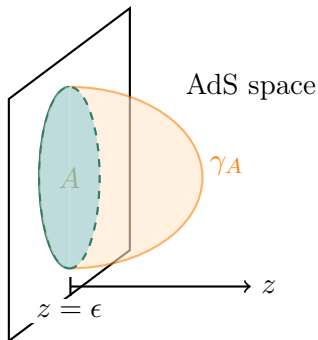


Following the GKP-W

$$S_A = \lim_{n \rightarrow 1} \partial_n (I_{\text{bulk}}[\mathcal{B}_n] - n I_{\text{bulk}}[\mathcal{B}])$$

Holographic formula [Ryu-Takayanagi 06]

$$S_A = \frac{\text{Area}(\gamma_A)}{4G_N}$$



Following the GKP-W

$$S_A = \lim_{n \rightarrow 1} \partial_n (I_{\text{bulk}}[\mathcal{B}_n] - n I_{\text{bulk}}[\mathcal{B}])$$

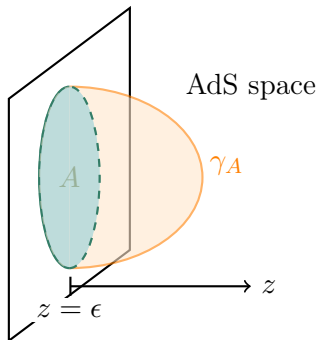
Holographic formula [Ryu-Takayanagi 06]

$$S_A = \frac{\text{Area}(\gamma_A)}{4G_N}$$

Reproduce the **area law divergence**

$$S_A = \frac{\text{Area}(\partial A)}{\epsilon^{d-2}} + \dots$$

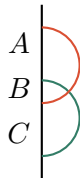
ϵ : UV cutoff at $z = \epsilon$



SSA follows from the **minimality**

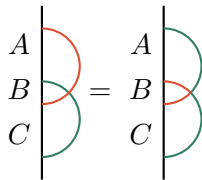
[Headrick-Takayanagi 07]

SSA follows from the **minimality**
[Headrick-Takayanagi 07]



SSA follows from the **minimality**

[Headrick-Takayanagi 07]



SSA follows from the **minimality**

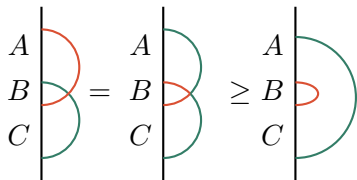
[Headrick-Takayanagi 07]

The diagram shows an equality and an inequality between tensor network diagrams. On the left, a vertical line is labeled with A , B , and C from top to bottom. A red arc connects the A and B regions, and a green arc connects the B and C regions. This is equal to a diagram where the red arc connects A and C , and the green arc connects B and C . This is greater than or equal to a diagram where the red arc connects A and B , and the green arc connects A and C .

SSA follows from the **minimality**

[Headrick-Takayanagi 07]

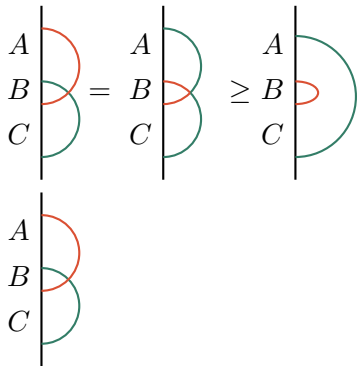
$$S_{A \cup B \cup C} + S_B \leq S_{A \cup B} + S_{B \cup C}$$



SSA follows from the **minimality**

[Headrick-Takayanagi 07]

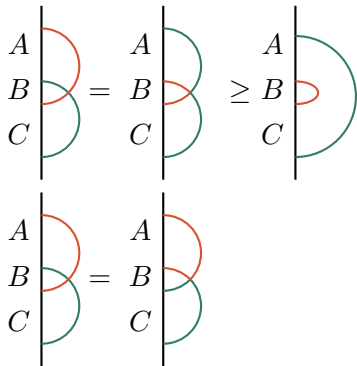
$$S_{A \cup B \cup C} + S_B \leq S_{A \cup B} + S_{B \cup C}$$



SSA follows from the **minimality**

[Headrick-Takayanagi 07]

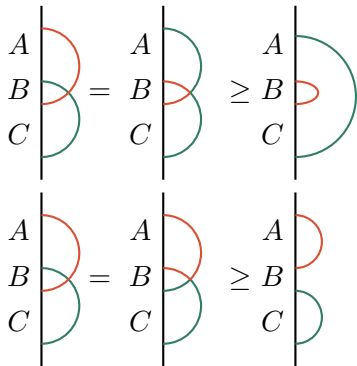
$$S_{A \cup B \cup C} + S_B \leq S_{A \cup B} + S_{B \cup C}$$



SSA follows from the **minimality**

[Headrick-Takayanagi 07]

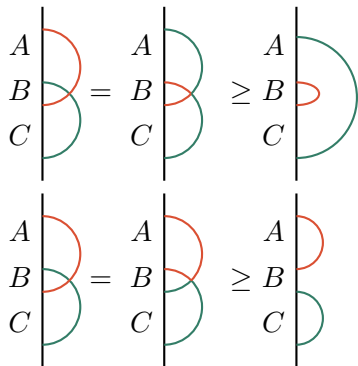
$$S_{A \cup B \cup C} + S_B \leq S_{A \cup B} + S_{B \cup C}$$



SSA follows from the **minimality**

[Headrick-Takayanagi 07]

$$S_{A \cup B \cup C} + S_B \leq S_{A \cup B} + S_{B \cup C}$$



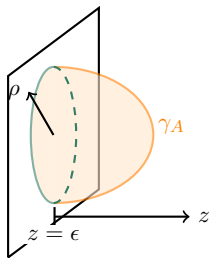
Spherical entangling surface in CFT_d

- In the Poincaré patch,
 $\Sigma = \{\rho = R, t = 0\}$

$$ds^2 = L^2 \frac{dz^2 + dt^2 + d\rho^2 + \rho^2 d\Omega_{d-2}^2}{z^2}$$

- The area functional for $z = z(\rho)$

$$\text{Area}(\gamma_A) = L^{d-1} \text{Vol}(\mathbb{S}^{d-2}) \int_0^R d\rho \frac{\rho^{d-2}}{z^{d-1}(\rho)} \sqrt{1 + (\partial_\rho z)^2}$$



- The minimal surface

$$z(\rho) = \sqrt{R^2 - \rho^2}$$

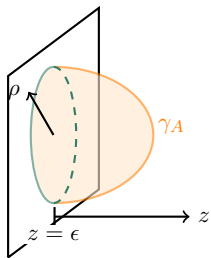
Spherical entangling surface in CFT_d

- In the Poincaré patch,
 $\Sigma = \{\rho = R, t = 0\}$

$$ds^2 = L^2 \frac{dz^2 + dt^2 + d\rho^2 + \rho^2 d\Omega_{d-2}^2}{z^2}$$

- The area functional for $z = z(\rho)$

$$\text{Area}(\gamma_A) = L^{d-1} \text{Vol}(\mathbb{S}^{d-2}) \int_0^R d\rho \frac{\rho^{d-2}}{z^{d-1}(\rho)} \sqrt{1 + (\partial_\rho z)^2}$$



- The minimal surface

$$z(\rho) = \sqrt{R^2 - \rho^2}$$

Spherical entangling surface in CFT_d

- In the Poincaré patch,
 $\Sigma = \{\rho = R, t = 0\}$

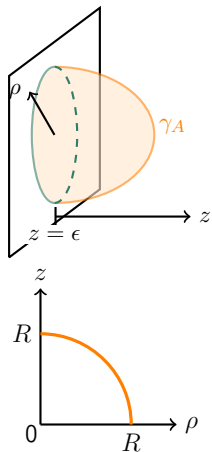
$$ds^2 = L^2 \frac{dz^2 + dt^2 + d\rho^2 + \rho^2 d\Omega_{d-2}^2}{z^2}$$

- The area functional for $z = z(\rho)$

$$\text{Area}(\gamma_A) = L^{d-1} \text{Vol}(\mathbb{S}^{d-2}) \int_0^R d\rho \frac{\rho^{d-2}}{z^{d-1}(\rho)} \sqrt{1 + (\partial_\rho z)^2}$$

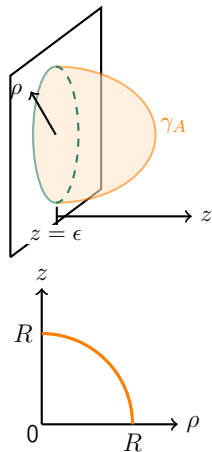
- The minimal surface

$$z(\rho) = \sqrt{R^2 - \rho^2}$$



■ Holographic EE

$$\begin{aligned}
 S_A &= \frac{L^{d-1} \text{Vol}(\mathbb{S}^{d-2})}{4G_N} \int_{\epsilon/R}^1 dy \frac{(1-y^2)^{\frac{d-3}{2}}}{y^{d-1}} \\
 &= \frac{L^{d-1} \text{Vol}(\mathbb{S}^{d-2})}{4G_N} \left[\frac{1}{d-2} \frac{R^{d-2}}{\epsilon^{d-2}} + \dots \right]
 \end{aligned}$$

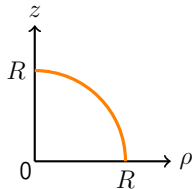
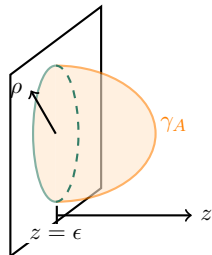


■ Holographic EE

$$\begin{aligned}
 S_A &= \frac{L^{d-1} \text{Vol}(\mathbb{S}^{d-2})}{4G_N} \int_{\epsilon/R}^1 dy \frac{(1-y^2)^{\frac{d-3}{2}}}{y^{d-1}} \\
 &= \frac{L^{d-1} \text{Vol}(\mathbb{S}^{d-2})}{4G_N} \left[\frac{1}{d-2} \frac{R^{d-2}}{\epsilon^{d-2}} + \dots \right]
 \end{aligned}$$

In odd dimensions

$$F = \frac{L^{d-1}}{4G_N} \frac{\pi^{\frac{d}{2}}}{\Gamma(\frac{d}{2})}$$



■ Holographic EE

$$S_A = \frac{L^{d-1} \text{Vol}(\mathbb{S}^{d-2})}{4G_N} \int_{\epsilon/R}^1 dy \frac{(1-y^2)^{\frac{d-3}{2}}}{y^{d-1}}$$

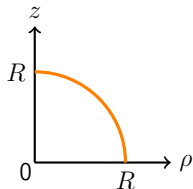
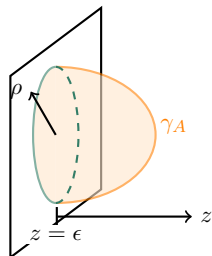
$$= \frac{L^{d-1} \text{Vol}(\mathbb{S}^{d-2})}{4G_N} \left[\frac{1}{d-2} \frac{R^{d-2}}{\epsilon^{d-2}} + \dots \right]$$

In odd dimensions

$$F = \frac{L^{d-1}}{4G_N} \frac{\pi^{\frac{d}{2}}}{\Gamma(\frac{d}{2})}$$

In even dimensions

$$c_0 = (-1)^{\frac{d}{2}+1} A = (-1)^{\frac{d}{2}+1} \frac{L^{d-2}}{2G_N} \frac{\pi^{\frac{d}{2}-1}}{\Gamma(\frac{d}{2})}$$

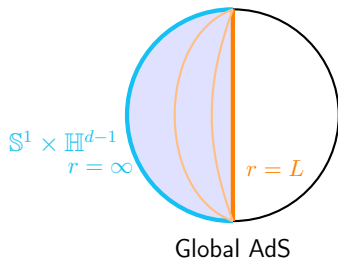


Viewpoint from the hyperbolic coordinates

- The entangling surface Σ is at the **spatial infinity** in the hyperbolic coordinates
- The minimal surface is anchored on Σ
- It coincides with the **BH horizon**!

Holographic EE for spherical entangling surface

$$S_A(R) = S_{\text{BH}}(T) = S_{\text{therm}}(T)$$

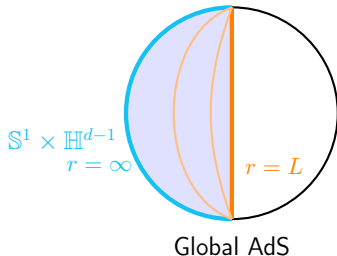


Viewpoint from the hyperbolic coordinates

- The entangling surface Σ is at the **spatial infinity** in the hyperbolic coordinates
- The minimal surface is anchored on Σ
- It coincides with the **BH horizon**!

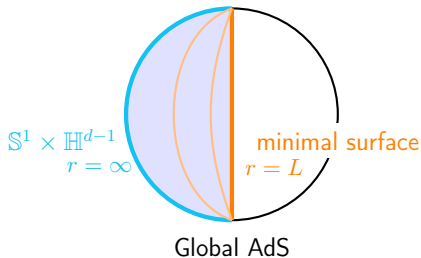
Holographic EE for spherical entangling surface

$$S_A(R) = S_{\text{BH}}(T) = S_{\text{therm}}(T)$$



Viewpoint from the hyperbolic coordinates

- The entangling surface Σ is at the **spatial infinity** in the hyperbolic coordinates
- The minimal surface is anchored on Σ
- It coincides with the **BH horizon!**

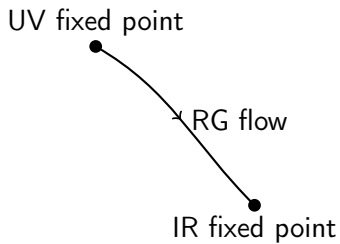


Holographic EE for spherical entangling surface

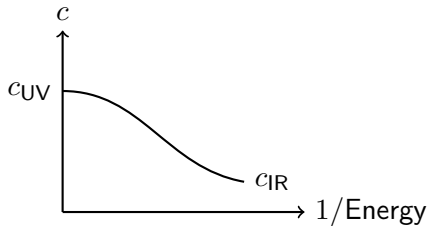
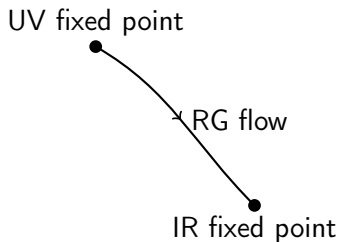
$$S_A(R) = S_{\text{BH}}(T) = S_{\text{therm}}(T)$$

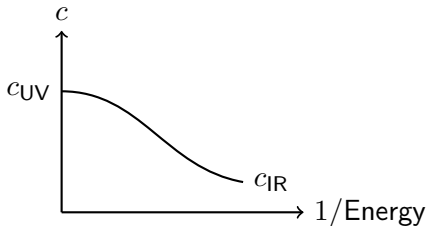
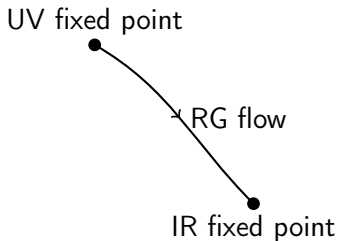
Outline

- 1 Basics of entanglement entropy
- 2 Field theoretic methods
- 3 Conformal field theory
- 4 Holographic method
- 5 Renormalization group flow
- 6 Perturbation
- 7 Summary



RG flow and c -function





c -function can be a measure of degrees of freedom!

[Zamolodchikov 86, Cardy 88, Komargodski-Shwimmer 11]

- 2d entropic c -function:

$$c(r) \equiv 3r \frac{dS_A(r)}{dr}$$

- Interpolate two fixed points

$$c(r) \rightarrow c_{UV} \quad (r \rightarrow 0), \quad c(r) \rightarrow c_{IR} \quad (r \rightarrow \infty)$$

- SSA + Lorentz invariance \Rightarrow **monotonicity** [Casini-Huerta 04]

$$c'(r) \leq 0$$

- 2d entropic c -function:

$$c(r) \equiv 3r \frac{dS_A(r)}{dr}$$

- Interpolate two fixed points

$$c(r) \rightarrow c_{UV} \quad (r \rightarrow 0), \quad c(r) \rightarrow c_{IR} \quad (r \rightarrow \infty)$$

- SSA + Lorentz invariance \Rightarrow monotonicity [Casini-Huerta 04]

$$c'(r) \leq 0$$

- 2d entropic c -function:

$$c(r) \equiv 3r \frac{dS_A(r)}{dr}$$

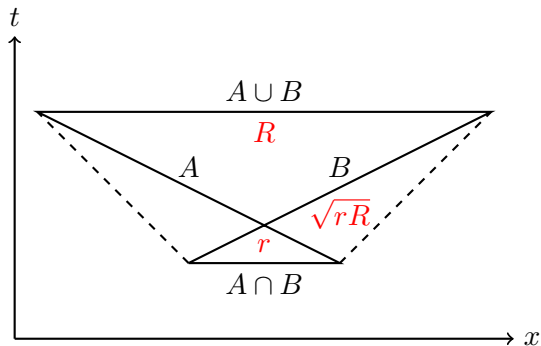
- Interpolate two fixed points

$$c(r) \rightarrow c_{UV} \quad (r \rightarrow 0), \quad c(r) \rightarrow c_{IR} \quad (r \rightarrow \infty)$$

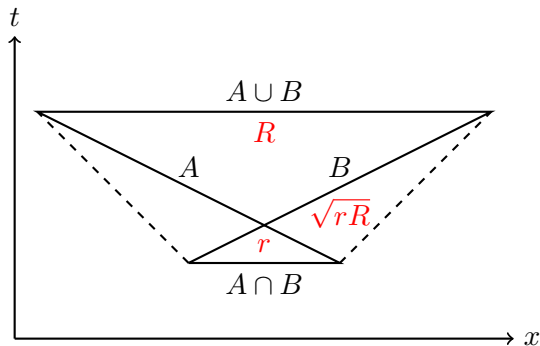
- SSA + Lorentz invariance \Rightarrow **monotonicity** [Casini-Huerta 04]

$$c'(r) \leq 0$$

Proof of entropic c -theorem



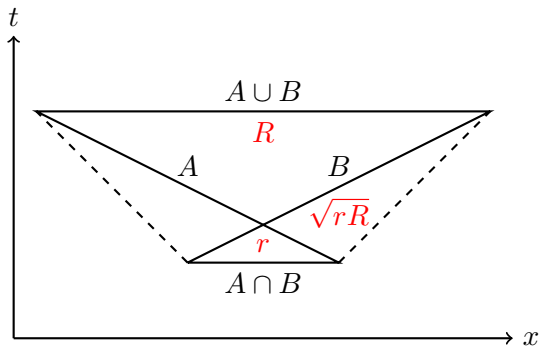
Proof of entropic c -theorem



SSA

$$S_A + S_B \geq S_{A \cup B} + S_{A \cap B}$$

Proof of entropic c -theorem



SSA

$$2S(\sqrt{rR}) \geq S(R) + S(r) \quad \Rightarrow \quad c'(r) \leq 0$$

- Entropic c -function (not stationary at a fixed point)

$$c(t) = c \quad \text{for CFT} , \quad c'(t) \leq 0$$

- Zamolodchikov's c -function (stationary at a fixed point)

$$c'(t) = -\frac{3}{2}G_{ij}\beta^i\beta^j \leq 0 , \quad \frac{\partial c}{\partial g^i} = G_{ij}\beta^j$$

- Thermal c -function

$$F_{\text{Therm}} \sim cT^2$$

- Every c -function coincides at a conformal fixed point

- Entropic c -function (not stationary at a fixed point)

$$c(t) = c \quad \text{for CFT} , \quad c'(t) \leq 0$$

- Zamolodchikov's c -function (stationary at a fixed point)

$$c'(t) = -\frac{3}{2}G_{ij}\beta^i\beta^j \leq 0 , \quad \frac{\partial c}{\partial g^i} = G_{ij}\beta^j$$

- Thermal c -function

$$F_{\text{Therm}} \sim cT^2$$

- Every c -function coincides at a conformal fixed point

- Entropic c -function (not stationary at a fixed point)

$$c(t) = c \quad \text{for CFT} , \quad c'(t) \leq 0$$

- Zamolodchikov's c -function (stationary at a fixed point)

$$c'(t) = -\frac{3}{2}G_{ij}\beta^i\beta^j \leq 0 , \quad \frac{\partial c}{\partial g^i} = G_{ij}\beta^j$$

- Thermal c -function

$$F_{\text{Therm}} \sim cT^2$$

- Every c -function coincides at a conformal fixed point

- Entropic c -function (not stationary at a fixed point)

$$c(t) = c \quad \text{for CFT} , \quad c'(t) \leq 0$$

- Zamolodchikov's c -function (stationary at a fixed point)

$$c'(t) = -\frac{3}{2}G_{ij}\beta^i\beta^j \leq 0 , \quad \frac{\partial c}{\partial g^i} = G_{ij}\beta^j$$

- Thermal c -function

$$F_{\text{Therm}} \sim cT^2$$

- Every c -function coincides at a conformal fixed point

- Thermal c -theorem:

$$F_{\text{Therm}} \sim c_{\text{Therm}} T^3$$

- C_T -theorem: [Petkou 94]

$$C_T|_{\text{UV}} \geq C_T|_{\text{IR}} , \quad \langle T_{\mu\nu}(x) T_{\rho\sigma}(0) \rangle_{\text{CFT}} = C_T \frac{I_{\mu\nu,\rho\sigma}(x)}{x^6}$$

- F -theorem: [Jafferis-Klebanov-Pufu-Safdi 11, Myers-Sinha 10]

$$F_{\text{UV}}(\mathbb{S}^3) \geq F_{\text{IR}}(\mathbb{S}^3) , \quad F = -\log Z(\mathbb{S}^3)$$

- Thermal c -theorem: Counter example by [Sachdev 93]

$$F_{\text{Therm}} \sim c_{\text{Therm}} T^3$$

- C_T -theorem: [Petkou 94]

$$C_T|_{\text{UV}} \geq C_T|_{\text{IR}} , \quad \langle T_{\mu\nu}(x) T_{\rho\sigma}(0) \rangle_{\text{CFT}} = C_T \frac{I_{\mu\nu,\rho\sigma}(x)}{x^6}$$

- F -theorem: [Jafferis-Klebanov-Pufu-Safdi 11, Myers-Sinha 10]

$$F_{\text{UV}}(\mathbb{S}^3) \geq F_{\text{IR}}(\mathbb{S}^3) , \quad F = -\log Z(\mathbb{S}^3)$$

- Thermal c -theorem: Counter example by [Sachdev 93]

$$F_{\text{Therm}} \sim c_{\text{Therm}} T^3$$

- C_T -theorem: [Petkou 94]

$$C_T|_{\text{UV}} \geq C_T|_{\text{IR}} , \quad \langle T_{\mu\nu}(x) T_{\rho\sigma}(0) \rangle_{\text{CFT}} = C_T \frac{I_{\mu\nu,\rho\sigma}(x)}{x^6}$$

- F -theorem: [Jafferis-Klebanov-Pufu-Safdi 11, Myers-Sinha 10]

$$F_{\text{UV}}(\mathbb{S}^3) \geq F_{\text{IR}}(\mathbb{S}^3) , \quad F = -\log Z(\mathbb{S}^3)$$

- Thermal c -theorem: Counter example by [Sachdev 93]

$$F_{\text{Therm}} \sim c_{\text{Therm}} T^3$$

- C_T -theorem: [Petkou 94] Counter example by [TN-Yonekura 13]

$$C_T|_{\text{UV}} \geq C_T|_{\text{IR}} , \quad \langle T_{\mu\nu}(x) T_{\rho\sigma}(0) \rangle_{\text{CFT}} = C_T \frac{I_{\mu\nu,\rho\sigma}(x)}{x^6}$$

- F -theorem: [Jafferis-Klebanov-Pufu-Safdi 11, Myers-Sinha 10]

$$F_{\text{UV}}(\mathbb{S}^3) \geq F_{\text{IR}}(\mathbb{S}^3) , \quad F = -\log Z(\mathbb{S}^3)$$

- Thermal c -theorem: Counter example by [Sachdev 93]

$$F_{\text{Therm}} \sim c_{\text{Therm}} T^3$$

- C_T -theorem: [Petkou 94] Counter example by [TN-Yonekura 13]

$$C_T|_{\text{UV}} \geq C_T|_{\text{IR}} , \quad \langle T_{\mu\nu}(x) T_{\rho\sigma}(0) \rangle_{\text{CFT}} = C_T \frac{I_{\mu\nu,\rho\sigma}(x)}{x^6}$$

- F -theorem: [Jafferis-Klebanov-Pufu-Safdi 11, Myers-Sinha 10]

$$F_{\text{UV}}(\mathbb{S}^3) \geq F_{\text{IR}}(\mathbb{S}^3) , \quad F = -\log Z(\mathbb{S}^3)$$

We use a **renormalized partition function** in the F -theorem

$$F(\mathbb{S}^3) \equiv -\log Z^{(\text{ren})}(\mathbb{S}^3) = \text{finite}$$

We use a **renormalized partition function** in the F -theorem

$$F(\mathbb{S}^3) \equiv -\log Z^{(\text{ren})}(\mathbb{S}^3) = \text{finite}$$

For CFT₃ [Casini-Huerta-Myers 11]

$$S_A(R) = \log Z[\mathbb{S}^3]$$

We use a **renormalized partition function** in the F -theorem

$$F(\mathbb{S}^3) \equiv -\log Z^{(\text{ren})}(\mathbb{S}^3) = \text{finite}$$

For CFT_3 [Casini-Huerta-Myers 11]

$$S_A(R) = \log Z[\mathbb{S}^3] = \alpha \frac{R}{\epsilon} - F(\mathbb{S}^3)$$

We use a **renormalized partition function** in the F -theorem

$$F(\mathbb{S}^3) \equiv -\log Z^{(\text{ren})}(\mathbb{S}^3) = \text{finite}$$

For CFT_3 [Casini-Huerta-Myers 11]

$$S_A(R) = \log Z[\mathbb{S}^3] = \alpha \frac{R}{\epsilon} - F(\mathbb{S}^3)$$

Proof of the F -theorem by using entanglement entropy?

- Interpolating function between F_{UV} and F_{IR}

- Interpolating function between F_{UV} and F_{IR}
- Monotonically decreasing under RG flow

- Interpolating function between F_{UV} and F_{IR}
- Monotonically decreasing under RG flow

Renormalized entanglement entropy [Liu-Mezzi 12]

$$\mathcal{F}(R) \equiv (R\partial_R - 1)S_A(R)$$

Renormalized entanglement entropy [Liu-Mezzi 12]

$$\mathcal{F}(R) \equiv (R\partial_R - 1)S_A(R)$$

- For CFT₃

$$S_A(R) = \alpha \frac{2\pi R}{\epsilon} - F(\mathbb{S}^3) \quad \Rightarrow \quad \mathcal{F}(R) = F(\mathbb{S}^3)$$

Renormalized entanglement entropy [Liu-Mezzi 12]

$$\mathcal{F}(R) \equiv (R\partial_R - 1)S_A(R)$$

- For CFT₃

$$S_A(R) = \alpha \frac{2\pi R}{\epsilon} - F(\mathbb{S}^3) \quad \Rightarrow \quad \mathcal{F}(R) = F(\mathbb{S}^3)$$

- Proof of monotonicity [Casini-Huerta 12]

$$\text{SSA} + \text{Lorentz invariance} \quad \Rightarrow \quad \mathcal{F}'(R) = R S''(R) \leq 0$$

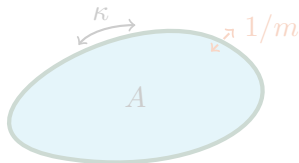
- Large m expansion: [cf. Grover-Turner-Vishwanath 11]

$$S_A(R) = \alpha \frac{\ell_\Sigma}{\epsilon} + \beta m \ell_\Sigma - \gamma + \sum_{l=0}^{\infty} \frac{c_{-1-2l}^\Sigma}{m^{2l+1}}$$

- ℓ_Σ : length of $\Sigma = \partial A$
- γ : topological entanglement entropy [Kitaev-Preskill 05, Levin-Wen 05]

- $c_{-1-2l}^\Sigma = \int_\Sigma f(\kappa, \partial_s \kappa, \partial_s^2 \kappa, \dots)$

f : even for $\kappa \rightarrow -\kappa$ ($S_A = S_{\bar{A}}$)



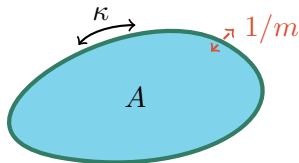
- Large m expansion: [cf. Grover-Turner-Vishwanath 11]

$$S_A(R) = \alpha \frac{\ell_\Sigma}{\epsilon} + \beta m \ell_\Sigma - \gamma + \sum_{l=0}^{\infty} \frac{c_{-1-2l}^\Sigma}{m^{2l+1}}$$

- ℓ_Σ : length of $\Sigma = \partial A$
- γ : topological entanglement entropy [Kitaev-Preskill 05, Levin-Wen 05]

- $c_{-1-2l}^\Sigma = \int_\Sigma f(\kappa, \partial_s \kappa, \partial_s^2 \kappa, \dots)$

f : even for $\kappa \rightarrow -\kappa$ ($S_A = S_{\bar{A}}$)



- Dimensional reduction:

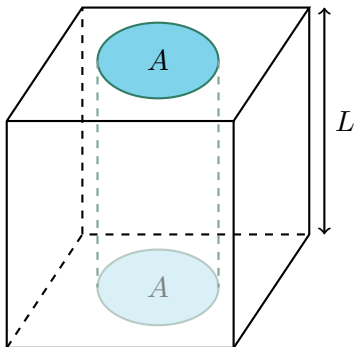
$$\mathbb{R}^{2,1} \times \mathbb{S}^1 \rightarrow \mathbb{R}^{2,1}$$

[Huerta 11, Klebanov-TN-Pufu-Safdi 12]

- Entangling surface: $\Sigma \times \mathbb{S}^1 \rightarrow \Sigma$

4d EE from 3d EE

$$S_{\Sigma \times \mathbb{S}^1}^{(3+1)} = \sum_{n \in \mathbb{Z}} S_{\Sigma}^{(2+1)} \left(m = \left\lfloor \frac{2\pi n}{L} \right\rfloor \right)$$



- Dimensional reduction:

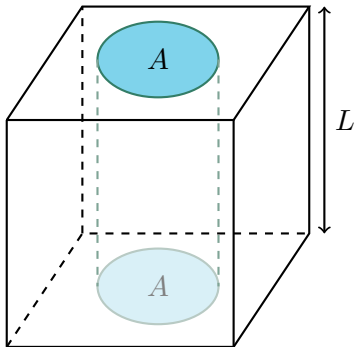
$$\mathbb{R}^{2,1} \times \mathbb{S}^1 \rightarrow \mathbb{R}^{2,1}$$

[Huerta 11, Klebanov-TN-Pufu-Safdi 12]

- Entangling surface: $\Sigma \times \mathbb{S}^1 \rightarrow \Sigma$

4d EE from 3d EE

$$S_{\Sigma \times \mathbb{S}^1}^{(3+1)} = \sum_{n \in \mathbb{Z}} S_{\Sigma}^{(2+1)} \left(m = \left\lfloor \frac{2\pi n}{L} \right\rfloor \right)$$



Shape dependence of EE in 3d

[Klebanov-TN-Pufu-Safdi 12]

- Log divergence in the large L limit

$$S_{\Sigma_2 = \Sigma \times S^1}^{(3+1)} = \sum_{n \in \mathbb{Z}} S_{\Sigma}^{(2+1)}(m_n)$$

- 4d EE has a logarithmic divergence

$$S_{\Sigma_2}^{(3+1)}|_{\log} = \frac{c}{2\pi} \int_{\Sigma_2} \left(\mathcal{K}_{\mu\nu}^a \mathcal{K}^{a\mu\nu} - \frac{1}{2} (\mathcal{K}_{\mu}^a)^2 \right) \log \epsilon ,$$

- From 6d anomaly, $c_{-3}^{\Sigma} = \# \oint_{\Sigma} ds \kappa^4 + \# \oint_{\Sigma} ds \left(\frac{d\kappa}{ds} \right)^2$

Shape dependence of EE in 3d

[Klebanov-TN-Pufu-Safdi 12]

- Log divergence in the large L limit

$$S_{\Sigma_2=\Sigma\times\mathbb{S}^1}^{(3+1)} \xrightarrow{L\rightarrow\infty} \frac{L}{\pi} \int_0^{1/\epsilon} dp S_{\Sigma}^{(2+1)}(m=p)$$

- 4d EE has a logarithmic divergence

$$S_{\Sigma_2}^{(3+1)}|_{\log} = \frac{c}{2\pi} \int_{\Sigma_2} \left(\mathcal{K}_{\mu\nu}^a \mathcal{K}^{a\mu\nu} - \frac{1}{2} (\mathcal{K}^a_{\mu}{}^{\mu})^2 \right) \log \epsilon ,$$

- From 6d anomaly, $c_{-3}^{\Sigma} = \# \oint_{\Sigma} ds \kappa^4 + \# \oint_{\Sigma} ds \left(\frac{d\kappa}{ds} \right)^2$

Shape dependence of EE in 3d

[Klebanov-TN-Pufu-Safdi 12]

- Log divergence in the large L limit

$$S_{\Sigma_2=\Sigma \times \mathbb{S}^1}^{(3+1)} \xrightarrow{L \rightarrow \infty} \frac{L}{\pi} \int_0^{1/\epsilon} dp S_{\Sigma}^{(2+1)}(m=p) \sim c_{-1}^{\Sigma} \log \epsilon$$

- 4d EE has a logarithmic divergence

$$S_{\Sigma_2}^{(3+1)} \Big|_{\log} = \frac{c}{2\pi} \int_{\Sigma_2} \left(\mathcal{K}_{\mu\nu}^a \mathcal{K}^{a\mu\nu} - \frac{1}{2} (\mathcal{K}^a_{\mu}{}^{\mu})^2 \right) \log \epsilon ,$$

- From 6d anomaly, $c_{-3}^{\Sigma} = \# \oint_{\Sigma} ds \kappa^4 + \# \oint_{\Sigma} ds \left(\frac{d\kappa}{ds} \right)^2$

- Log divergence in the large L limit

$$S_{\Sigma_2 = \Sigma \times \mathbb{S}^1}^{(3+1)} \xrightarrow{L \rightarrow \infty} \frac{L}{\pi} \int_0^{1/\epsilon} dp S_{\Sigma}^{(2+1)}(m=p) \sim c_{-1}^{\Sigma} \log \epsilon$$

- 4d EE has a logarithmic divergence

$$S_{\Sigma_2}^{(3+1)} \Big|_{\log} = \frac{c}{2\pi} \int_{\Sigma_2} \left(\mathcal{K}_{\mu\nu}^a \mathcal{K}^{a\mu\nu} - \frac{1}{2} (\mathcal{K}^a_{\mu}{}^{\mu})^2 \right) \log \epsilon ,$$

- From 6d anomaly, $c_{-3}^{\Sigma} = \# \oint_{\Sigma} ds \kappa^4 + \# \oint_{\Sigma} ds \left(\frac{d\kappa}{ds} \right)^2$

Shape dependence of EE in 3d

[Klebanov-TN-Pufu-Safdi 12]

- Log divergence in the large L limit

$$S_{\Sigma_2 = \Sigma \times S^1}^{(3+1)} \xrightarrow{L \rightarrow \infty} \frac{L}{\pi} \int_0^{1/\epsilon} dp S_{\Sigma}^{(2+1)}(m=p) \sim c_{-1}^{\Sigma} \log \epsilon$$

- 4d EE has a logarithmic divergence

$$S_{\Sigma_2}^{(3+1)} \Big|_{\log} = \frac{c}{2\pi} \int_{\Sigma_2} \left(\mathcal{K}_{\mu\nu}^a \mathcal{K}^{a\mu\nu} - \frac{1}{2} (\mathcal{K}_{\mu}^a)^2 \right) \log \epsilon ,$$

For free n_0 scalar fields and $n_{1/2}$ Dirac fermions

$$c_{-1}^{\Sigma} = \frac{1}{480} (n_0 + 3n_{1/2}) \oint_{\Sigma} ds \kappa^2$$

- From 6d anomaly, $c_{-3}^{\Sigma} = \# \oint_{\Sigma} ds \kappa^4 + \# \oint_{\Sigma} ds \left(\frac{d\kappa}{ds} \right)^2$

- Log divergence in the large L limit

$$S_{\Sigma_2 = \Sigma \times S^1}^{(3+1)} \xrightarrow{L \rightarrow \infty} \frac{L}{\pi} \int_0^{1/\epsilon} dp S_{\Sigma}^{(2+1)}(m=p) \sim c_{-1}^{\Sigma} \log \epsilon$$

- 4d EE has a logarithmic divergence

$$S_{\Sigma_2}^{(3+1)} \Big|_{\log} = \frac{c}{2\pi} \int_{\Sigma_2} \left(\mathcal{K}_{\mu\nu}^a \mathcal{K}^{a\mu\nu} - \frac{1}{2} (\mathcal{K}_{\mu}^a)^2 \right) \log \epsilon ,$$

For free n_0 scalar fields and $n_{1/2}$ Dirac fermions

$$c_{-1}^{\Sigma} = \frac{1}{480} (n_0 + 3n_{1/2}) \oint_{\Sigma} ds \kappa^2$$

- From 6d anomaly, $c_{-3}^{\Sigma} = \# \oint_{\Sigma} ds \kappa^4 + \# \oint_{\Sigma} ds \left(\frac{d\kappa}{ds} \right)^2$

- REE: $\Sigma =$ a circle of radius $R \Rightarrow \kappa = \frac{1}{R}$
- REE is monotonically decreasing to zero in IR!
- What happens in small mass region?

- REE: $\Sigma =$ a circle of radius $R \Rightarrow \kappa = \frac{1}{R}$

$$\mathcal{F}(R) = \frac{\pi}{120mR} + O(1/(mR)^3)$$

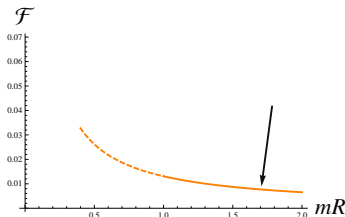
- REE is monotonically decreasing to zero in IR!
- What happens in small mass region?

REE of free massive scalar in IR region

- REE: $\Sigma =$ a circle of radius $R \Rightarrow \kappa = \frac{1}{R}$

$$\mathcal{F}(R) = \frac{\pi}{120mR} + O(1/(mR)^3)$$

- REE is monotonically decreasing to zero in IR!
- What happens in small mass region?

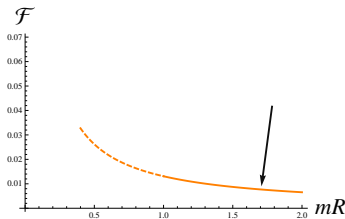


REE of free massive scalar in IR region

- REE: $\Sigma =$ a circle of radius $R \Rightarrow \kappa = \frac{1}{R}$

$$\mathcal{F}(R) = \frac{\pi}{120mR} + O(1/(mR)^3)$$

- REE is monotonically decreasing to zero in IR!
- What happens in small mass region?



- Perturbation around $m = 0$ doesn't work for this case (will be discussed from the holographic viewpoint)
- Numerical method [Huerta 11]: $\mathcal{F}(0) \simeq 0.0638 = F_{UV}(\mathbb{S}^3)$
- \mathcal{F} is **not stationary** at UV fixed point!
 $(\partial_{(mR)^2} \mathcal{F})|_{(mR)^2=0} \sim \langle \phi^2 \rangle \neq 0$
[Klebanov-TN-Pufu-Safdi 12, TN 14]

$$\mathcal{F} = F_{UV} - 0.13(mR)^2$$

- IR divergence?

- Perturbation around $m = 0$ doesn't work for this case (will be discussed from the holographic viewpoint)
- Numerical method [Huerta 11]: $\mathcal{F}(0) \simeq 0.0638 = F_{UV}(\mathbb{S}^3)$
- \mathcal{F} is **not stationary** at UV fixed point!
 $(\partial_{(mR)^2} \mathcal{F})|_{(mR)^2=0} \sim \langle \phi^2 \rangle \neq 0$
[Klebanov-TN-Pufu-Safdi 12, TN 14]

$$\mathcal{F} = F_{UV} - 0.13(mR)^2$$

- IR divergence?

Numerical study of REE for free massive scalar

- Perturbation around $m = 0$ doesn't work for this case (will be discussed from the holographic viewpoint)

- Numerical method [Huerta 11]: $\mathcal{F}(0) \simeq 0.0638 = F_{UV}(\mathbb{S}^3)$

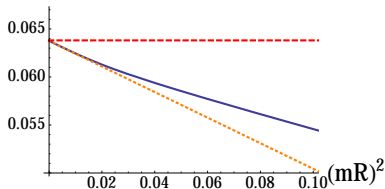
- \mathcal{F} is **not stationary** at UV fixed point!

$$(\partial_{(mR)^2} \mathcal{F})|_{(mR)^2=0} \sim \langle \phi^2 \rangle \neq 0$$

[Klebanov-TN-Pufu-Safdi 12, TN 14] \mathcal{F}

$$\mathcal{F} = F_{UV} - 0.13(mR)^2$$

- IR divergence?



Numerical study of REE for free massive scalar

- Perturbation around $m = 0$ doesn't work for this case (will be discussed from the holographic viewpoint)

- Numerical method [Huerta 11]: $\mathcal{F}(0) \simeq 0.0638 = F_{UV}(\mathbb{S}^3)$

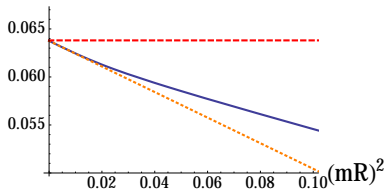
- \mathcal{F} is **not stationary** at UV fixed point!

$$(\partial_{(mR)^2} \mathcal{F})|_{(mR)^2=0} \sim \langle \phi^2 \rangle \neq 0$$

[Klebanov-TN-Pufu-Safdi 12, TN 14] \mathcal{F}

$$\mathcal{F} = F_{UV} - 0.13(mR)^2$$

- IR divergence?



- Asymptotically AdS space

$$ds^2 = \frac{L^2}{z^2} \left[\frac{dz^2}{f(z)} - dt^2 + d\vec{x}_{d-1}^2 \right]$$

if $f(z) \rightarrow 1$ as $z \rightarrow 0$

- Consider the Einstein gravity coupled to matters

$$I = \frac{1}{16\pi G_N} \int d^{d+1}x \sqrt{-g} [\mathcal{R} + \mathcal{L}_{\text{matter}}]$$

- Asymptotically AdS space

$$ds^2 = \frac{L^2}{z^2} \left[\frac{dz^2}{f(z)} - dt^2 + d\vec{x}_{d-1}^2 \right]$$

if $f(z) \rightarrow 1$ as $z \rightarrow 0$

- Consider the Einstein gravity coupled to matters

$$I = \frac{1}{16\pi G_N} \int d^{d+1}x \sqrt{-g} [\mathcal{R} + \mathcal{L}_{\text{matter}}]$$

- Asymptotically AdS space

$$ds^2 = \frac{L^2}{z^2} \left[\frac{dz^2}{f(z)} - dt^2 + d\vec{x}_{d-1}^2 \right]$$

if $f(z) \rightarrow 1$ as $z \rightarrow 0$

- Consider the Einstein gravity coupled to matters

$$I = \frac{1}{16\pi G_N} \int d^{d+1}x \sqrt{-g} [\mathcal{R} + \mathcal{L}_{\text{matter}}]$$

Null energy condition

$$T_{\mu\nu}^{\text{matter}} \xi^\mu \xi^\nu \geq 0 \quad \text{for any null vector} \quad (\xi_\mu \xi^\mu = 0)$$

- The null vector

$$\xi^z = \sqrt{f(z)}, \quad \xi^t = 1, \quad \xi^i = 0 \quad (i \neq t, z)$$

- The Einstein equation

$$T_{\mu\nu}^{\text{matter}} = \mathcal{R}_{\mu\nu} - \frac{\mathcal{R}}{2} g_{\mu\nu}$$

- Applying the NEC

$$f'(z) \geq 0$$

- The null vector

$$\xi^z = \sqrt{f(z)}, \quad \xi^t = 1, \quad \xi^i = 0 \quad (i \neq t, z)$$

- The Einstein equation

$$T_{\mu\nu}^{\text{matter}} = \mathcal{R}_{\mu\nu} - \frac{\mathcal{R}}{2} g_{\mu\nu}$$

- Applying the NEC

$$f'(z) \geq 0$$

- The null vector

$$\xi^z = \sqrt{f(z)}, \quad \xi^t = 1, \quad \xi^i = 0 \quad (i \neq t, z)$$

- The Einstein equation

$$T_{\mu\nu}^{\text{matter}} = \mathcal{R}_{\mu\nu} - \frac{\mathcal{R}}{2} g_{\mu\nu}$$

- Applying the NEC

$$f'(z) \geq 0$$

- The null vector

$$\xi^z = \sqrt{f(z)}, \quad \xi^t = 1, \quad \xi^i = 0 \quad (i \neq t, z)$$

- The Einstein equation

$$T_{\mu\nu}^{\text{matter}} = \mathcal{R}_{\mu\nu} - \frac{\mathcal{R}}{2}g_{\mu\nu}$$

- Applying the NEC

$$f'(z) \geq 0$$

A counterpart of SSA?

- A solution interpolating two AdS spaces of radius L and L_{IR}

$$f(z) = 1 + \eta g(z) , \quad \eta = \frac{L^2}{L_{\text{IR}}^2} - 1 \ll 1$$

where $g'(z) \geq 0$, $g(0) = 0$, $g(\infty) = 1$

- Perturbatively calculate the entropy across \mathbb{S}^{d-2} with $\rho_0(z) = \sqrt{R^2 - z^2}$

$$\rho(z) = \rho_0(z) + \eta \rho_1(z) + O(\eta^2)$$

which gives the expansion of the area functional

$$\text{Area} = A_0 + \eta A_1 + O(\eta^2)$$

- A solution interpolating two AdS spaces of radius L and L_{IR}

$$f(z) = 1 + \eta g(z) , \quad \eta = \frac{L^2}{L_{\text{IR}}^2} - 1 \ll 1$$

where $g'(z) \geq 0$, $g(0) = 0$, $g(\infty) = 1$

- Perturbatively calculate the entropy across \mathbb{S}^{d-2} with $\rho_0(z) = \sqrt{R^2 - z^2}$

$$\rho(z) = \rho_0(z) + \eta \rho_1(z) + O(\eta^2)$$

which gives the expansion of the area functional

$$\text{Area} = A_0 + \eta A_1 + O(\eta^2)$$

- The variation of the \mathcal{F} -function [Liu-Mezei 12]

$$\Delta\mathcal{F}(R) = -\eta \frac{\pi L^2}{2G_N} \int_0^1 dz g(zR)$$

- The difference between UV and IR

$$\Delta\mathcal{F}(R \rightarrow \infty) = -\eta \frac{\pi L^2}{2G_N}$$

- The variation of the \mathcal{F} -function [Liu-Mezei 12]

$$\Delta\mathcal{F}(R) = -\eta \frac{\pi L^2}{2G_N} \int_0^1 dz g(zR)$$

Monotonically decreases as R increases!

- The difference between UV and IR

$$\Delta\mathcal{F}(R \rightarrow \infty) = -\eta \frac{\pi L^2}{2G_N}$$

- The variation of the \mathcal{F} -function [Liu-Mezei 12]

$$\Delta\mathcal{F}(R) = -\eta \frac{\pi L^2}{2G_N} \int_0^1 dz g(zR)$$

Monotonically decreases as R increases!

- The difference between UV and IR

$$\Delta\mathcal{F}(R \rightarrow \infty) = -\eta \frac{\pi L^2}{2G_N}$$

- The variation of the \mathcal{F} -function [Liu-Mezei 12]

$$\Delta\mathcal{F}(R) = -\eta \frac{\pi L^2}{2G_N} \int_0^1 dz g(zR)$$

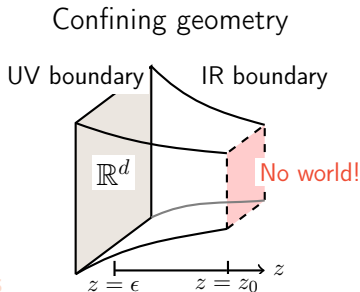
Monotonically decreases as R increases!

- The difference between UV and IR

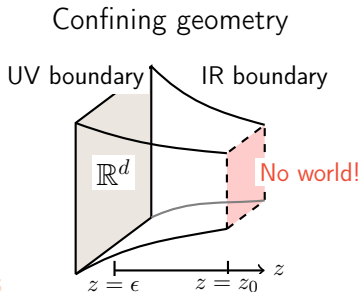
$$\begin{aligned}\Delta\mathcal{F}(R \rightarrow \infty) &= -\eta \frac{\pi L^2}{2G_N} \\ &= F_{\text{IR}} - F_{\text{UV}} + O(\eta^2)\end{aligned}$$

- Cap off in IR of AdS space
⇔ confining gauge theory
- Two minimal surfaces:
 - disk type
 - cylinder type
- A phase transition happens from the disk type to the cylinder type as the radius is increased
(~ a confinement/deconfinement transition)

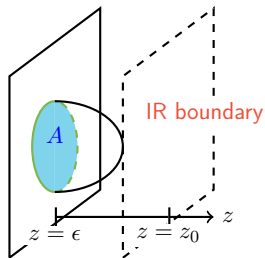
- Cap off in IR of AdS space
⇔ confining gauge theory
- Two minimal surfaces:
 - disk type
 - cylinder type
- A phase transition happens from the disk type to the cylinder type as the radius is increased
(~ a confinement/deconfinement transition)



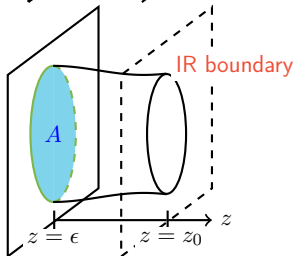
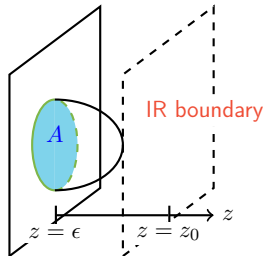
- Cap off in IR of AdS space
⇔ confining gauge theory
- Two minimal surfaces:
 - disk type
 - cylinder type
- A phase transition happens from the disk type to the cylinder type as the radius is increased
(~ a confinement/deconfinement transition)



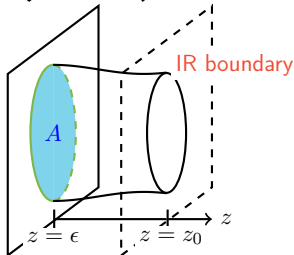
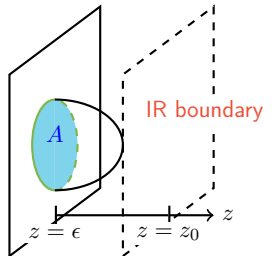
- Cap off in IR of AdS space
⇔ confining gauge theory
- Two minimal surfaces:
 - disk type
 - cylinder type
- A phase transition happens from the disk type to the cylinder type as the radius is increased
(~ a confinement/deconfinement transition)



- Cap off in IR of AdS space
⇔ confining gauge theory
- Two minimal surfaces:
 - disk type
 - cylinder type
- A phase transition happens from the disk type to the cylinder type as the radius is increased
(~ a confinement/deconfinement transition)



- Cap off in IR of AdS space
⇔ confining gauge theory
- Two minimal surfaces:
 - disk type
 - cylinder type
- A phase transition happens from the disk type to the cylinder type as the radius is increased
(~ a confinement/deconfinement transition)



[Klebanov-TN-Pufu-Safdi 12]

- Dual to a gapped $(2 + 1)$ -dim QFT [Cvetic-Gibbons-Lu-Pope 00]
- Relevant deformation at UV: $S = S_{UV} + g \int d^3x \mathcal{O}(x)$
 $\Delta[\mathcal{O}] = \frac{7}{3}, \Delta[g] = \frac{2}{3}$

[Klebanov-TN-Pufu-Safdi 12]

- Dual to a gapped $(2 + 1)$ -dim QFT [Cvetic-Gibbons-Lu-Pope 00]
- Relevant deformation at UV: $S = S_{UV} + g \int d^3x \mathcal{O}(x)$
 $\Delta[\mathcal{O}] = \frac{7}{3}, \Delta[g] = \frac{2}{3}$

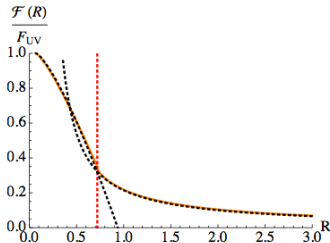
Holographic example: CGLP solution

[Klebanov-TN-Pufu-Safdi 12]

- Dual to a gapped $(2 + 1)$ -dim QFT [Cvetic-Gibbons-Lu-Pope 00]
- Relevant deformation at UV: $S = S_{UV} + g \int d^3x \mathcal{O}(x)$
 $\Delta[\mathcal{O}] = \frac{7}{3}, \Delta[g] = \frac{2}{3}$

REE is stationary at UV fixed point

$$\partial_g \mathcal{F}(g^{3/2} R) \Big|_{g=0} = 0$$



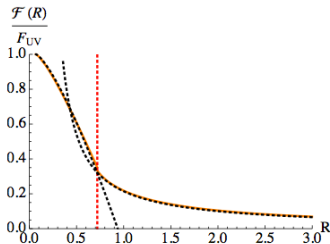
[Klebanov-TN-Pufu-Safdi 12]

- Dual to a gapped $(2 + 1)$ -dim QFT [Cvetic-Gibbons-Lu-Pope 00]
- Relevant deformation at UV: $S = S_{UV} + g \int d^3x \mathcal{O}(x)$
 $\Delta[\mathcal{O}] = \frac{7}{3}, \Delta[g] = \frac{2}{3}$

REE is stationary at UV fixed point

$$\partial_g \mathcal{F}(g^{3/2} R) \Big|_{g=0} = 0$$

When is REE stationary for a relevant perturbation?



Outline

- 1 Basics of entanglement entropy
- 2 Field theoretic methods
- 3 Conformal field theory
- 4 Holographic method
- 5 Renormalization group flow
- 6 Perturbation**
- 7 Summary

- Perturbation of CFT_d : $S = S_{\text{CFT}} + g \int d^d x \mathcal{O}$
- Holographically described by a free massive scalar Φ of mass M in AdS_{d+1}

$$\Delta_{\pm} = \frac{d}{2} \pm \sqrt{(ML_{\text{AdS}})^2 + \frac{d^2}{4}}$$

- Two boundary conditions near $z \rightarrow 0$ [Klebanov-Witten 99]

$$\Phi(z, \vec{x}) \rightarrow z^{\Delta_+} [A(\vec{x}) + \dots] + z^{\Delta_-} [B(\vec{x}) + \dots]$$

- $\Delta = \Delta_+$: $A = \langle \mathcal{O} \rangle$, $B = g$ (standard quantization)
- $\Delta = \Delta_-$: $A = g$, $B = \langle \mathcal{O} \rangle$ (alternative quantization)

- Perturbation of CFT_d : $S = S_{\text{CFT}} + g \int d^d x \mathcal{O}$
- Holographically described by a free massive scalar Φ of mass M in AdS_{d+1}

$$\Delta_{\pm} = \frac{d}{2} \pm \sqrt{(ML_{\text{AdS}})^2 + \frac{d^2}{4}}$$

- Two boundary conditions near $z \rightarrow 0$ [Klebanov-Witten 99]

$$\Phi(z, \vec{x}) \rightarrow z^{\Delta_+} [A(\vec{x}) + \dots] + z^{\Delta_-} [B(\vec{x}) + \dots]$$

- $\Delta = \Delta_+$: $A = \langle \mathcal{O} \rangle$, $B = g$ (standard quantization)
- $\Delta = \Delta_-$: $A = g$, $B = \langle \mathcal{O} \rangle$ (alternative quantization)

- Perturbation of CFT_d : $S = S_{\text{CFT}} + g \int d^d x \mathcal{O}$
- Holographically described by a free massive scalar Φ of mass M in AdS_{d+1}

$$\Delta_{\pm} = \frac{d}{2} \pm \sqrt{(ML_{\text{AdS}})^2 + \frac{d^2}{4}}$$

- Two boundary conditions near $z \rightarrow 0$ [Klebanov-Witten 99]

$$\Phi(z, \vec{x}) \rightarrow z^{\Delta_+} [A(\vec{x}) + \dots] + z^{\Delta_-} [B(\vec{x}) + \dots]$$

- $\Delta = \Delta_+$: $A = \langle \mathcal{O} \rangle$, $B = g$ (standard quantization)
- $\Delta = \Delta_-$: $A = g$, $B = \langle \mathcal{O} \rangle$ (alternative quantization)

- The backreacted geometry by the scalar field

$$f(z) = 1 + \begin{cases} \#z^{2\Delta-} + \dots, & \Delta \neq d/2 \\ \#z^d(\log z)^2 + \dots, & \Delta = d/2 \end{cases}$$

- HEE of a disk in the backreacted geometry ($t \equiv gR^{d-\Delta}$)

$$\frac{dS}{dt} = \begin{cases} -\#t^{2\Delta-/(d-\Delta)-1} + \dots, & \Delta \neq d/2 \\ -\#t \log^2 t + \dots, & \Delta = d/2 \end{cases}$$

- Near the UV fixed point ($t = 0$)

- The backreacted geometry by the scalar field

$$f(z) = 1 + \begin{cases} \#z^{2\Delta-} + \dots, & \Delta \neq d/2 \\ \#z^d(\log z)^2 + \dots, & \Delta = d/2 \end{cases}$$

- HEE of a disk in the backreacted geometry ($t \equiv gR^{d-\Delta}$)

$$\frac{dS}{dt} = \begin{cases} -\#t^{2\Delta-/(d-\Delta)-1} + \dots, & \Delta \neq d/2 \\ -\#t \log^2 t + \dots, & \Delta = d/2 \end{cases}$$

- Near the UV fixed point ($t = 0$)

- The backreacted geometry by the scalar field

$$f(z) = 1 + \begin{cases} \#z^{2\Delta_-} + \dots, & \Delta \neq d/2 \\ \#z^d(\log z)^2 + \dots, & \Delta = d/2 \end{cases}$$

- HEE of a disk in the backreacted geometry ($t \equiv gR^{d-\Delta}$)

$$\frac{dS}{dt} = \begin{cases} -\#t^{2\Delta_-/(d-\Delta)-1} + \dots, & \Delta \neq d/2 \\ -\#t \log^2 t + \dots, & \Delta = d/2 \end{cases}$$

- Near the UV fixed point ($t = 0$)

Classification of EE for a relevant perturbation

- (1) $d/2 < \Delta < d$: stationary ($dS/dt|_{t=0} = 0$)
- (2) $d/3 < \Delta \leq d/2$: stationary, but the perturbation fails
- (3) $d/2 - 1 < \Delta \leq d/3$: neither stationary nor perturbative

- Free massive scalar in 3d: $\Delta = 1 \Rightarrow (3)$

$$S(mR) = S(0) - \#(mR)^2 + \dots$$

Consistent with the numerical computation of the \mathcal{F} -function!

[Klebanov-TN-Pufu-Safdi 12, TN 14]

- Free massive fermion in 2d (an interval): $\Delta = 1 \Rightarrow (2)$

$$S(mR) = S(0) - \#(mR)^2 \log^2(mR) + \dots$$

Agree with the known results! [Casini-Fosco-Huerta 05, Herzog-TN 13]

- Under the variation $\rho_A \rightarrow \rho_A + \delta\rho_A$

1st law of entanglement

$$\delta S_A = -\text{tr}_A(\delta\rho_A \log \rho_A)$$

where $\delta\langle\mathcal{O}\rangle \equiv \text{tr}_A(\delta\rho_A\mathcal{O})$

- The **modular Hamiltonian** $H_A \equiv -\log \rho_A$ is given by the stress-energy tensor
- For a spherical entangling surface in CFT [Casini-Huerta-Myers 11]

$$H_A = 2\pi \int_{r \leq R} d^{d-1}x \frac{R^2 - r^2}{2R} T_{00}(x)$$

- Under the variation $\rho_A \rightarrow \rho_A + \delta\rho_A$

1st law of entanglement

$$\delta S_A = -\text{tr}_A(\delta\rho_A \log \rho_A) = \delta\langle H_A \rangle$$

where $\delta\langle \mathcal{O} \rangle \equiv \text{tr}_A(\delta\rho_A \mathcal{O})$

- The **modular Hamiltonian** $H_A \equiv -\log \rho_A$ is given by the stress-energy tensor
- For a spherical entangling surface in CFT [Casini-Huerta-Myers 11]

$$H_A = 2\pi \int_{r \leq R} d^{d-1}x \frac{R^2 - r^2}{2R} T_{00}(x)$$

- Under the variation $\rho_A \rightarrow \rho_A + \delta\rho_A$

1st law of entanglement

$$\delta S_A = -\text{tr}_A(\delta\rho_A \log \rho_A) = \delta\langle H_A \rangle$$

where $\delta\langle \mathcal{O} \rangle \equiv \text{tr}_A(\delta\rho_A \mathcal{O})$

- The **modular Hamiltonian** $H_A \equiv -\log \rho_A$ is given by the stress-energy tensor
- For a spherical entangling surface in CFT [Casini-Huerta-Myers 11]

$$H_A = 2\pi \int_{r \leq R} d^{d-1}x \frac{R^2 - r^2}{2R} T_{00}(x)$$

- Under the variation $\rho_A \rightarrow \rho_A + \delta\rho_A$

1st law of entanglement

$$\delta S_A = -\text{tr}_A(\delta\rho_A \log \rho_A) = \delta\langle H_A \rangle$$

where $\delta\langle \mathcal{O} \rangle \equiv \text{tr}_A(\delta\rho_A \mathcal{O})$

- The **modular Hamiltonian** $H_A \equiv -\log \rho_A$ is given by the stress-energy tensor
- For a spherical entangling surface in CFT [Casini-Huerta-Myers 11]

$$H_A = 2\pi \int_{r \leq R} d^{d-1}x \frac{R^2 - r^2}{2R} T_{00}(x)$$

- For a spherical entangling surface in CFT

The variation of the density matrix $\rho_A \rightarrow \rho_A + \delta\rho_A$

$$\delta S_A = 2\pi \int_{r \leq R} d^{d-1}x \frac{R^2 - r^2}{2R} \delta \langle T_{00}(x) \rangle$$

- In the gravity theory with the holographic stress tensor $\delta \langle T_{00}(x) \rangle \propto h_{\mu\nu}(x, z \rightarrow 0)$

The variation of the metric $g_{\mu\nu} \rightarrow g_{\mu\nu} + h_{\mu\nu}$

$$\frac{\delta \text{Area}}{4G_N} = 2\pi \int_{r \leq R} d^{d-1}x \frac{R^2 - r^2}{2R} \delta \langle T_{00}(x) \rangle$$

- For a spherical entangling surface in CFT

The variation of the density matrix $\rho_A \rightarrow \rho_A + \delta\rho_A$

$$\delta S_A = 2\pi \int_{r \leq R} d^{d-1}x \frac{R^2 - r^2}{2R} \delta \langle T_{00}(x) \rangle$$

- In the gravity theory with the holographic stress tensor
 $\delta \langle T_{00}(x) \rangle \propto h_{\mu\nu}(x, z \rightarrow 0)$

The variation of the metric $g_{\mu\nu} \rightarrow g_{\mu\nu} + h_{\mu\nu}$

$$\frac{\delta \text{Area}}{4G_N} = 2\pi \int_{r \leq R} d^{d-1}x \frac{R^2 - r^2}{2R} \delta \langle T_{00}(x) \rangle$$

- For a spherical entangling surface in CFT

The variation of the density matrix $\rho_A \rightarrow \rho_A + \delta\rho_A$

$$\delta S_A = 2\pi \int_{r \leq R} d^{d-1}x \frac{R^2 - r^2}{2R} \delta \langle T_{00}(x) \rangle$$

- In the gravity theory with the holographic stress tensor $\delta \langle T_{00}(x) \rangle \propto h_{\mu\nu}(x, z \rightarrow 0)$

The variation of the metric $g_{\mu\nu} \rightarrow g_{\mu\nu} + h_{\mu\nu}$

$$\frac{\delta \text{Area}}{4G_N} = 2\pi \int_{r \leq R} d^{d-1}x \frac{R^2 - r^2}{2R} \delta \langle T_{00}(x) \rangle$$

It yields the **linearized Einstein equation** $\delta E_{\mu\nu}[h] = 0!$

[Lashkari-McDermott-Raamsdonk 13,

Faulkner-Guica-Hartman-Myers-Raamsdonk 13]

- 1 Basics of entanglement entropy
- 2 Field theoretic methods
- 3 Conformal field theory
- 4 Holographic method
- 5 Renormalization group flow
- 6 Perturbation
- 7 Summary**

- EE a useful measure of degrees of freedom defined in arbitrary dimensions
 - For even d , central charge dependence in the logarithmic term
 - For odd d , the finite part as an analogue of central charge
- The entropic c -theorem in two dimensions and the F -theorem in three dimensions
- REE not a c -function in the Zamolodchikov's sense (non-stationarity \simeq IR divergence)
- 1st law of entanglement = the linearized Einstein equation through the holographic formula

- Proof of the a - and F -theorem with SSA in higher dimensions? (The holographic c -theorem [Myers-Sinha 10, Freedman-Gubser-Pilch-Warner 99, ...])
- Perturbative computation of EE in QFT? IR divergence? [Rosenhaus-Smolkin 14]
- Holographic Rényi entropy formula? (For a spherical entangling surface, [Hung-Myers-Smolkin-Yale 11])

- Is SSA equivalent to the Null energy condition?
[Lashkari-Rabideau-Sabella-Garnier-Raamsdonk 14,
Bhattacharya-Hubeny-Rangamani-Takayanagi 14]
- Einstein gravity from entanglement at non-linear level? (with
MERA [Swingle 09, Raamsdonk 09, Nozaki-Ryu-Takayanagi 12, ...])
- ... and more!

See the slides of the workshop
"Quantum Information in Quantum Gravity"
<http://www.maths.dur.ac.uk/~dma0mr/qiqg-ubc/>

- Is SSA equivalent to the Null energy condition?
[Lashkari-Rabideau-Sabella-Garnier-Raamsdonk 14,
Bhattacharya-Hubeny-Rangamani-Takayanagi 14]
- Einstein gravity from entanglement at non-linear level? (with
MERA [Swingle 09, Raamsdonk 09, Nozaki-Ryu-Takayanagi 12, ...])
- ... and more!

See the slides of the workshop
“Quantum Information in Quantum Gravity”
<http://www.maths.dur.ac.uk/dma0mr/qiqg-ubc/>