

# Leptophilic DM model for LFV Higgs decay

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based on

SB, Zhaofeng Kang, arXiv:1510.00100

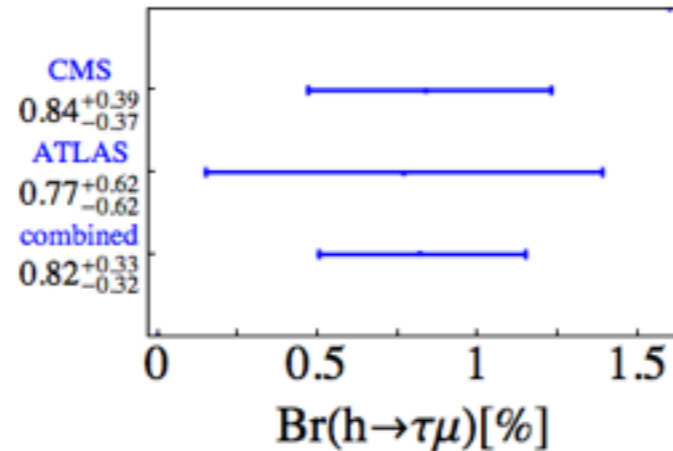
# Outline

- LFV Higgs decay  $h \rightarrow \tau\mu$
- Leptophilic model
- $h \rightarrow \tau\mu$
- $\tau \rightarrow \mu\gamma$ ,  $h \rightarrow \gamma\gamma$  constraints,  
relic density and direct detection of DM
- Conclusions

# LFV Higgs decay

- Discovery of 125 GeV Higgs
- Measurement of Higgs couplings to the SM particles
- LF is accidental symmetry in the minimal SM.  
Neutrino oscillation experiments show it is broken.  
Although LFV Higgs decay is allowed in the SM, this loop-induced process is highly suppressed by the neutrino mass.
- Measurement of LFV Higgs decay → New Physics

# LFV Higgs decay



- CMS reports  $2.4\sigma$  excess from the SM, using  $19.7\text{fb}^{-1}$  at  $\sqrt{s} = 8 \text{ TeV}$ , [CMS cn, arXiv:1502.07400](#)

$$\mathcal{B}(h \rightarrow \mu\tau) = (0.84_{-0.37}^{+0.39})\%$$

- ATLAS result with similar dataset is consistent with CMS

$$\mathcal{B}(h \rightarrow \mu\tau) = (0.77 \pm 0.62)\%, \quad \text{ATLAS cn, arXiv:1508.03372}$$

- LHC Run II can probe down to  $\sim 10^{-3}$

# Leptophilic DM

- $N(1,1,0,-)$ : Majorana fermion which we assume DM candidate
- $\varphi_l(1,2,-1/2,-)$ ,  $\varphi_e(1,1,-1,-)$ : Scalar doublets, and singlet

$$\begin{aligned}
 -\mathcal{L} = & -\mathcal{L}_{\text{SM}} + m_{\phi_l}^2 |\phi_l|^2 + m_{\phi_e}^2 |\phi_e|^2 + \frac{1}{2} M \bar{N} N \\
 & + \left( -y_{La} \bar{l}_a P_R N \tilde{\phi}_l + y_{Ra} \bar{e}_a P_L N \phi_e + h.c. \right) \\
 & + \left( -\mu H^\dagger \tilde{\phi}_l \phi_e^* + h.c. \right) + \lambda_{-1} |\phi_e|^2 |\phi_l|^2 + \lambda_0 |H|^2 |\phi_e|^2 + V_{2\text{HDM}},
 \end{aligned}$$

$$V_{2\text{HDM}} = \frac{\lambda_1}{2} |\phi_l|^4 + \frac{\lambda_2}{2} |H|^4 + \lambda_3 |\phi_l|^2 |H|^2 + \lambda_4 \left( \phi_l^\dagger H \right) \left( H^\dagger \phi_l \right) + \left( \frac{\lambda_5}{2} \left( \phi_l^\dagger H \right)^2 + h.c. \right)$$

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 & \quad \quad \quad \text{H} \rightarrow \tau \mu
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# Neutrino masses

- The model has all the ingredients for radiative neutrino mass generation a la Ma model

$$m_\nu \sim \lambda_5 \frac{y_{La}^2}{16\pi^2} \left( \frac{v}{m_{\phi_\ell}} \right)^2 M.$$

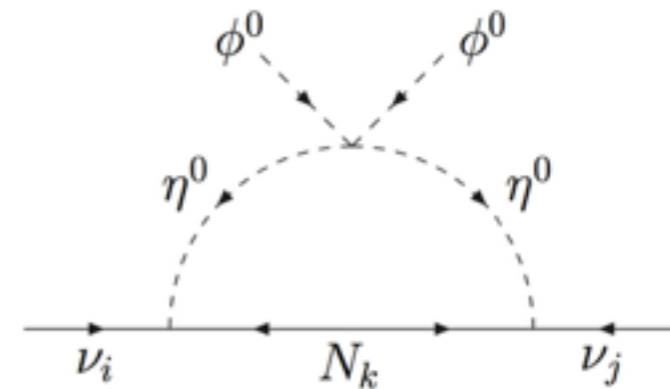


FIG. 5: One-loop scotogenic neutrino mass.

Ma, 0905.0221

# Mass spectrum

- Charged scalars

mixing between  $\phi_I, \phi_e$  after ew symmetry breaking

$$\phi_\ell = (\phi_\ell^+, (\phi_R + i\phi_I)/\sqrt{2})^T$$

$$\tilde{e}_1 = \cos \theta (\phi_\ell^+)^* - \sin \theta \phi_e, \quad \tilde{e}_2 = \sin \theta (\phi_\ell^+)^* + \cos \theta \phi_e,$$

$$m_{\tilde{e}_{1,2}}^2 = \frac{1}{2} \left[ (m_{\phi_\ell}^2 + m_{\phi_e}^2) \mp \sqrt{(m_{\phi_\ell}^2 - m_{\phi_e}^2)^2 + 2\mu^2 v^2} \right]$$

- Neutral scalars

$$m_{\phi_R}^2 \approx m_{\phi_\ell}^2 + (\lambda_4 + \lambda_5) v^2 / 2, \quad m_{\phi_I}^2 \approx m_{\phi_\ell}^2 + (\lambda_4 - \lambda_5) v^2 / 2.$$

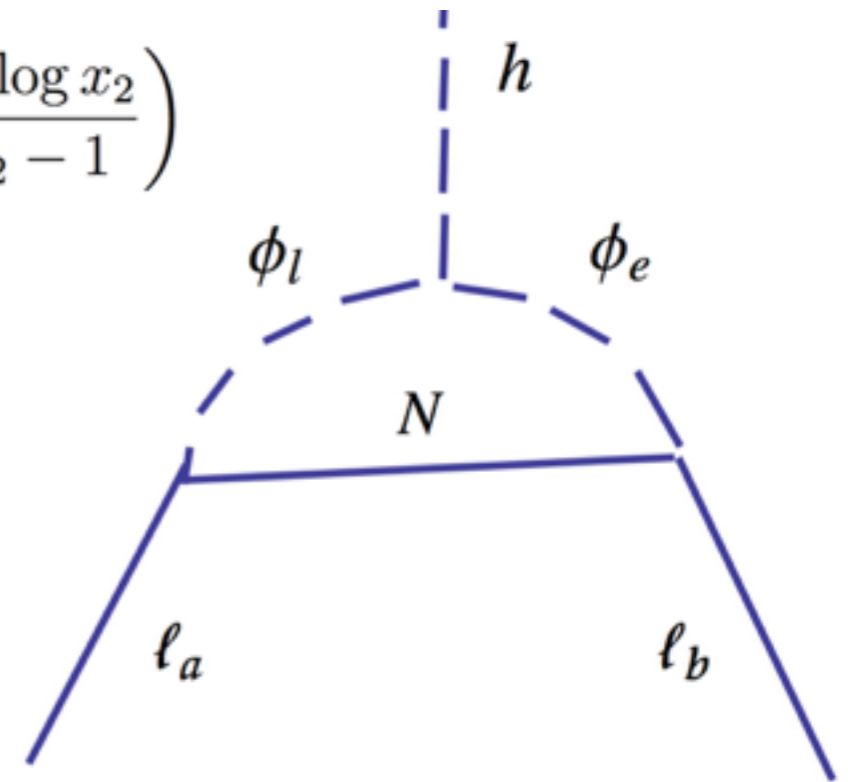
# LFV Higgs decay, $H \rightarrow \tau\mu$

$$i\mathcal{M} = +i\bar{u}_b (F_L P_L + F_R P_R) v_a$$

$$F_L \simeq \frac{1}{16\pi^2} \frac{\mu}{\sqrt{2}M} y_{Rb} y_{La}^* \left[ \frac{1}{2} \sin^2 2\theta (G(x_1) + G(x_2)) + \cos^2 2\theta G(x_1, x_2) \right]$$

$$x_i \equiv m_{\tilde{e}_i}^2 / M^2 \quad G(x_1, x_2) \equiv \frac{1}{x_1 - x_2} \left( \frac{x_1 \log x_1}{x_1 - 1} - \frac{x_2 \log x_2}{x_2 - 1} \right)$$

$$\Gamma(h \rightarrow \bar{\ell}_a \ell_b) = \frac{m_h}{16\pi} (|F_L|^2 + |F_R|^2)$$



Large  $\mu$  can enhance the LFV H decay both in the decoupling regime (eg.  $\varphi_l \ll \varphi_e$  and  $\theta \ll 1$ ) or in the large mixing regime ( $\theta \approx \pi/4$ )

# LFV Higgs decay, $H \rightarrow \tau\mu$

- Decoupling limit

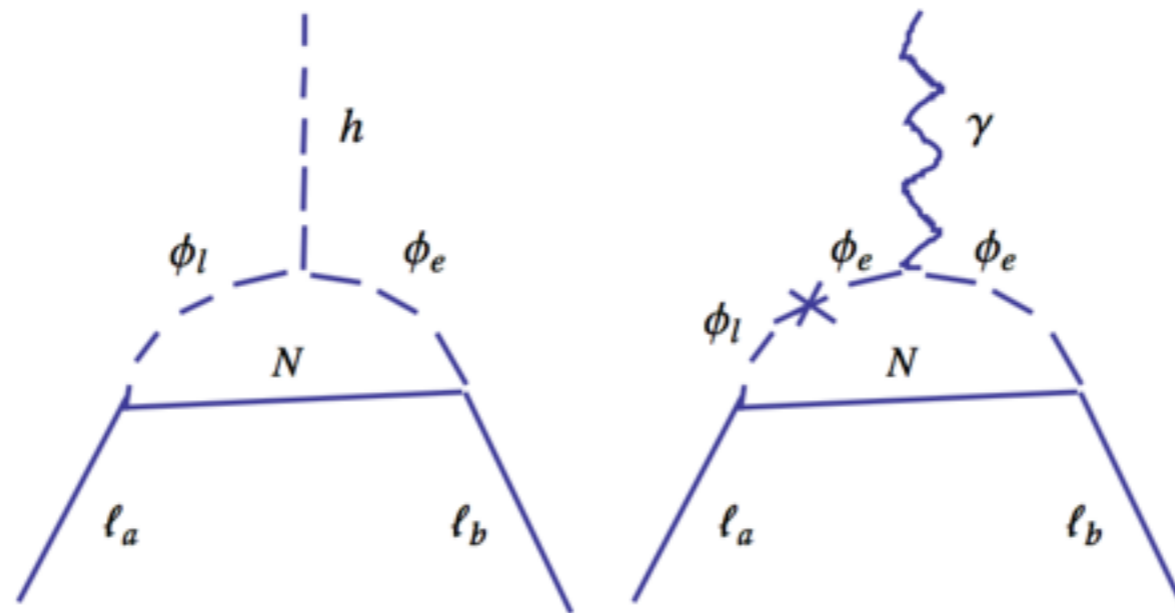
$$\text{Br}(h \rightarrow \bar{\tau}\mu) = 1.2 \times 10^{-2} \left(\frac{\mu}{5\text{TeV}}\right)^2 \left(\frac{1\text{TeV}}{M}\right)^2 \left(\frac{G(x_1, x_2)}{0.2}\right)^2 \left(\frac{|y_{R\tau}y_{L\mu}^*|}{1}\right)^2$$

- Maximal mixing

$$\text{Br}(h \rightarrow \bar{\tau}\mu) = 1.2 \times 10^{-2} \left(\frac{\mu}{10\text{TeV}}\right)^2 \left(\frac{1\text{TeV}}{M}\right)^2 \left(\frac{G(x_1) + G(x_2)}{0.4}\right)^2 \left(\frac{|y_{R\tau}y_{L\mu}^*|}{0.5}\right)^2$$

# $\tau \rightarrow \mu \gamma$ constraint

- Generally very stringent in models with large  $H \rightarrow \tau \mu$



- We can evade the  $\tau \rightarrow \mu \gamma$  constraint in the decoupling regime without fine-tuning

# $\tau \rightarrow \mu \gamma$ constraint

- $B(\tau \rightarrow \mu \gamma) < 4.4 \times 10^{-8}$

$$\mathcal{H}_{\text{eff}} = C_L \bar{\mu}_L \sigma^{\mu\nu} \tau_R F_{\mu\nu} + C_R \bar{\mu}_R \sigma^{\mu\nu} \tau_L F_{\mu\nu}.$$

$$C_L = \frac{e}{32\pi^2 M} y_{L\mu} y_{R\tau}^* s_\theta c_\theta (F_2(x_1) - F_2(x_2))$$

$$F_2(x) = \frac{-1 + x^2 - 2x \log x}{2x(1-x)^2}$$

- $\tau \rightarrow \mu \gamma$  is easily suppressed in the decoupling regime,  $C_L \rightarrow 0$  if  $\theta \rightarrow 0$ .
- Fine-tuning is required in the maximal mixing case

# $h \rightarrow \gamma\gamma$ constraint

$$\mathcal{L}_{\text{eff}} = c_\gamma \frac{\alpha}{\pi v} h F_{\mu\nu} F^{\mu\nu}$$

$$r_\gamma = r_{\text{SM},\gamma} + \delta r_\gamma \approx -0.81 + \frac{1}{24} \frac{v\mu \sin 2\theta}{2m_{\tilde{e}_1}^2}$$

$$-0.05 \lesssim \delta r_\gamma / r_{\text{SM},\gamma} \lesssim 0.20, \quad -2.20 \lesssim \delta r_\gamma / r_{\text{SM},\gamma} \lesssim -1.95 \quad \text{at 68.3\% C.L.}$$

$$-2.8 \times \left( \frac{m_{\tilde{e}_1}}{300 \text{GeV}} \right)^2 \text{TeV} \lesssim \mu \sin 2\theta \lesssim 0.7 \times \left( \frac{m_{\tilde{e}_1}}{300 \text{GeV}} \right)^2 \text{TeV},$$

$$28 \times \left( \frac{m_{\tilde{e}_1}}{300 \text{GeV}} \right)^2 \text{TeV} \lesssim \mu \sin 2\theta \lesssim 31 \times \left( \frac{m_{\tilde{e}_1}}{300 \text{GeV}} \right)^2 \text{TeV},$$



# Relic density of DM

- The relic density is achieved by the annihilation of DM,  $N$ , into lepton pair via t-channel slepton exchange.

$$\sigma v_r \approx a + bv_r^2 \qquad \Omega h^2 \approx \frac{0.88 \times 10^{-10} x_f \text{GeV}^{-2}}{g_*^{1/2} (a + 3b/x_f)}$$

$$a = \frac{1}{16\pi M^2} \frac{1}{(1+x_i)^2} (|\lambda_{ia}^L \lambda_{ib}^R|^2 + |\lambda_{ia}^R \lambda_{ib}^L|^2),$$

$$b = \frac{1}{192\pi M^2} \frac{1}{(1+x_i)^4} [|\lambda_{ia}^L \lambda_{ib}^L|^2 (1-6x_i+x_i^2) + |\lambda_{ia}^L \lambda_{ib}^R|^2 (1-2x_i) + (L \leftrightarrow R)]$$

$$\lambda_{1a}^L = -\sin \theta y_{Ra}, \quad \lambda_{2a}^L = \cos \theta y_{Ra}; \quad \lambda_{1a}^R = \cos \theta y_{La}, \quad \lambda_{2a}^R = \sin \theta y_{La}$$

# Relic density of DM

- If  $y_L=0$ ,  $a=0$ . [M. Garny, A. Ibarra, S. Vogl, 1503.01500](#)

The annihilation cross section is p-wave suppressed, and EW scale slepton is required to get the correct relic density.

- In our case, s-wave annihilation is allowed

$$a \approx \frac{1}{64\pi M^2} \sin^2 2\theta (|y_{La}y_{Rb}|^2 + |y_{Ra}y_{Lb}|^2) \sum_i \frac{1}{(1+x_i)^2}$$

$$a \simeq 0.8 \times \left(\frac{400 \text{ GeV}}{M}\right)^2 \left(\frac{\sin^2 \theta}{0.1}\right) \frac{(|y_{La}y_{Rb}|^2 + |y_{Ra}y_{Lb}|^2)}{1.0} \text{pb}$$

and the correct relic density is obtained in the decoupling regime

# Direct DM detection

- DM-nucleon scattering is absent at tree-level.
- At one-loop, DM can interact with nucleon via Higgs or photon exchange.

$$\mathcal{O}_h = \lambda_{hN}(0) h \bar{N} N$$

$$\mathcal{O}_A = \mathcal{A} \bar{N} \gamma^\mu \gamma^5 N \partial^\nu F_{\mu\nu}$$

$$\lambda_{hN}(0) \approx 0.01 \times \left(\frac{\sin \theta}{0.2}\right) \left(\frac{|y_{La}|^2 + |y_{Ra}|^2}{1}\right) \left(\frac{\mu}{5\text{TeV}}\right) \left(\frac{0.3\text{TeV}}{M}\right) \mathcal{A} / (|\lambda_{ia}^L|^2 + |\lambda_{ia}^R|^2) \sim \mathcal{O}(10^{-7}) \text{GeV}^{-2}$$

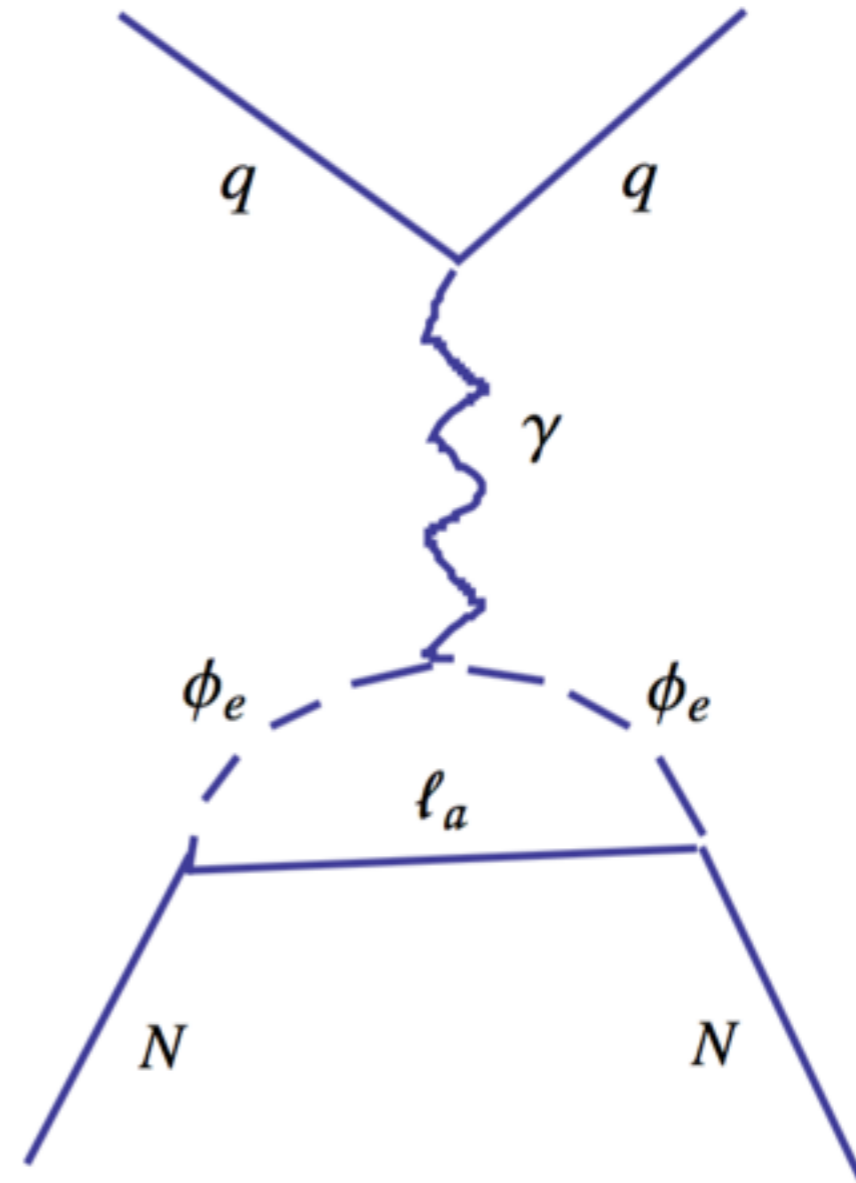
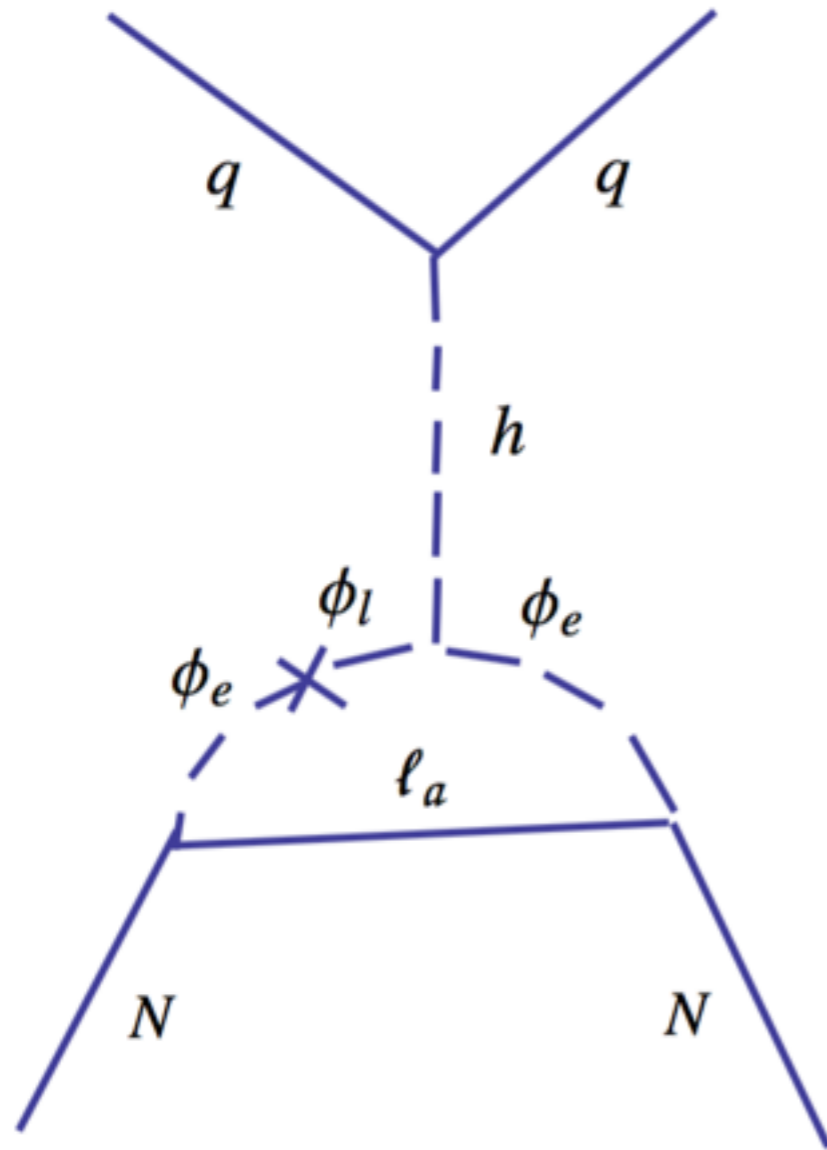
$$O(10^{-2})$$

$$O(10^{-4})$$

of current LUX bound

$$\sigma_{\text{SI}}^p \approx 4.0 \times 10^{-8} \text{pb}$$

# Direct DM detection



# Conclusions

- Leptophilic dark matter model with coupling to both  $SU(2)_L$  doublet and singlet scalars can explain the large  $h \rightarrow \tau\mu$  signal
- $\tau \rightarrow \mu\gamma$ ,  $h \rightarrow \gamma\gamma$  constraints can be evaded easily in the decoupling regime
- The correct relic density can be obtained. Direct detection cross section is typically two-orders of magnitude smaller than the current LUX bound

**Thank you very much!**