

Connecting Large Scale Structures to Galaxy morphology

Can we predict the morphology of galaxies **on** the cosmic web **from first principles**?

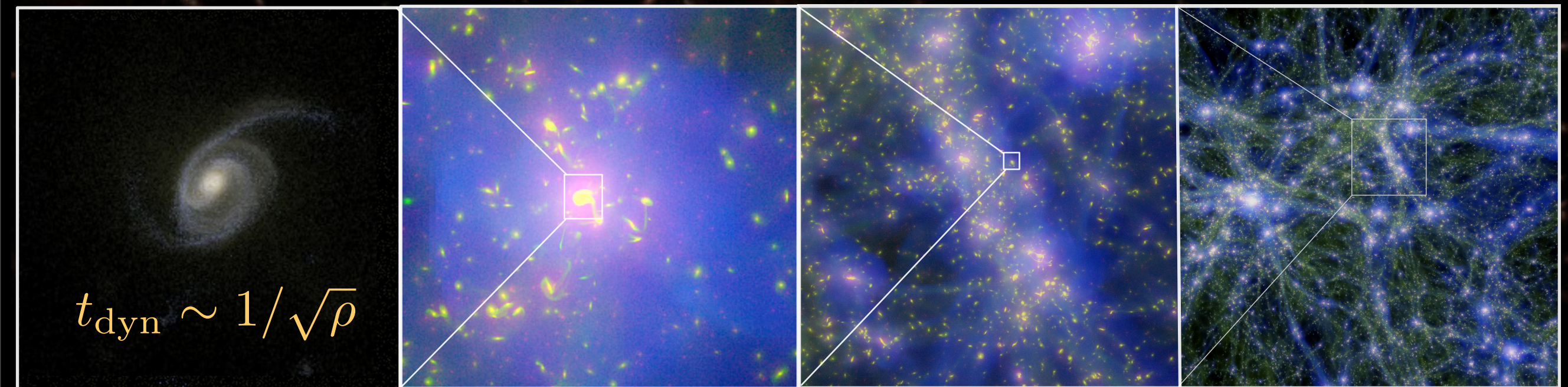
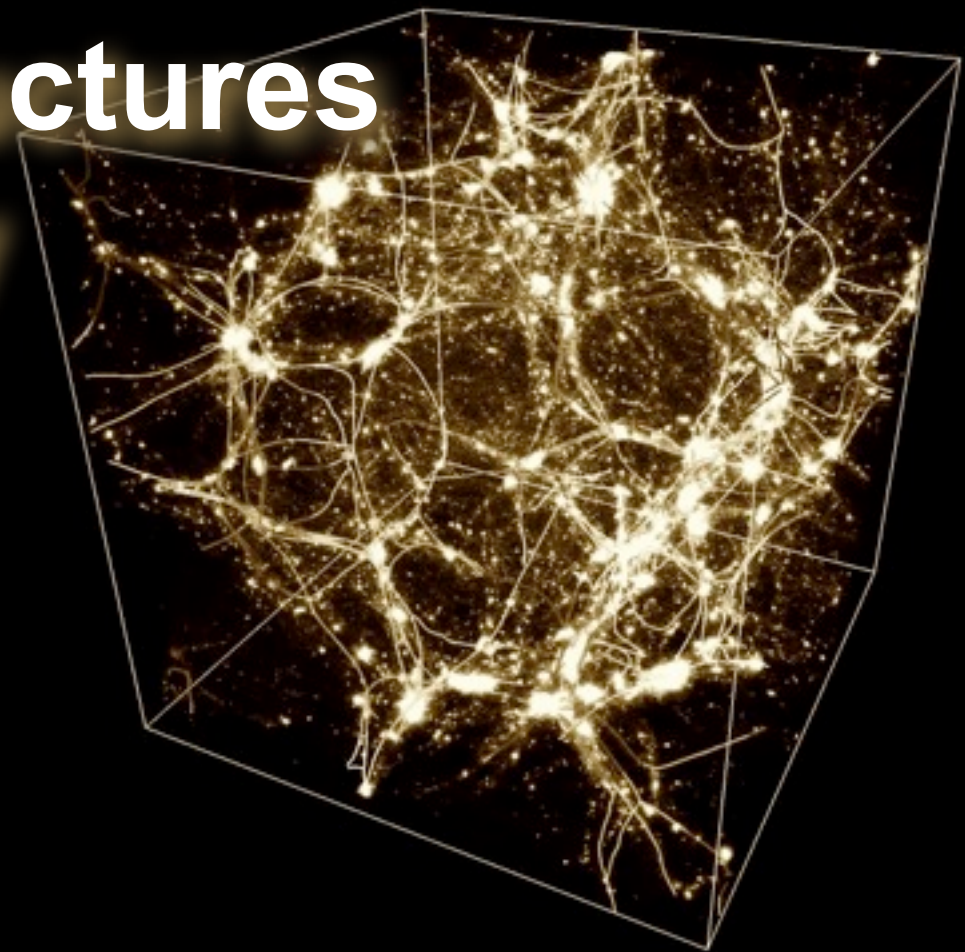
Is the cosmic web **driving** the Hubble sequence?

Christophe Pichon

Institut d'astrophysique de paris

S. Codis, C. Laigle, C. Welker T. Kimm D. Pogosyan, J. Devriendt, Y Dubois+ Horizon Collaboration

MareNostrum z=1.55



Outline

- How discs build up from persistent cosmic web?
- How dark halo's spin flip relative to filament?
- Why are they initially aligned with filaments?
Why the transition mass? **Eulerian** view
- What is the corresponding **Lagrangian** theory?

Outline

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Galactic morphology is driven by AM acquisition through anisotropic secondary infall, coming from larger scales, which are less dense, hence more steady; cold flows provide the link.

Outline

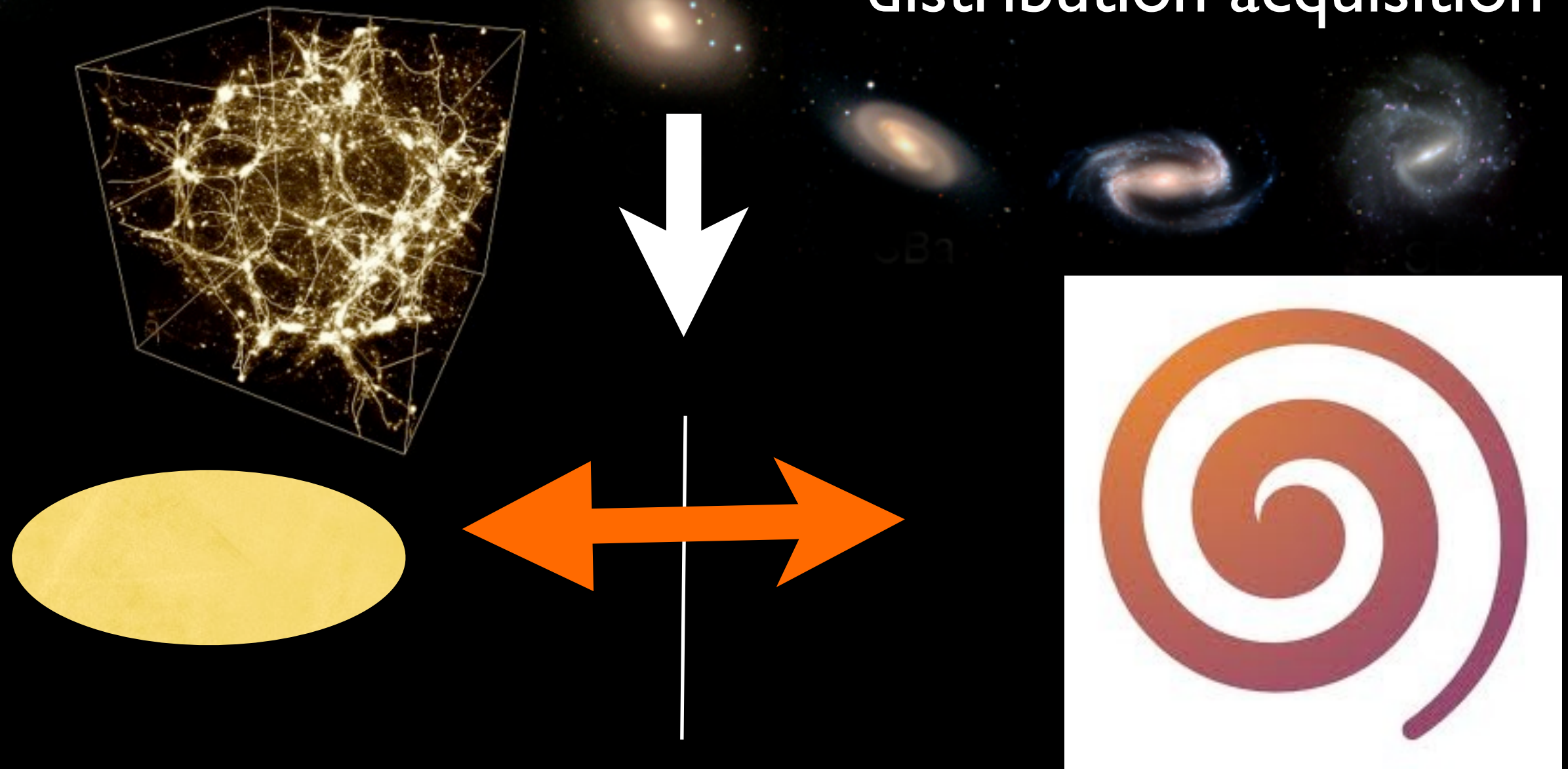
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- **What is the corresponding *Lagrangian* theory?**

Galactic morphology is driven by AM acquisition through anisotropic secondary infall, coming from larger scales, which are less dense, hence more steady; cold flows provide the link.

*Where galaxies form does matter, and can be traced back to ICs
Flattened filaments generate point-reflection-symmetric AM/vorticity distribution
they induce the observed spin transition mass & the helicity of cold flows*

The Hubble diagram: a crude theorist's view

driven by
angular momentum
distribution acquisition



What drives coherent secondary infall?

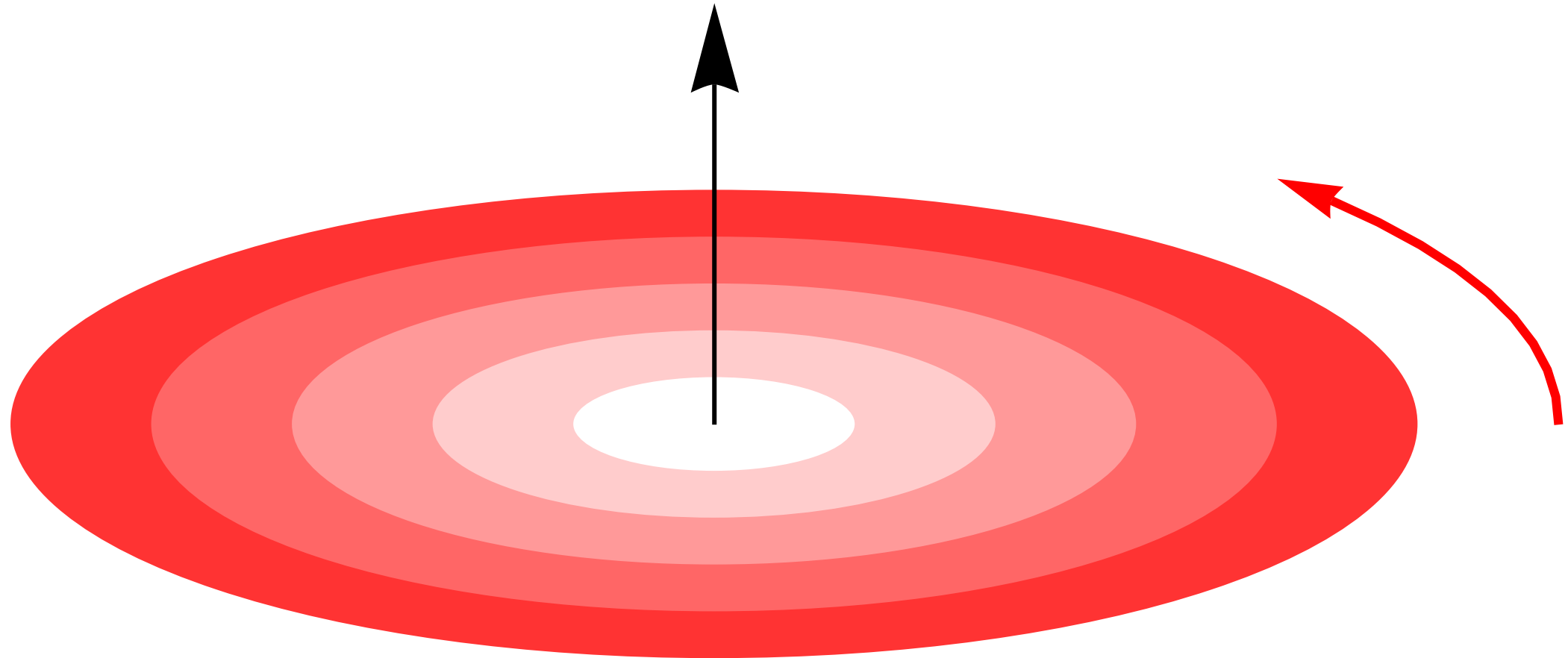
- **How discs build up from persistent cosmic web?**

Part I Outline

- 4 trivial facts about galaxies in their web
 - what's a disc?
 - what's a void?
 - what's a shock?
 - what do numerical hydro suggest?
- The proposition
- Various proofs of various value?

Fact number one

“theoretically”, a galactic disc:



An ensemble of ring made of gas,

- turning around the same axis
- whose outer parts rotate with **more** angular momentum (flat rotation curve)

Fact number **two**

The *Virtual (dark matter)* universe

Voids become more void

Filaments drifts...

Log density



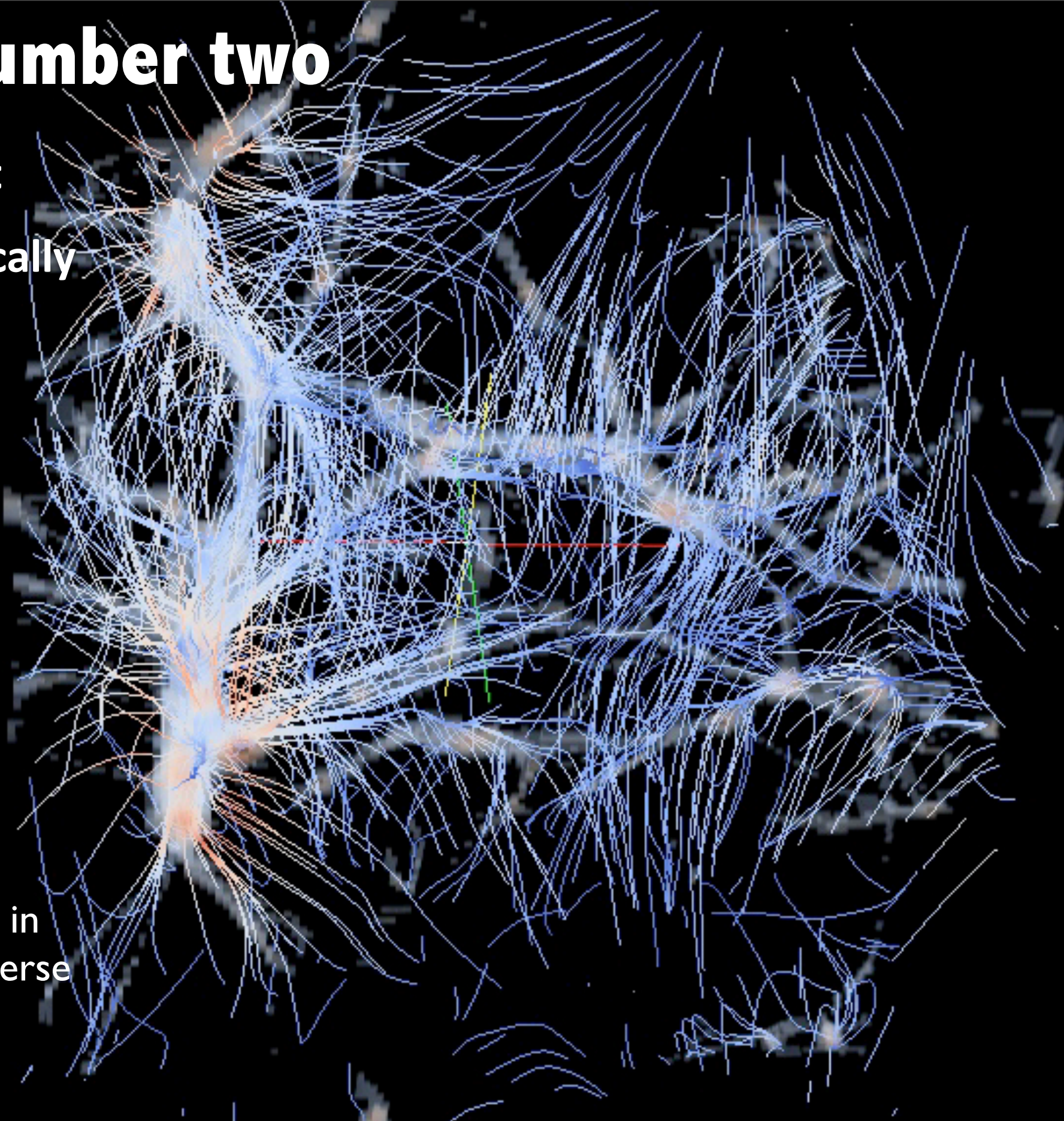
... and get **distorted**

$$t_{\text{dyn}} \sim 1/\sqrt{\rho}$$

not much happens on LS: which is good & expected

Fact number two

Peak attract
catastrophically



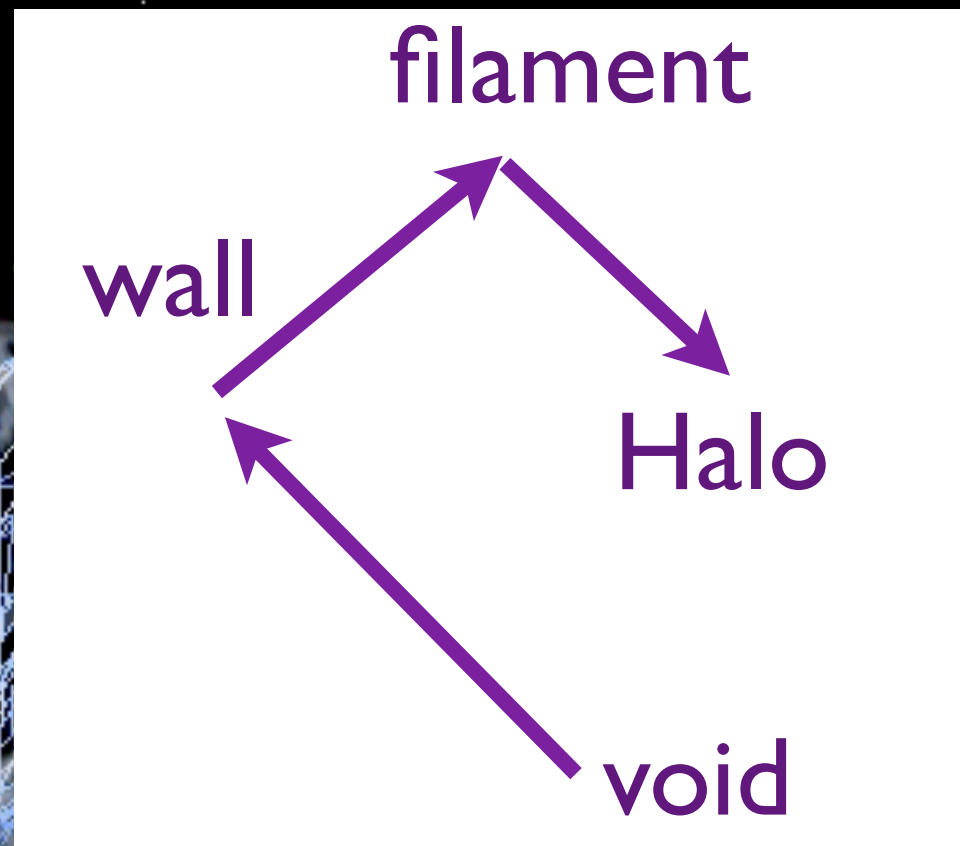
Velocity flow in
expanding universe

BUT surrounding void repel (contrast<0) & contribute to secondary infall.

Fact number two

Peak attract
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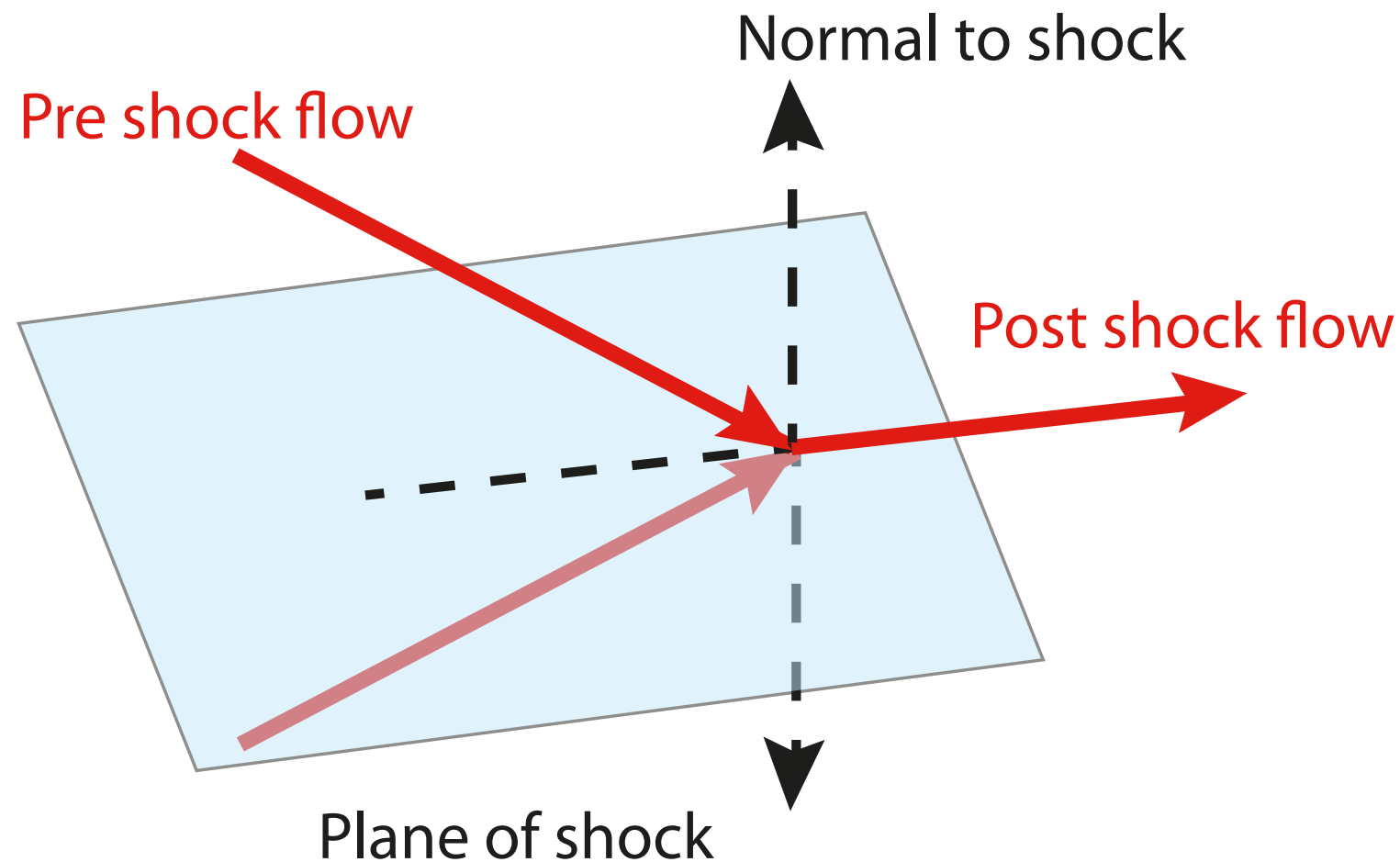
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Fact number **three**

“theoretically”, a shock:

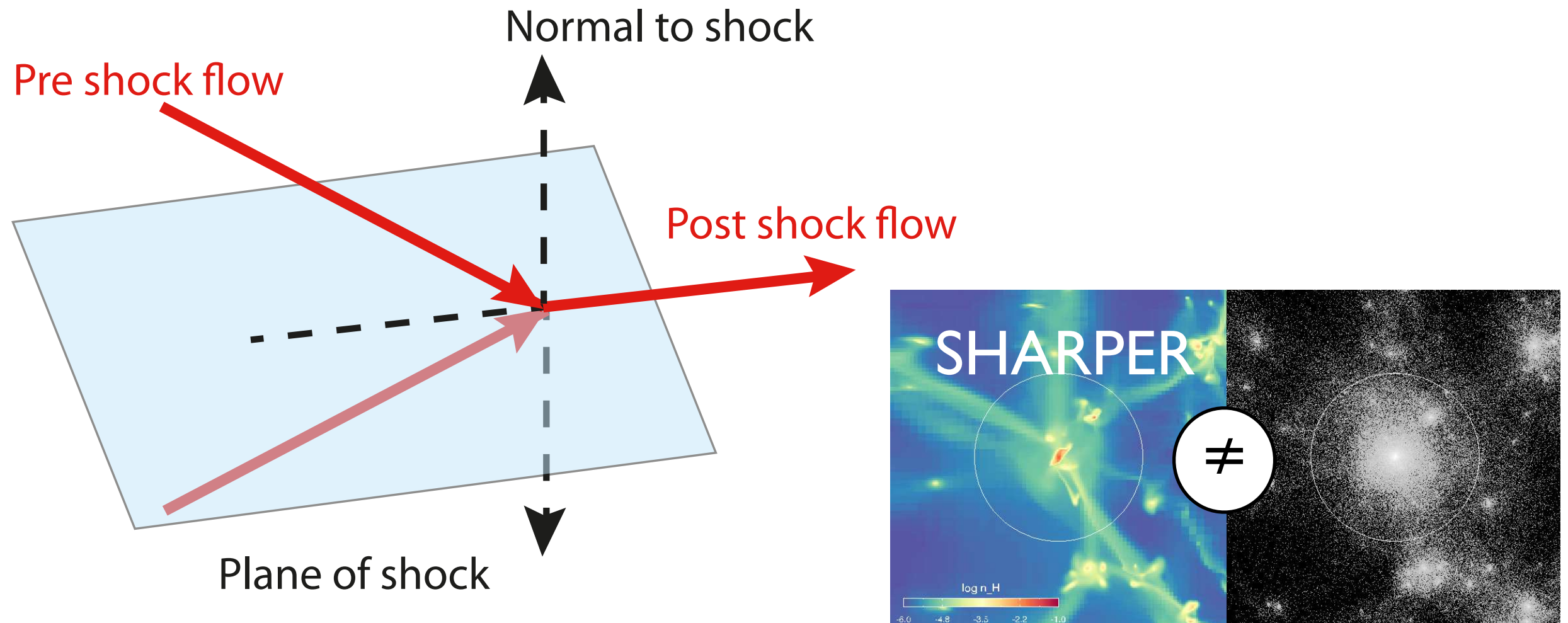


Gas, unlike dark matters, shocks (iso-T) and
follows closely the cosmic web

→ cosmic web is important for galaxy morphology

Fact number **three**

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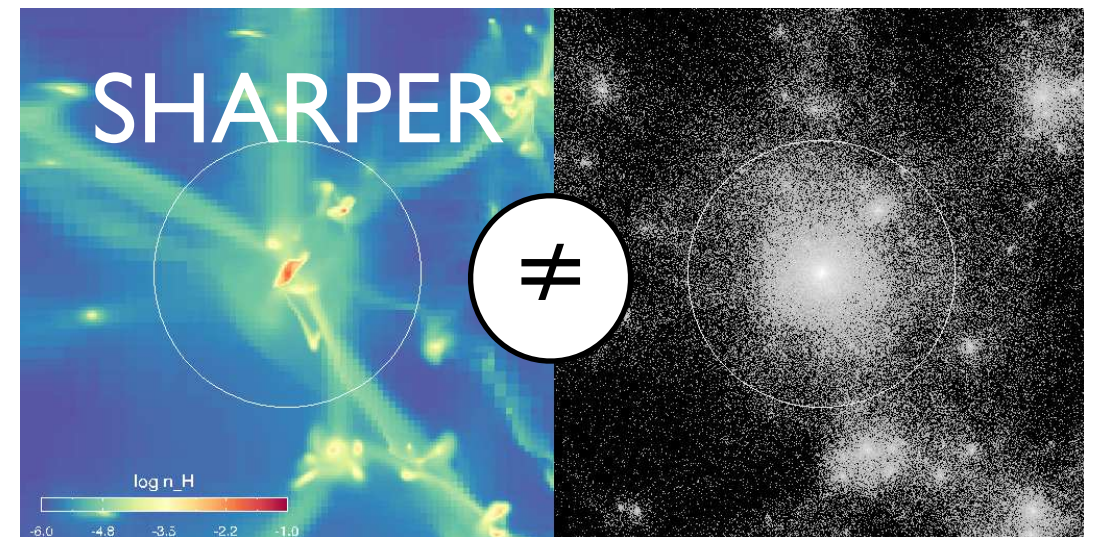
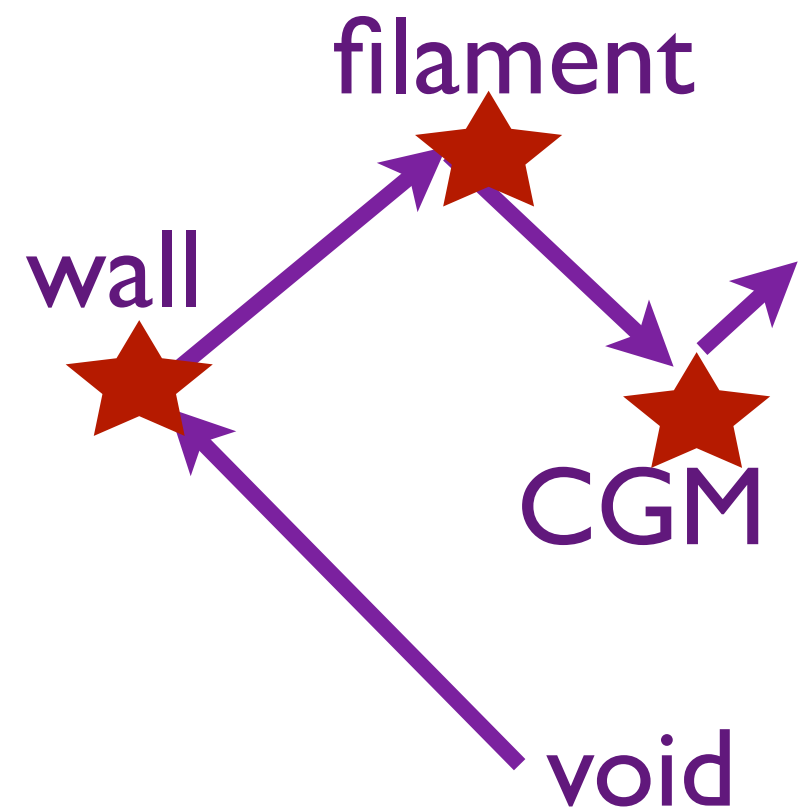
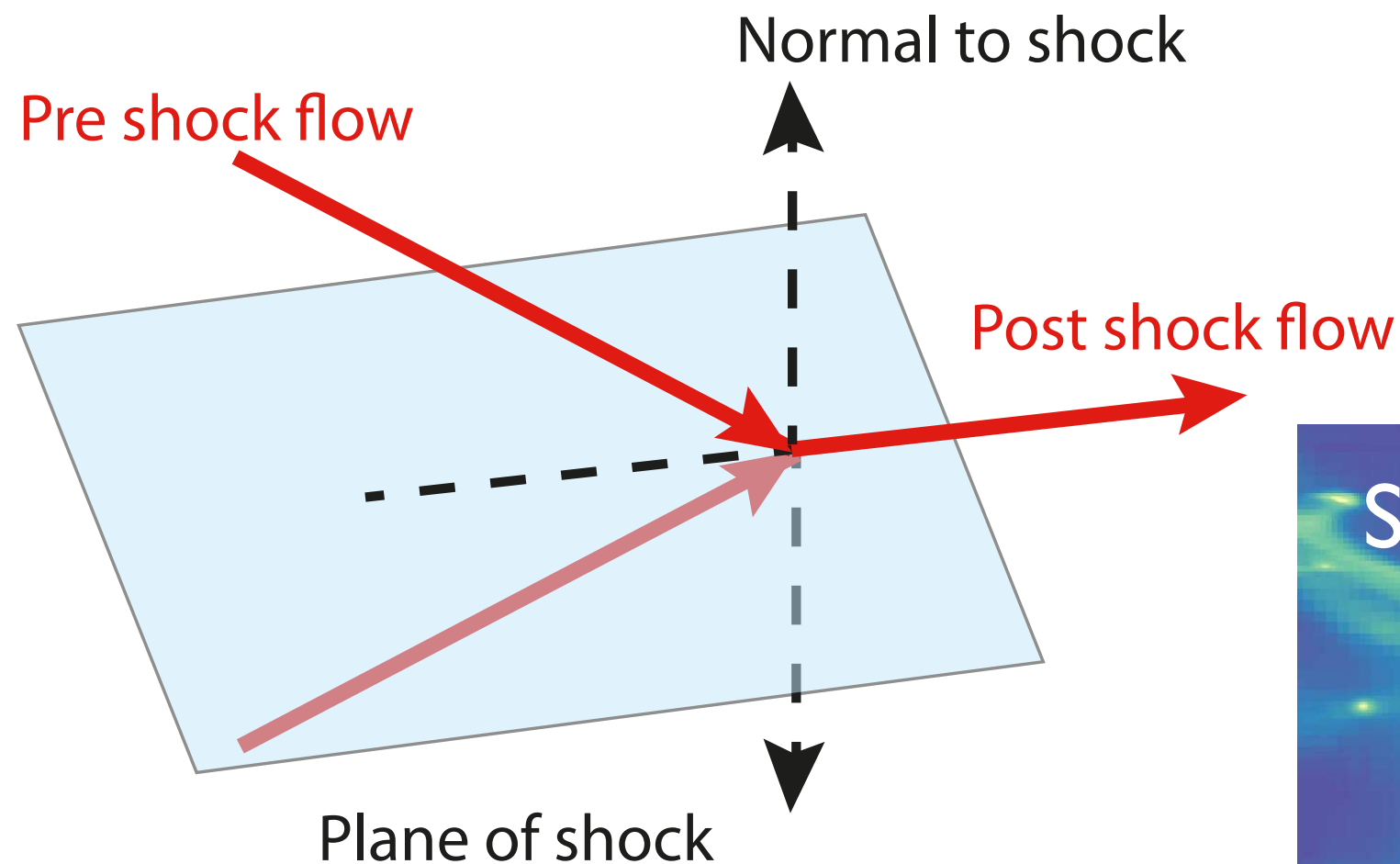


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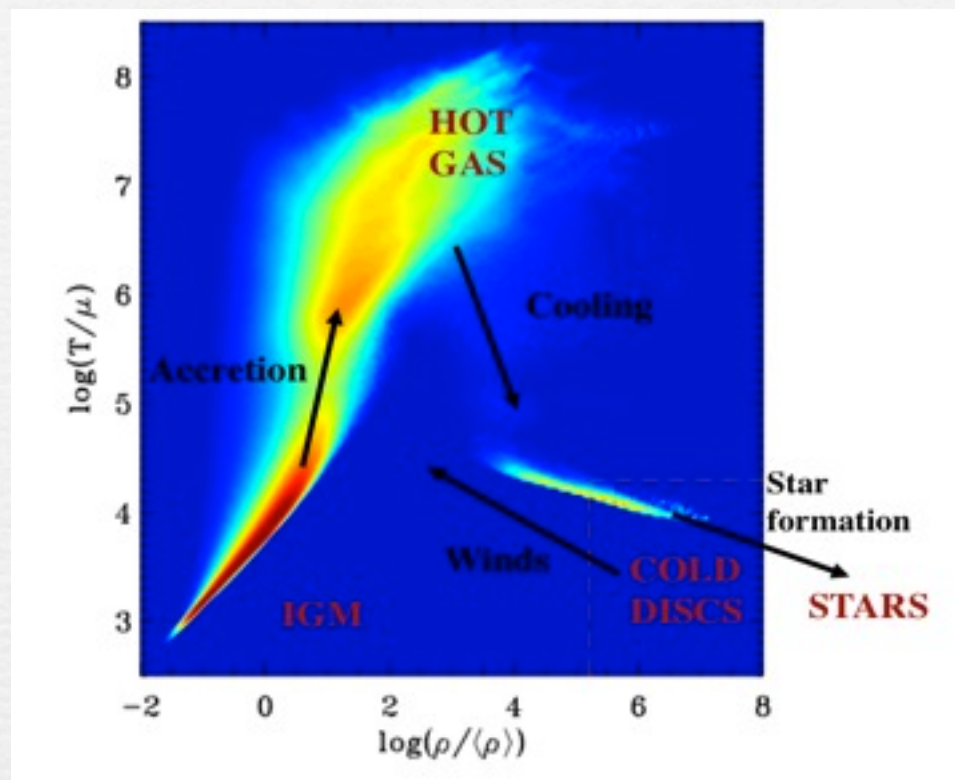
“theoretically”, a shock:



Gas, unlike dark matters, shocks (iso-T) and **follows closely the cosmic web**

→ cosmic web is important for galaxy morphology

@ high z / low mass



Paradigm
shift

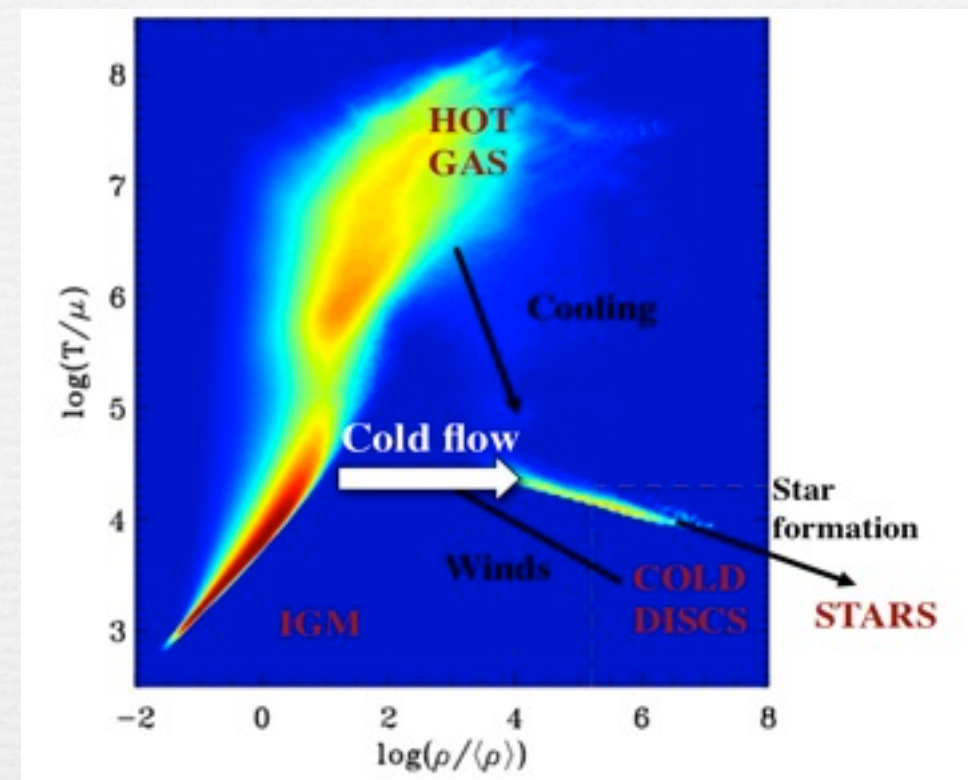


cf Binney 77 !!

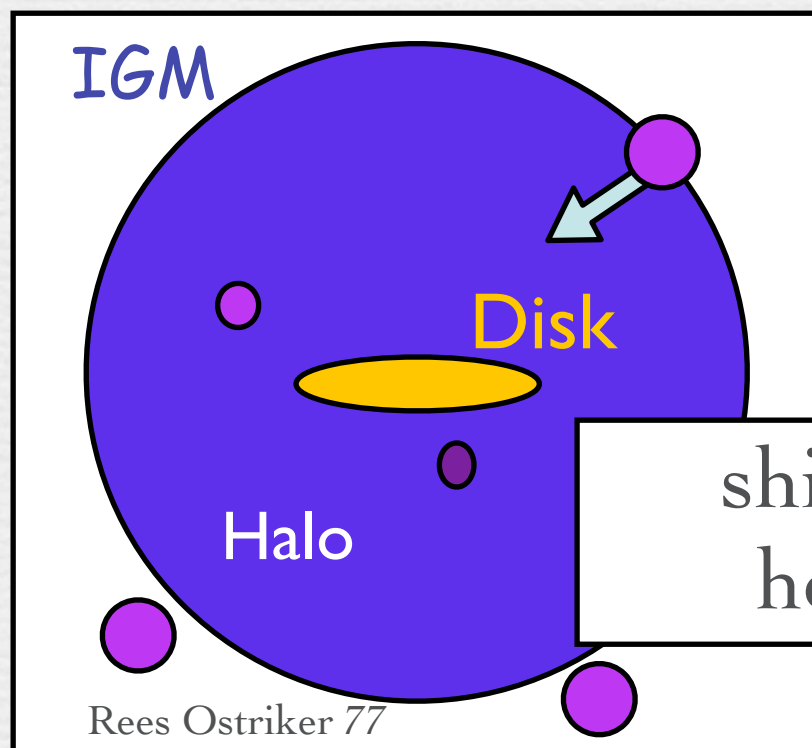
Gas shocks

isothermally

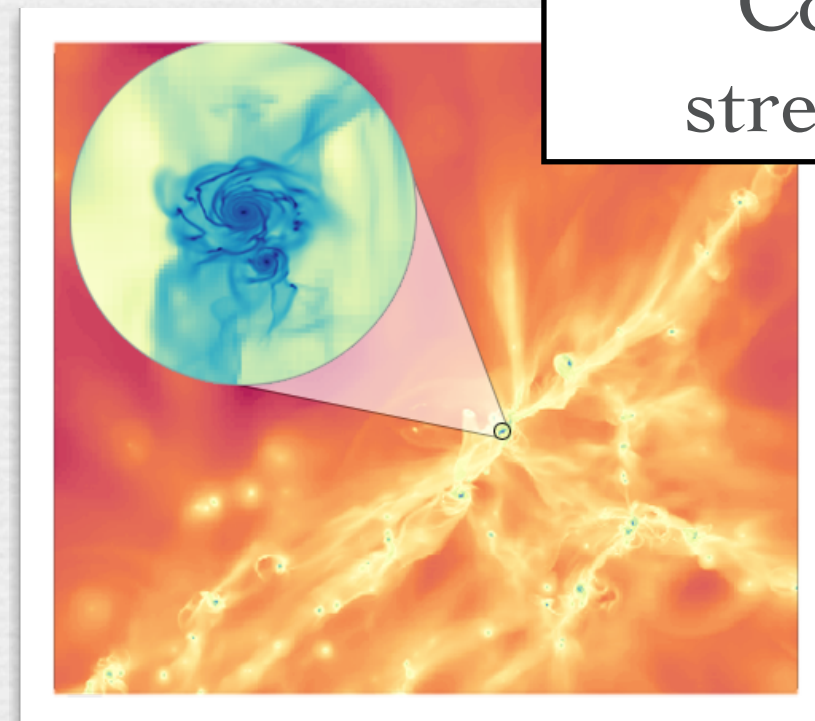
in LSS



« Interface » @ Virial Radius



shielding (?)
hot corona



Cold
streams

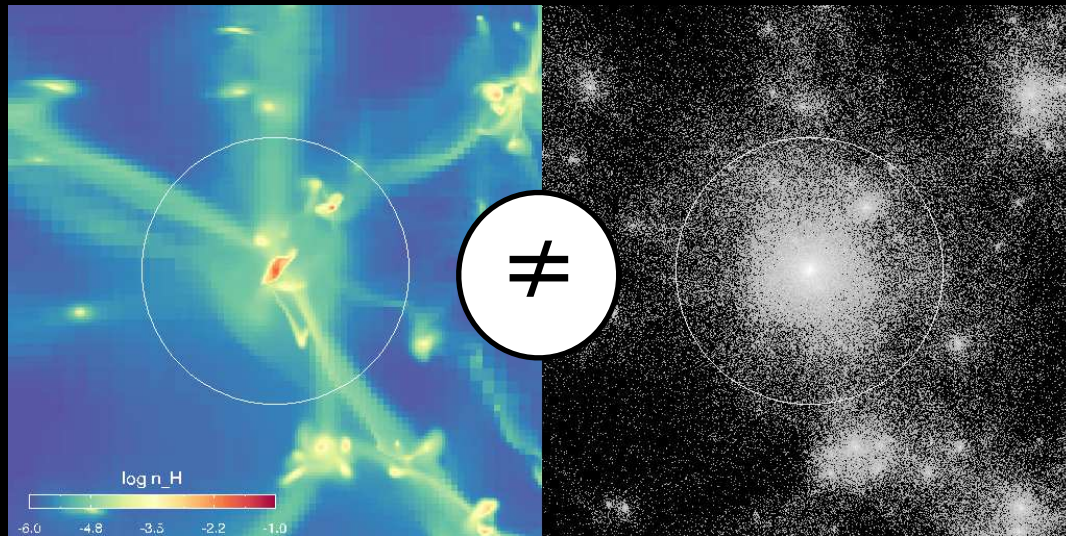
Katz 2002

@ high redshift

Fact number **four**

The *Virtual (hydrodynamical)* universe

$z=99.00$



Cosmic web SHARPER

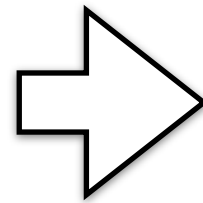
2 kpc

Agertz et al. (2009)

we see cold flows + recurrent disk reformation
LSS drives secondary infall & SPIN ALIGNMENT

Context & clues

standard
hierarchical
clustering picture



halo rigging



aka
completely
useless
(nautical) analogy
that probably only
the author
understands

- ✓ disc must have a coherent stratified angular momentum
- ✓ surrounding void/wall repel ($\text{contrast} < 0$) contribute to secondary infall
- ✓ gas shocks isothermally during shell crossing, follows filaments closely
- ✓ there are discs on the sky and in numerical simulations
- ✓ galaxies form and evolve on the cosmic web (anisotropic PBS)

Part I Outline

- 4 trivial facts about galaxies in their web
- **the proposition**
- various proofs of various value?

The proposition in one sentence

Disks form because LSS are large (*dynamically young*) and (*partially*) an-isotropic :
they induce persistent angular momentum
advection of cold gas **along** filaments
which stratifies
accordingly so as to (re)build discs
continuously.

Part I Outline

- 4 trivial facts about galaxies in their web
- The proposition
- **Various proofs of various value?**
 - smoking gun?
 - lots of hand waving ??
 - robust statistics?

Clues from LSS

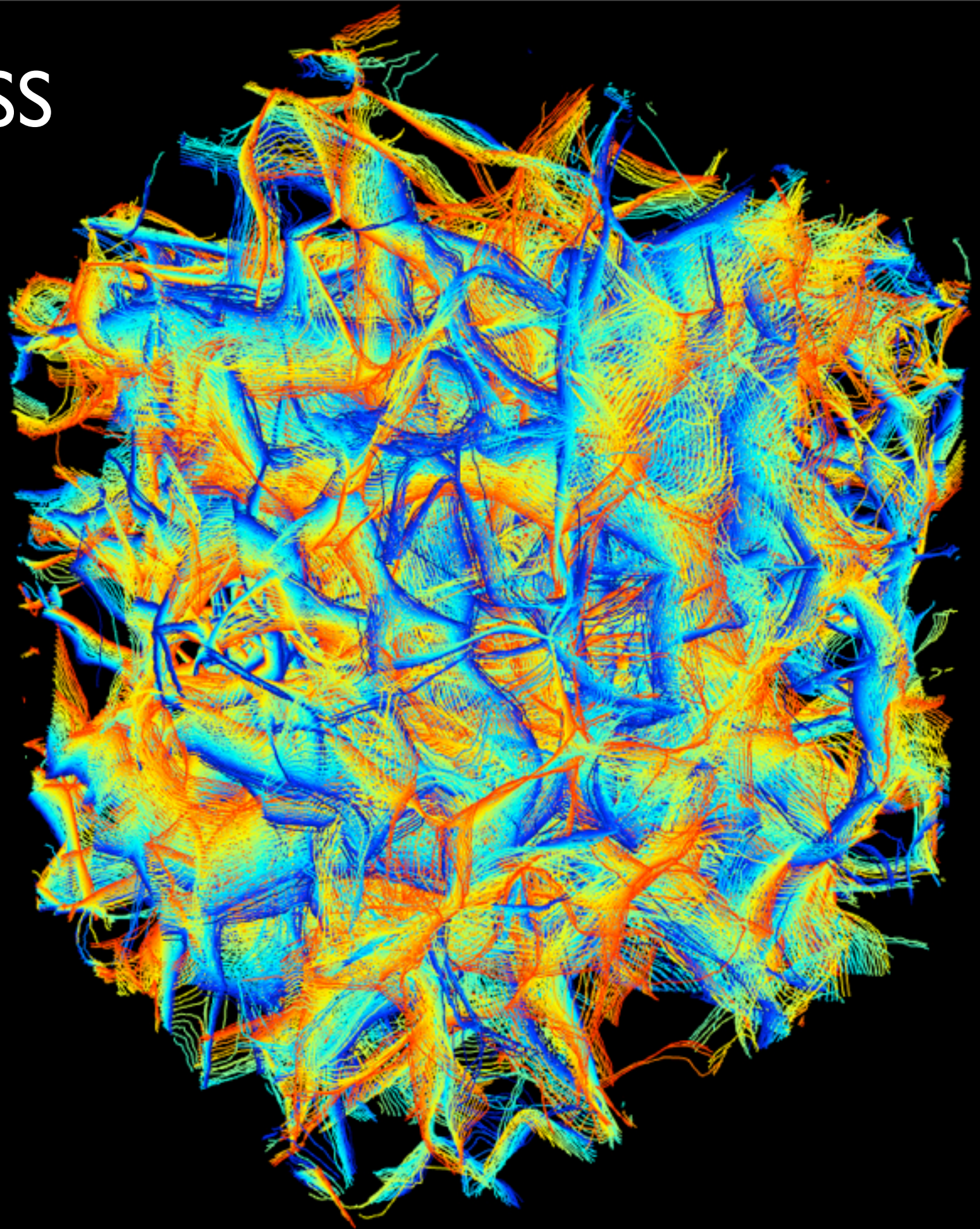
"Proof by halo centric environment"

a.k.a

proof by hypnosis,
fishy analogy &
mathematical jargon

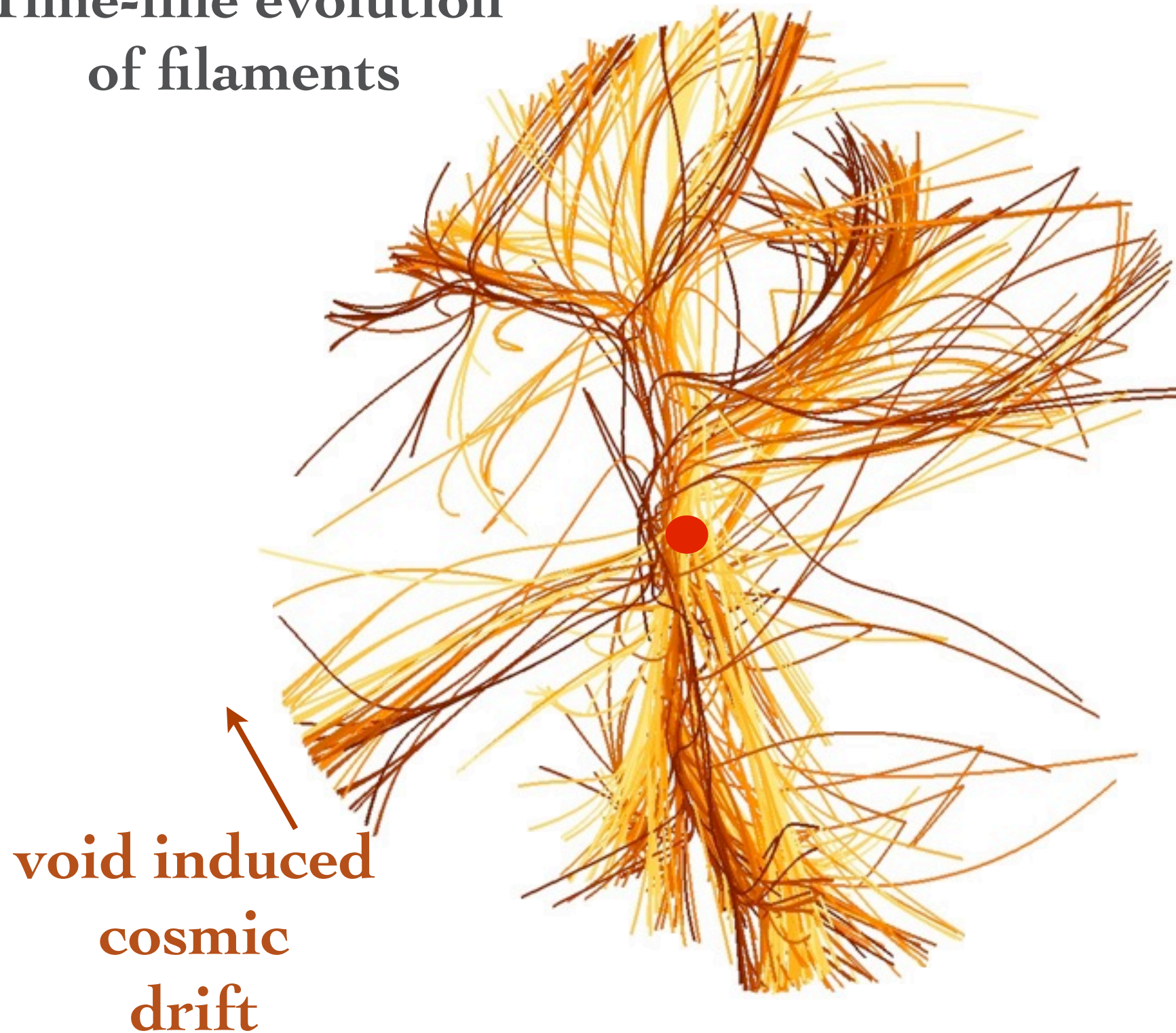
Time line of LSS

full history of universe
at once !



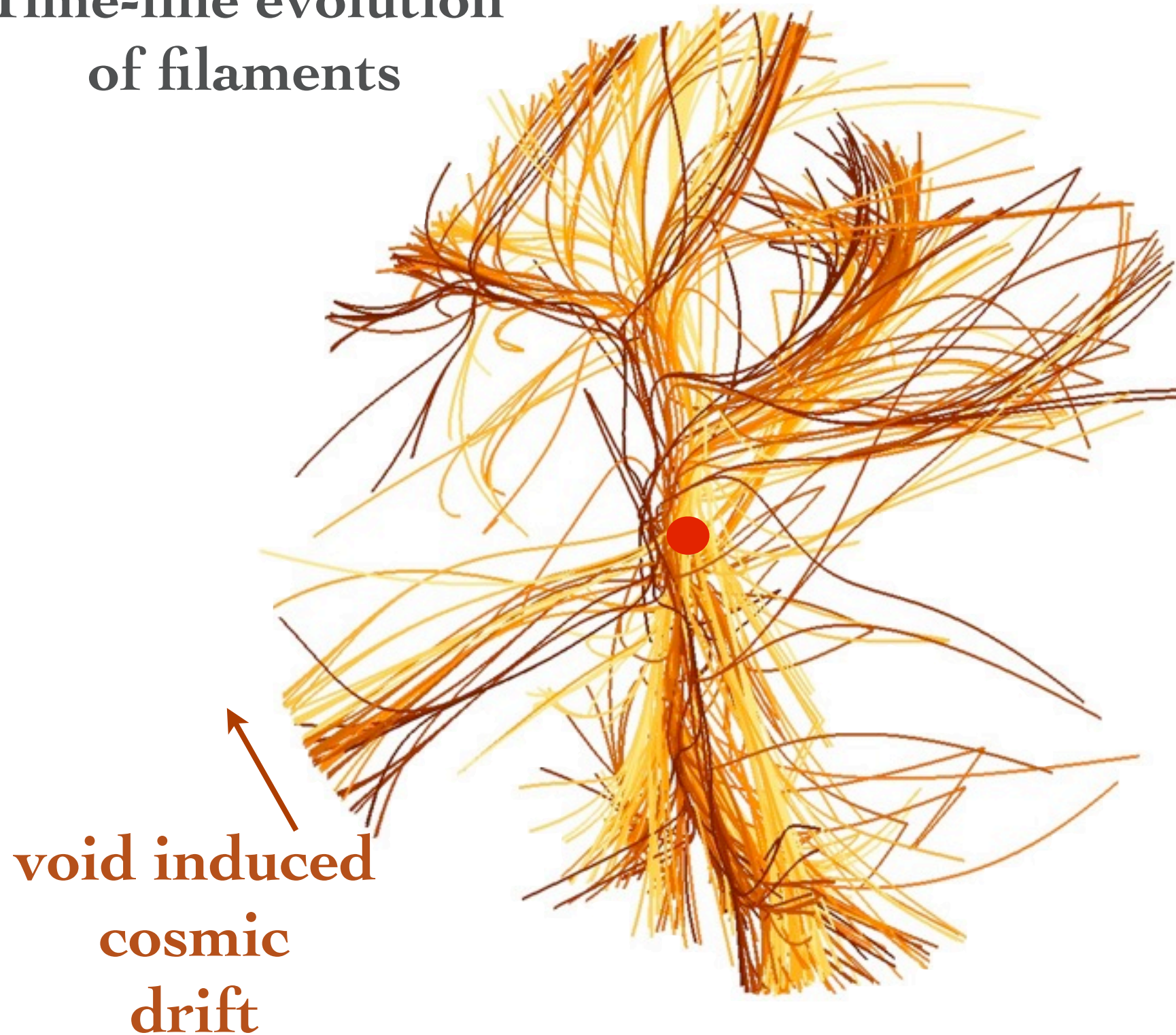
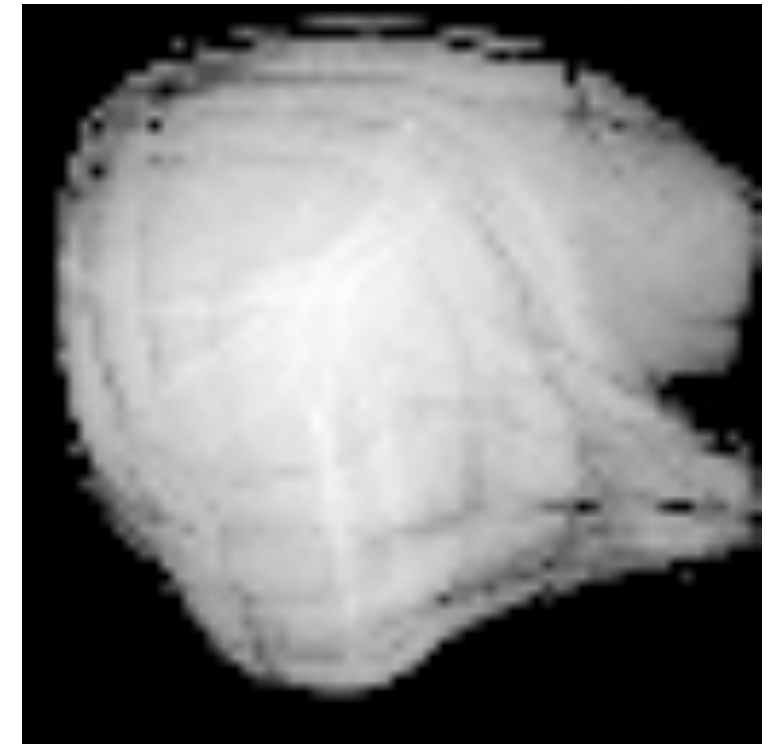
Drift of filaments

Time-line evolution
of filaments



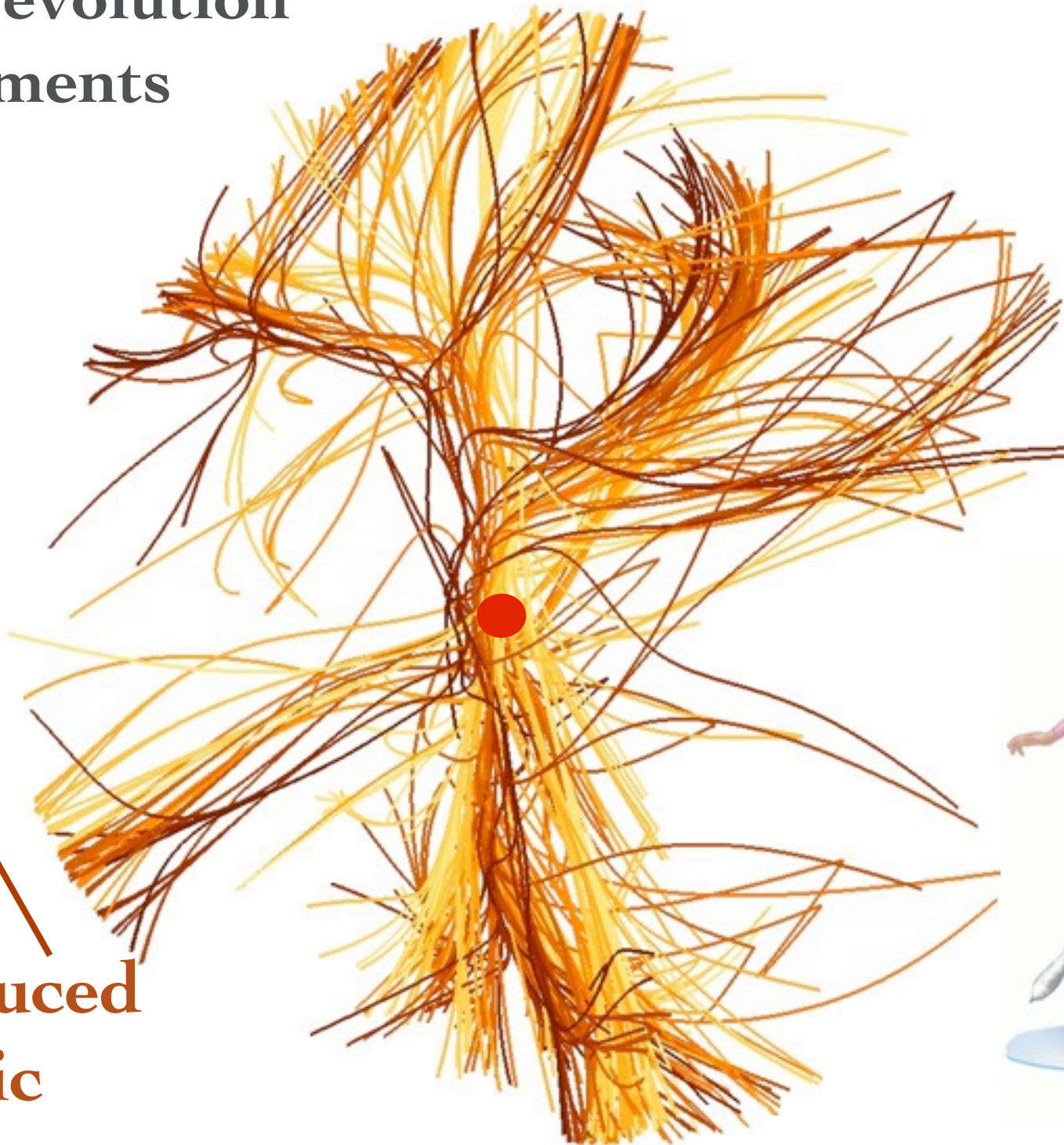
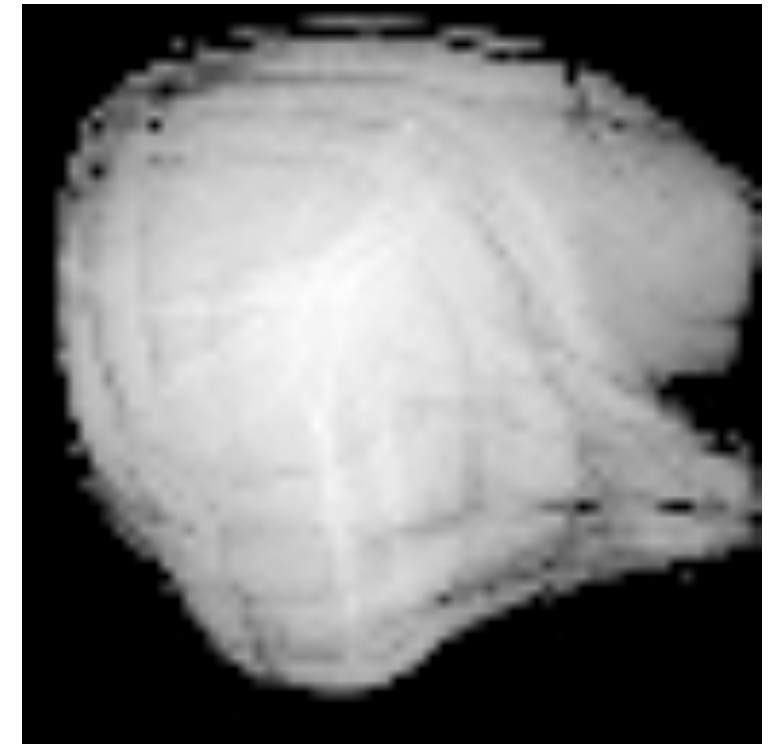
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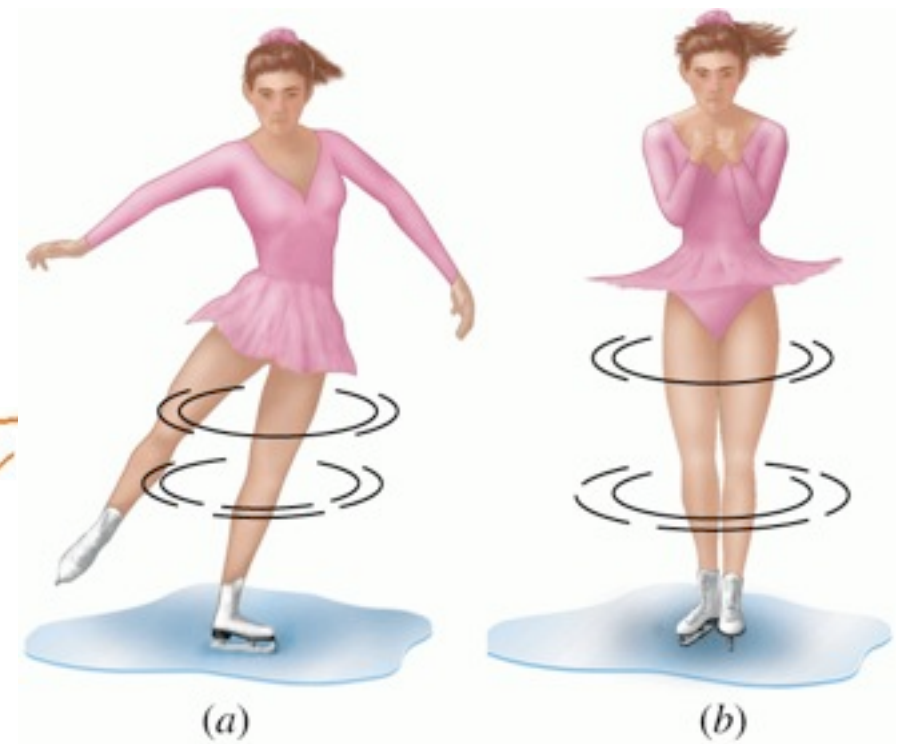


Drift of filaments

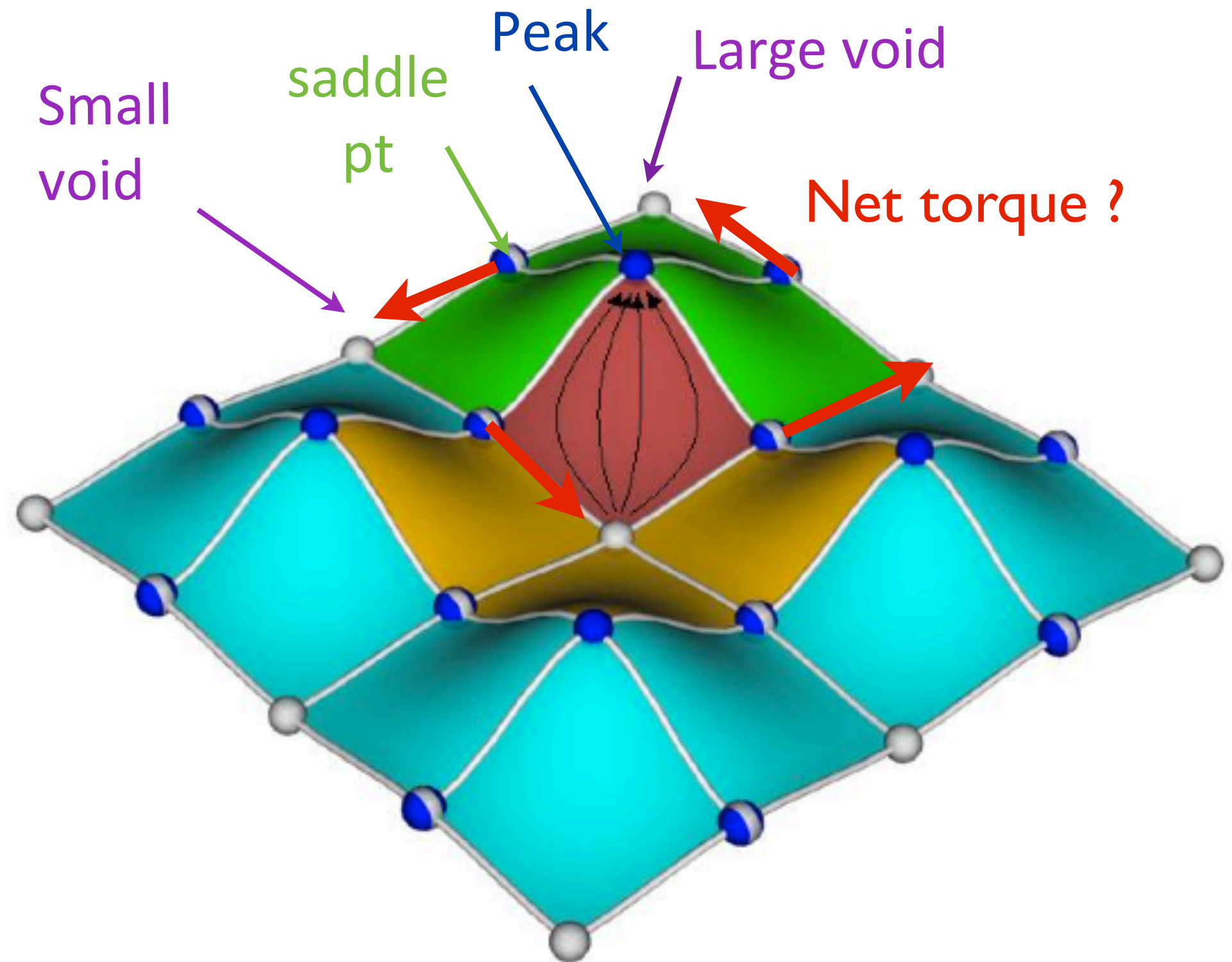
Time-line evolution
of filaments



void induced
cosmic
drift



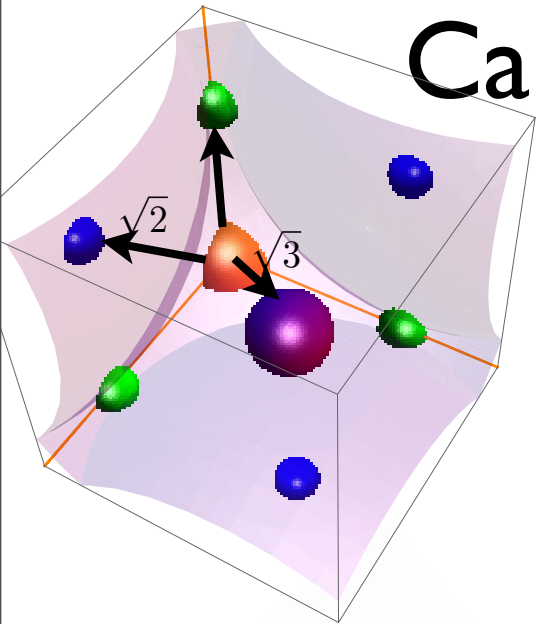
2D Cartoon of "ideal" cosmic environment :



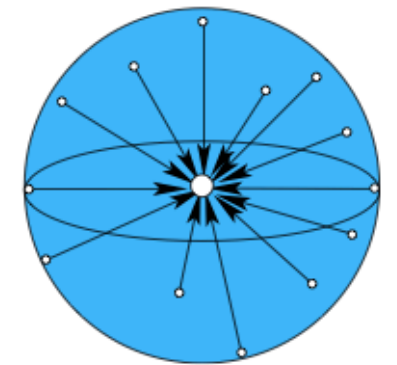
Mean local cosmic initial condition
homeomorphic to such crystal

3D "ideal" cosmic crystal

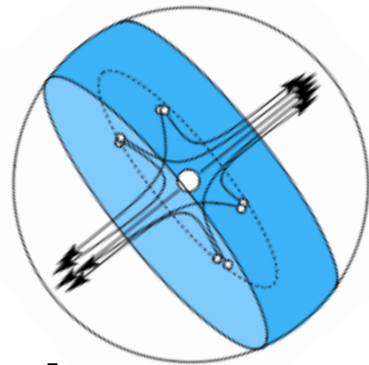
Cartoon:



Mean local cosmic initial condition
homeomorphic to such crystal

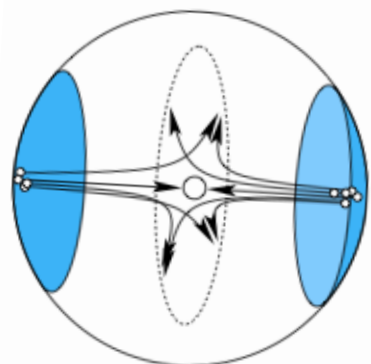


maximum



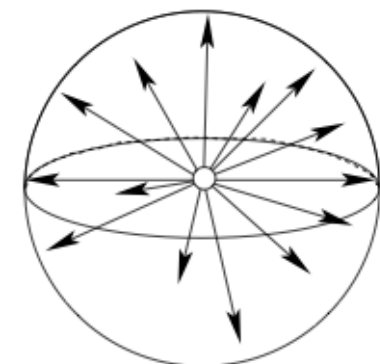
tube type saddle

big void



wall type saddle

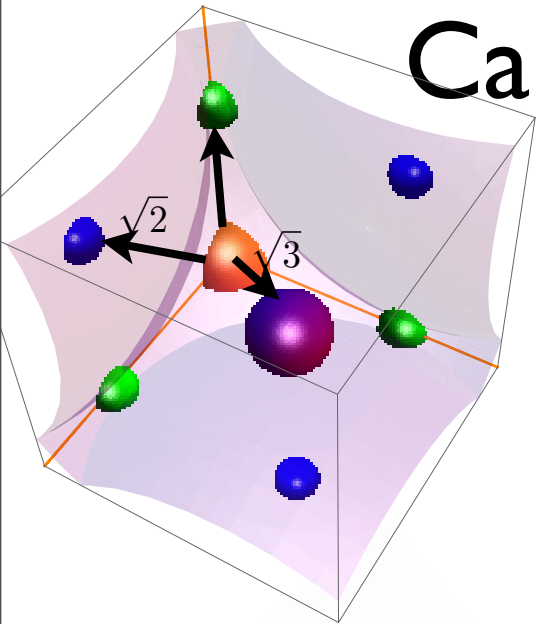
Net torque ?



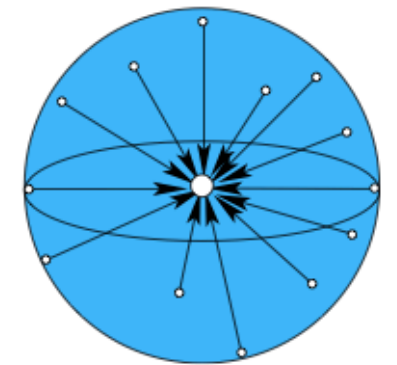
minimum

3D "ideal" cosmic crystal

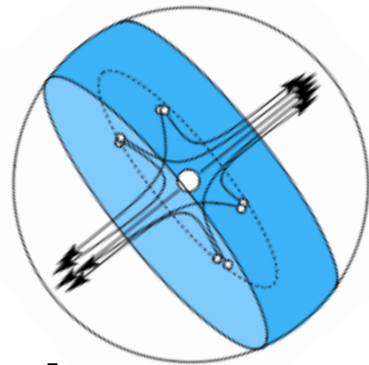
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Mean local cosmic initial condition
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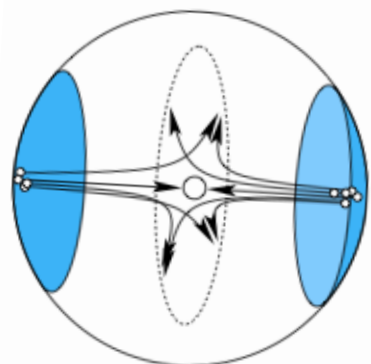


maximum



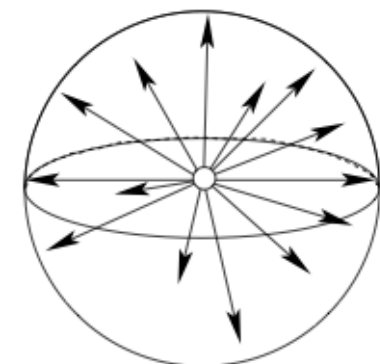
tube type saddle

big void



wall type saddle

Net torque ?



minimum

²⁰
biased by assumed isotropy: generically one fil+wall

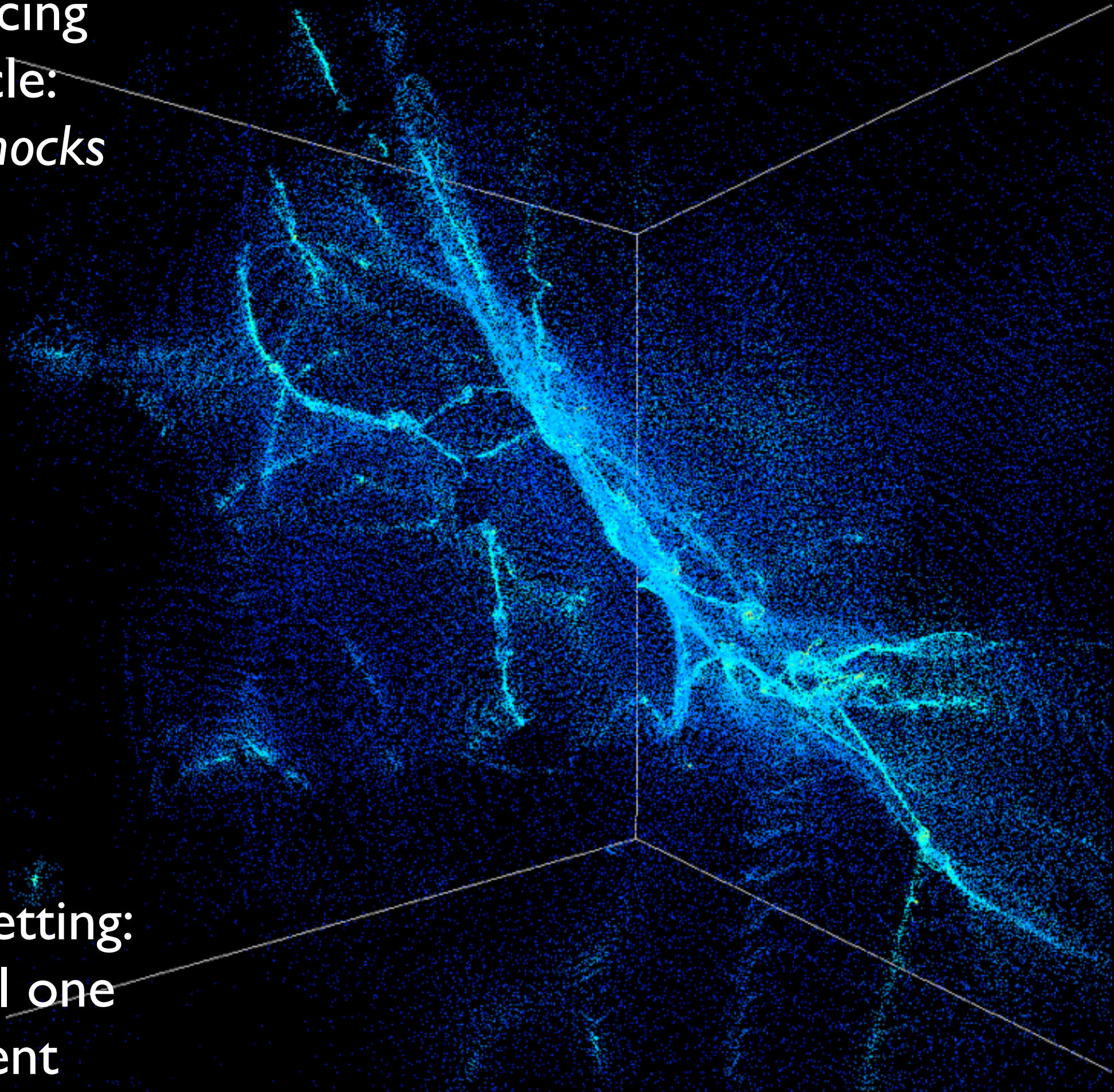
Do we see this?

"Proof" by visualisation of
hydrodynamical simulation

a.k.a

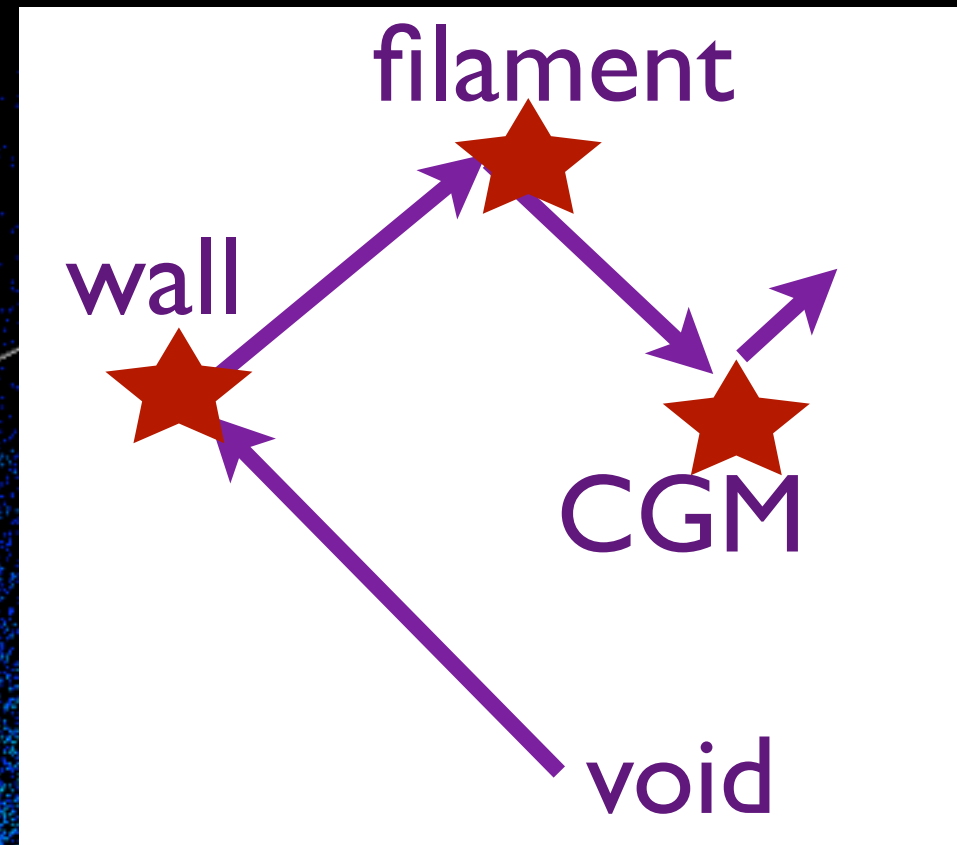
proof by pretty pictures

gas tracing
particle:
follow shocks



typical setting:
one wall one
filament

gas tracing
particle:
follow shocks



typical setting:
one wall one
filament



locus of 3rd
shock

Note the high **helicity** of inflow:
AM rich quasi-**polar** accretion

Explain this !

Can we trace this back in time?

"Proof" by tagging

a.k.a "Proof" by looking at **ONE** object !

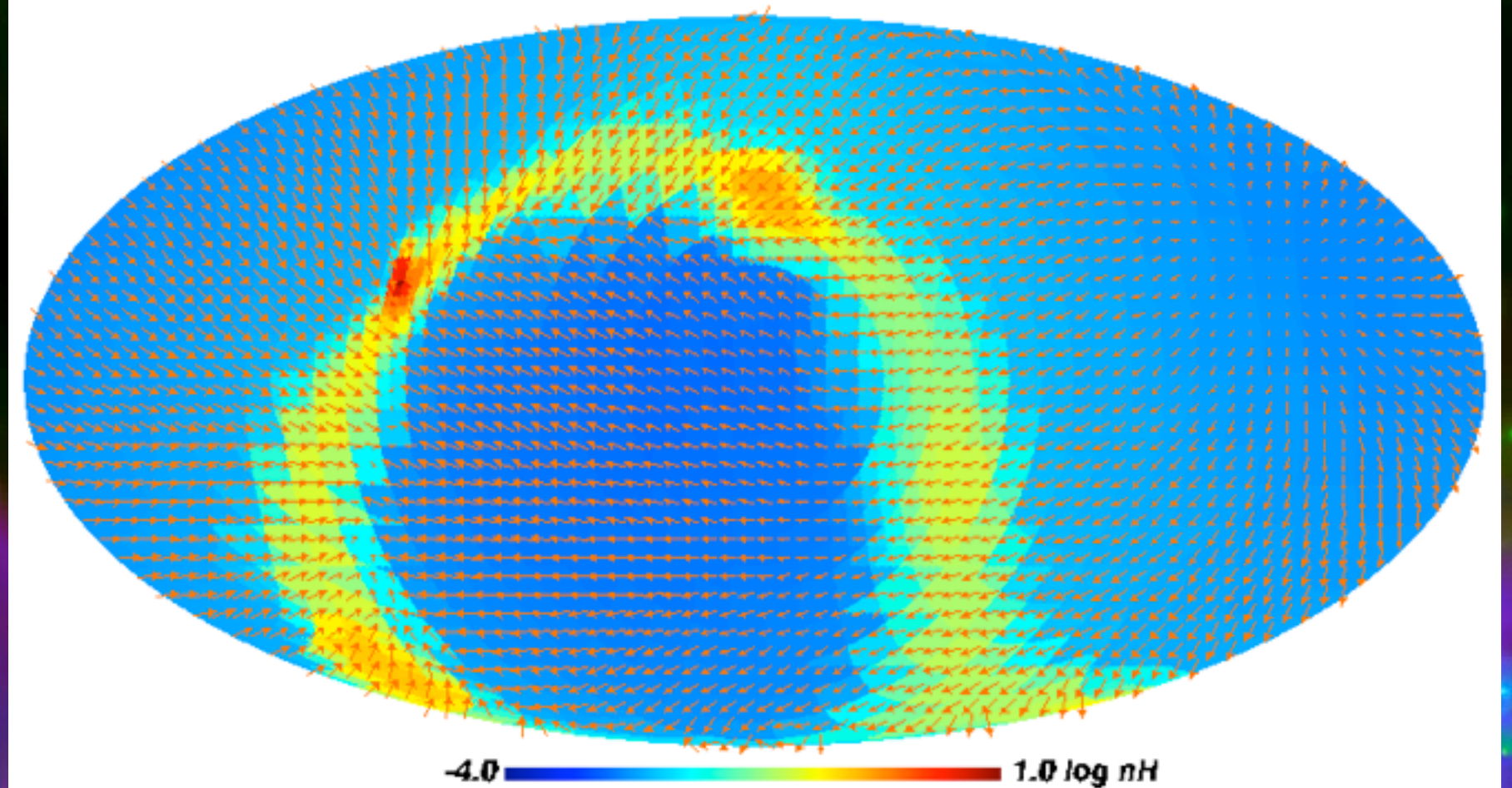
Angular momentum rich filamentary cold flows: progenitor of thin discs?



Nut Simulation
0.5 pc resolution
"full physics"

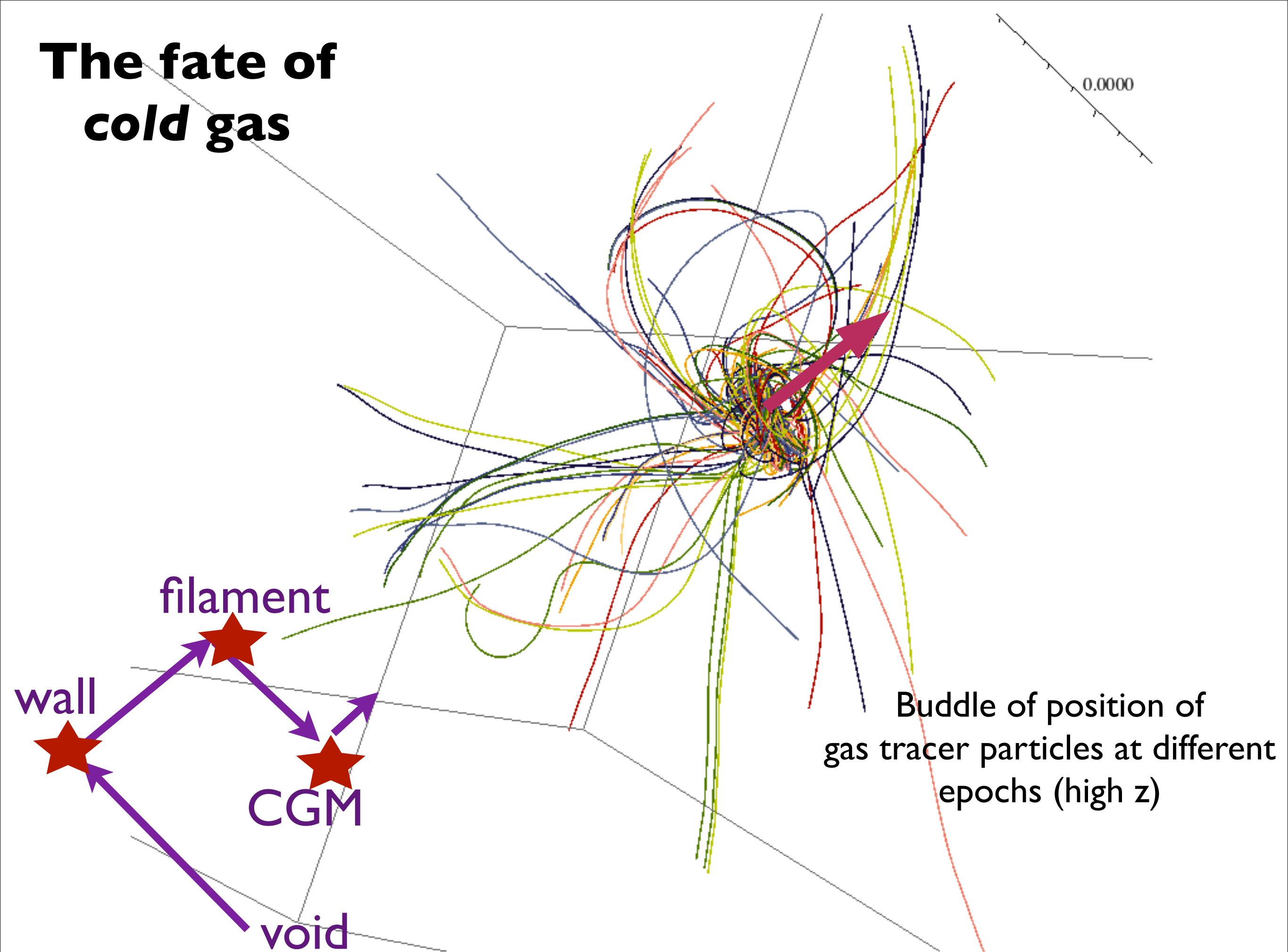
Angular momentum rich filamentary cold flows: progenitor of thin discs?

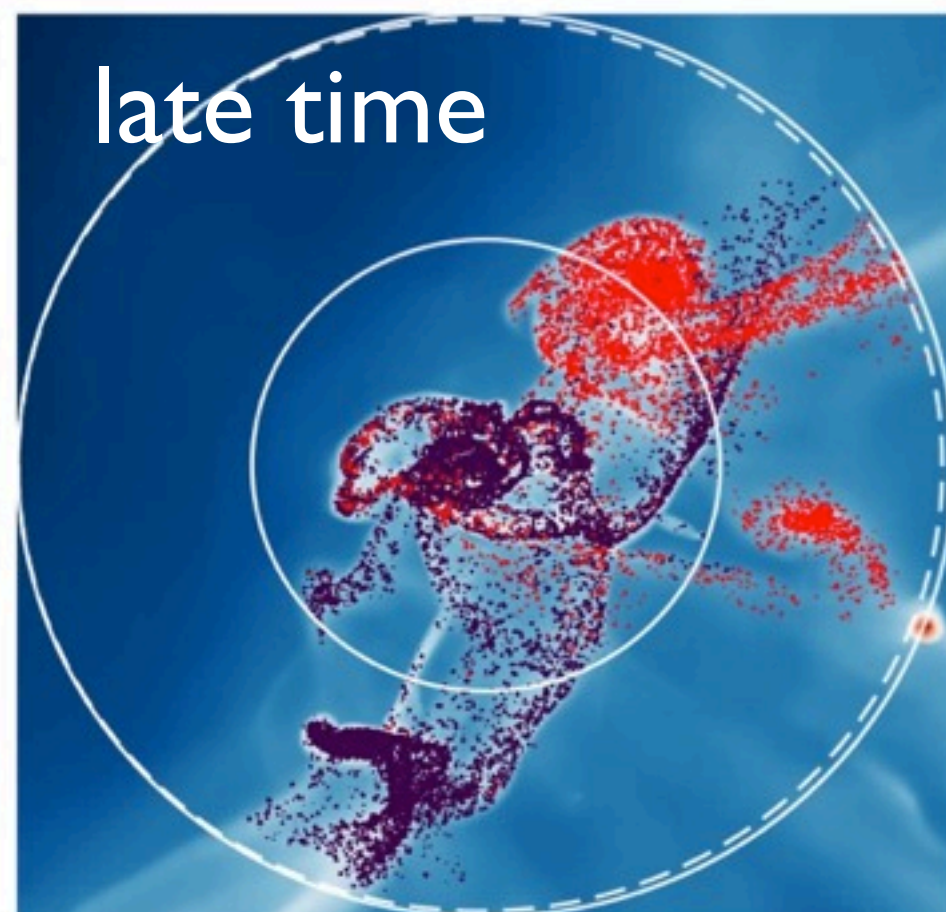
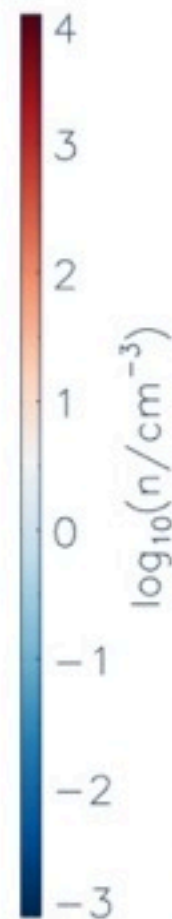
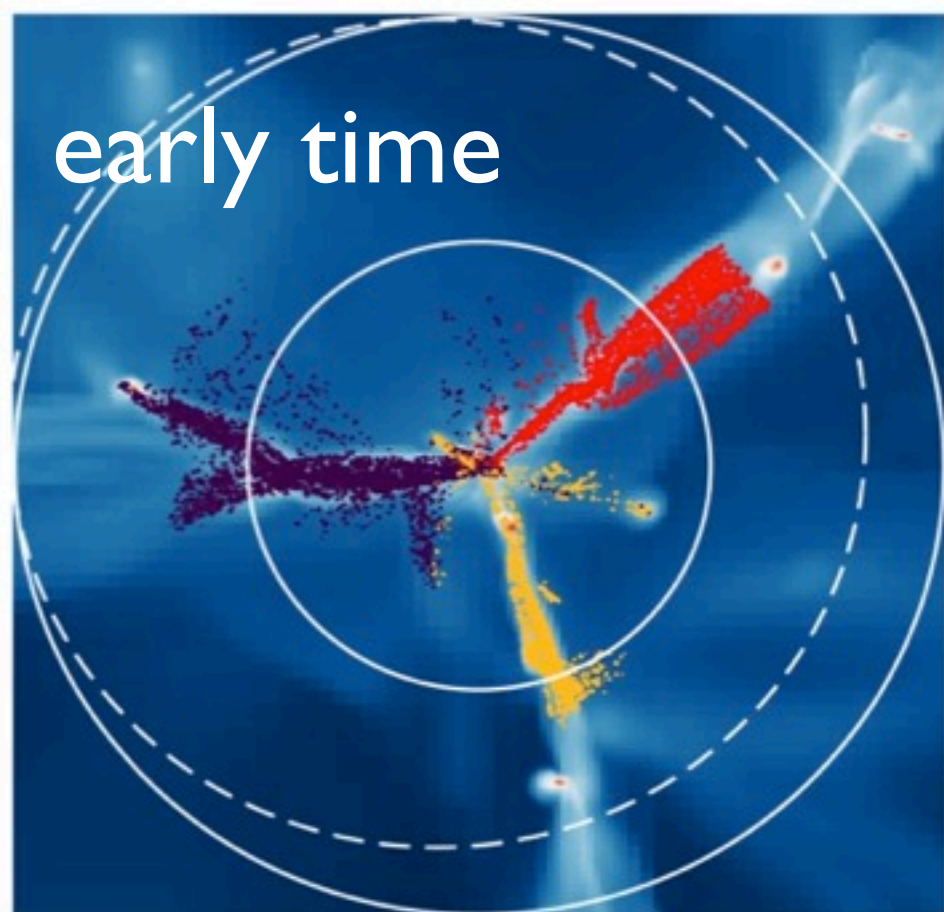
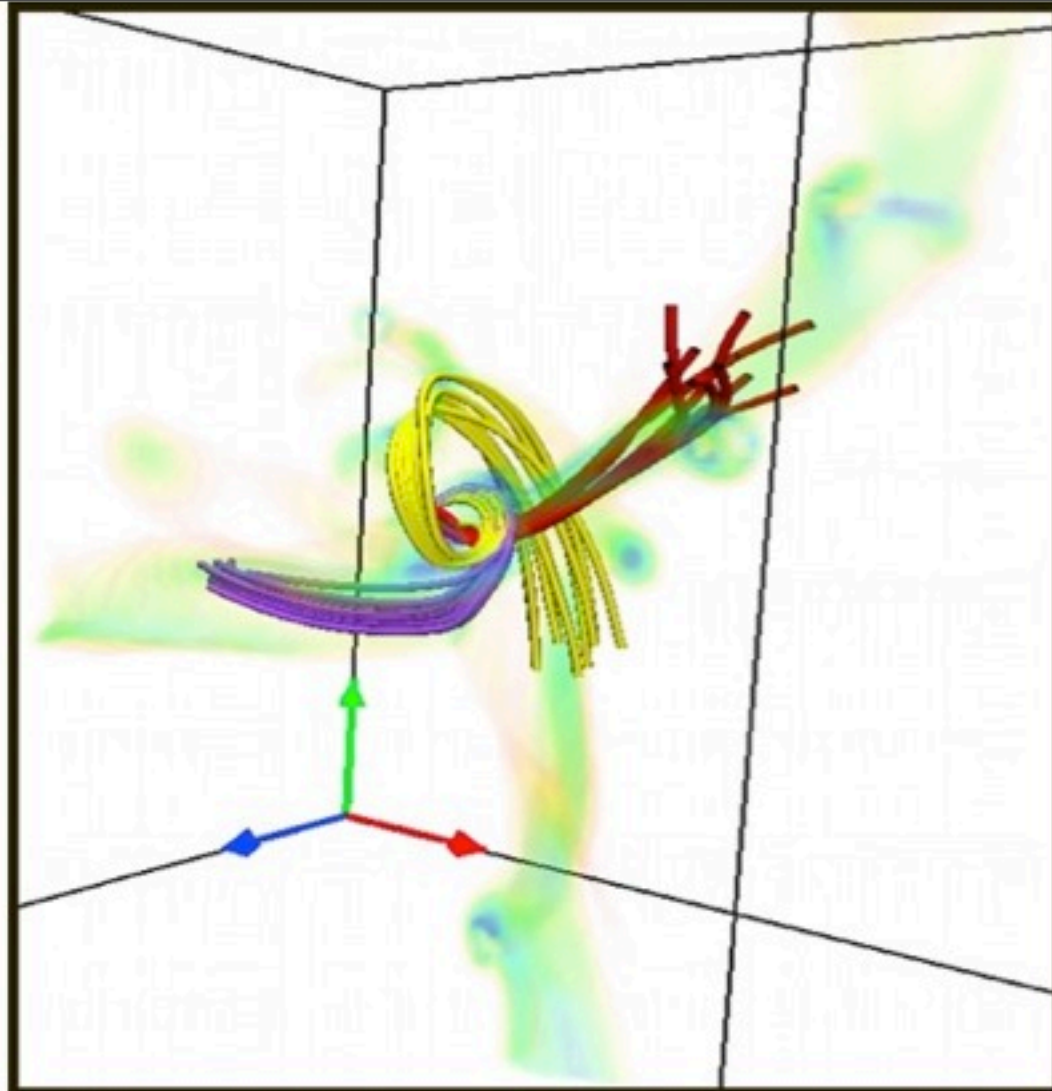
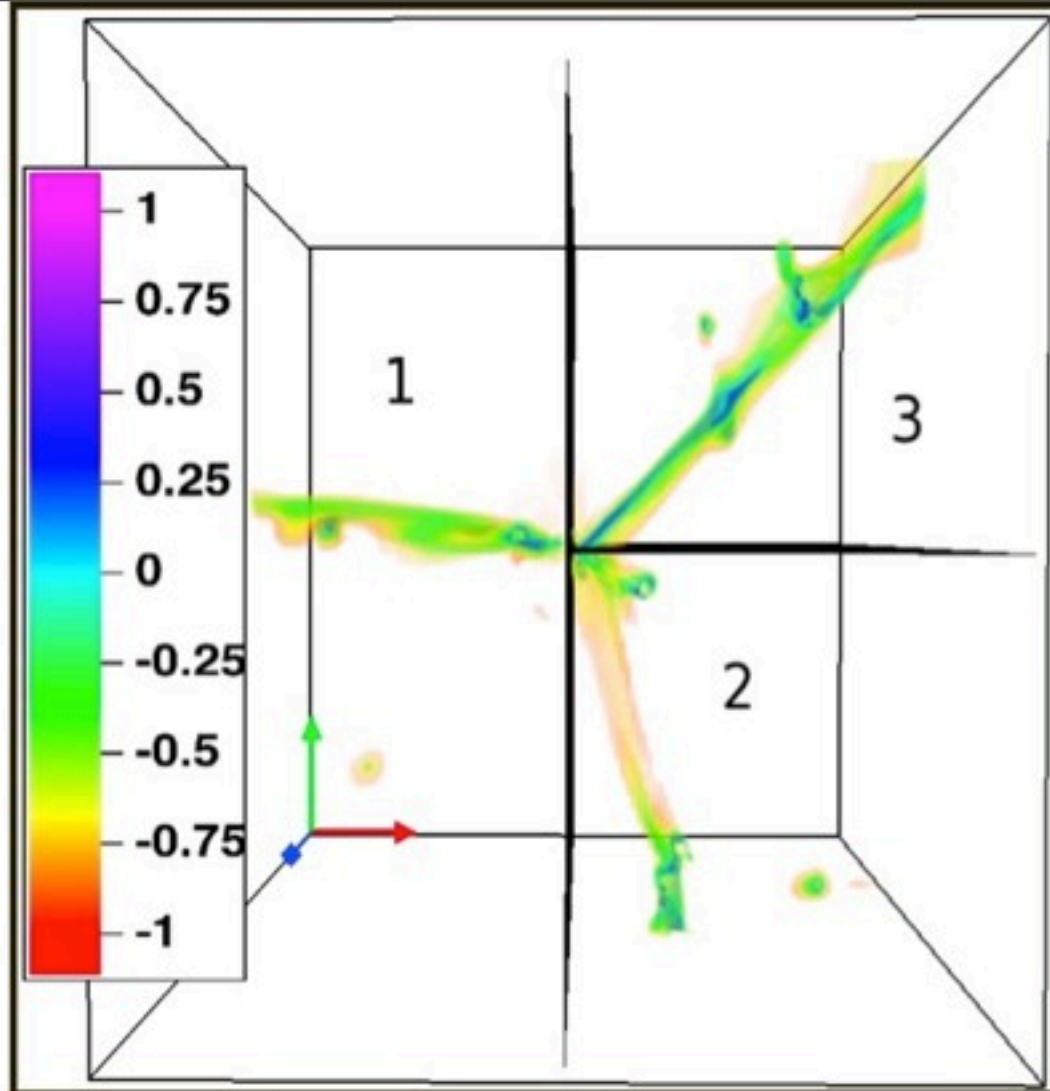
$z=11.20$ ($r=1.0 R_{\text{vir}}$)



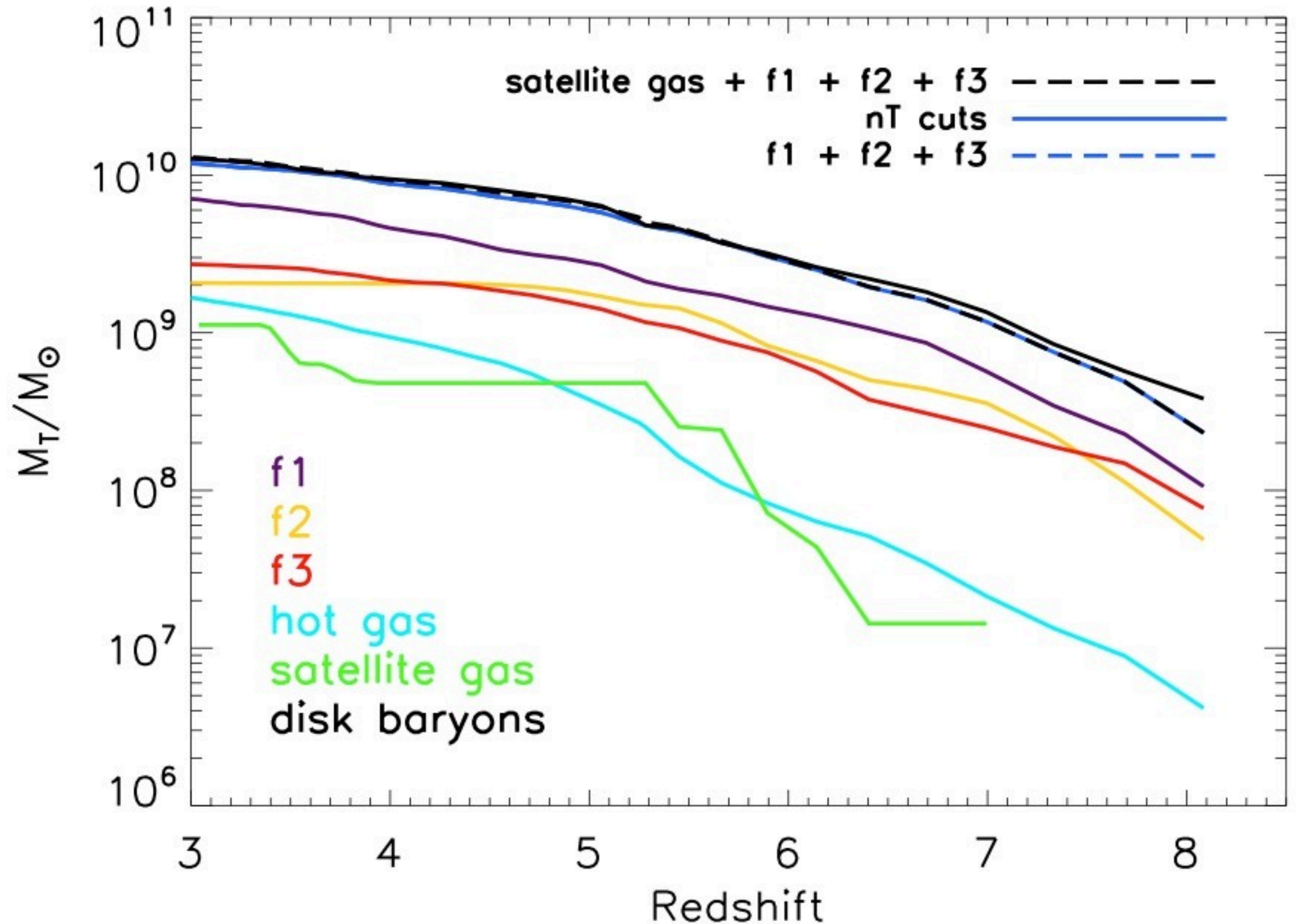
Nut Simulation
0.5 pc resolution
"full physics"

The fate of *cold* gas

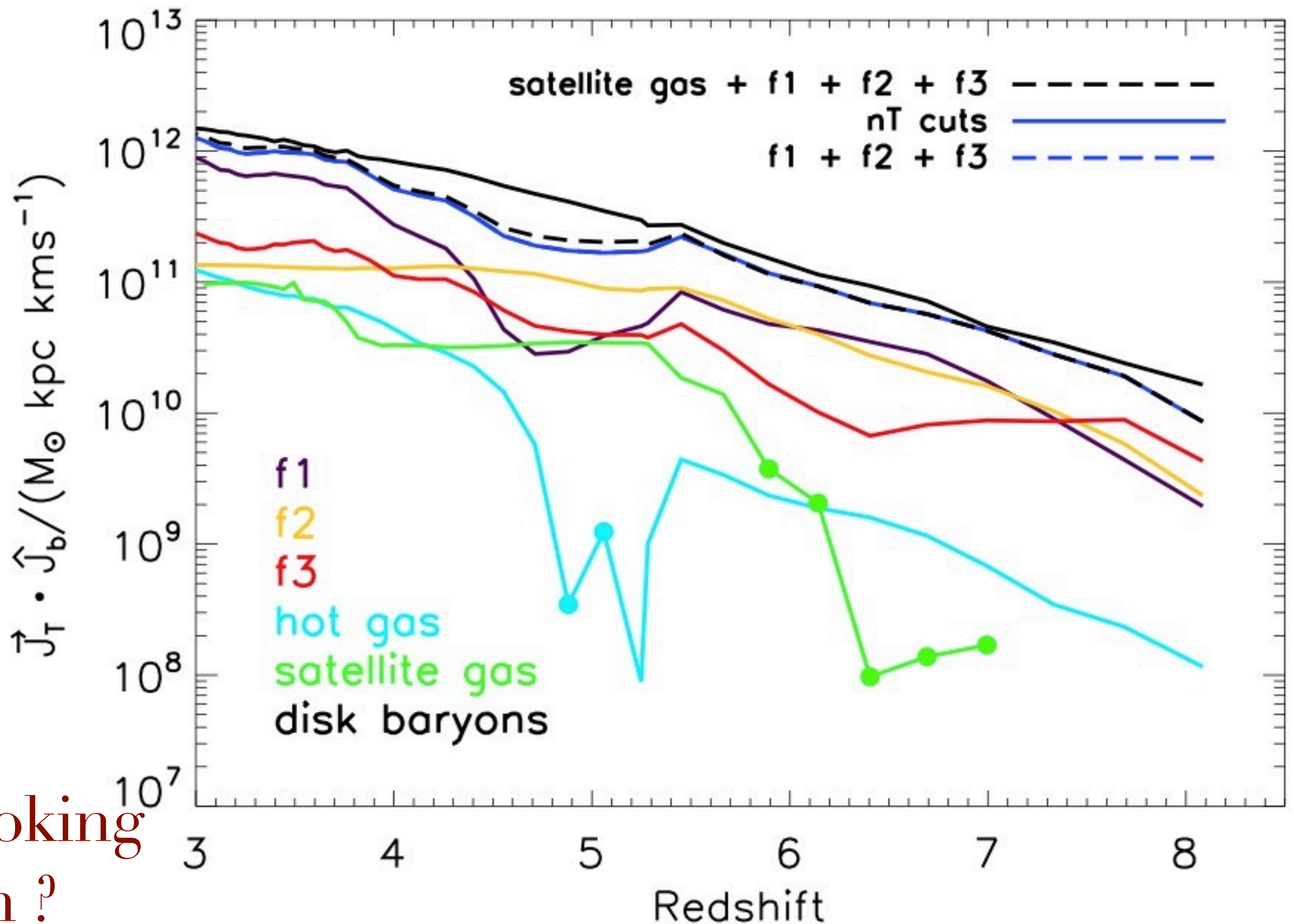




Mass in disc originate from filaments



Angular momentum in disc originate from filaments



Smoking
gun ?

Can it be made quantitative?

"Proof" by robust statistical analysis

a.k.a

lies, damn lies and statistics

Anisotropic accretion: cold flows driven by LSS

MareNostrum $z \sim 2$

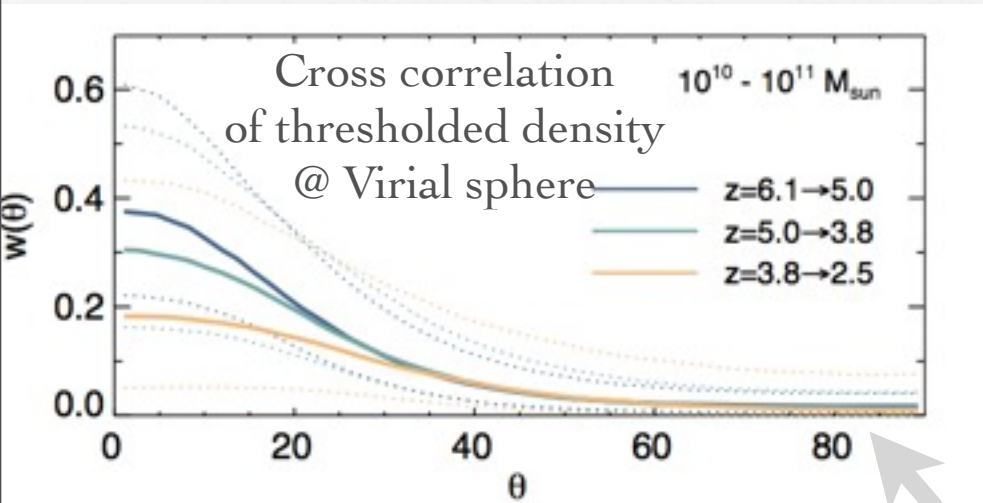
130 000 galaxies

Red: density
Blue: temperature
Green: metal

$M = 10^{12} M_{\odot}$ $z=3$

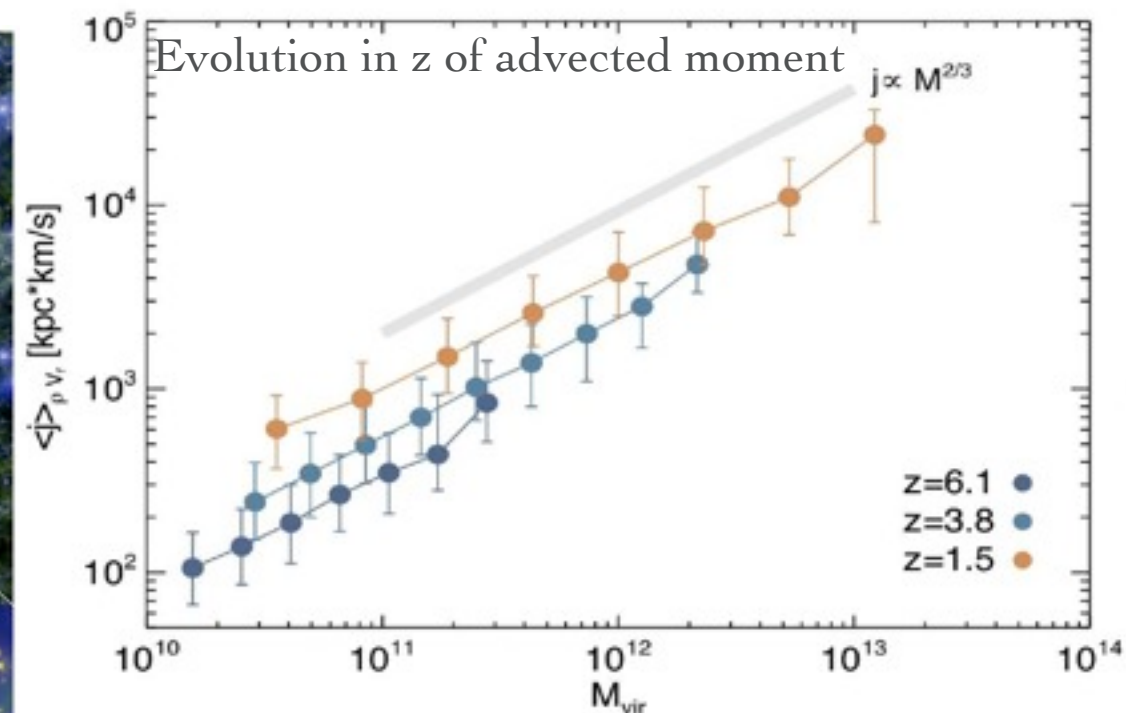
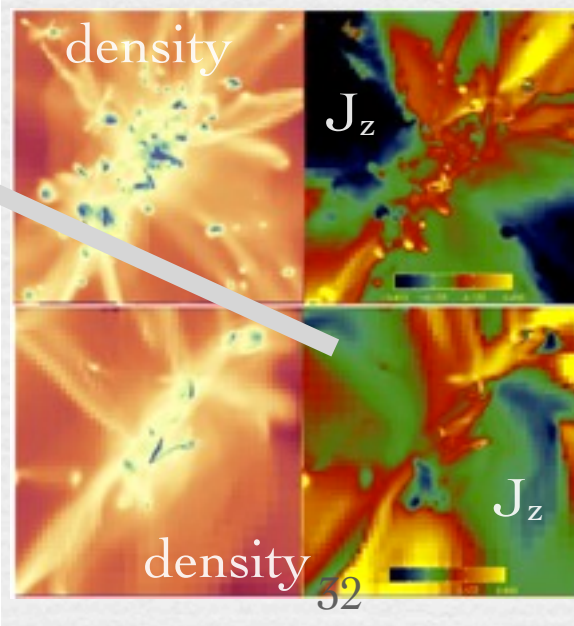
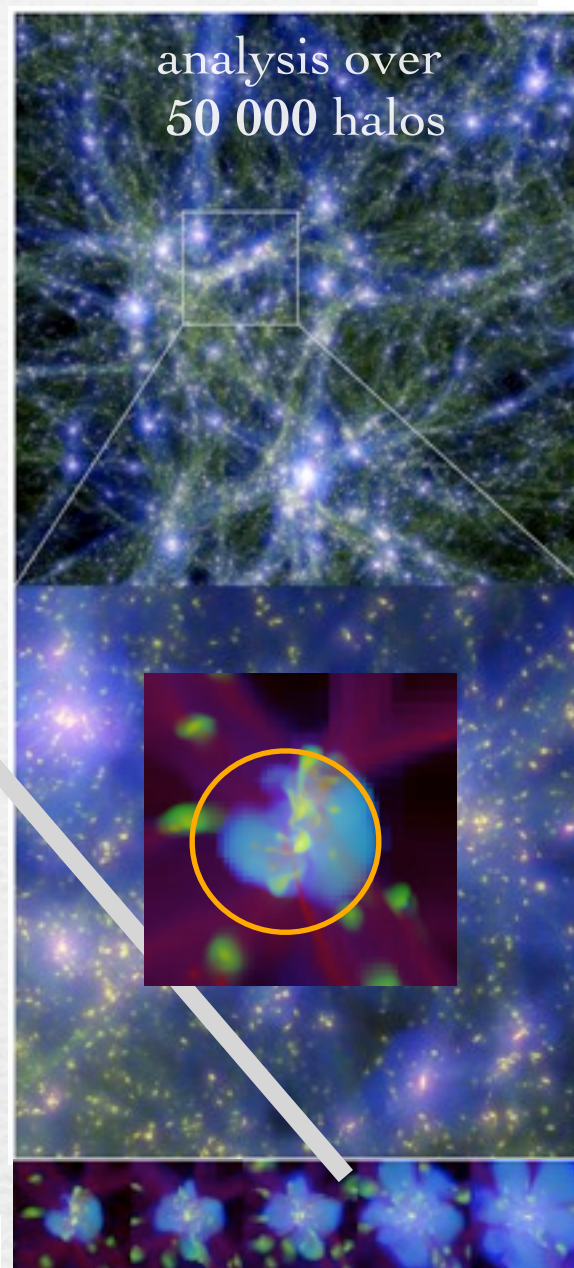
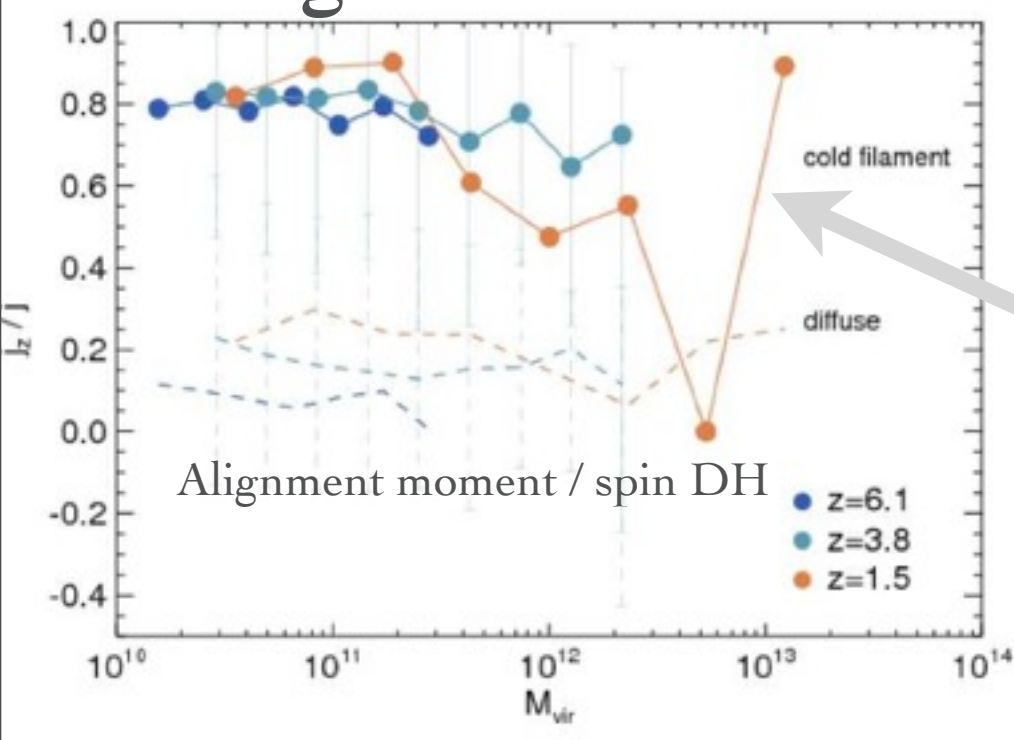
- Use LSS dynamics to statistically analyse AM infall @ R_{vir}

$$w(\theta) = \sum_{l,m} \langle a_{lm} b_{lm}^* \rangle P_l(\cos(\theta)) / \sqrt{(\sum_{l,m} \langle |a_{lm}|^2 \rangle \sum_{l,m} \langle |b_{lm}|^2 \rangle)},$$



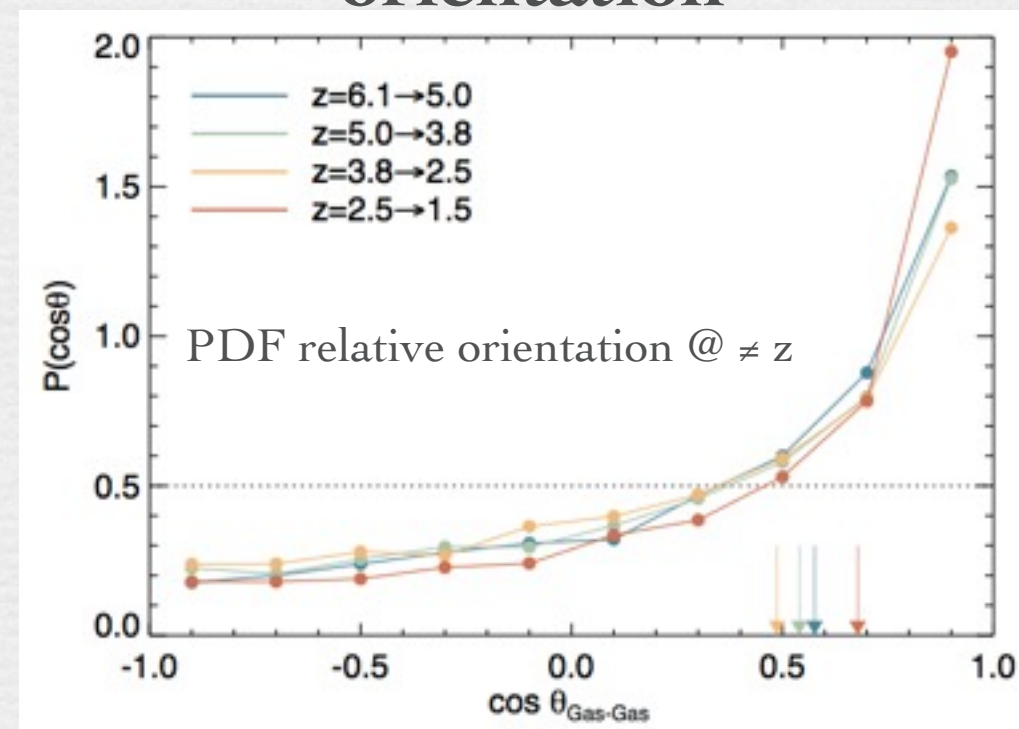
Filamentary Accretion:
coherent orientation

Advected Ang Moment
Alignment / halo

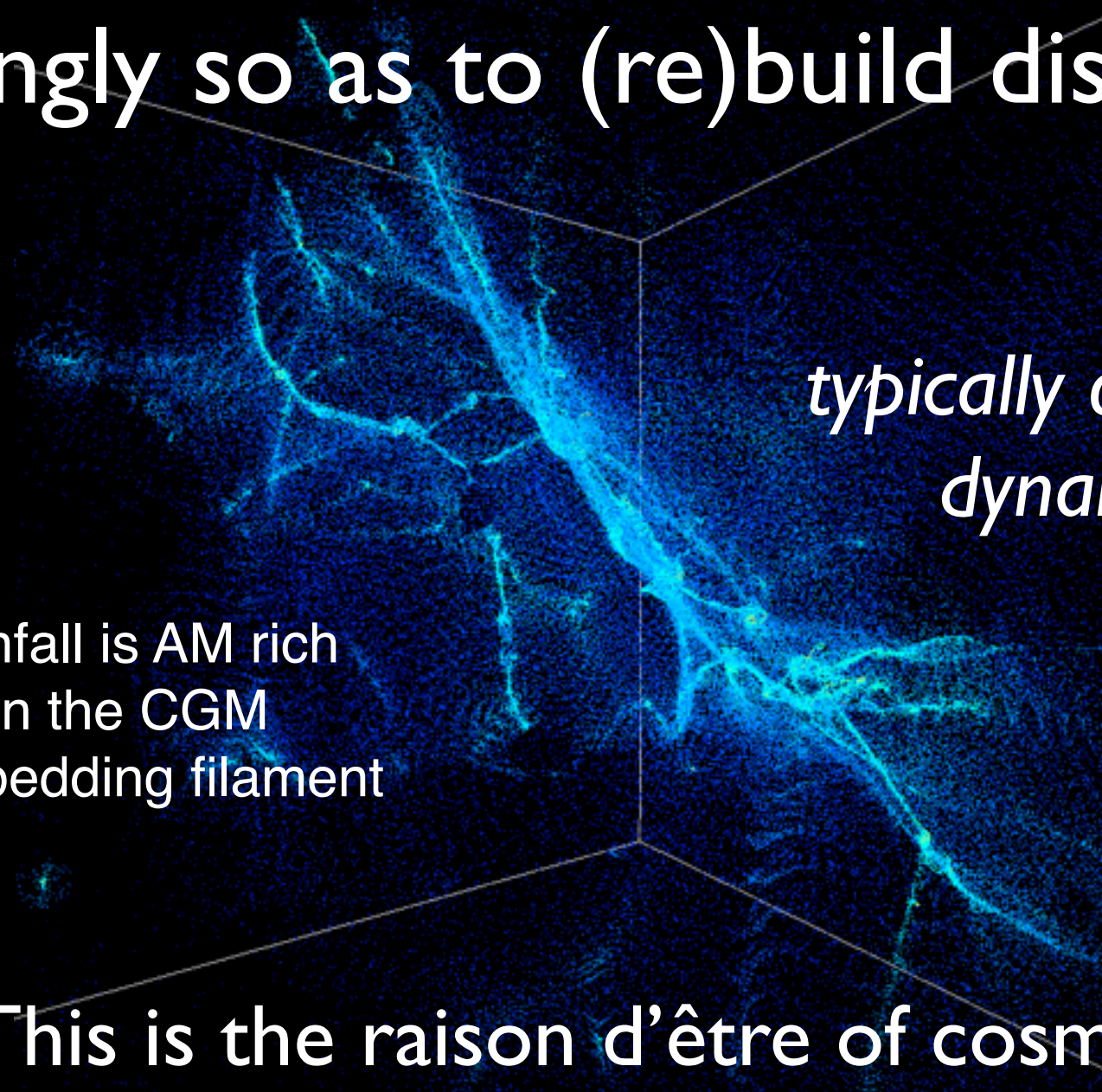


growth of lang moment!

Coherent ang. moment
orientation



Disks form because LSS are large (*dynamically young*)
and (*partially*) an-isotropic :
they induce persistent angular momentum
advection of cold gas **along** filaments
which stratifies
accordingly so as to (re)build discs continuously.



*typically one wall one filament:
dynamical implication?*

initial galactic infall is AM rich
but quasi polar in the CGM
-> spin // to embedding filament

This is the raison d'être of cosmic web :-)

PART II

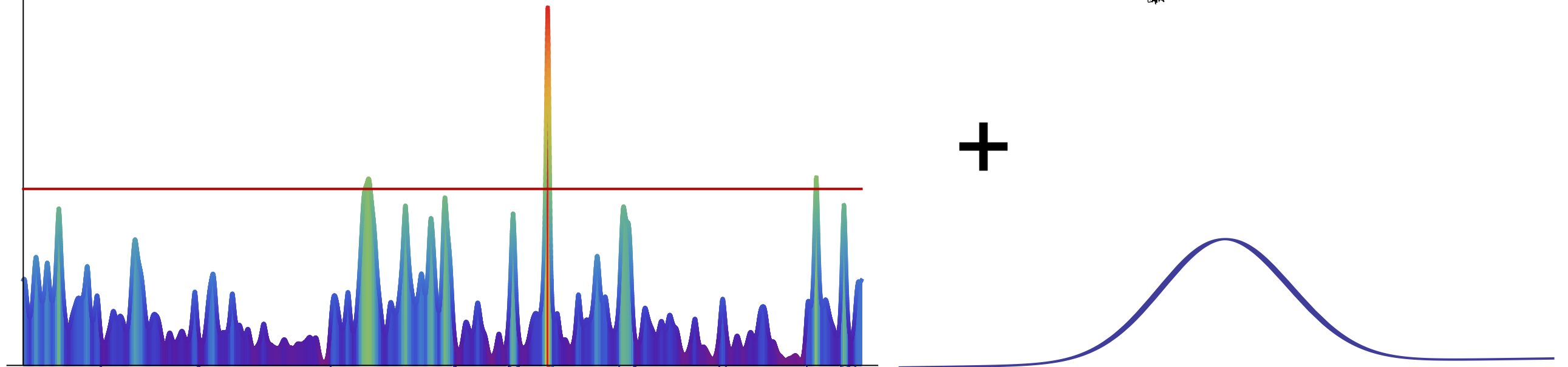
What's happening on larger scales?

How is the cosmic web woven?
i.e Where do galaxies form in our Universe?
What are the dynamical implications?
Why?

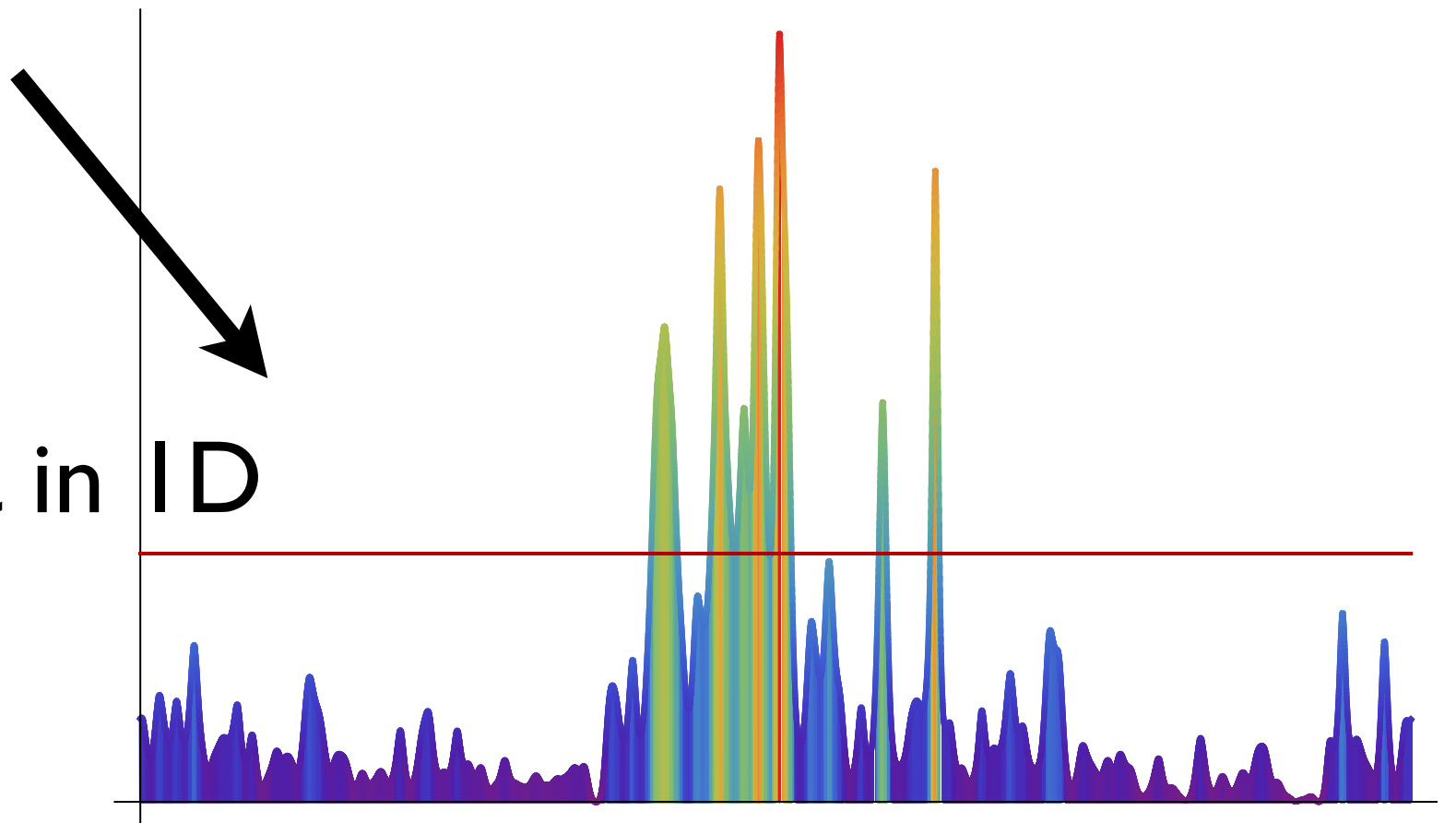


- How discs build up from persistent cosmic web?
- How dark halo's spin flip relative to filament?
Why are they initially aligned with filaments?
Why the transition mass? **Eulerian view**
- What is the corresponding Lagrangian theory?

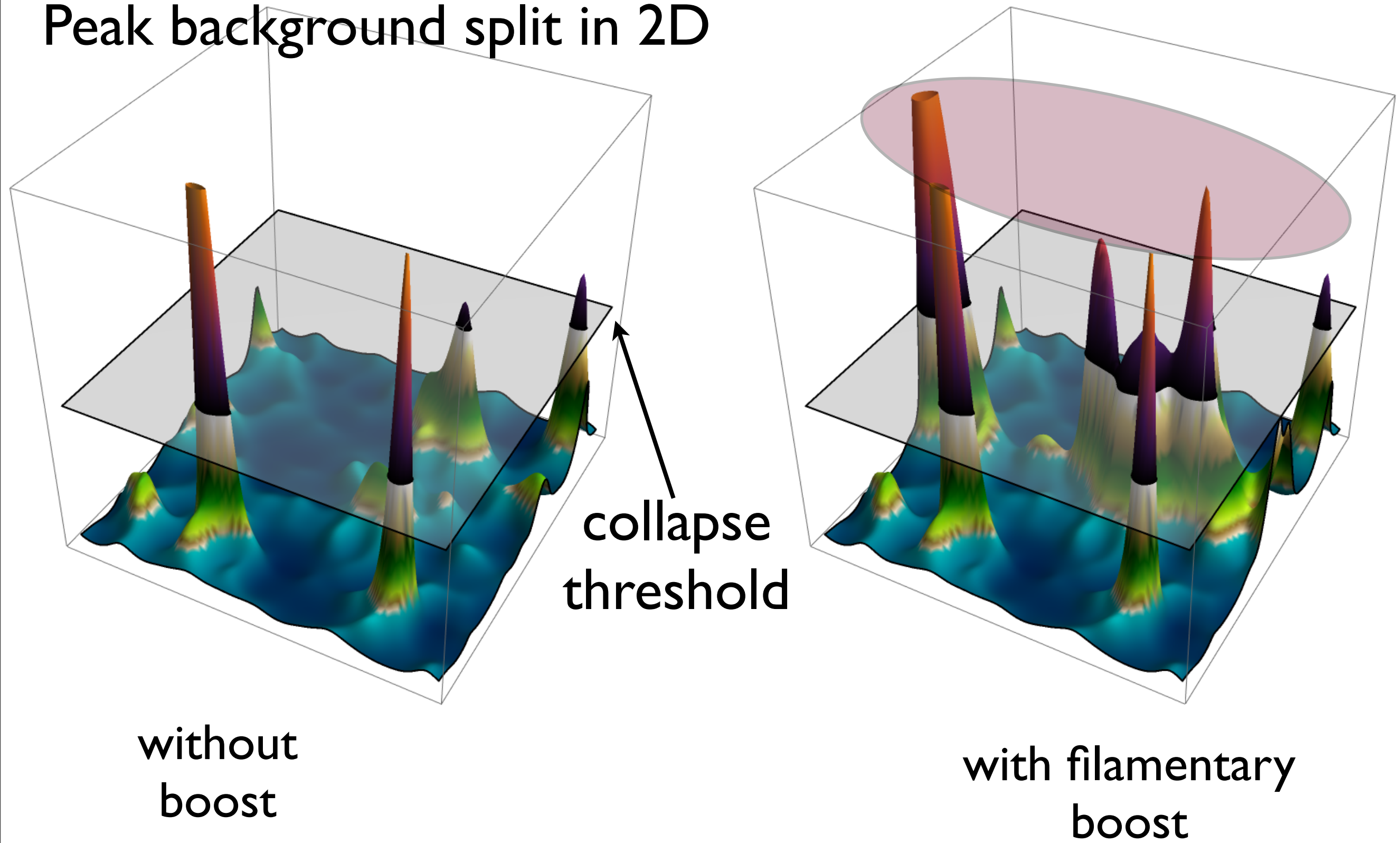
dark halos don't form anywhere



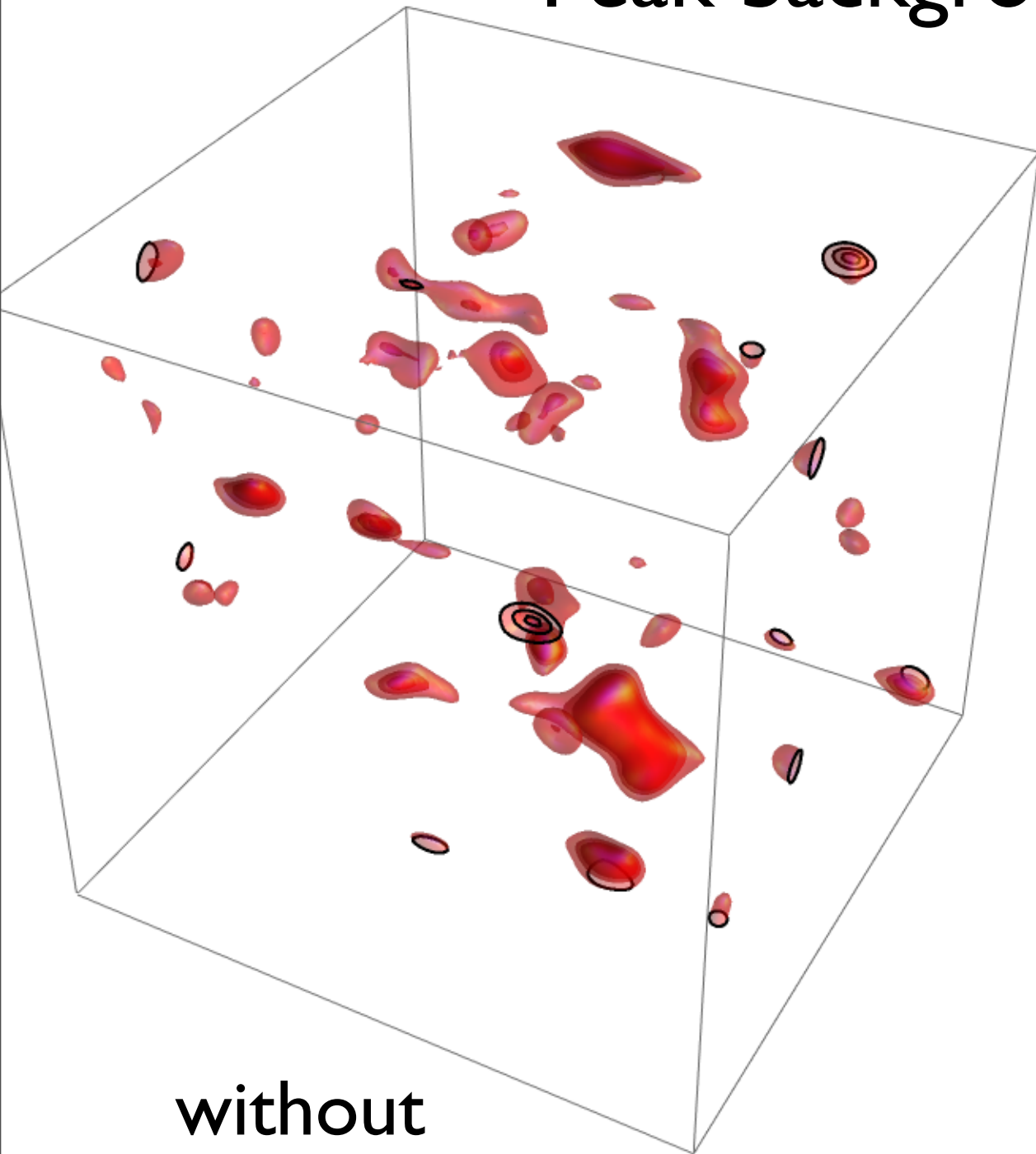
Peak background split in 1D



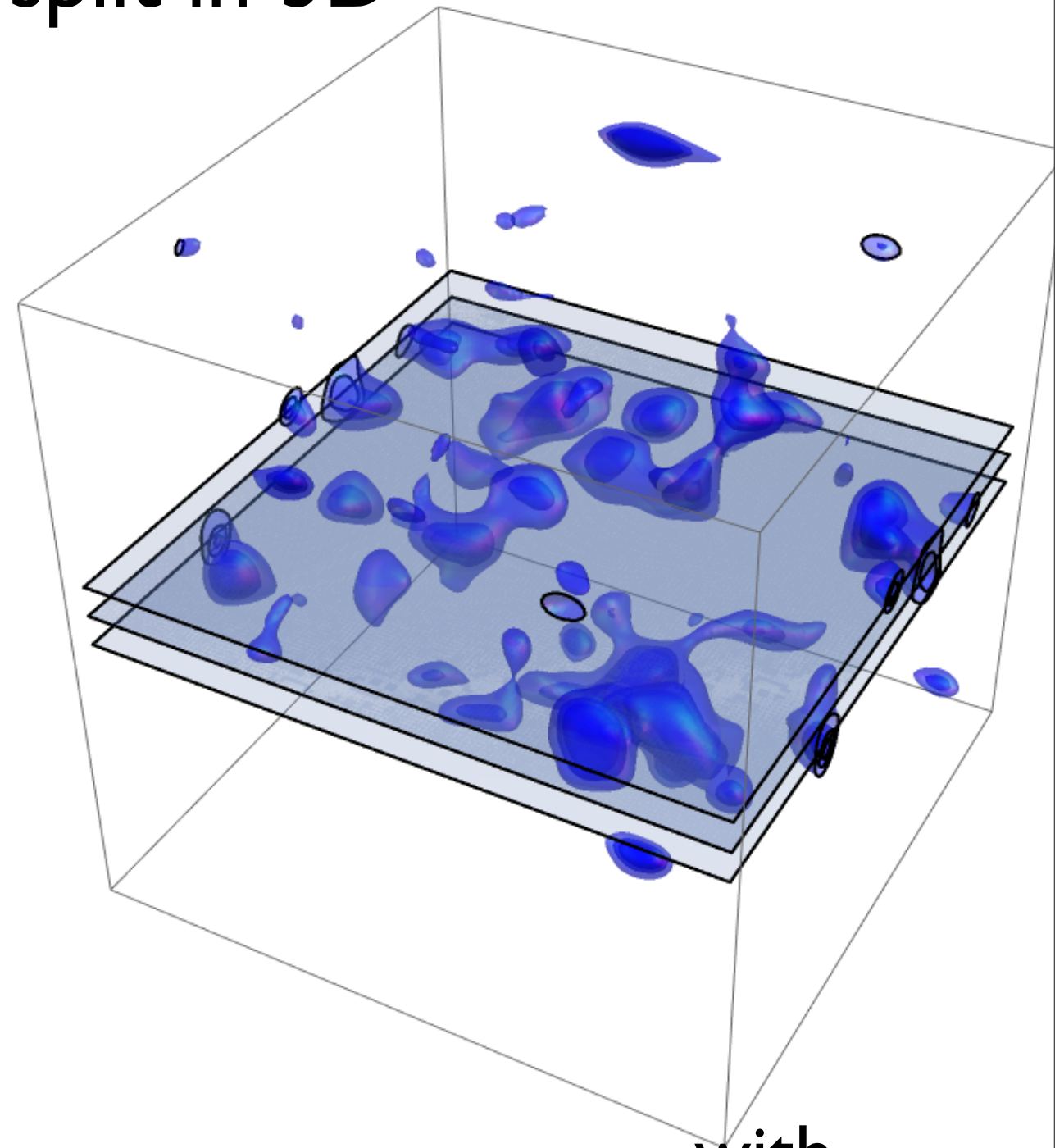
Peak background split in 2D



Peak background split in 3D



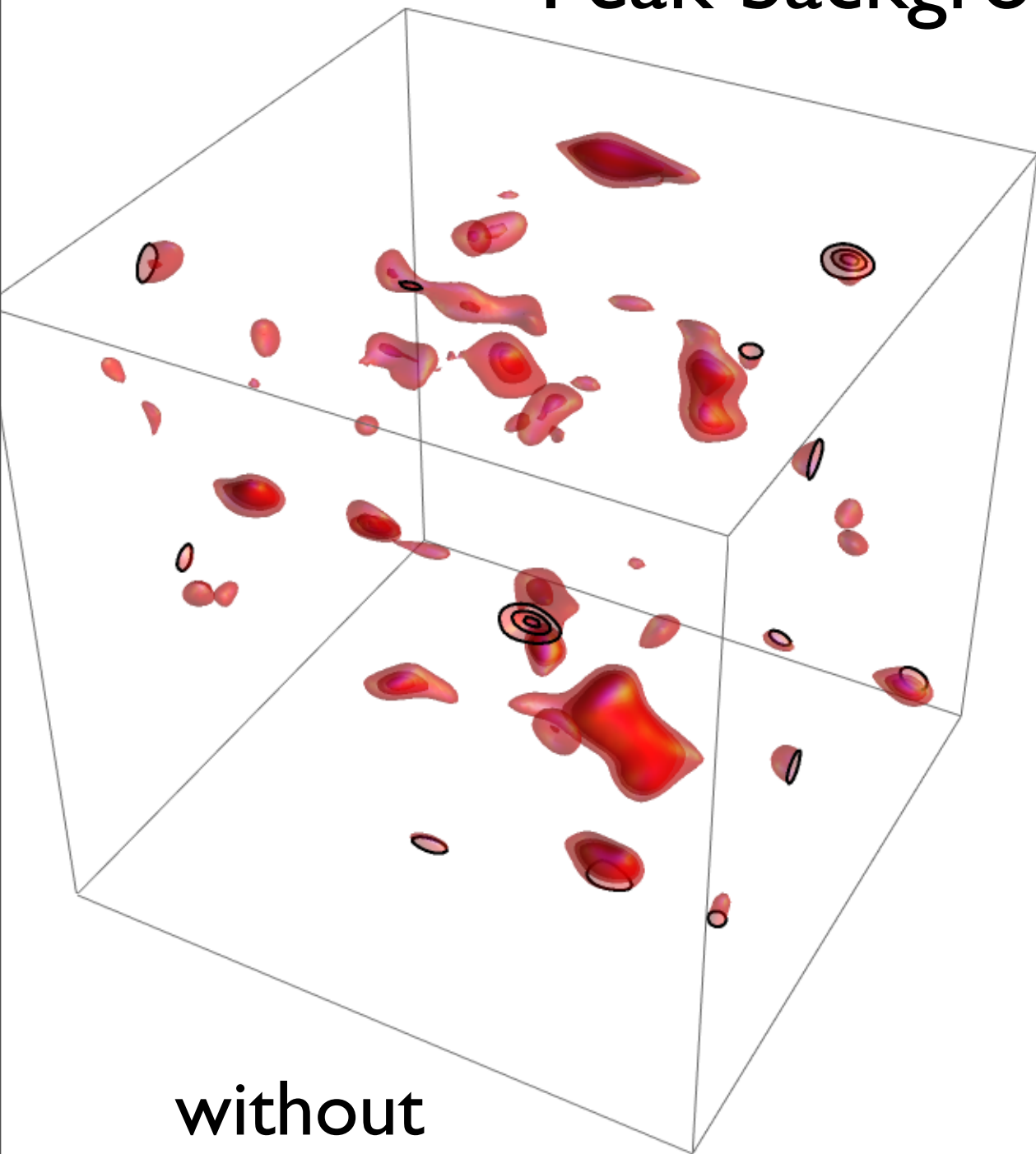
without
boost



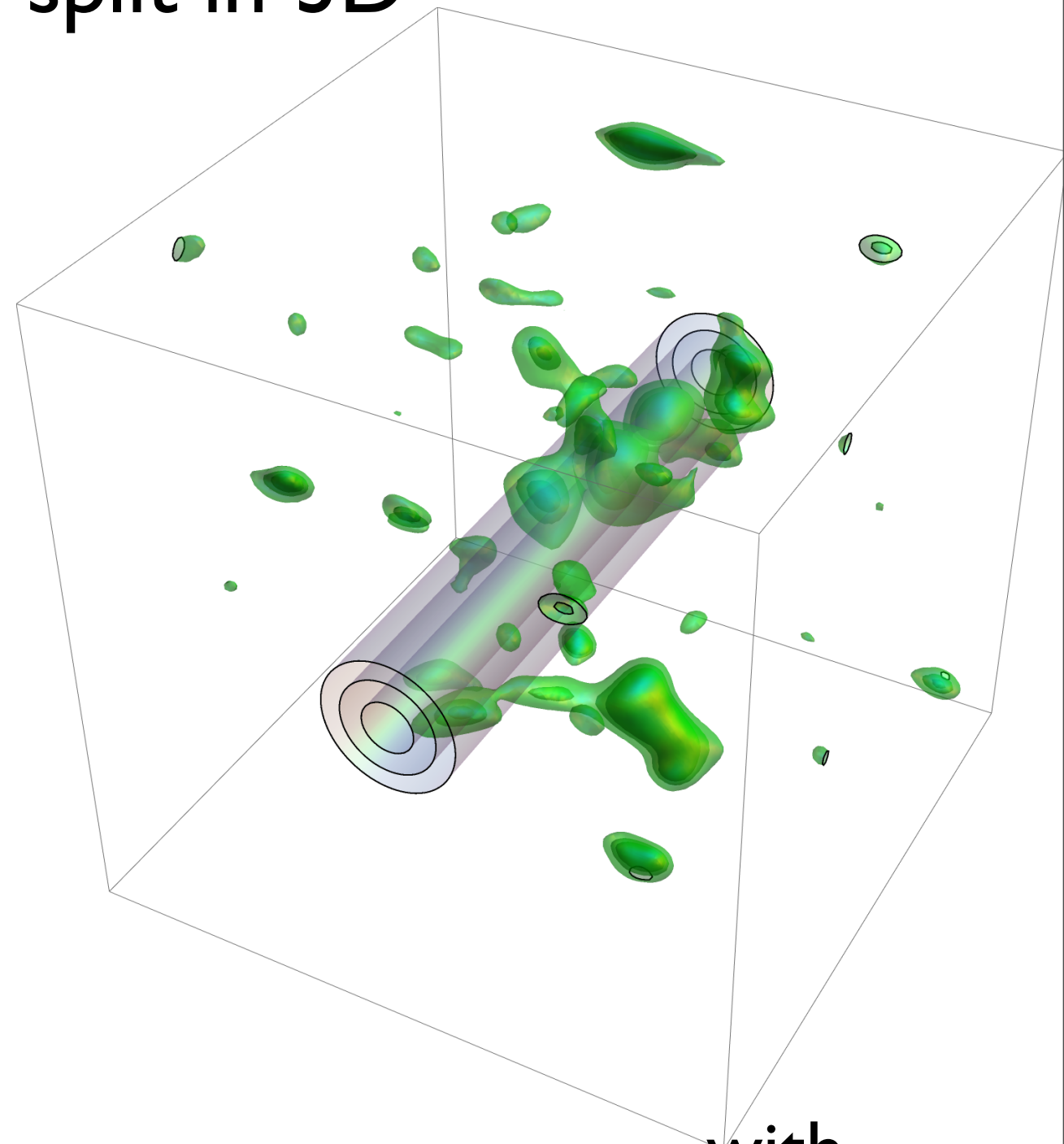
with
boost

**Does this anisotropic biassing have
a dynamical signature? yes!**

Peak background split in 3D



without
boost



with
boost

**Does this anisotropic biasing have
a dynamical signature? yes!**

Orientation of the spins w.r.t the filaments

Horizon 4Pi:

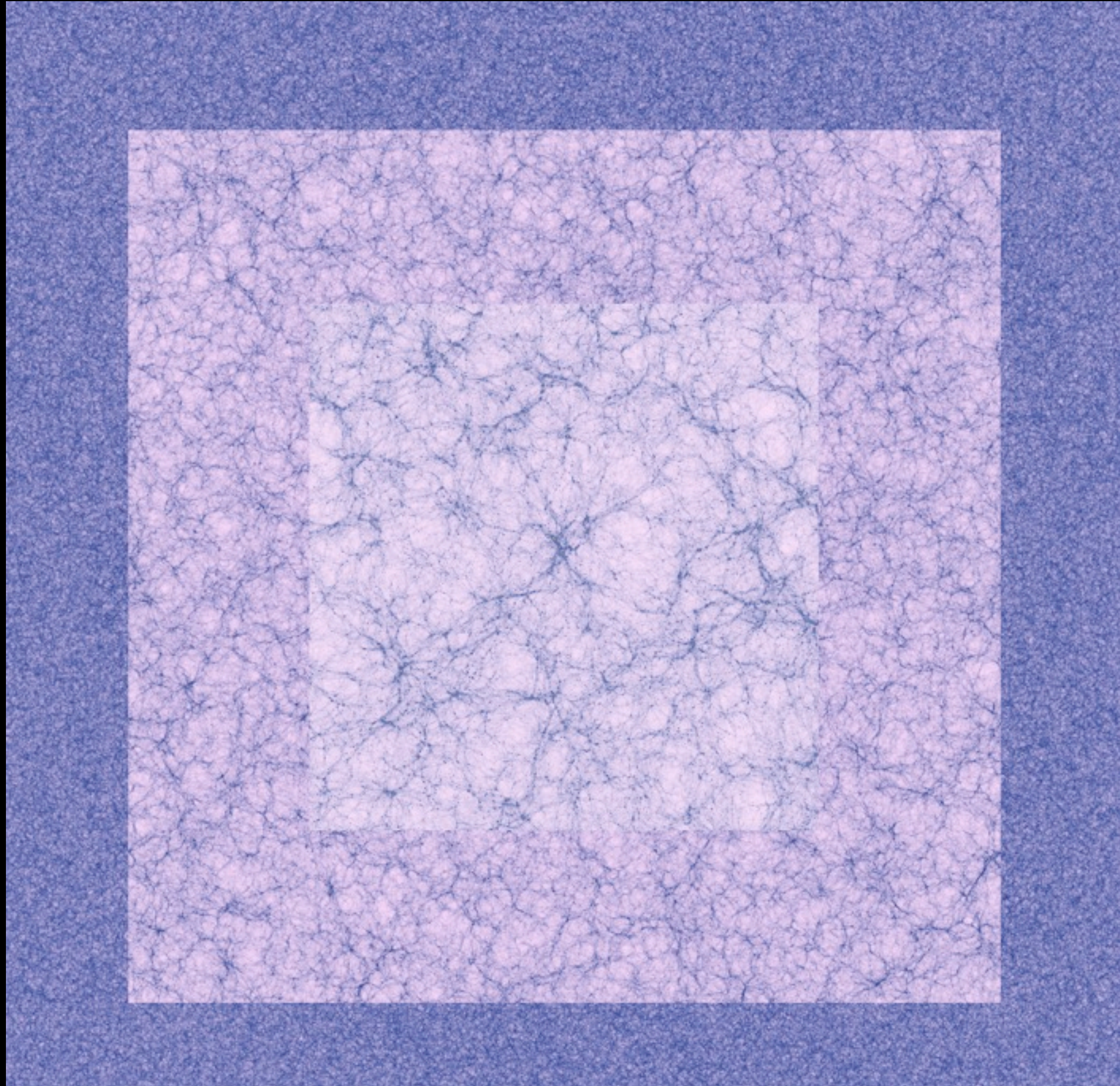
DM only

2 Gpc/h periodic box

4096^3 DM part.

43 million dark halos at
 $z=0$

(Teyssier et al, 2009)



Orientation of the spins w.r.t the filaments

Horizon 4Pi:

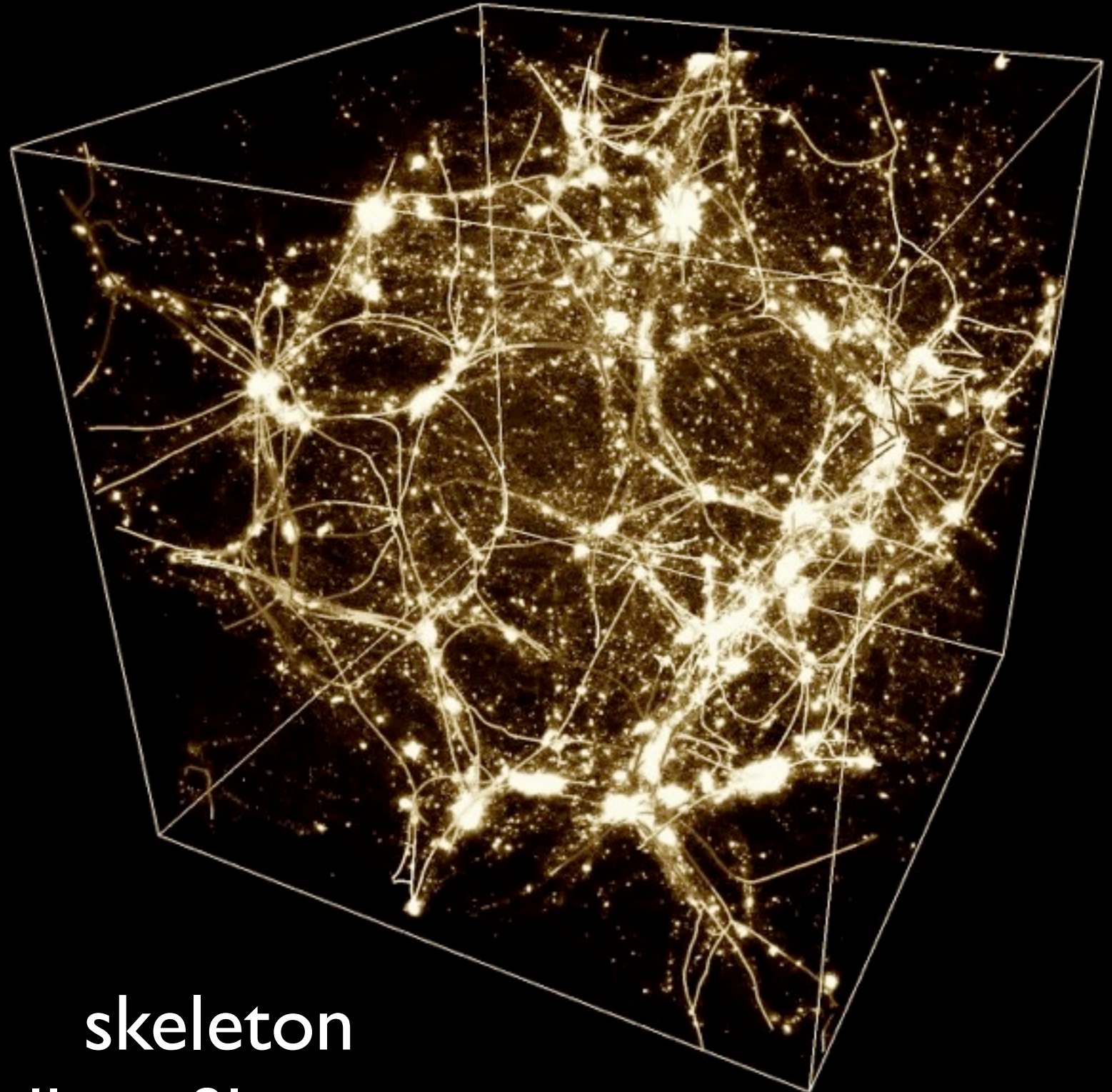
DM only

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(Teyssier et al, 2009)



skeleton
follow filaments

Orientation of the spins w.r.t the filaments

Horizon 4Pi:

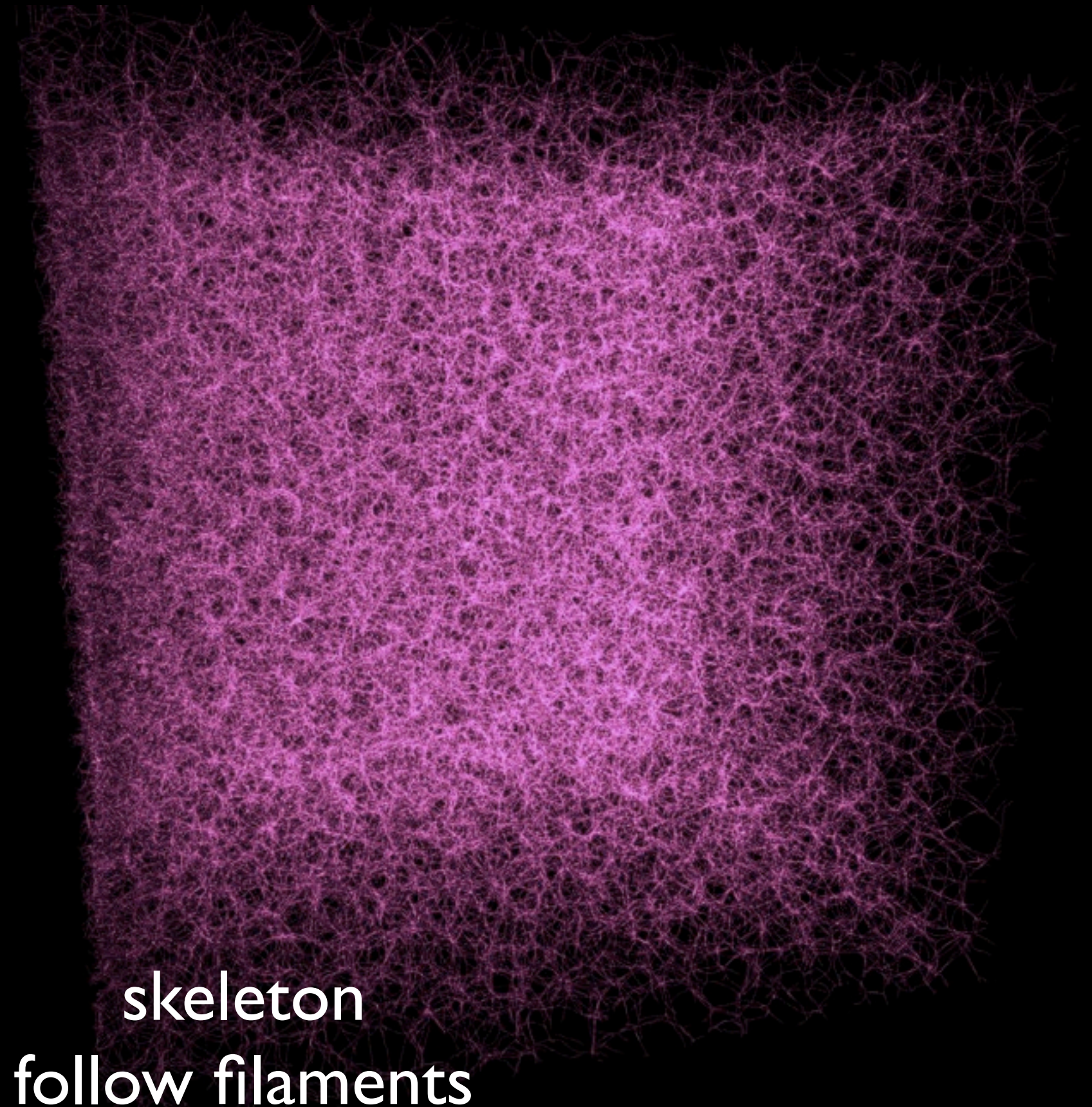
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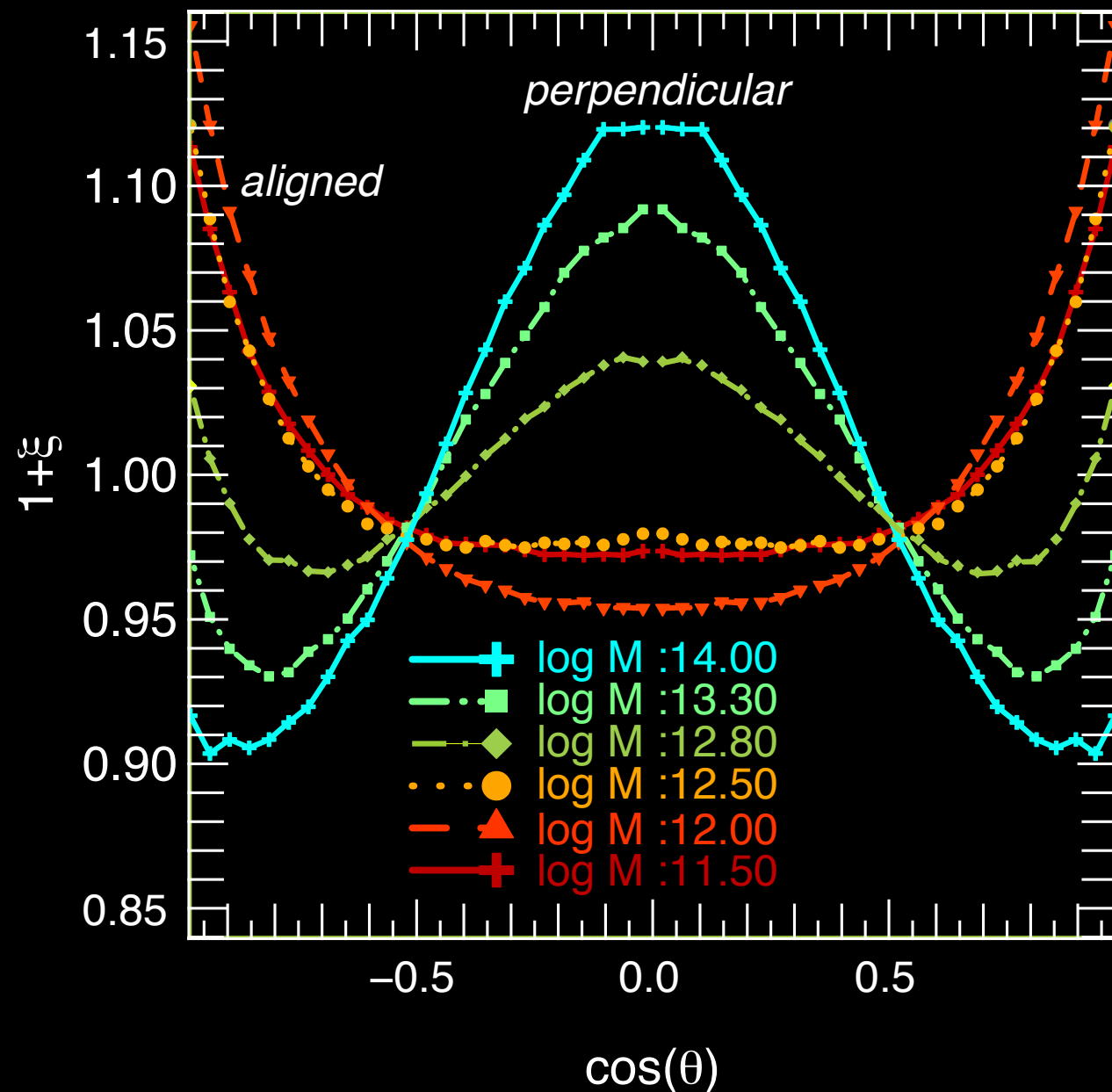
Excess probability of alignment between the spins and their host filament

mass transition:

$M < M_{\text{crit}}$: aligned

$M > M_{\text{crit}}$: perpendicular

Excess probability of alignment between the spins and their host filament



mass transition:

$$M_{\text{crit}} = 4 \cdot 10^{12} M_{\odot}$$

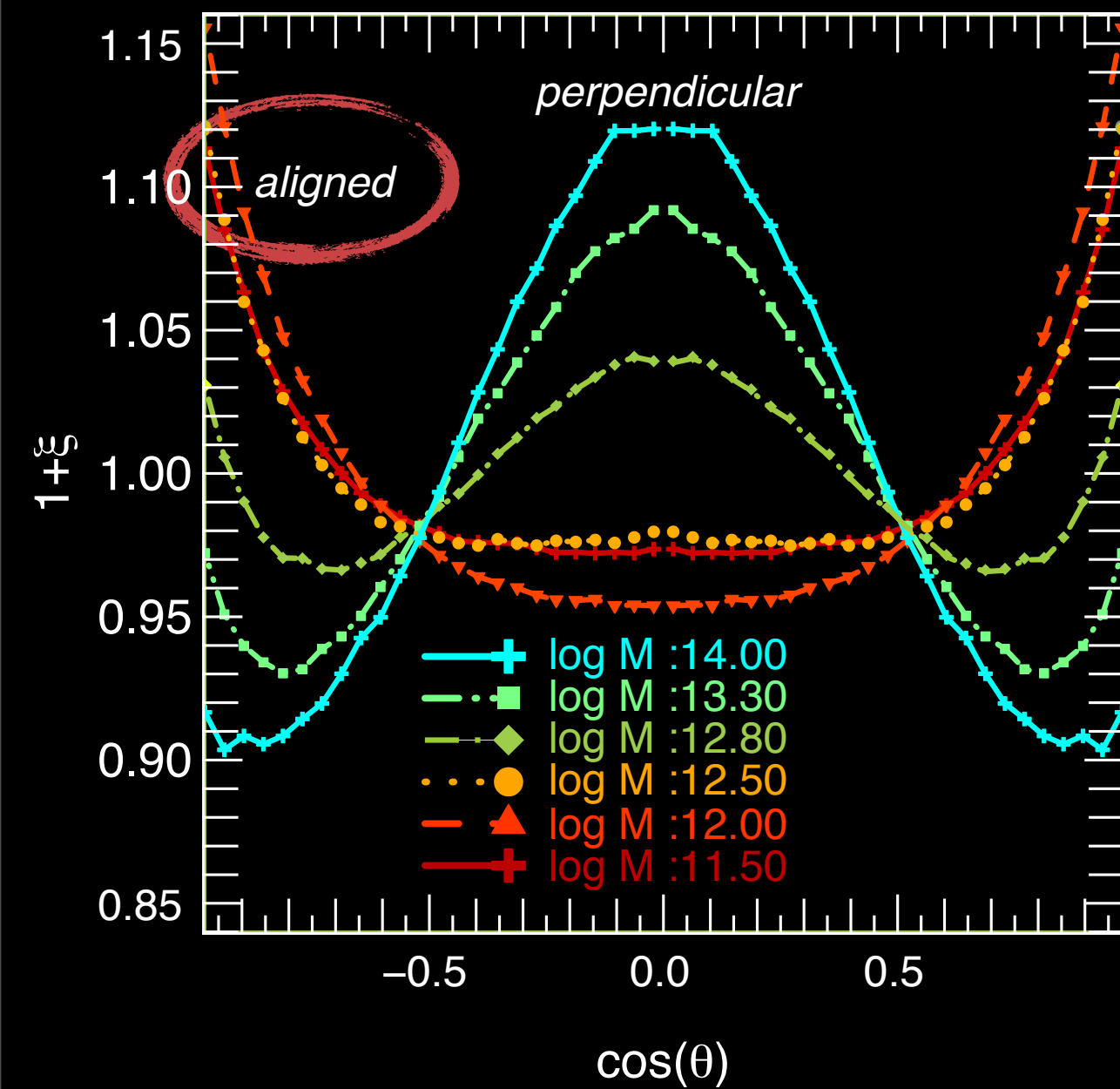
$M < M_{\text{crit}} : \text{aligned}$

$M > M_{\text{crit}} : \text{perpendicular}$

- In agreement with other numerical studies e.g Bailin & Steinmetz (2005); Aragon-Calvo et al. (2007,2013); Hahn et al. (2007); Paz et al. (2008)
- Confirmed by observations e.g Tempel et al 2013 using the SDSS data

Low-mass haloes:

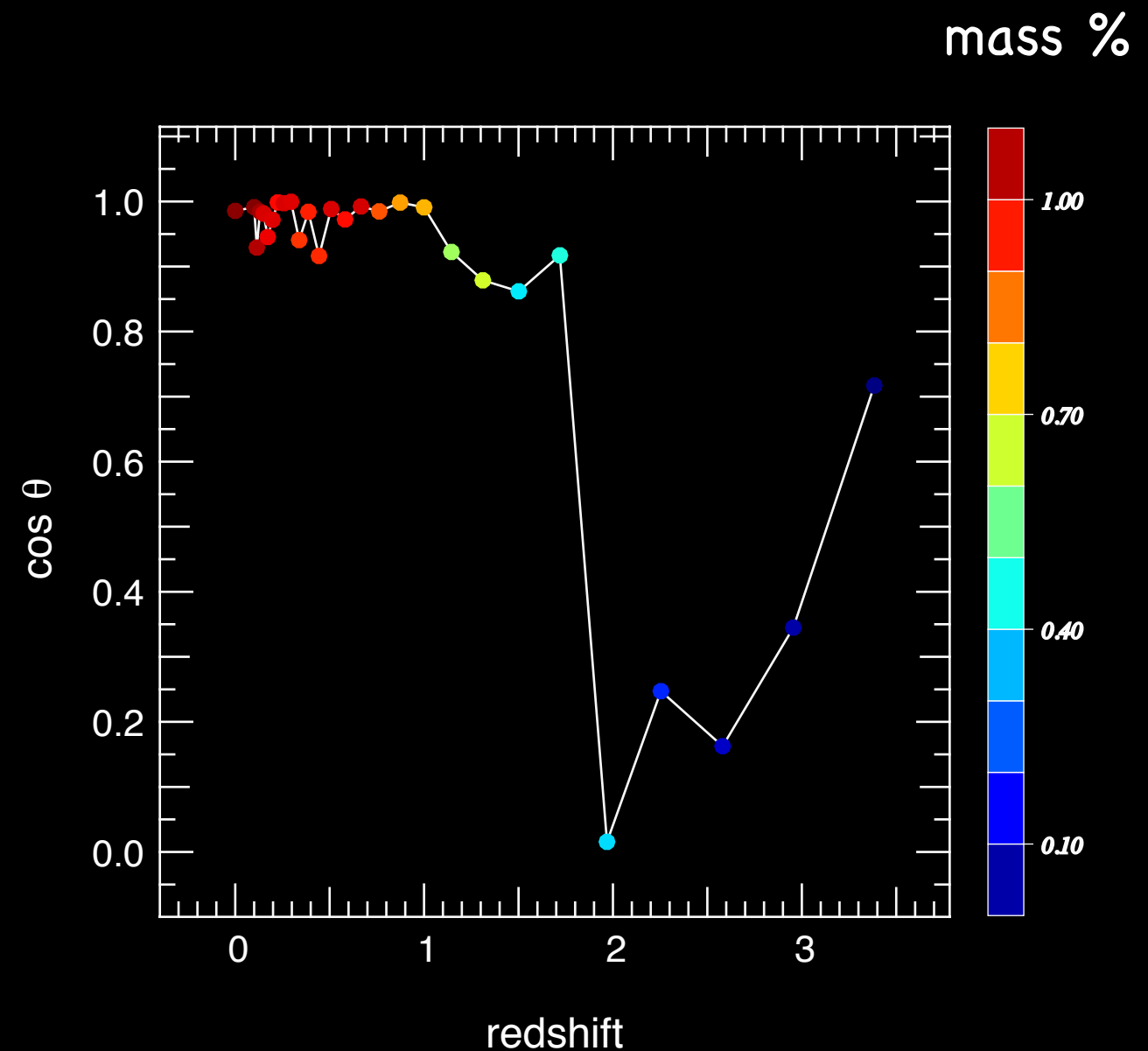
$$M < M_{\text{crit}}$$



$$M_{\text{crit}} = 4 \cdot 10^{12} M_{\odot}$$

mass transition:

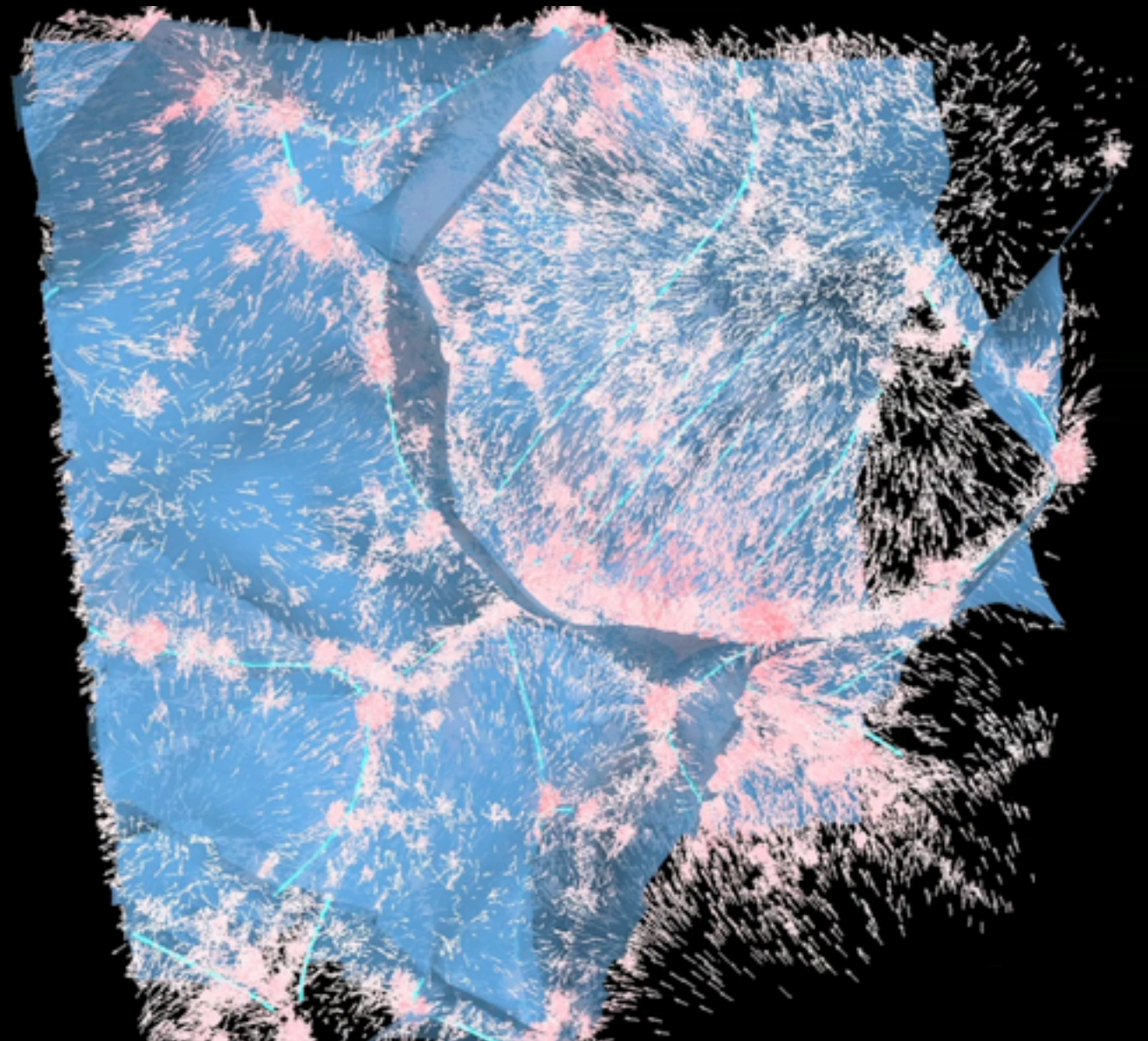
$M < M_{\text{crit}}$: aligned



How does the formation of the filaments
generate spin parallel to them?

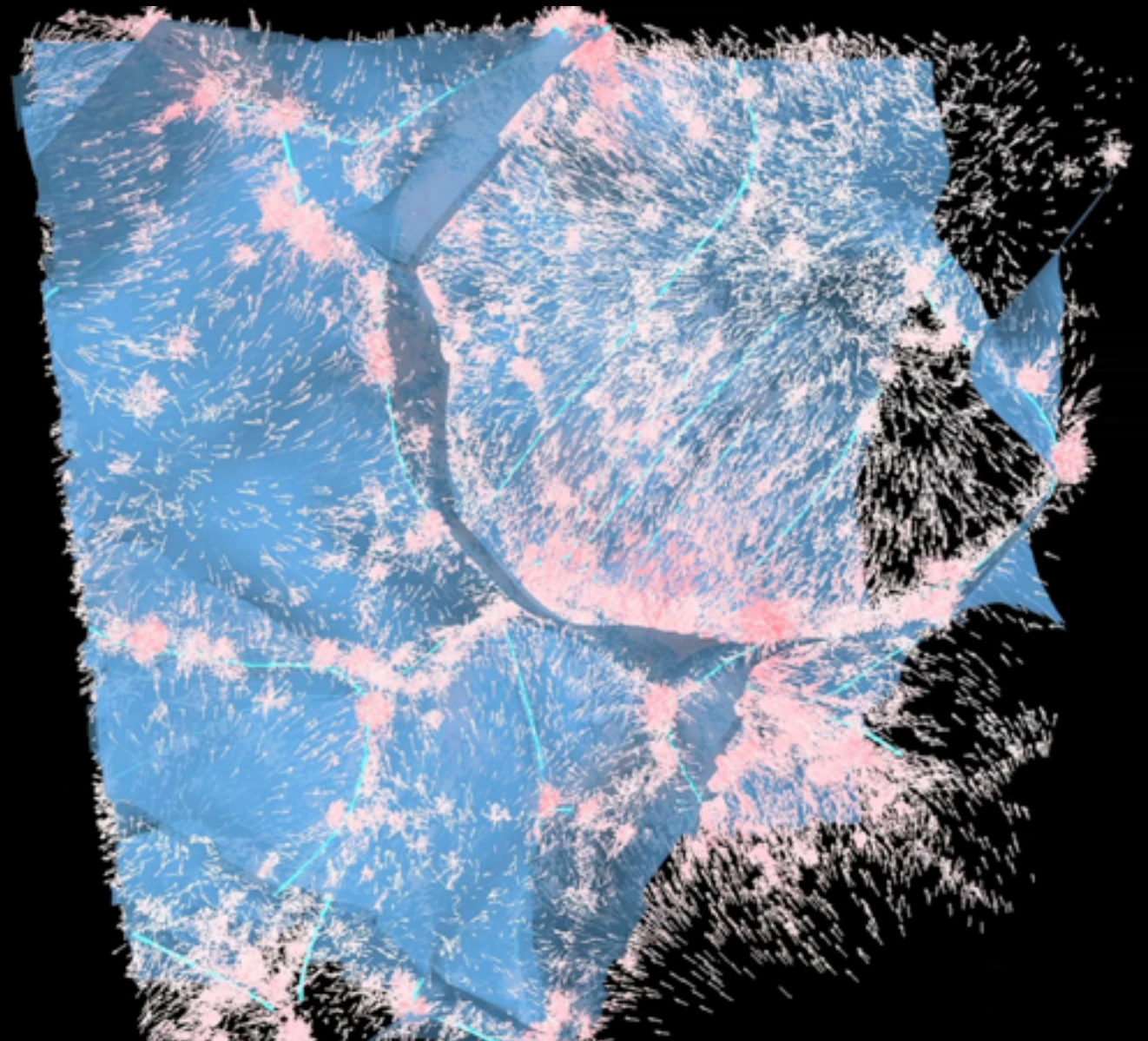
Voids/wall saddle
repel...

winding of walls



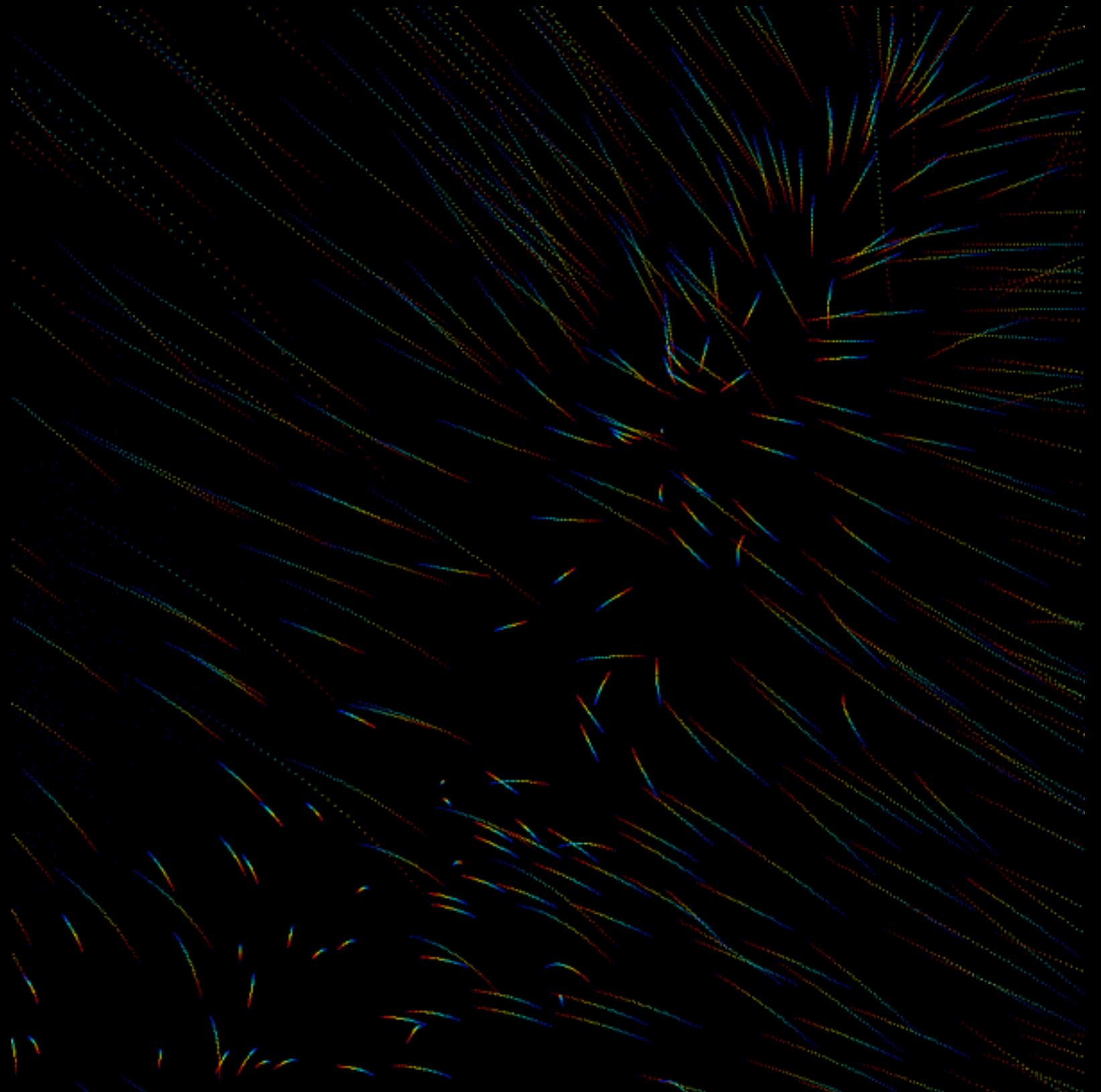
How does the formation of the filaments
generate spin parallel to them?

winding of walls



How does the formation of the filaments
generate spin parallel to them?

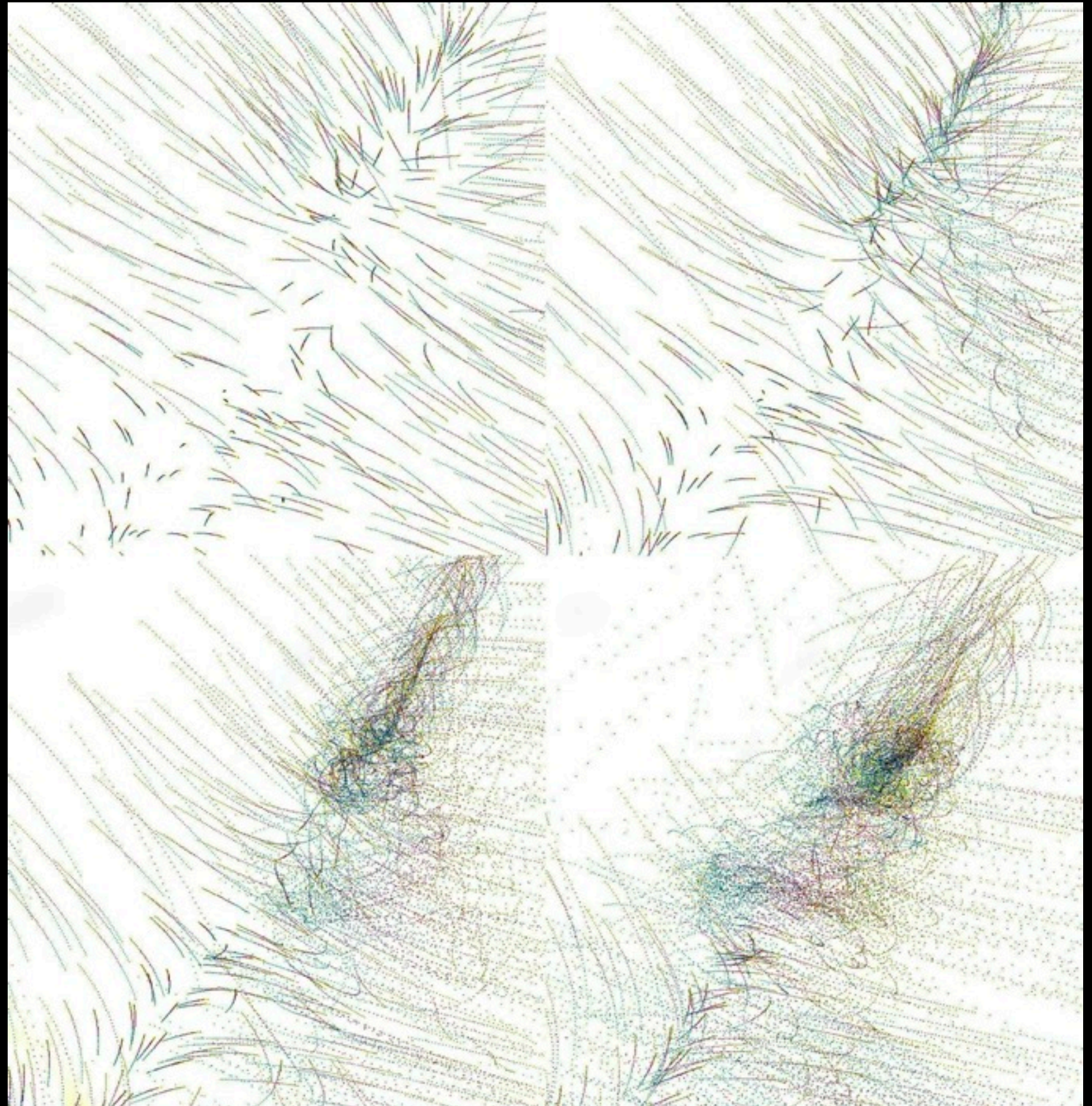
winding of walls



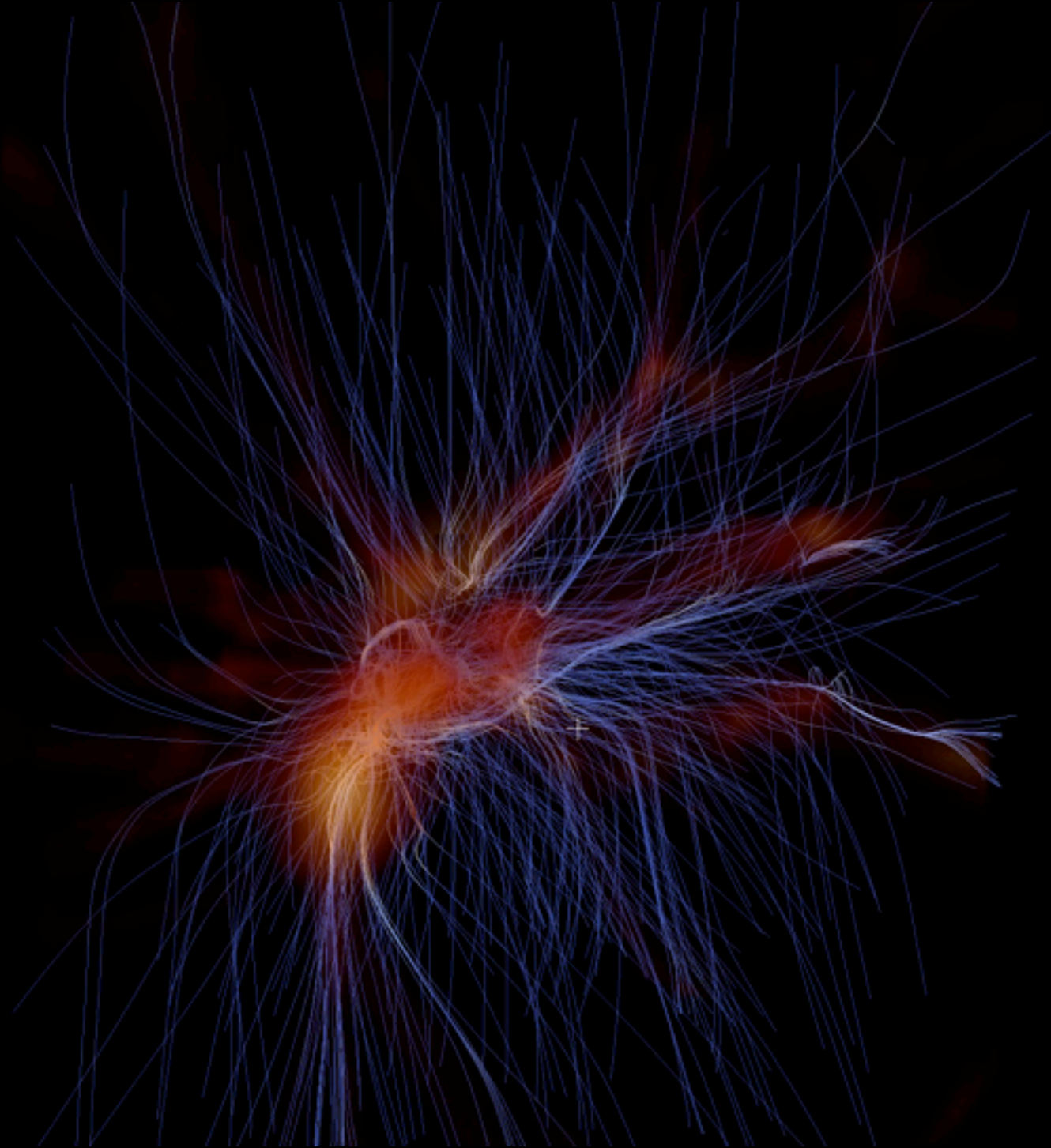
Winding of walls
onto filaments
generate spin
//
to filament

->

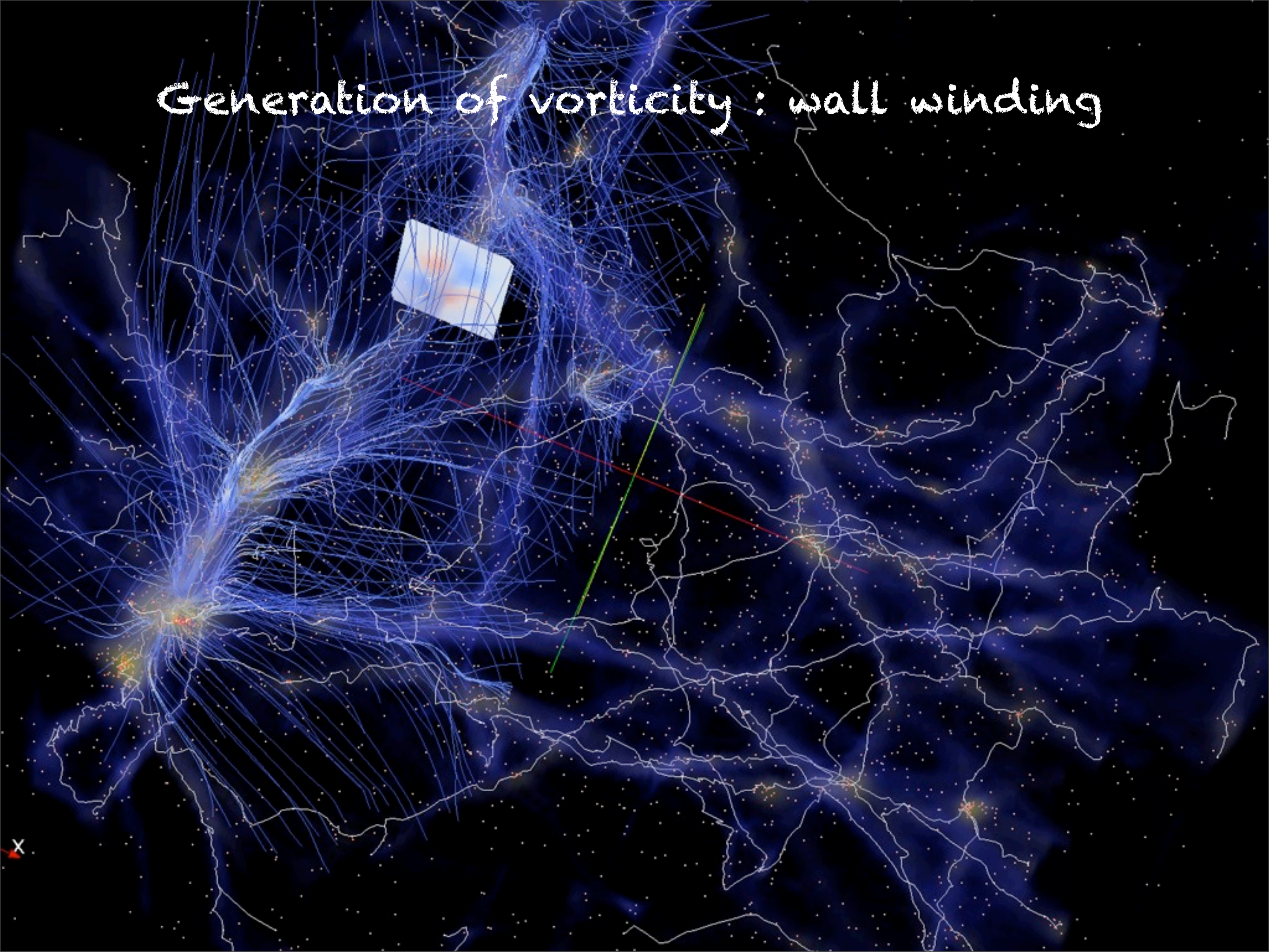
Vorticity?



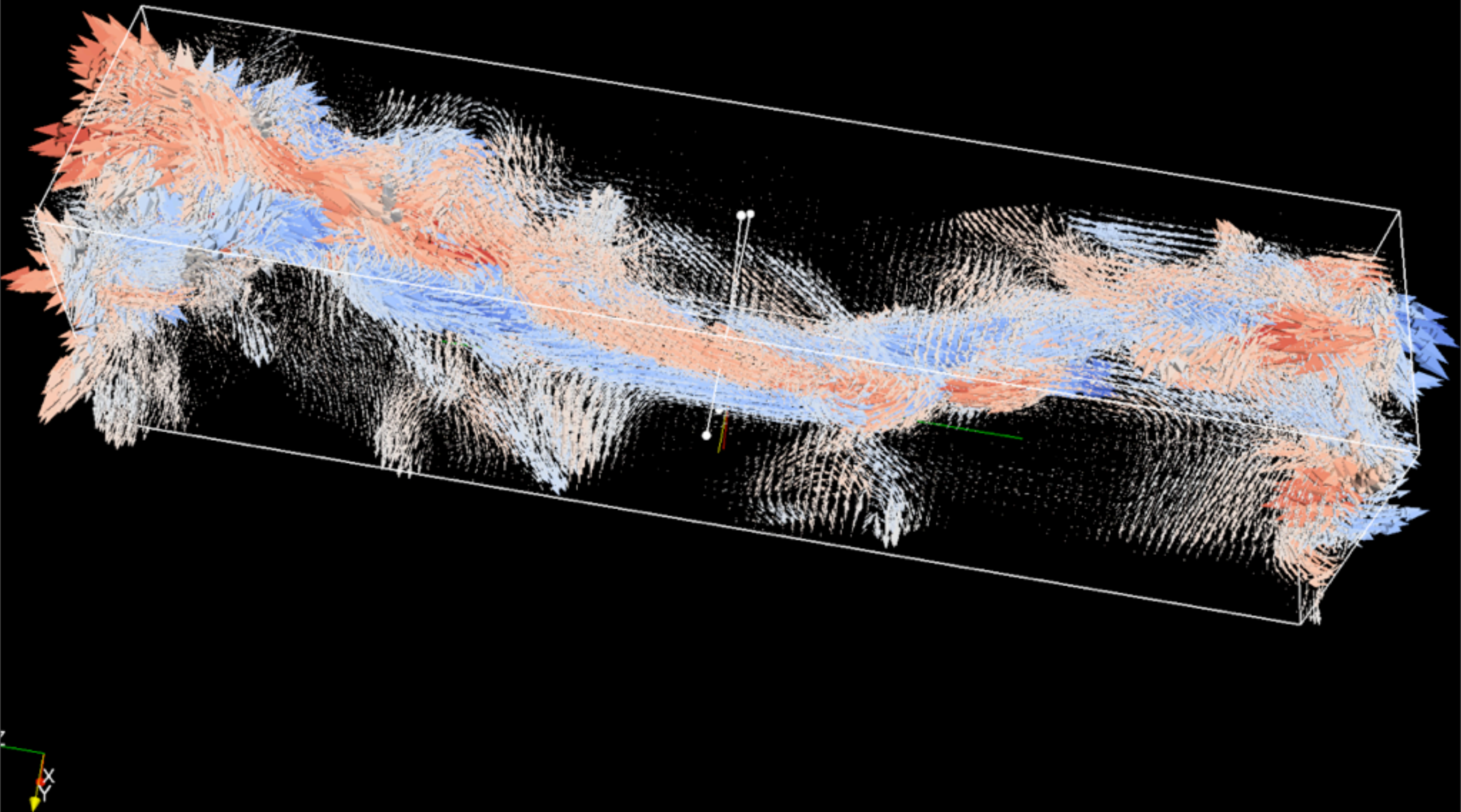
Focussing on main filament



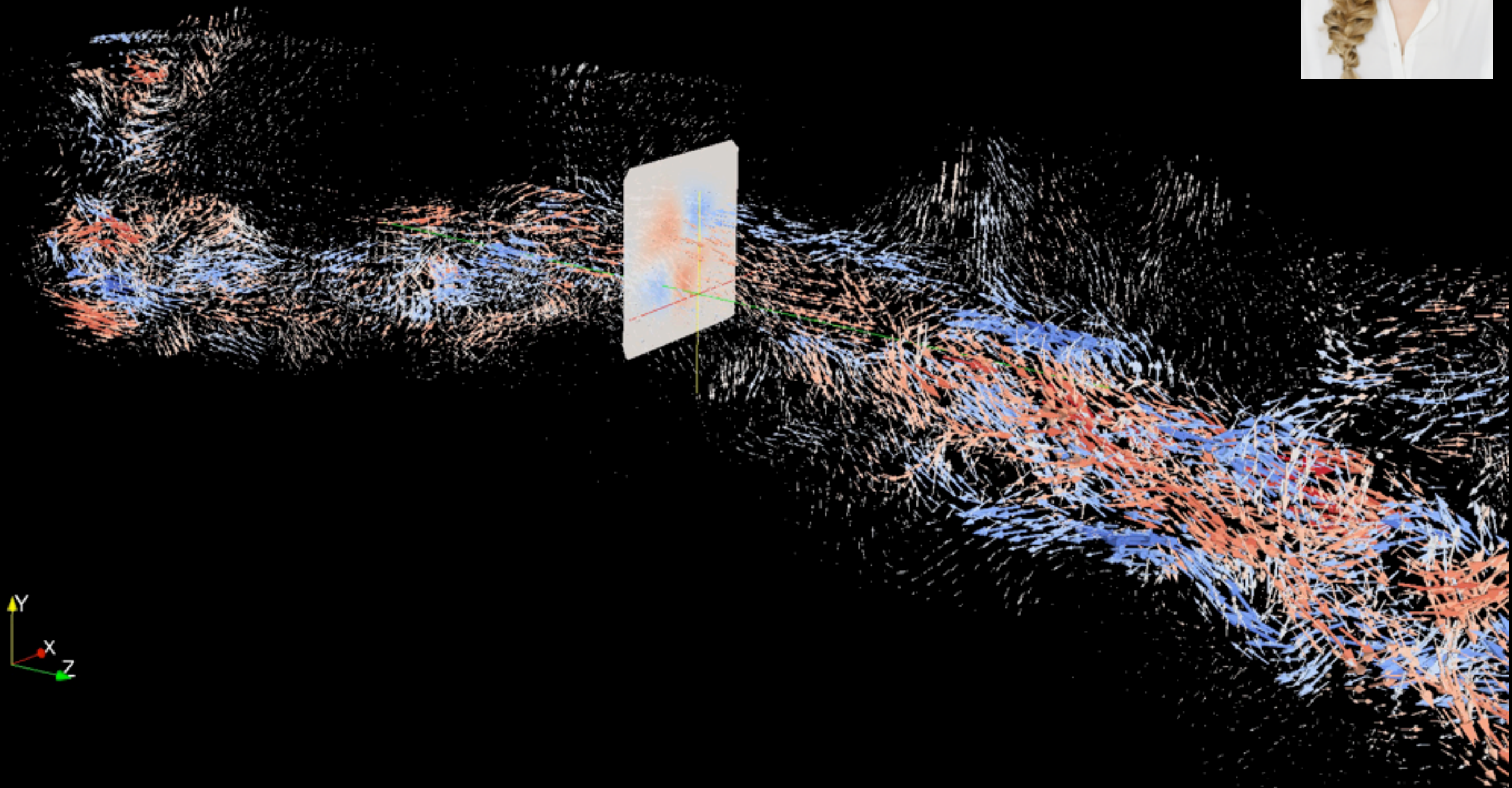
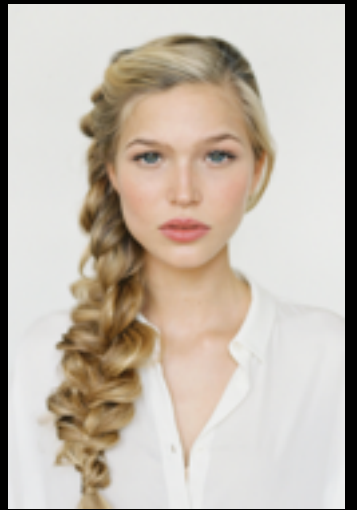
Generation of vorticity : wall winding



Alignment of vorticity with cosmic web



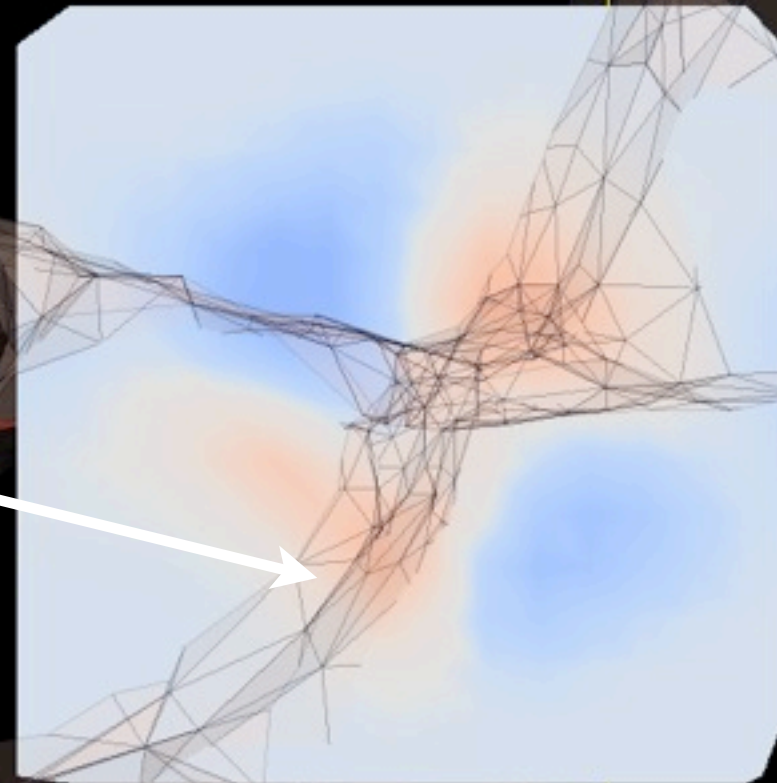
Alignment of vorticity with cosmic web



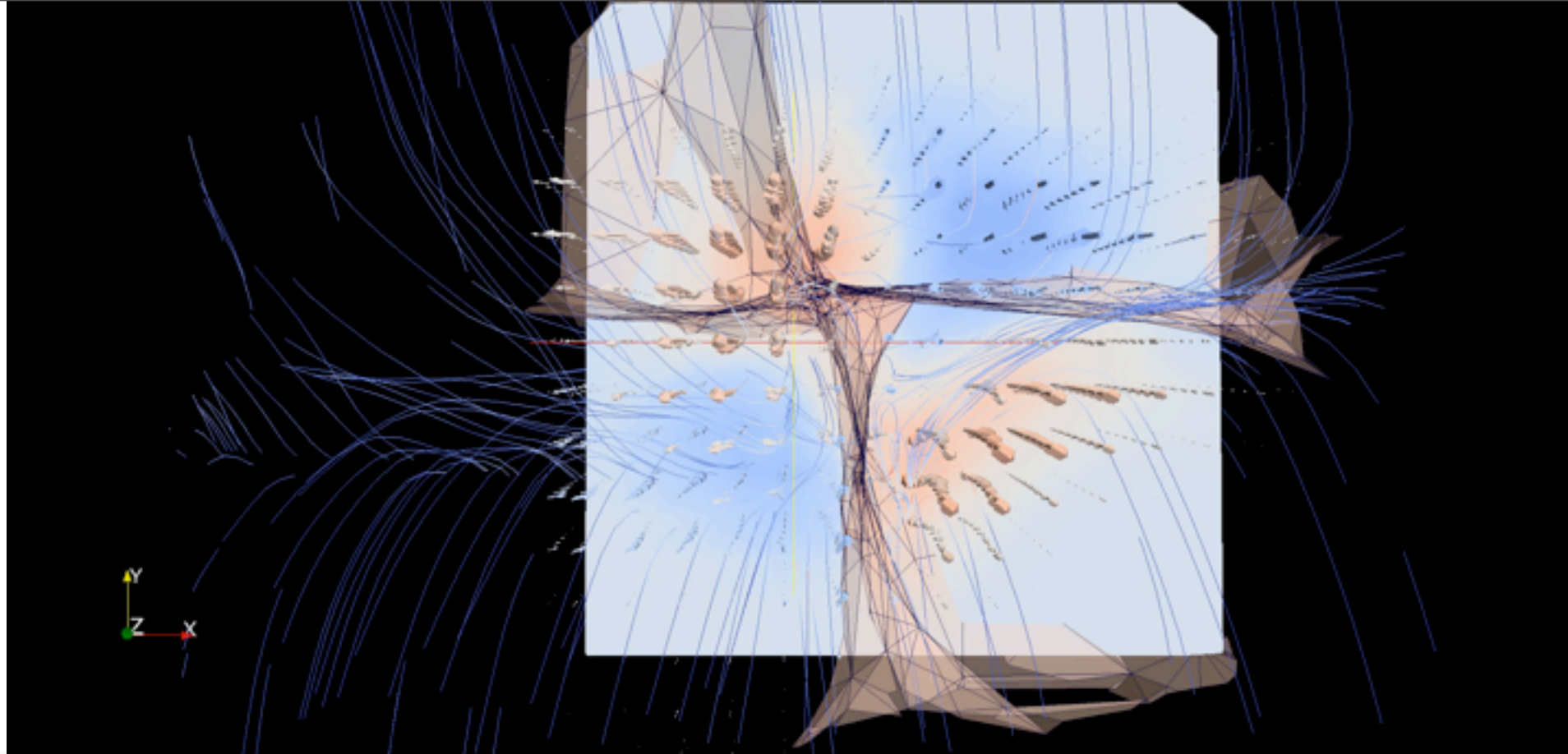
braids structure of vorticity.

vorticity cross section

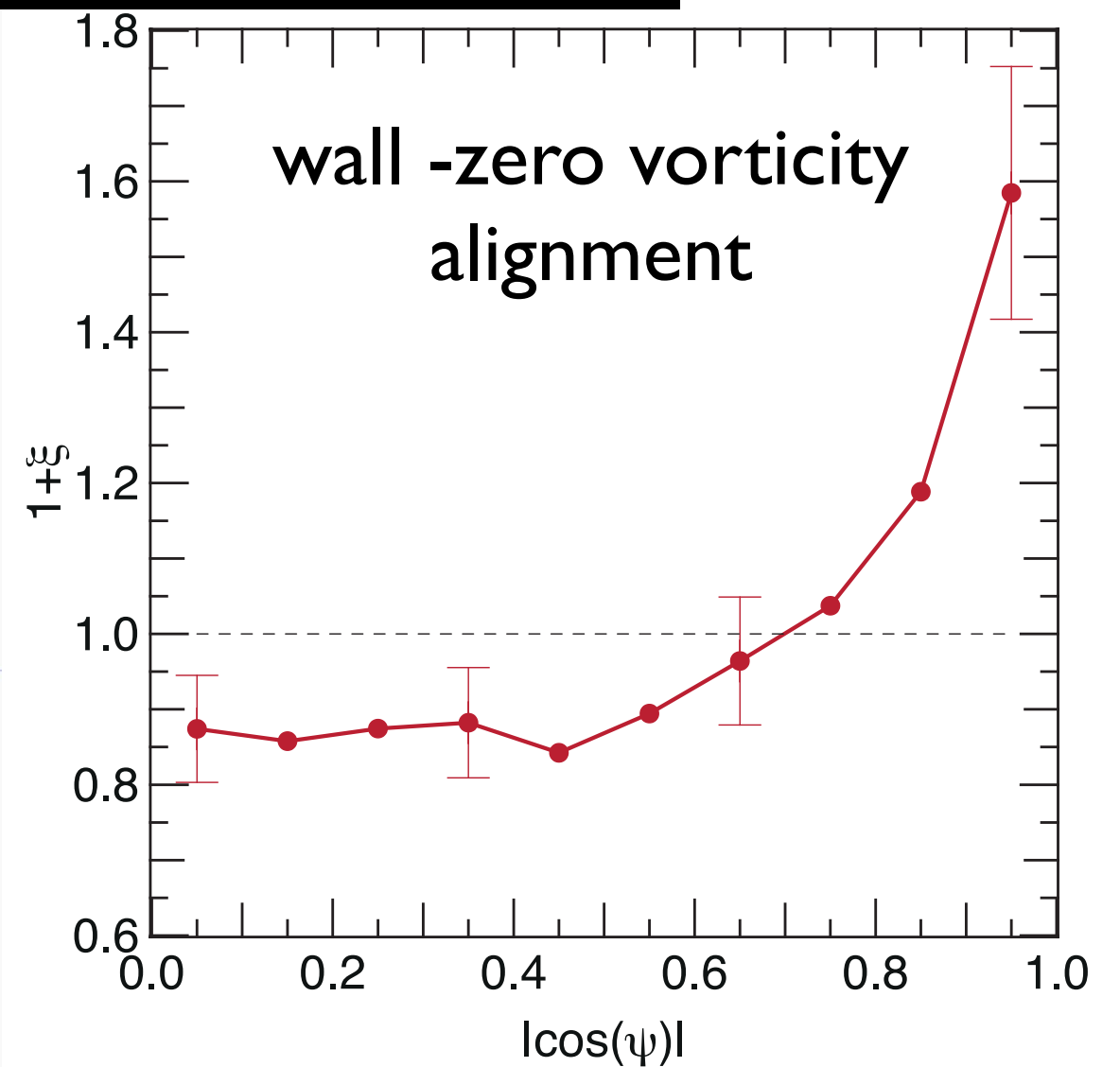
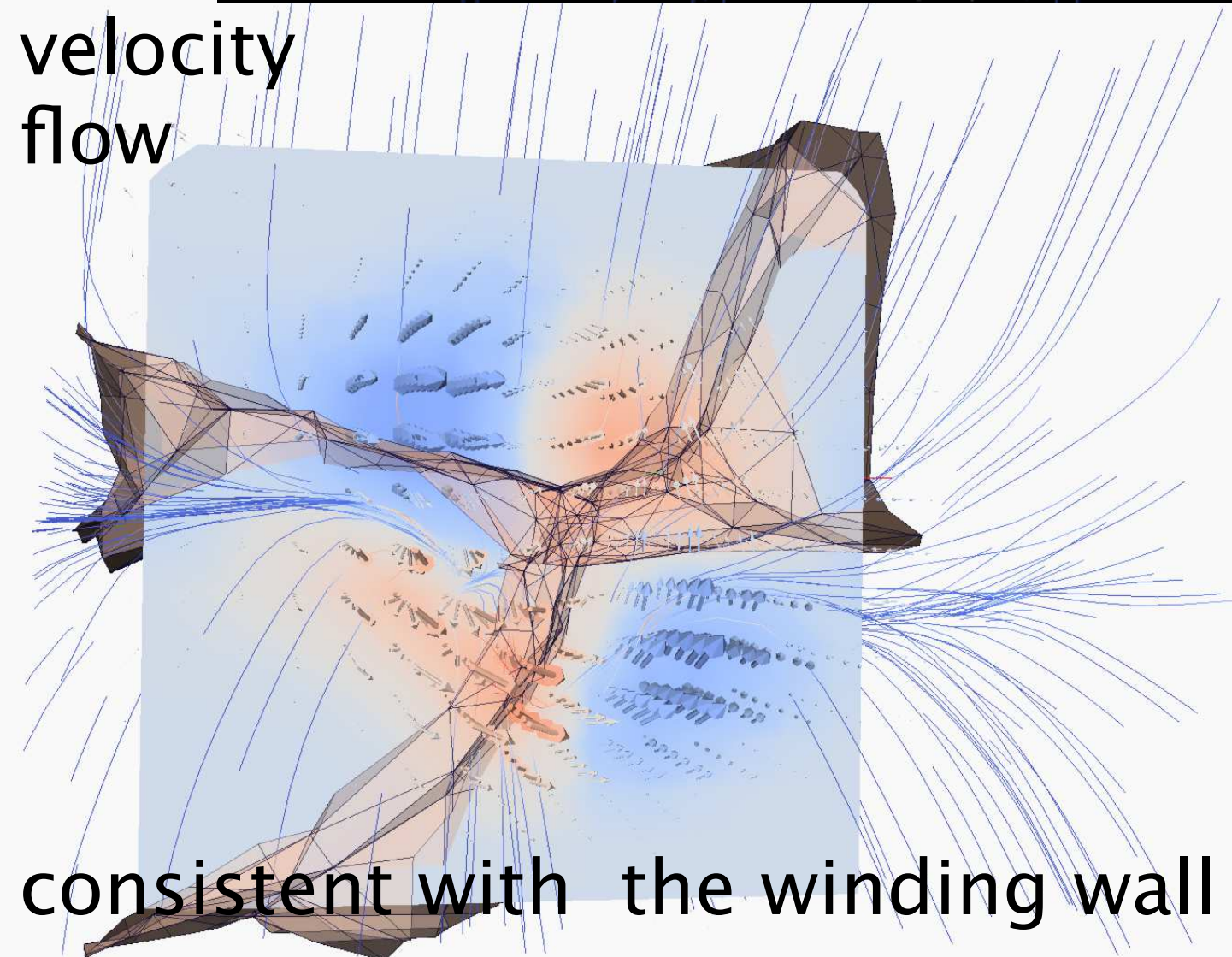
wall



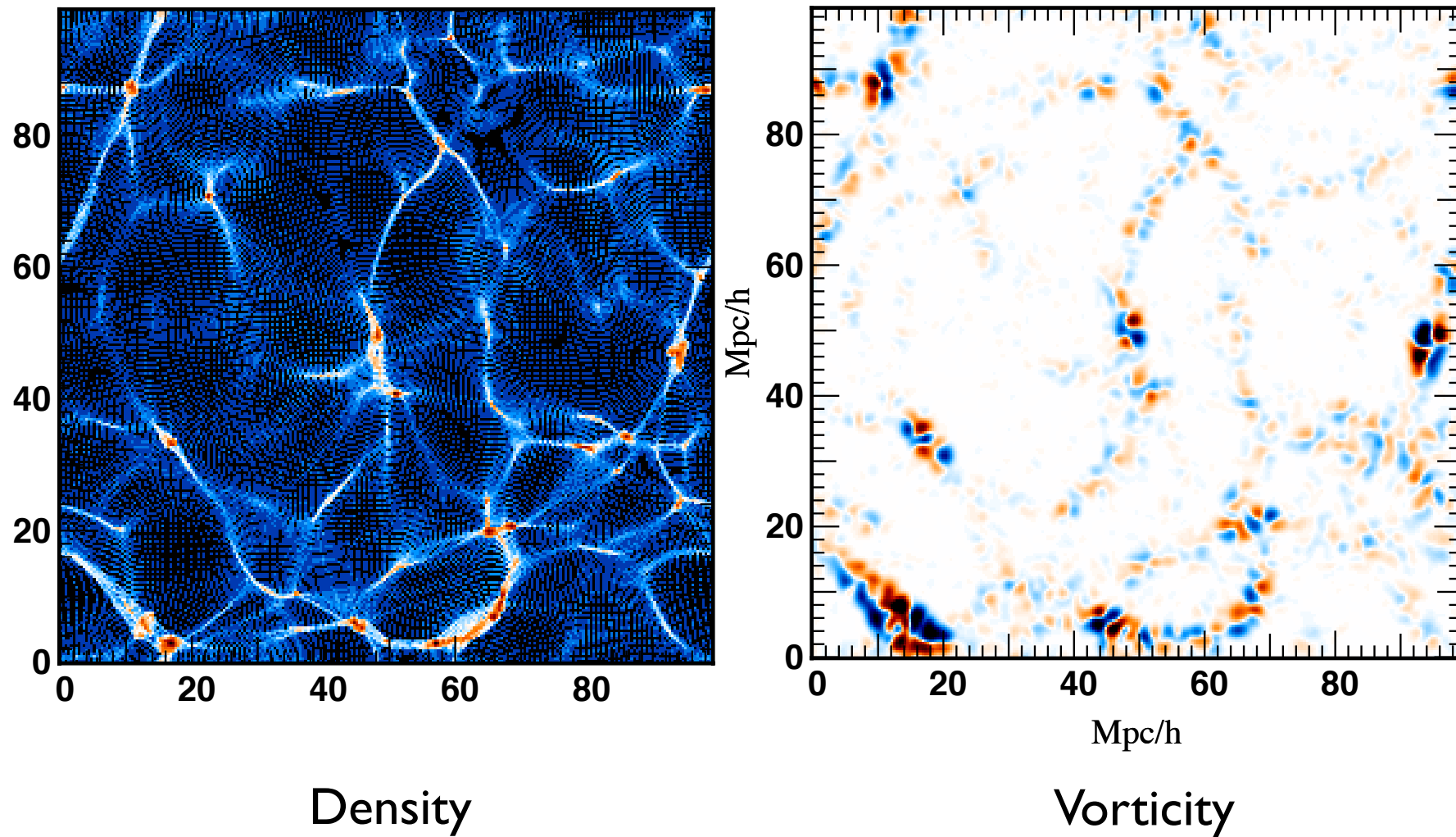
fully consistent with the winding wall scenario



velocity
flow

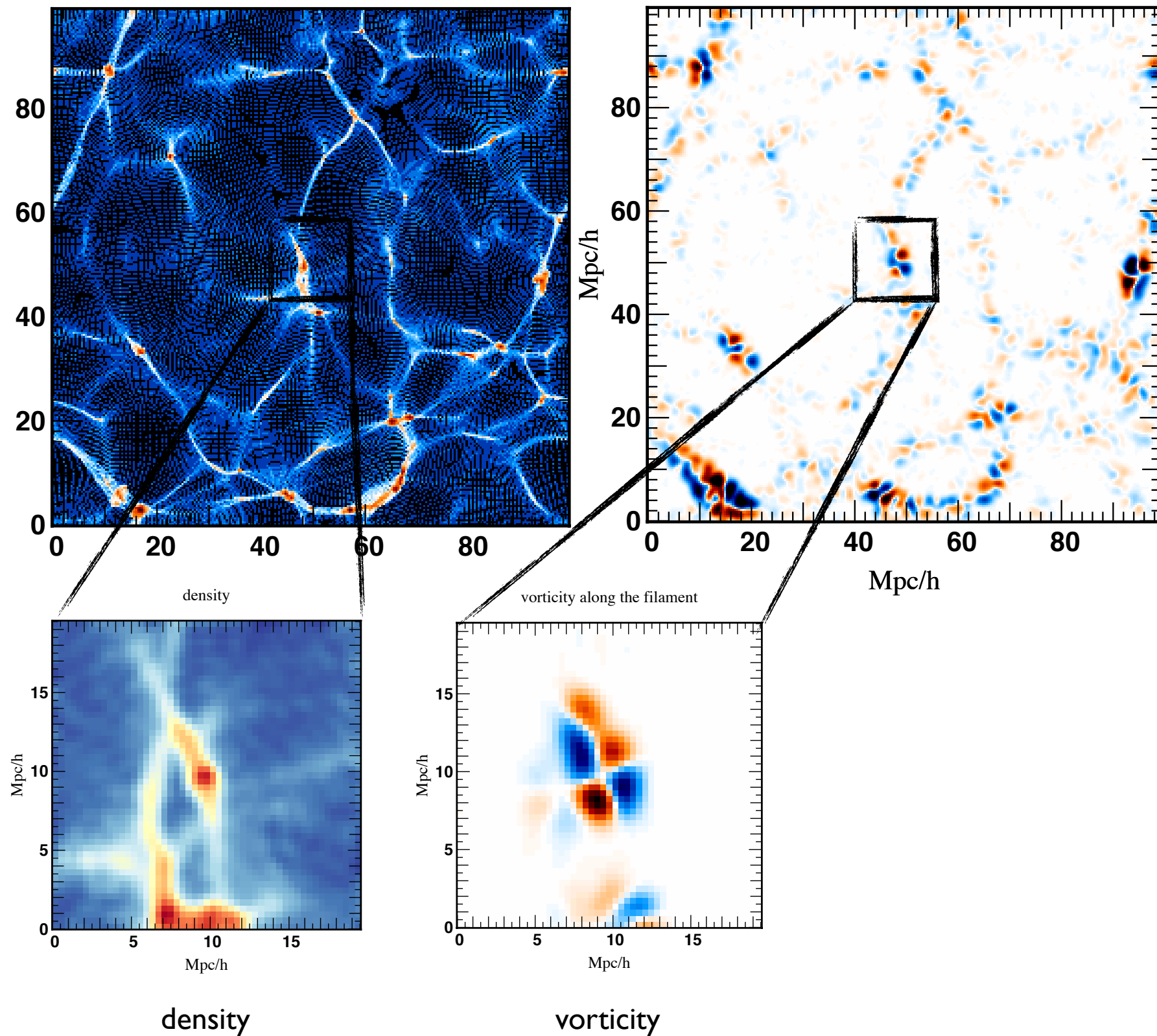


Locus of vorticity



Density & vorticity slice in a DM simulation.

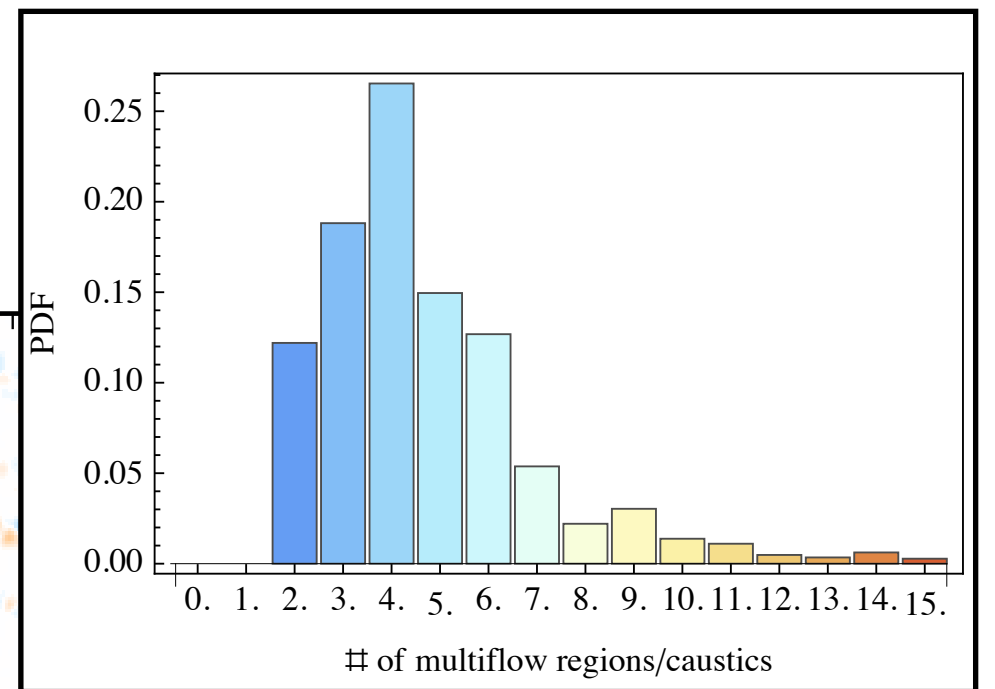
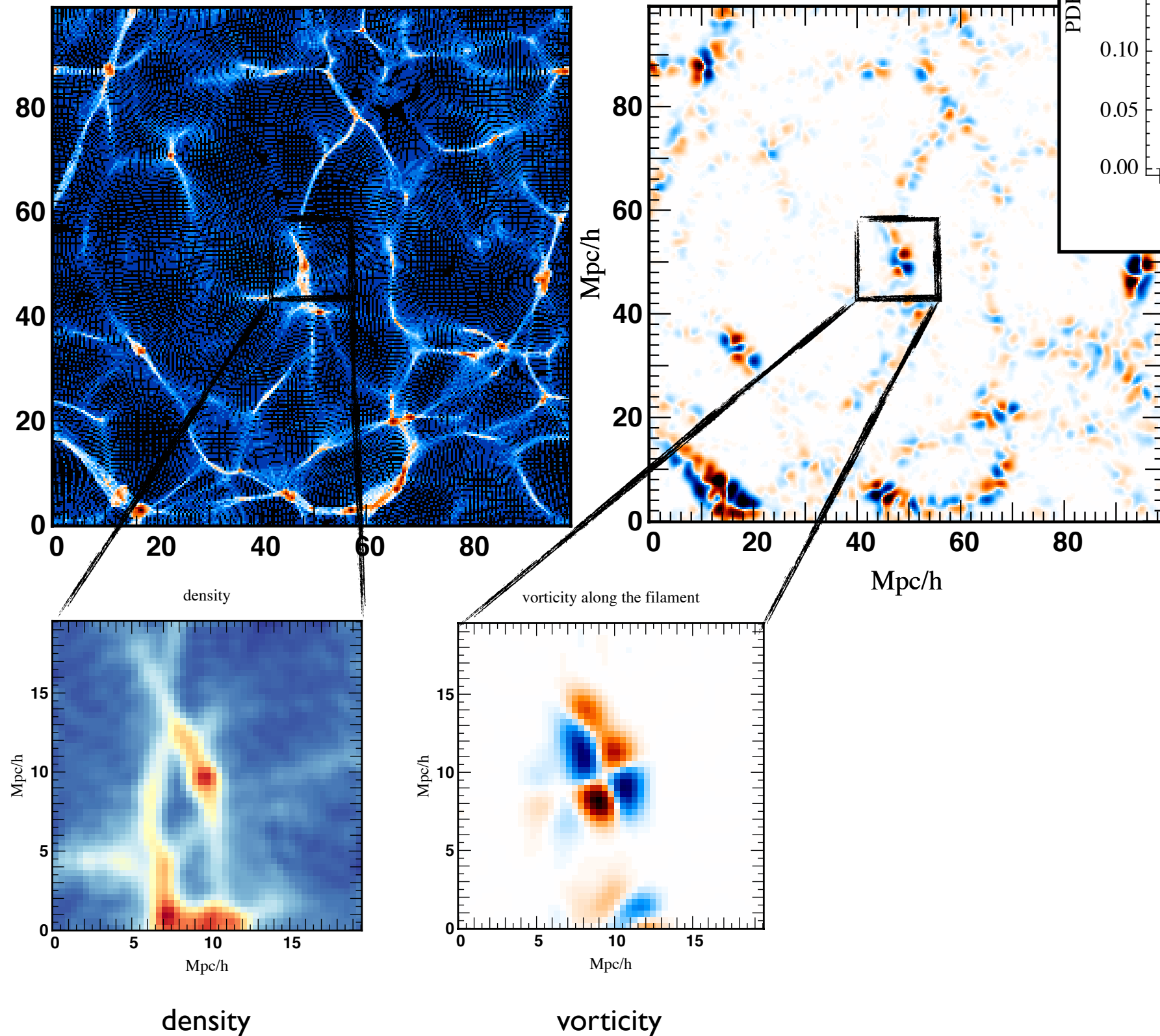
Locus of vorticity



- * vorticity is confined to filaments.
- * Vorticity is aligned with filament

region of constant polarity are smaller !

Locus of vorticity

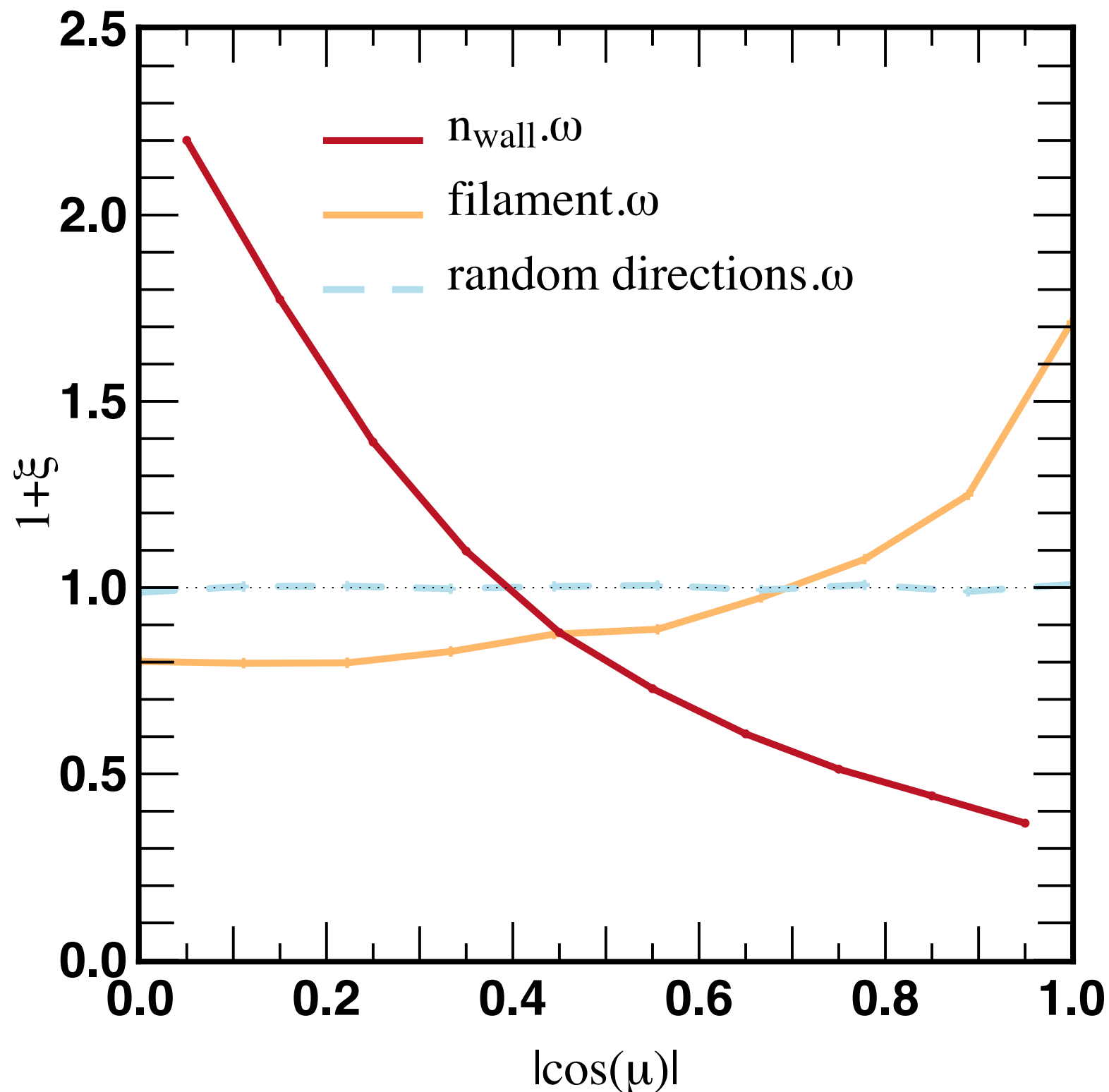


- * vorticity is confined to filaments.
- * Vorticity is aligned with filament
- * caustics are \sim quadripolar

region of constant polarity are smaller !

Alignment of vorticity with cosmic web

Alignment of vorticity and cosmic web



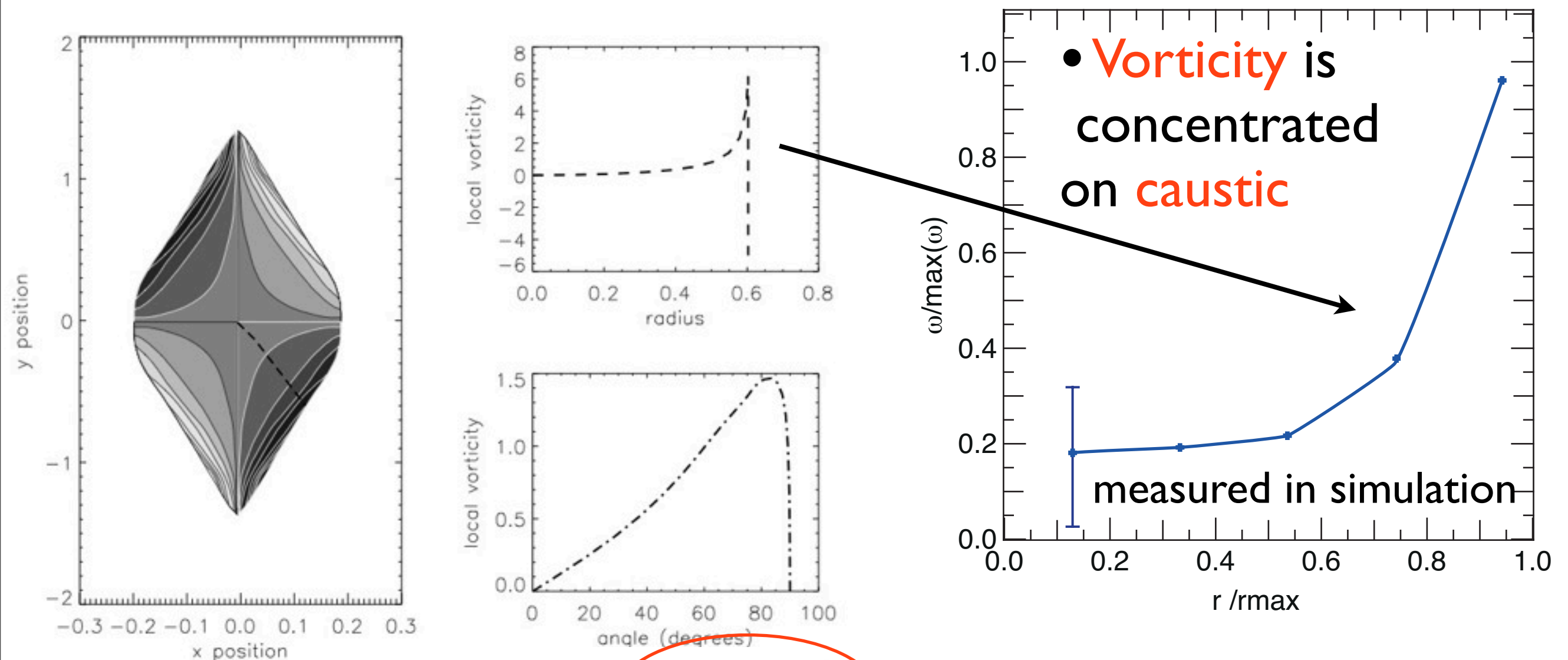
Vorticity is aligned with filaments

In walls, vorticity is perpendicular to the normal of walls

Cross section of vorticity in caustic

$$\omega_k(\mathbf{x}) = \sum_{i,j} \epsilon^{k,j,i} \frac{\partial \mathbf{u}_i(\mathbf{x})}{\partial x_j} = \sum_{i,j} \epsilon^{k,j,i} \left(\left[\sum_{\text{flow } s} \frac{\partial \rho(\mathbf{q}_s)}{\partial \mathbf{q}_{sl}} (D^{-1})_{j,l} \mathbf{u}_i(\mathbf{q}_s) \right] \left[\sum_{\text{flow } s} \rho(\mathbf{q}_s) \right] - \left[\sum_{\text{flow } s} \rho(\mathbf{q}_s) \mathbf{u}_i(\mathbf{q}_s) \right] \times \right. \\ \left. \left[\sum_{\text{flow } s} \frac{\partial \rho(\mathbf{q}_s)}{\partial \mathbf{q}_{sl}} (D^{-1})_{j,l} \right] \right) / \left[\sum_{\text{flow } s} \rho(\mathbf{q}_s) \mathbf{u}_i(\mathbf{q}_s) \right]^2,$$

Pichon Bernardeau, 1999



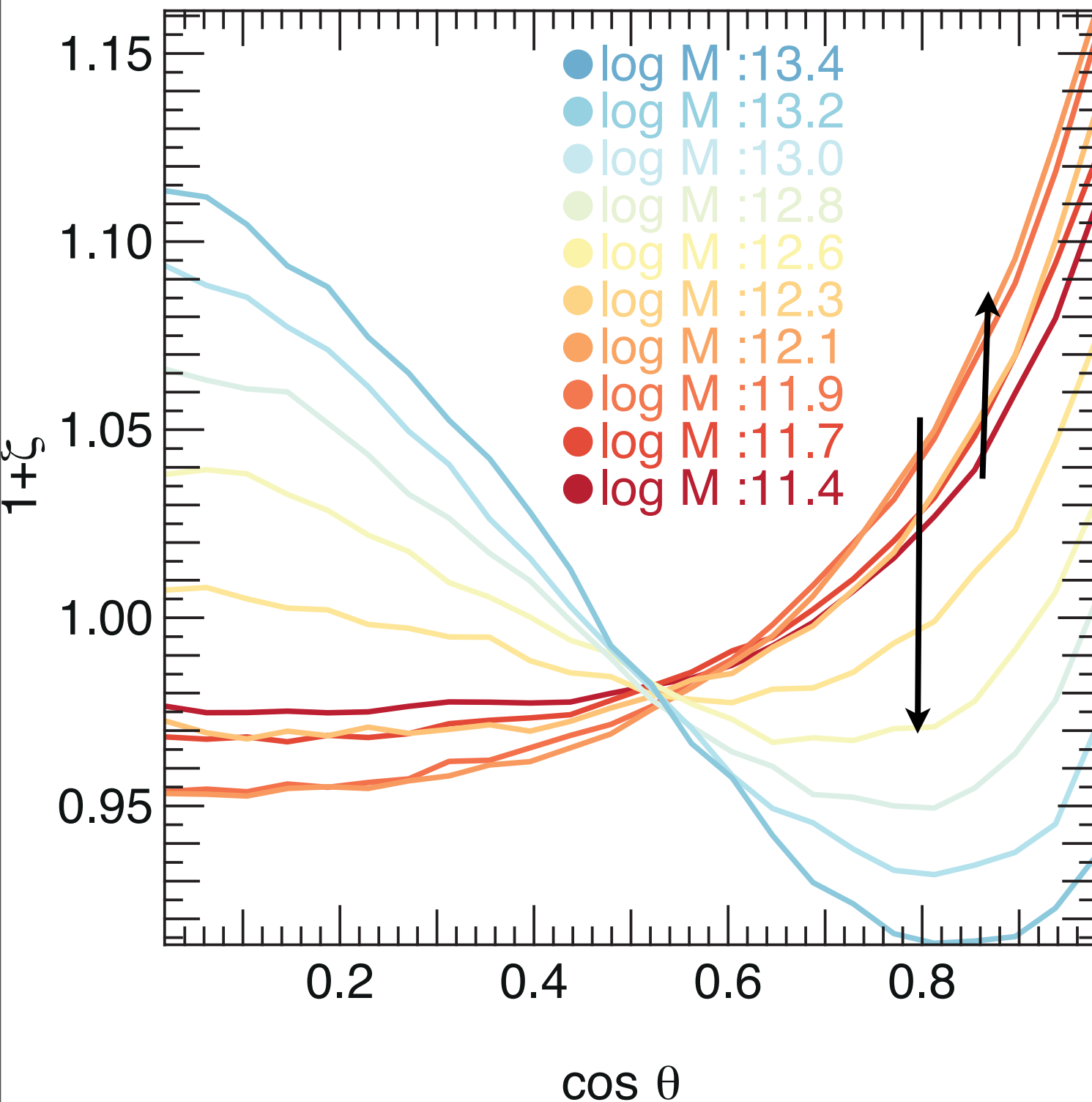
$$\omega(\mathbf{x}) \approx -\rho(\mathbf{q}_3) \sqrt{\eta} (\mathbf{x}_0 - \mathbf{x})_{\perp}^{-1/2} (\mathbf{u}(\mathbf{q}_3) - \mathbf{u}(\mathbf{q}_0))_{\parallel} / 2.$$

Revisit

Alignment of spin with cosmic web

vorticity
max on edge

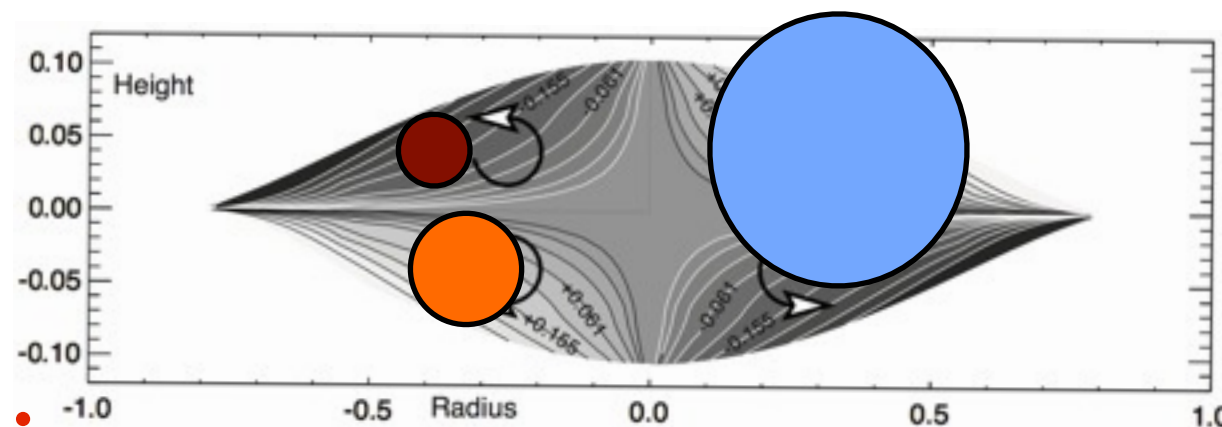
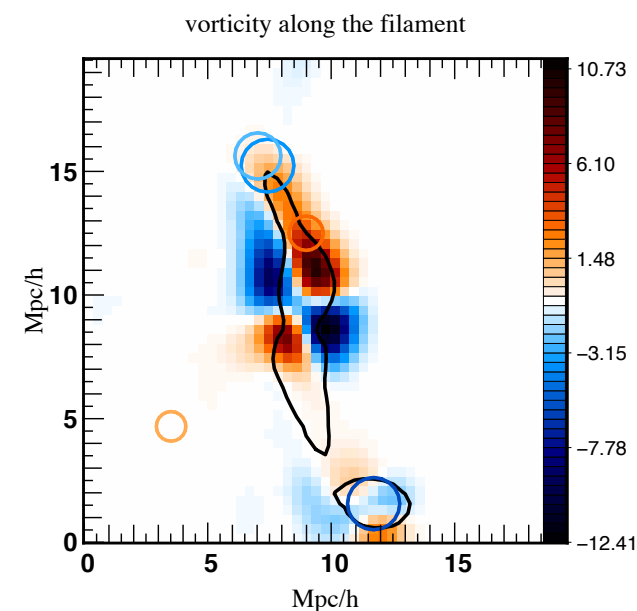
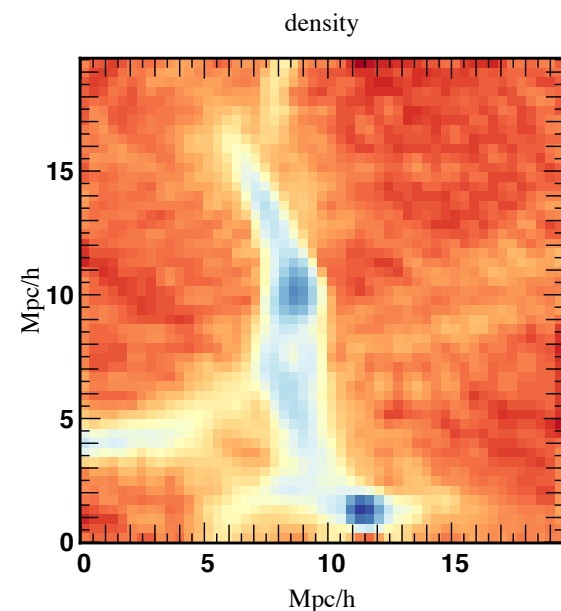
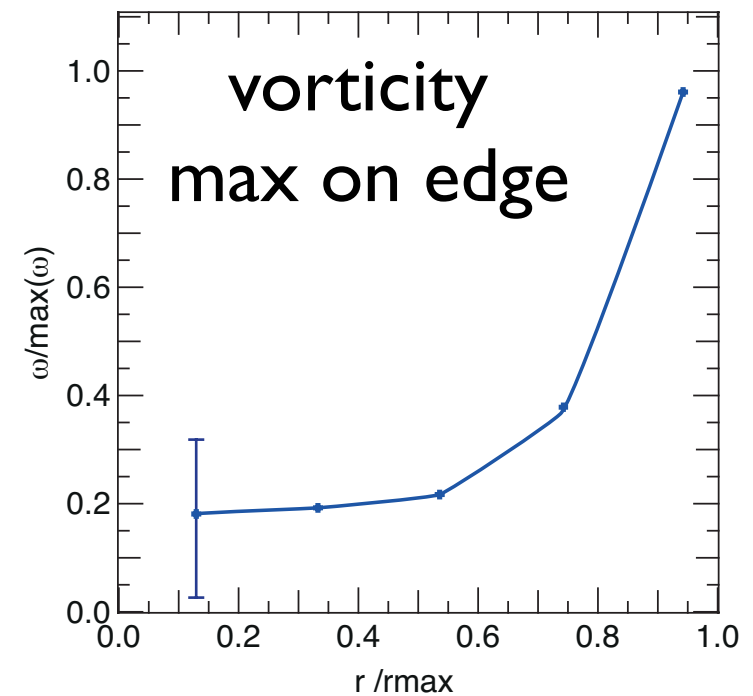
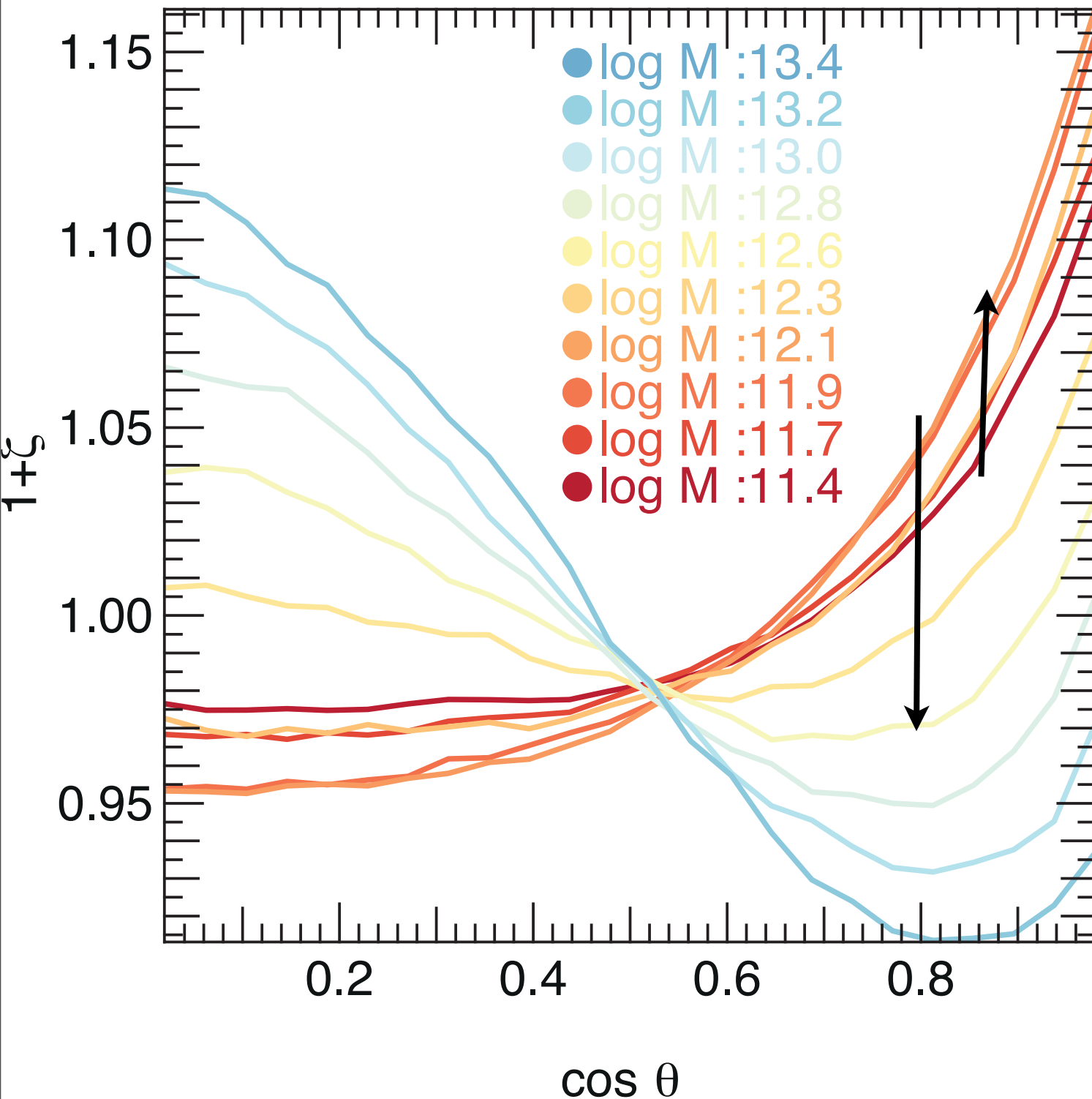
Spin alignment first INCREASES with mass !!!



Revisit

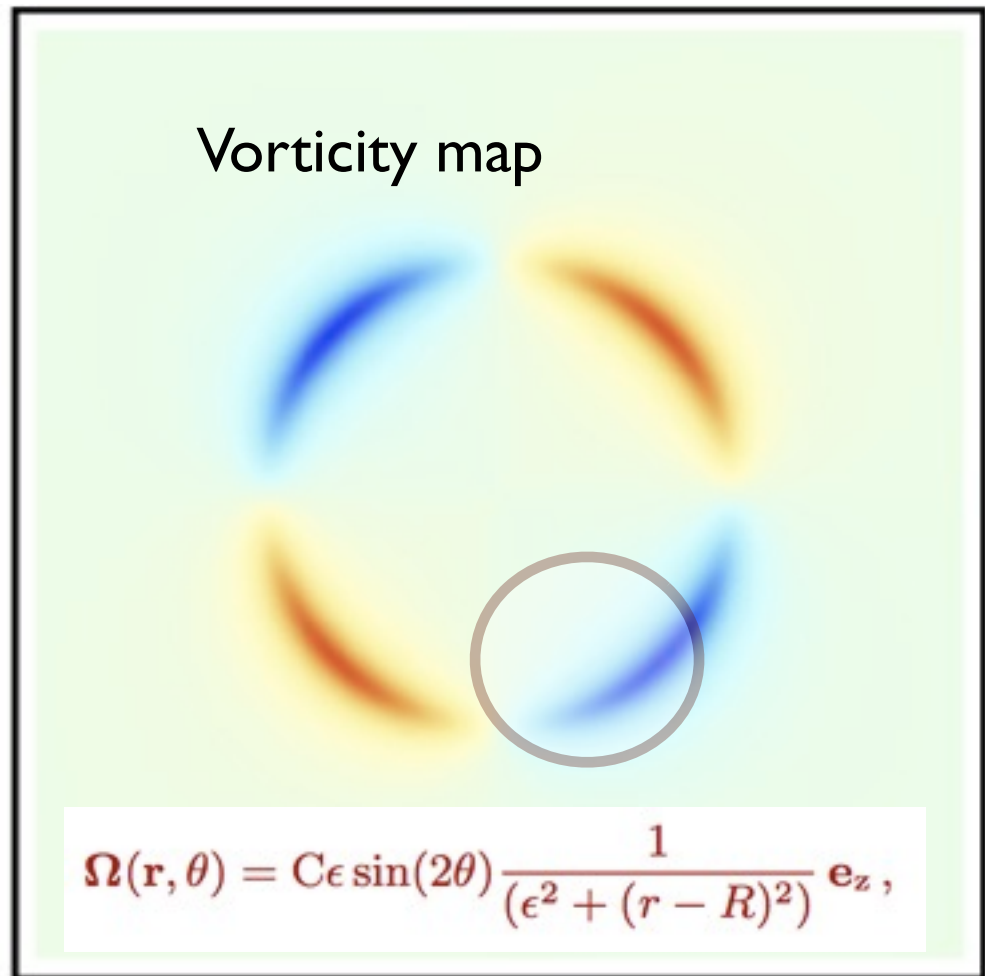
Alignment of spin with cosmic web

Spin alignment first INCREASES with mass !!!

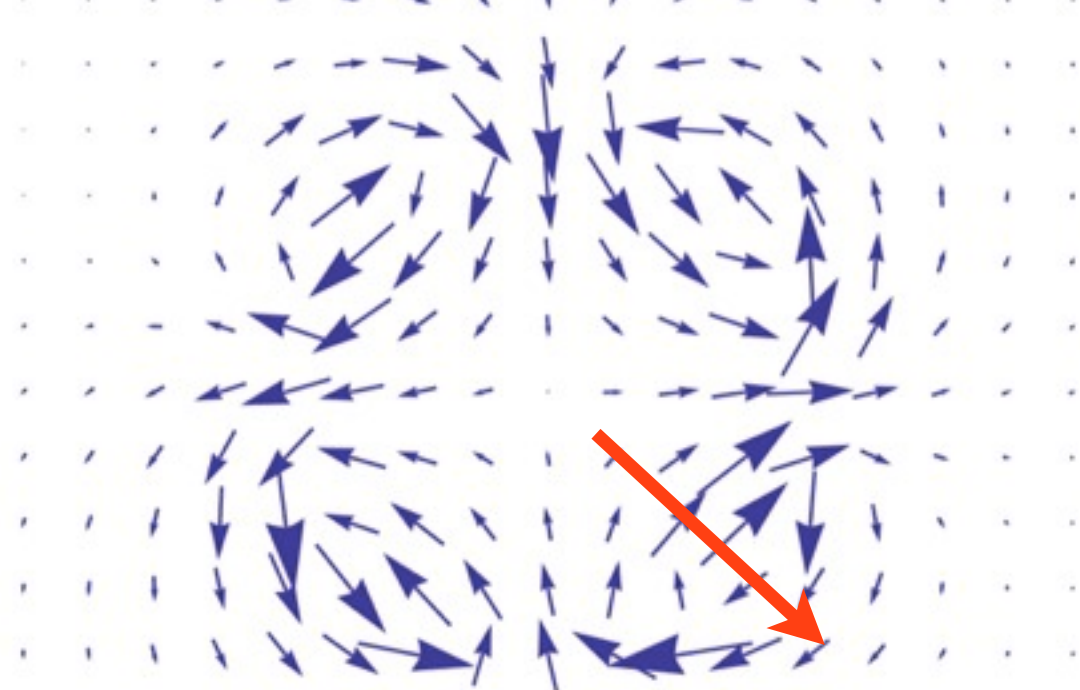


Spin flip imposed by caustic size

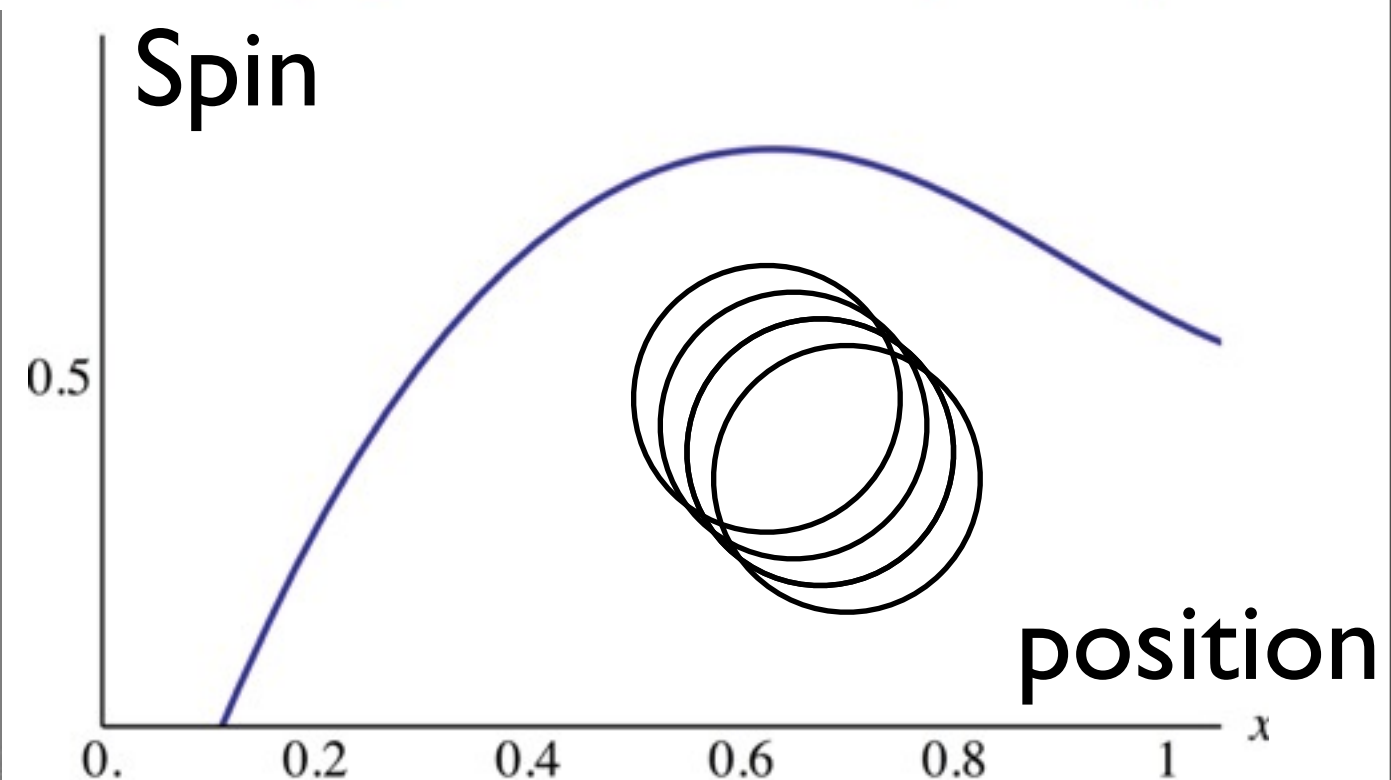
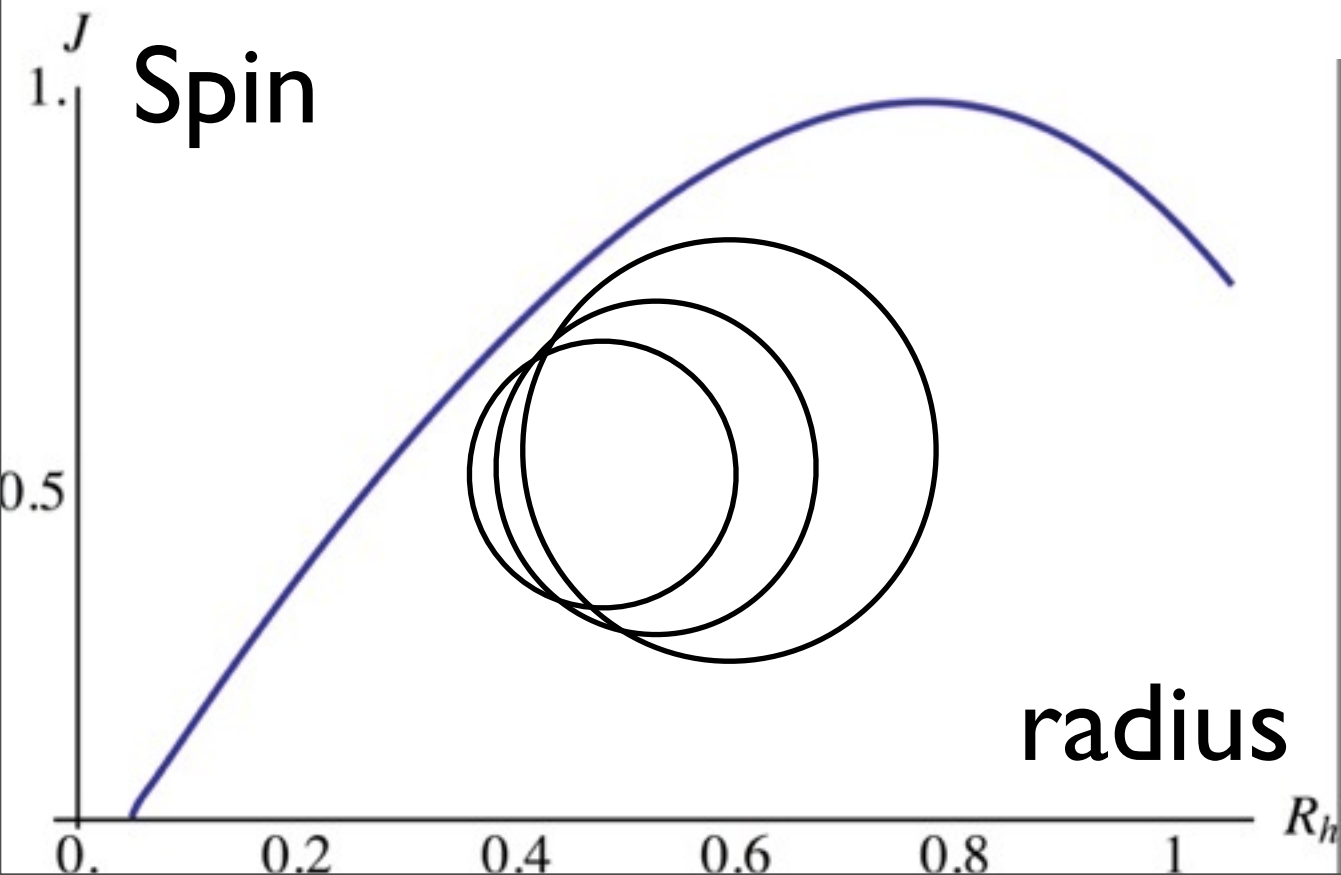
A toy model for vorticity section & spin flip



Velocity flow



$$\mathbf{v}(\mathbf{r}, \theta) = \frac{1}{4\pi} \int_{\mathcal{V}} \nabla \times \boldsymbol{\Omega}(\mathbf{r}', \theta') \frac{1}{|\mathbf{r} - \mathbf{r}'|} d\mathcal{V}.$$



Horizon-AGN simulation

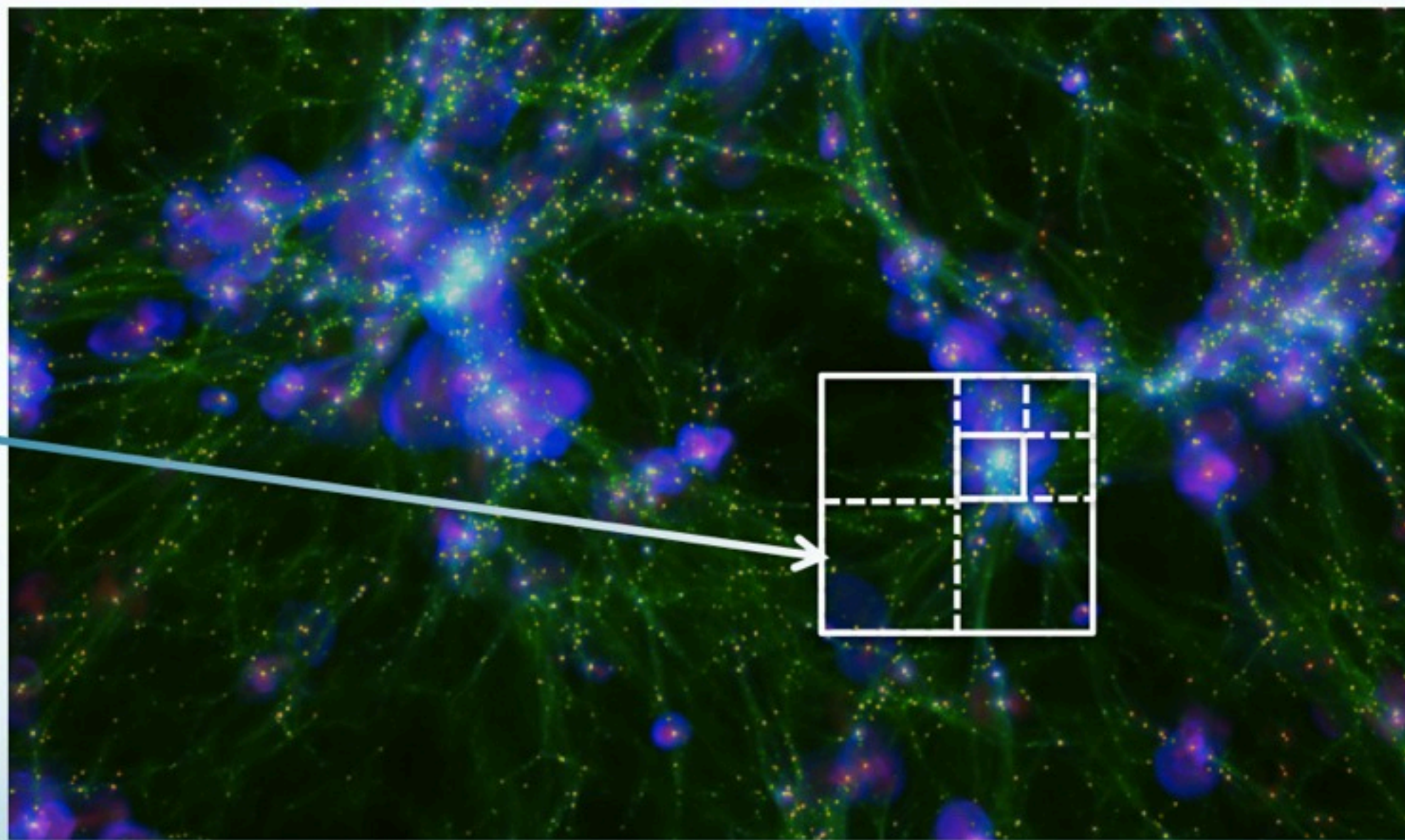
II — (1) Full physics large-scale hydrodynamical cosmological AMR simulation with AGN

Run with
RAMSES code

(R. Teyssier)

DM particles:
 1024^3

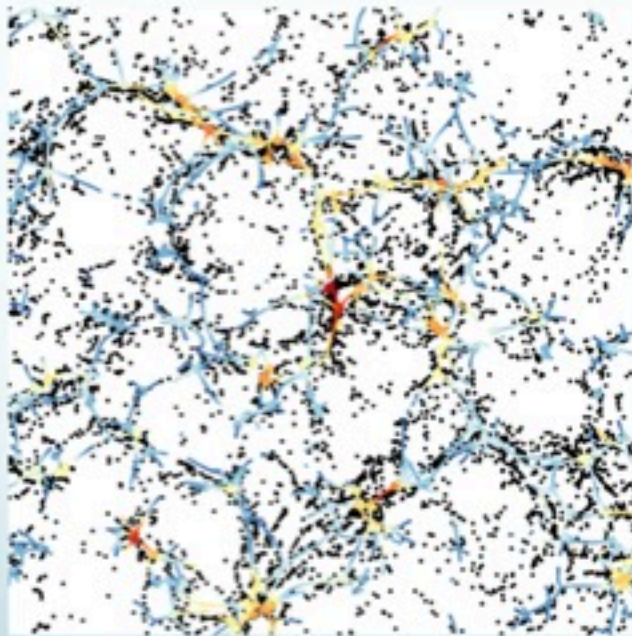
Max. refinement:
 2^{17}



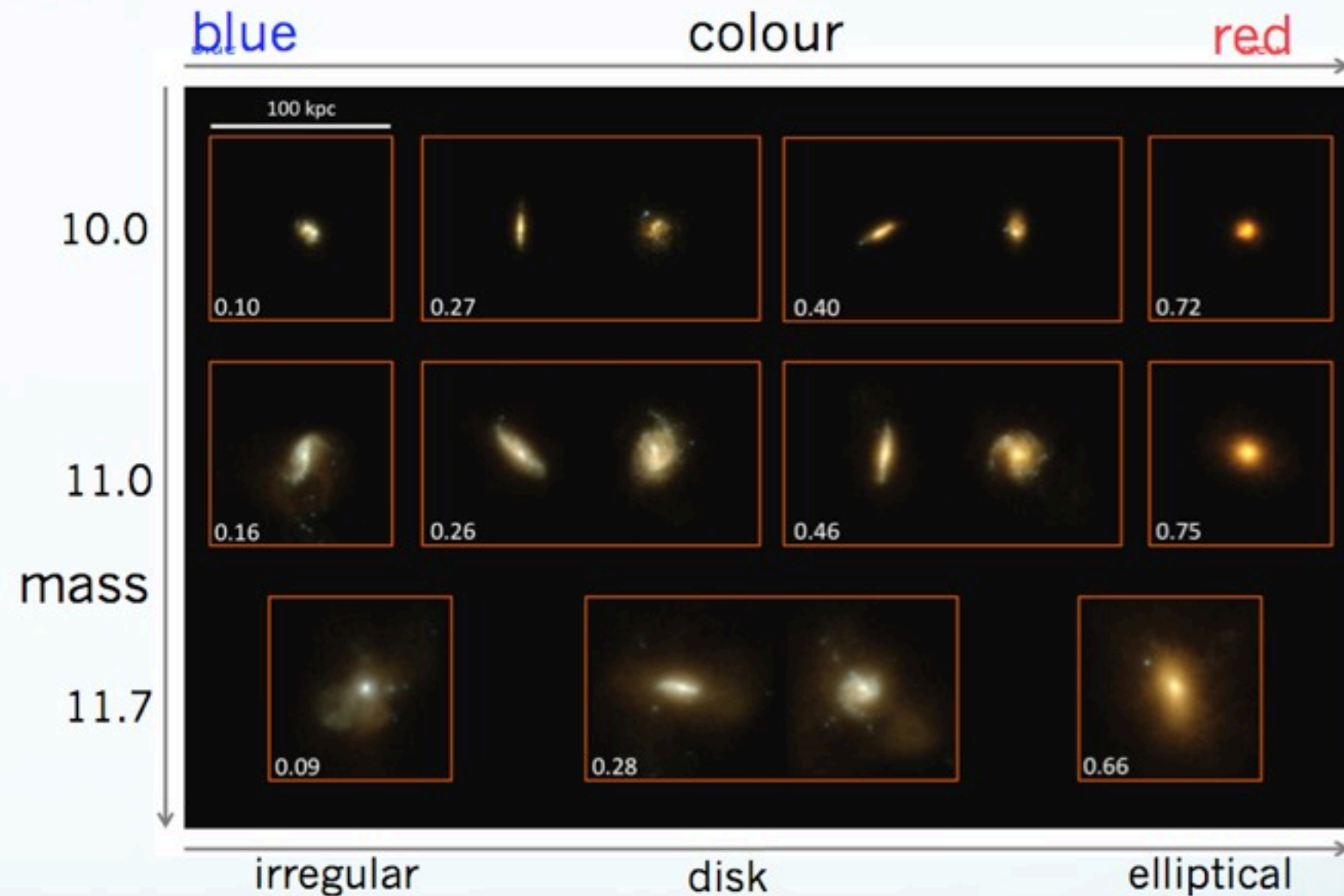
Dubois et al. 2014

- 100 Mpc.h⁻¹ square box.
run down to $z=1.2$ with
LCDM cosmology:

- 150 000 galaxies and
- 300 000 dark haloes
in a snapshot.

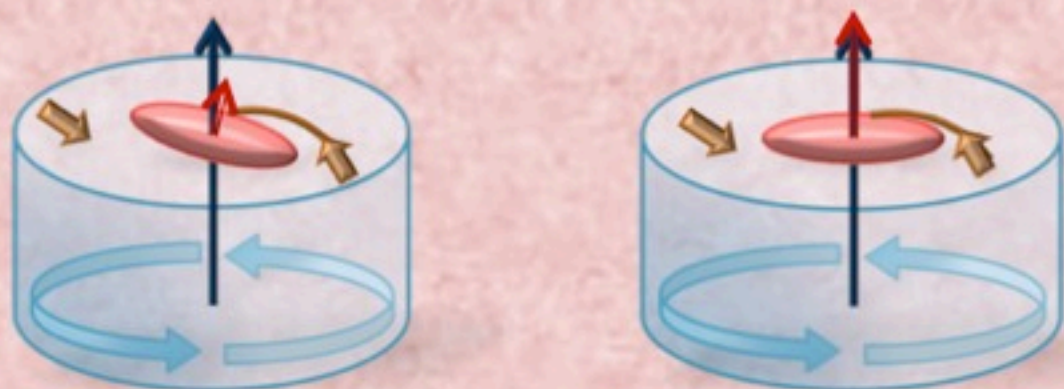


- Horizon-AGN:
skeleton

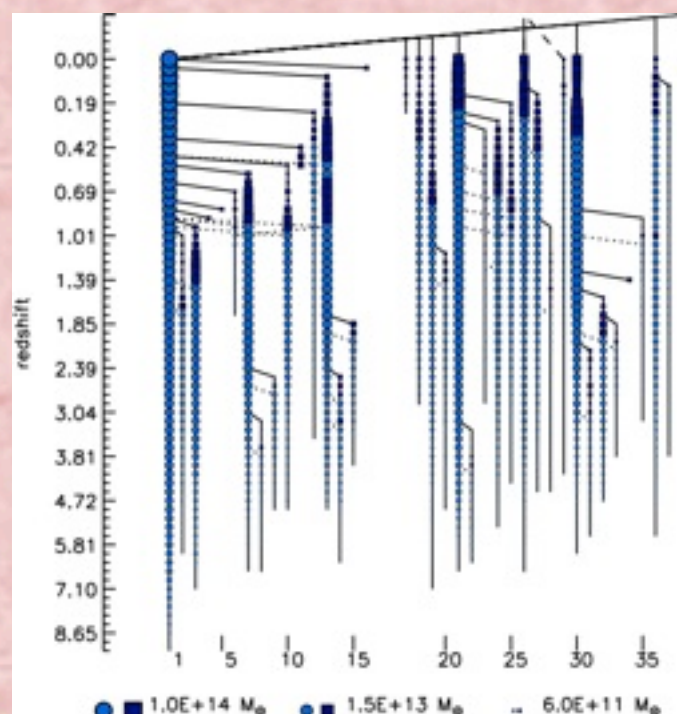


- Various physical features:**
 - Star formation
 - Feedback: supernovae, AGN, ...
 - Cooling: H, He+metal contribution
 - Metals

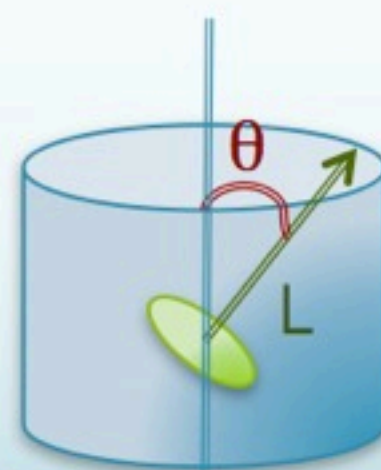
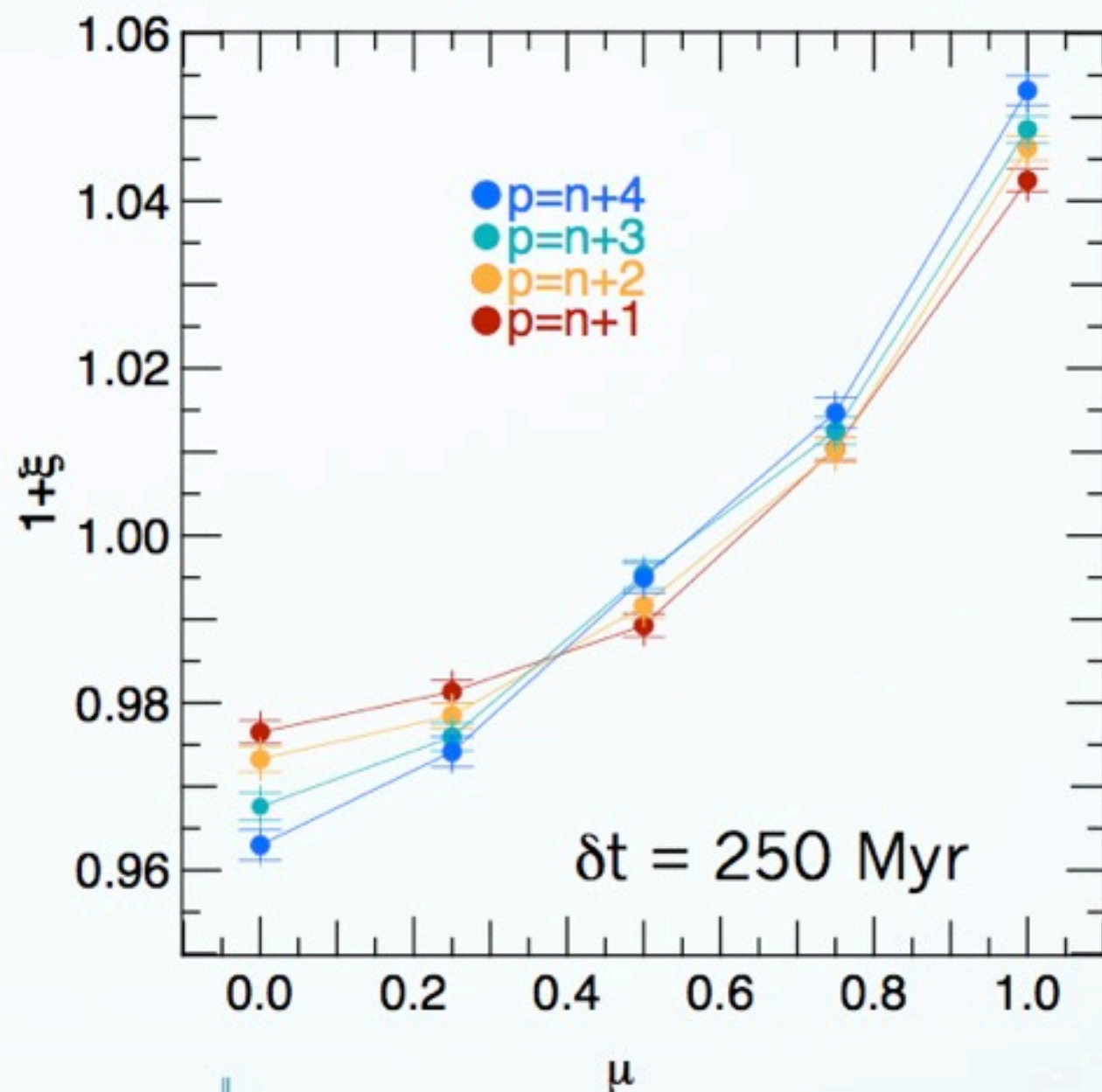
SMOOTH ACCRETION



- Gas inflows (re)-align galaxies with their filament



PDF of μ over 4 timesteps δt



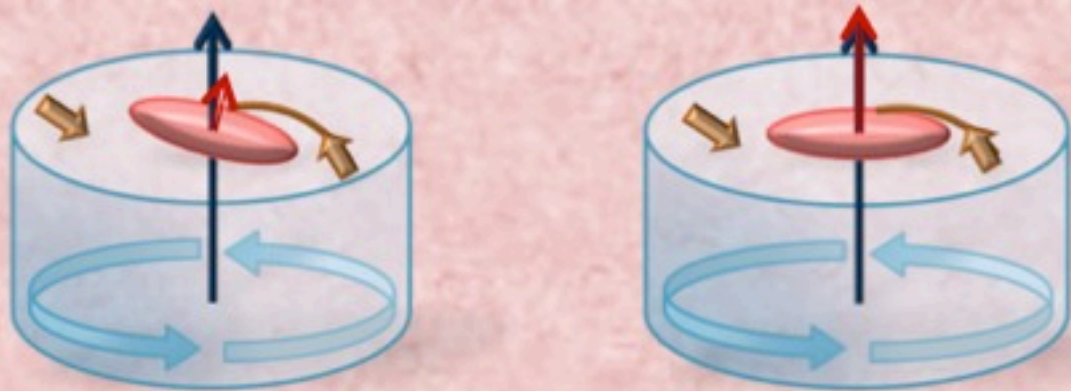
$$\mu = |\cos(\theta)|$$

Spin-filament angle

ξ : excess probability

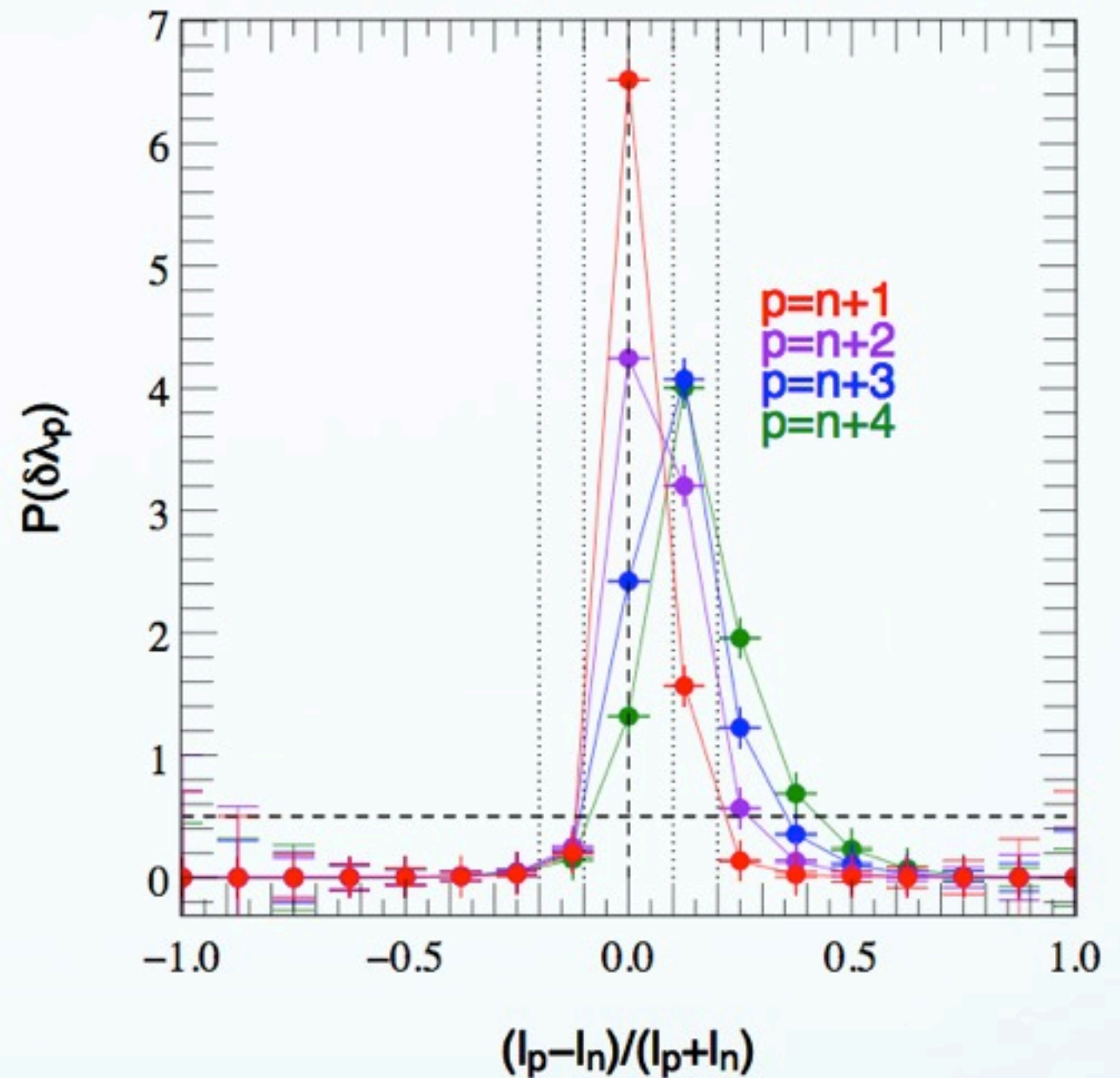
no merger : orientation versus look-back time

SMOOTH ACCRETION



- Gas inflows (re)-align galaxies with their filament
- Gas inflows build up the galactic spin

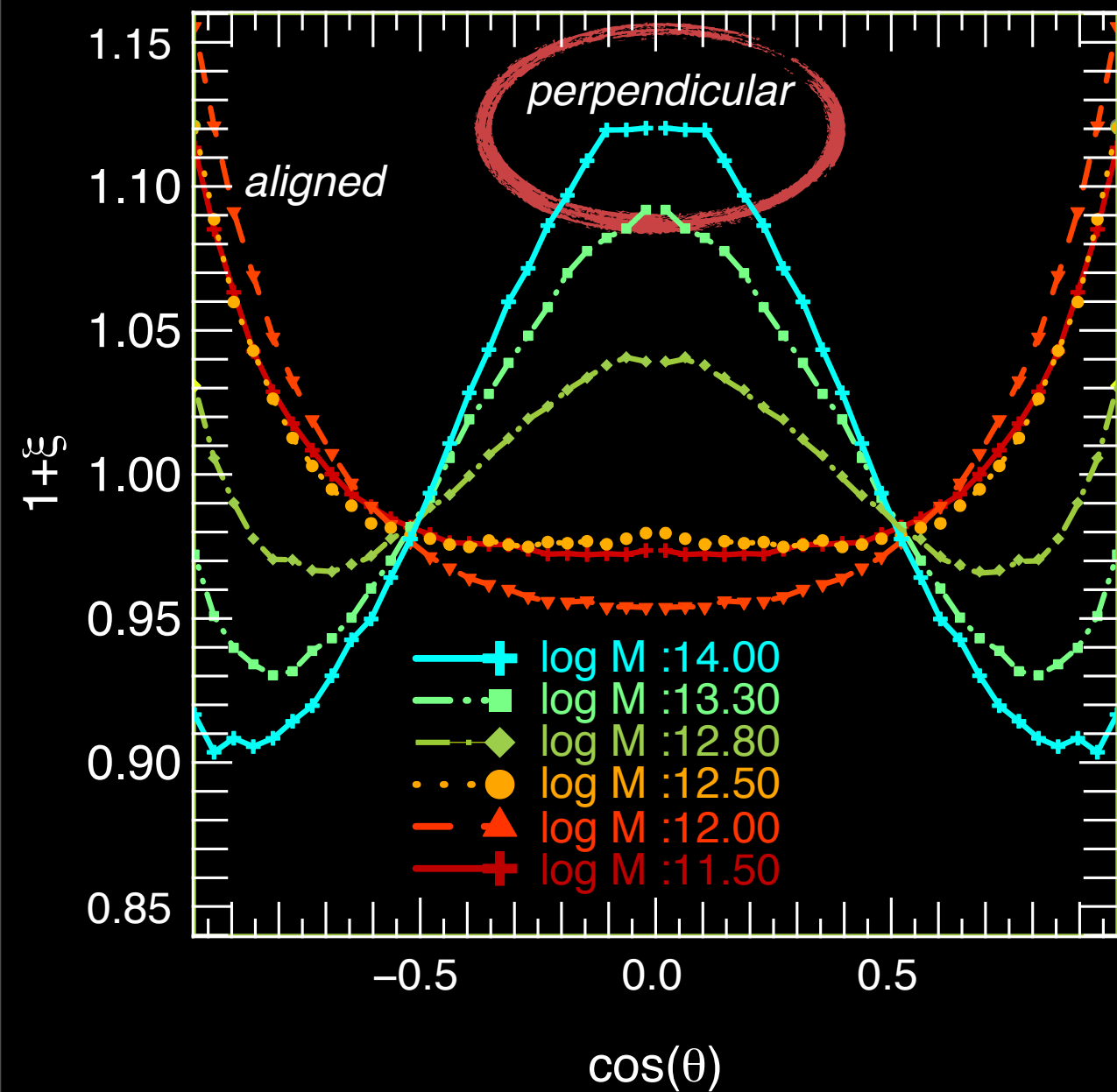
PDF of $\delta\lambda(p)$ for 4 different δt_p



- $\delta\lambda(p)$: angular momentum contrast over timestep $\delta t_p = (p-n)*\delta t$

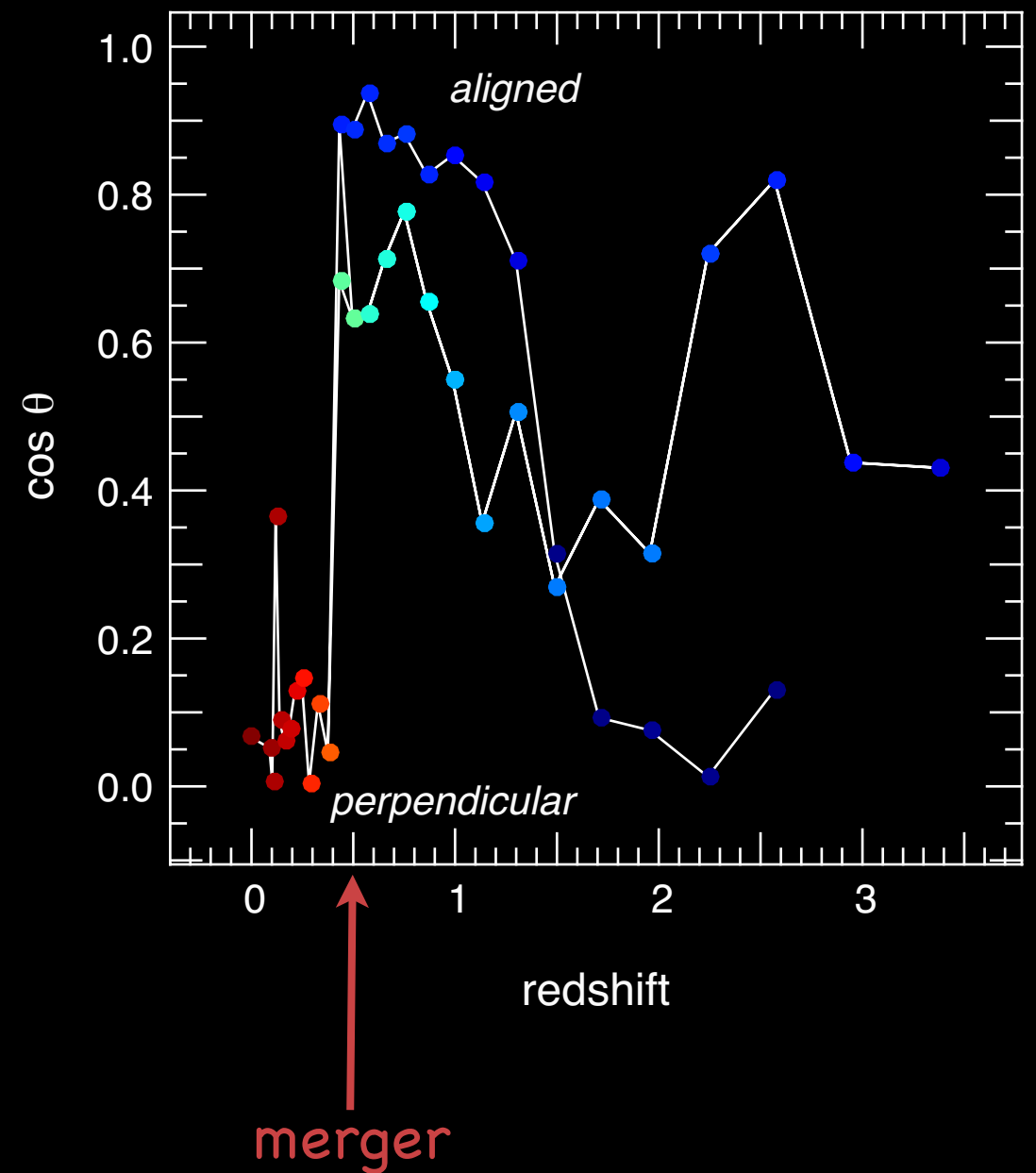
no merger : spin amplitude versus look-back time

High-mass haloes: $M > M_{\text{crit}}$



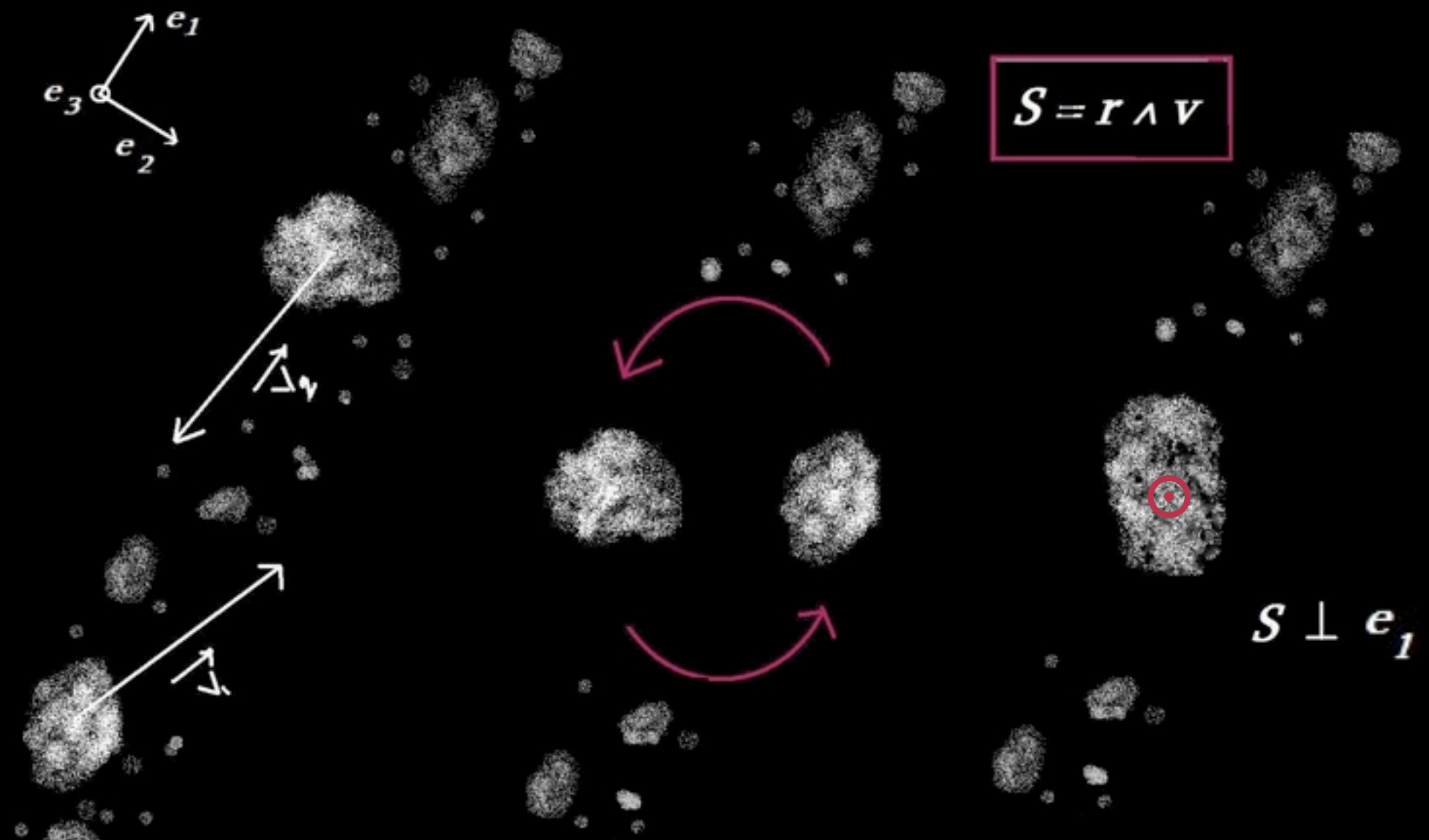
$$M_{\text{crit}} = 4 \cdot 10^{12} M_{\odot}$$

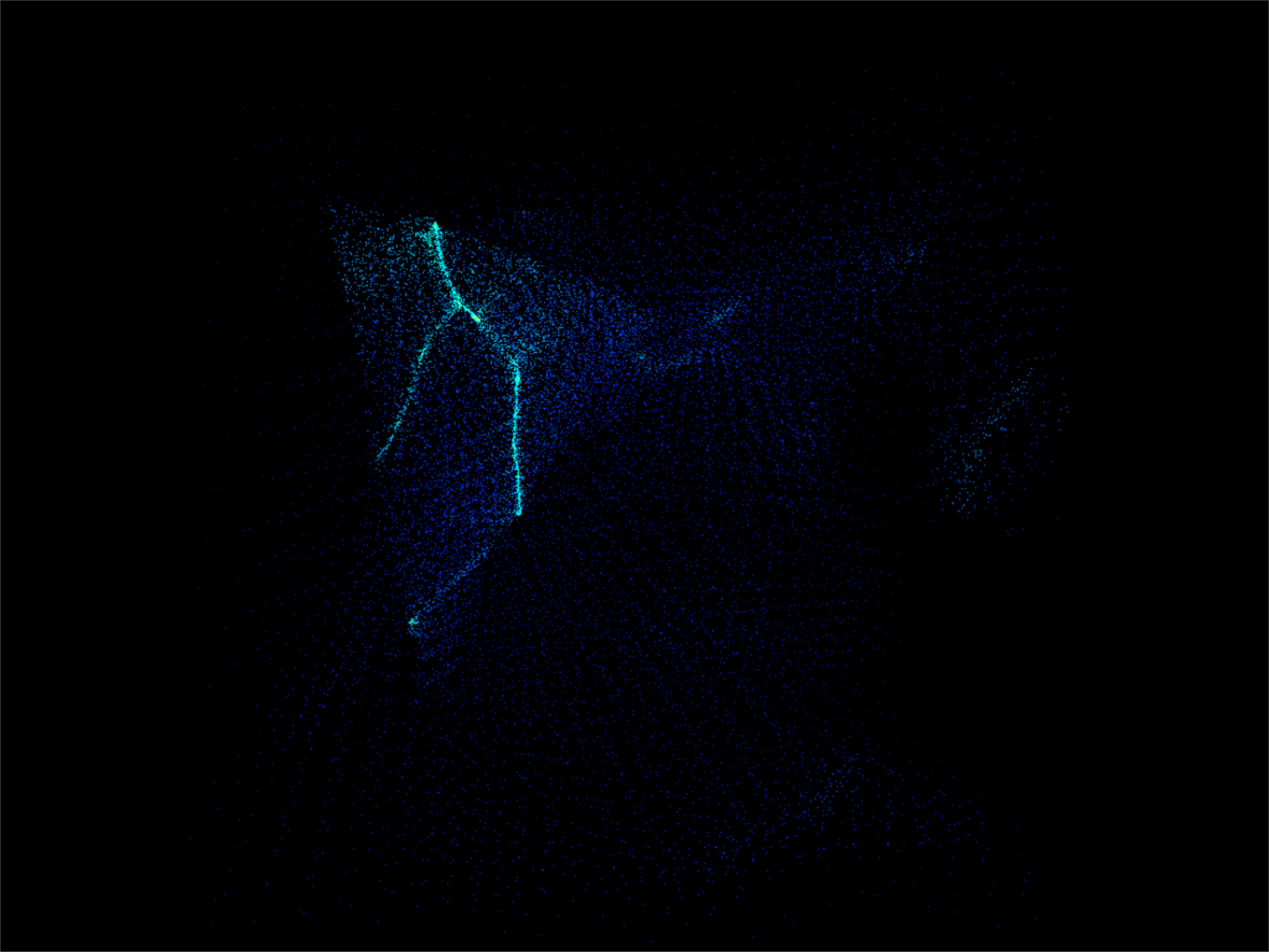
formed at low z by mergers inside the filaments

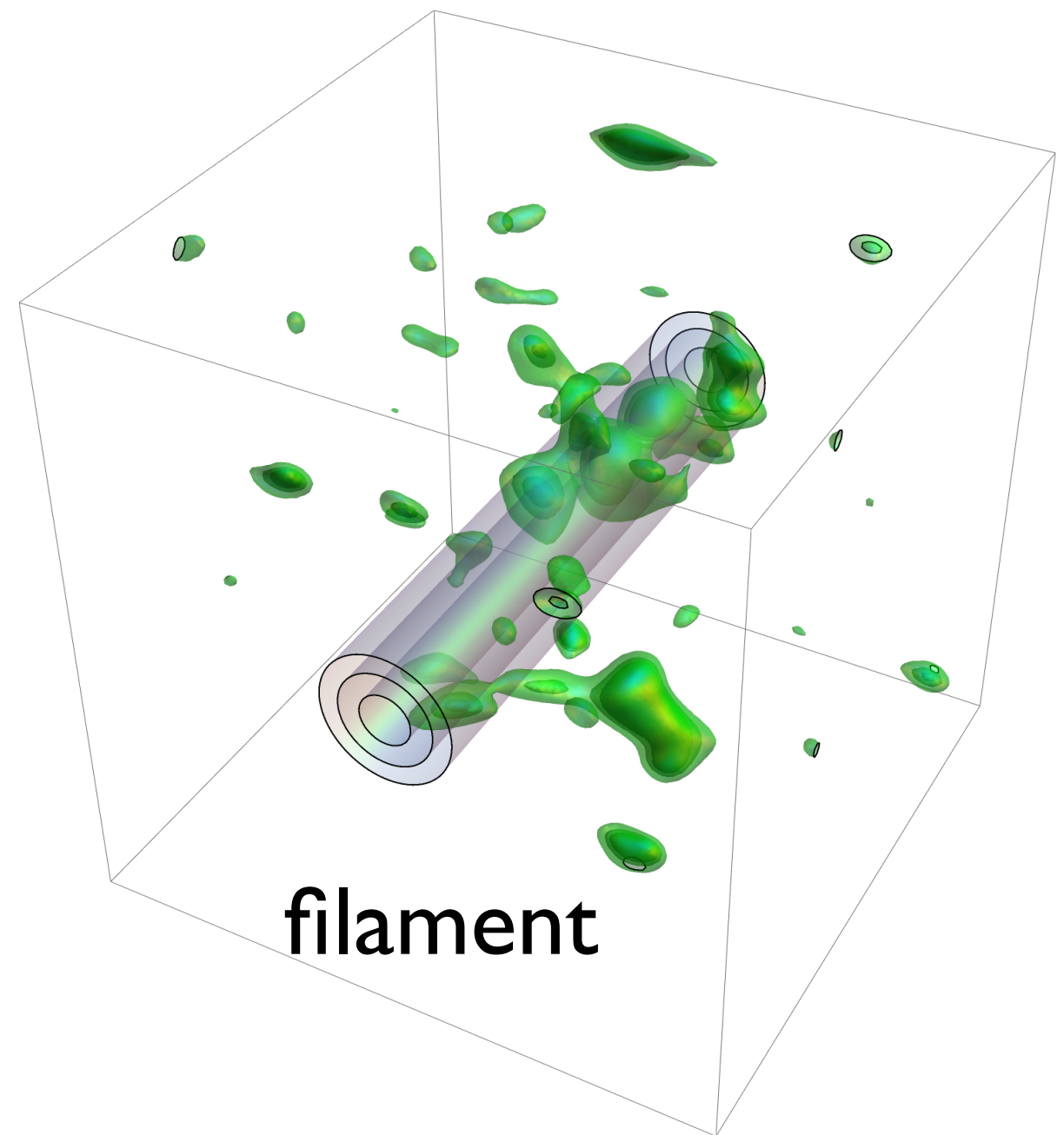
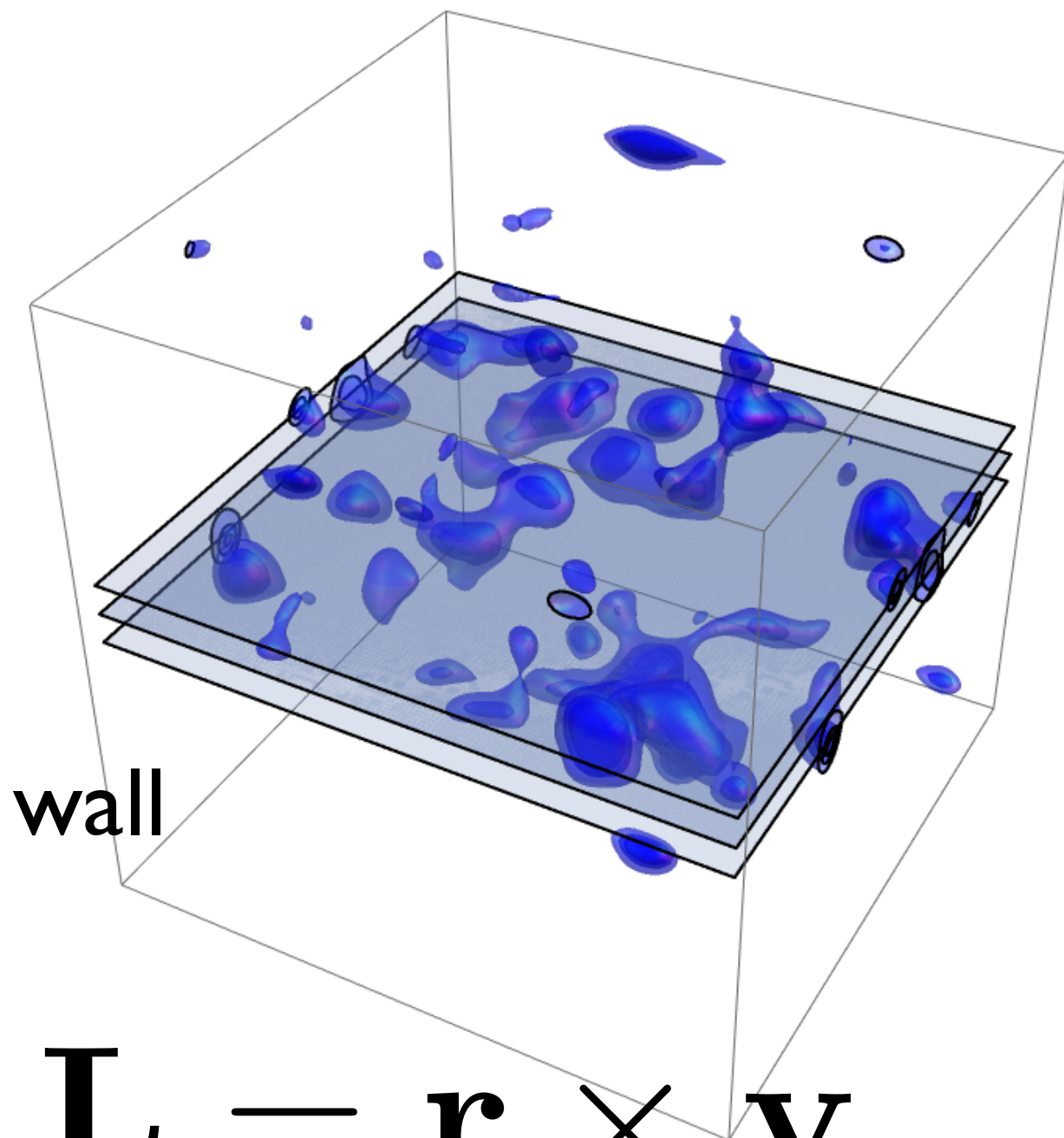


How do mergers along the filaments create spin perpendicular to them?

halos catch up with each other along the filaments



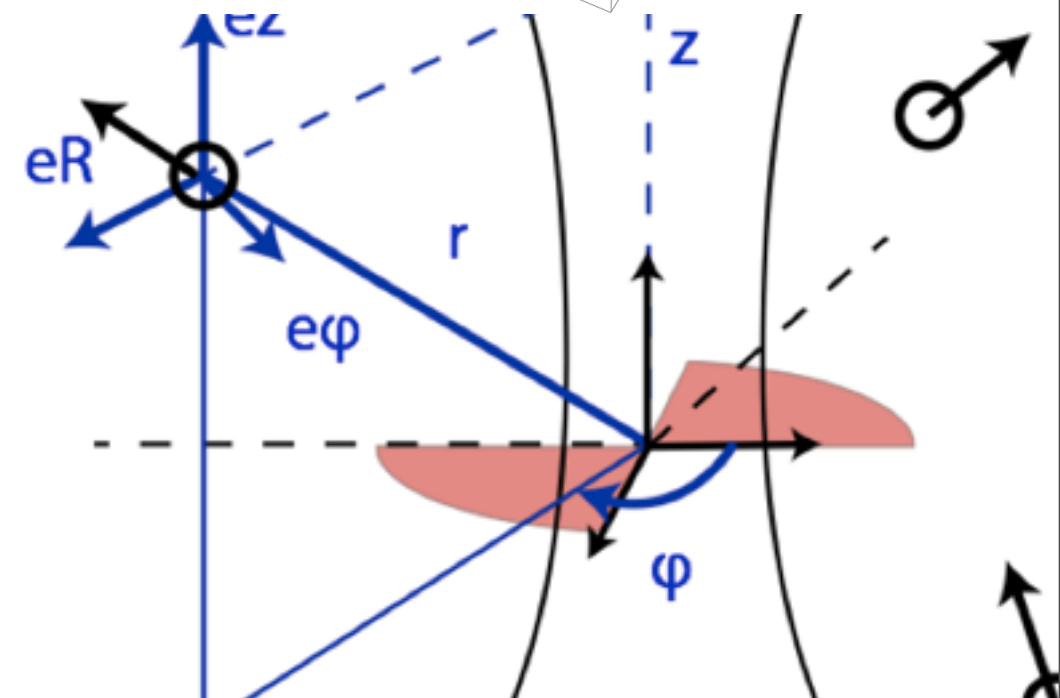




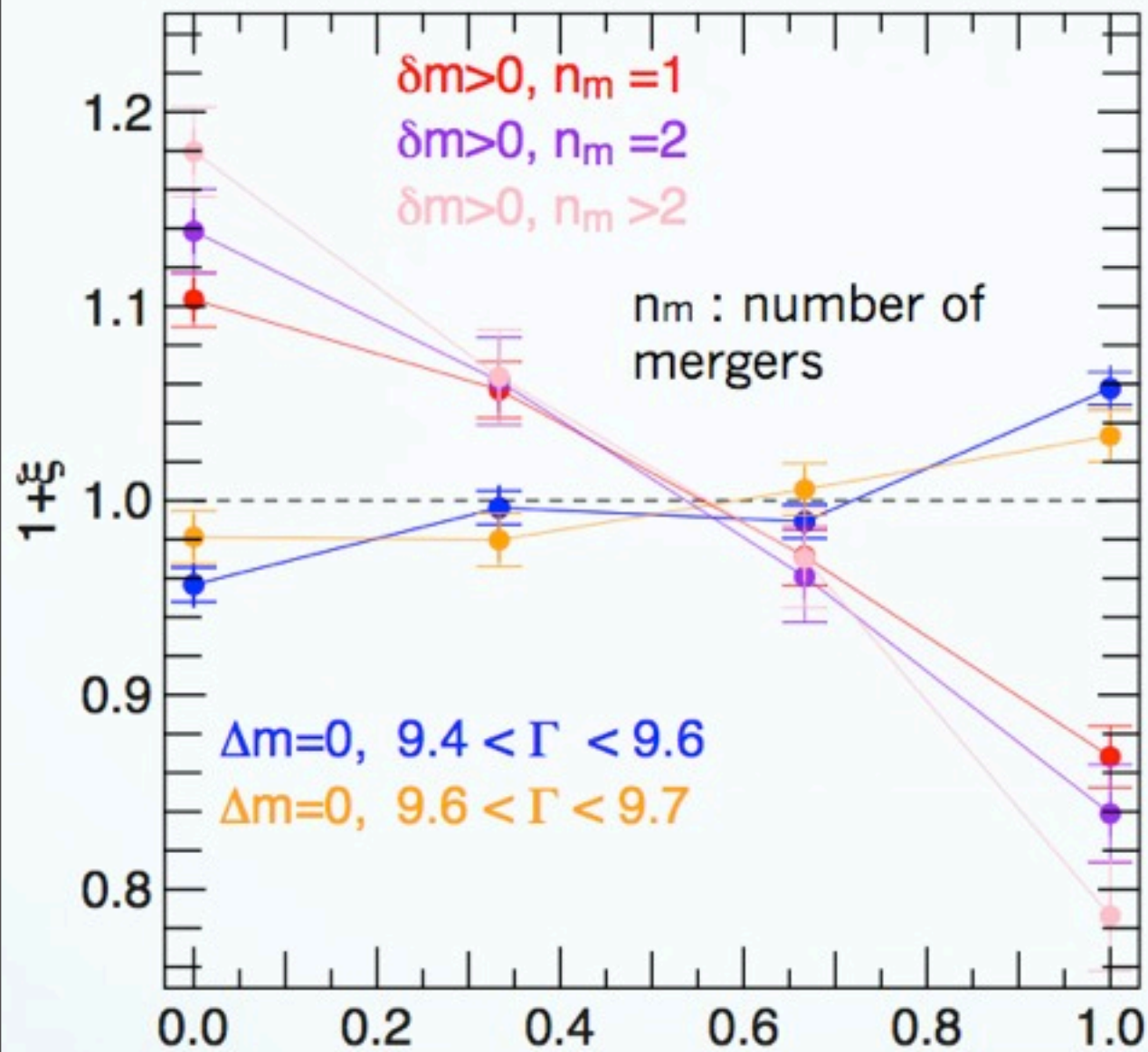
$$\mathbf{L} = \mathbf{r} \times \mathbf{v}$$

If pancake \ni filament

→ mergers along e_ϕ



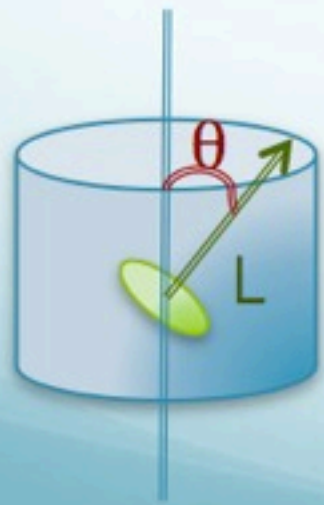
PDF of μ for different merging histories



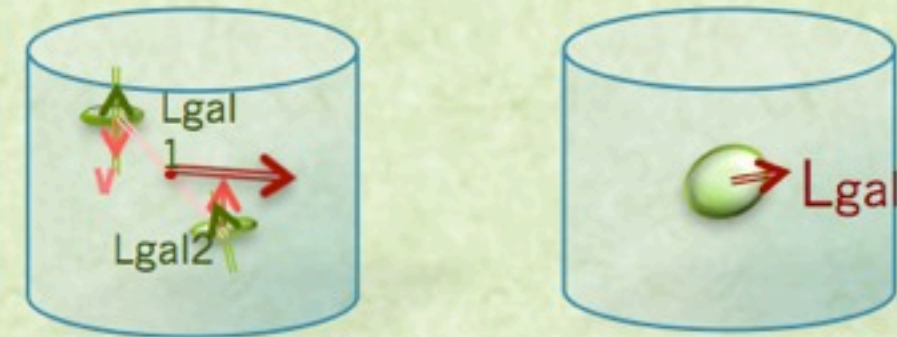
μ
 Merger rate

better tracer of flip

$$\mu = |\cos(\theta)|$$



MERGERS



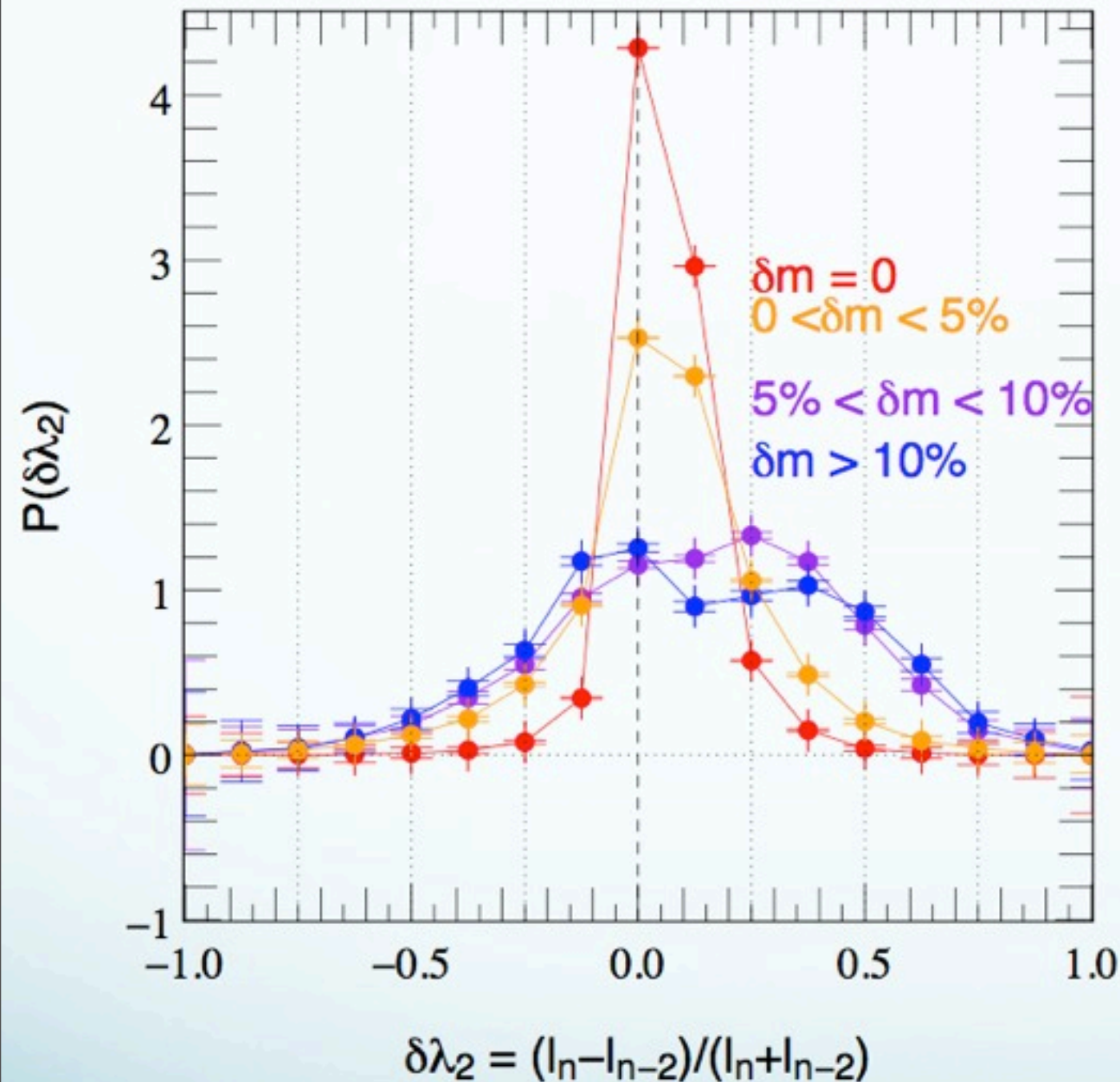
- Mergers drive galactic spin flips.
- Mergers flip the spin perpendicular to the filament



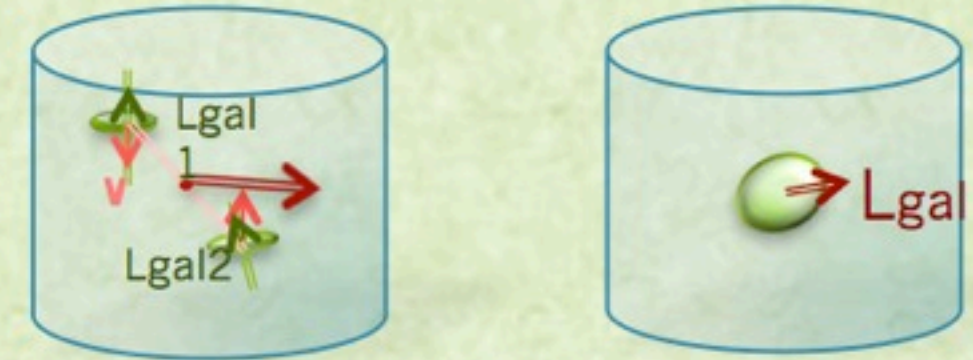
○ Red to pink : merged between $z=5.2$ and $z=1.8$

○ Blue to yellow : never merged, $\Gamma = \log(M/M_{\text{sun}})$

PDF of the angular momentum contrast for different merging histories



MERGERS



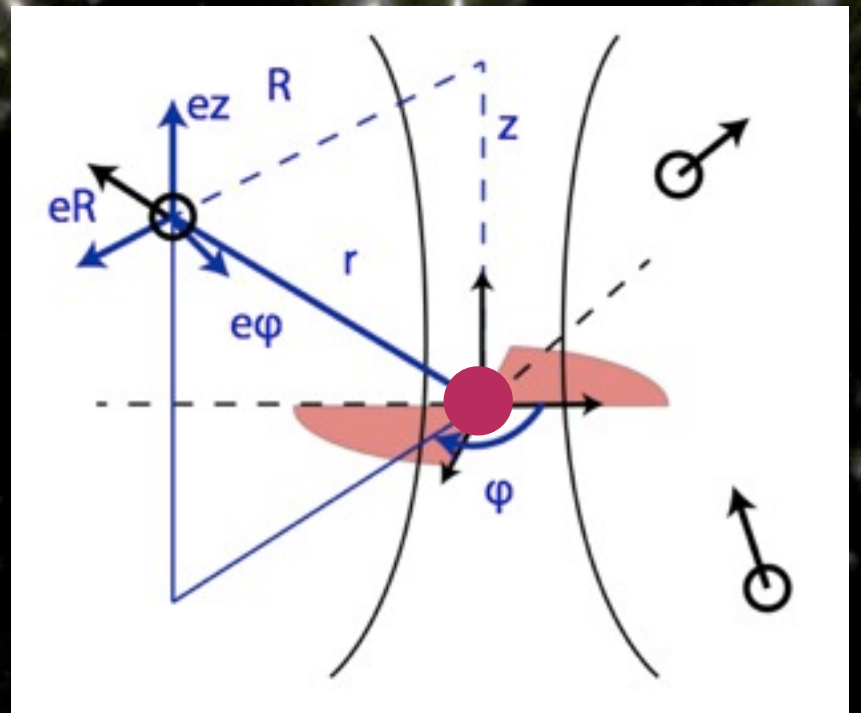
- Mergers drive galactic spin flips.
- Mergers flip the spin perpendicular to the filament
- Mergers build up the spin on average



- How discs build up from persistent cosmic web?
- How dark halo's spin flip relative to filament?
Why are they initially aligned with filaments?
Why the transition mass? Eulerian view
- What is the corresponding Lagrangian theory?

Part III

Tidal torque theory with a peak background split near a **saddle**



Part III Outline

• The Idea

- walls/filament/peak locally bias differentially
tidal and inertia tensor: spin alignment reflect this in TTT

• The picture

- Geometry of spin near saddle: point reflection
symmetric distribution

• The Maths

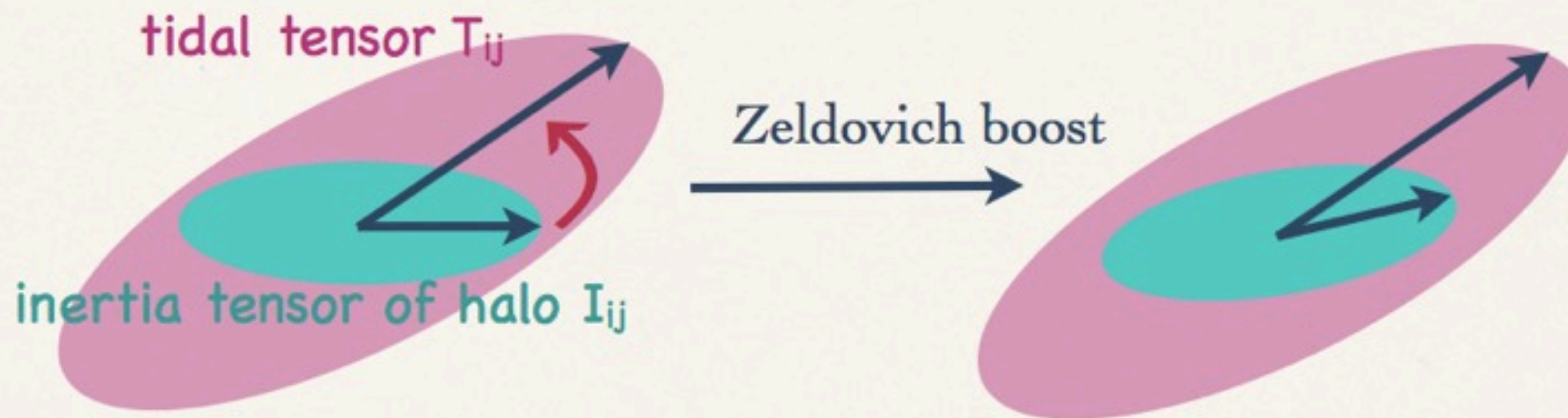
- Very simple **ab initio** prediction for mass transition

The Lagrangian view of spin/LSS connection

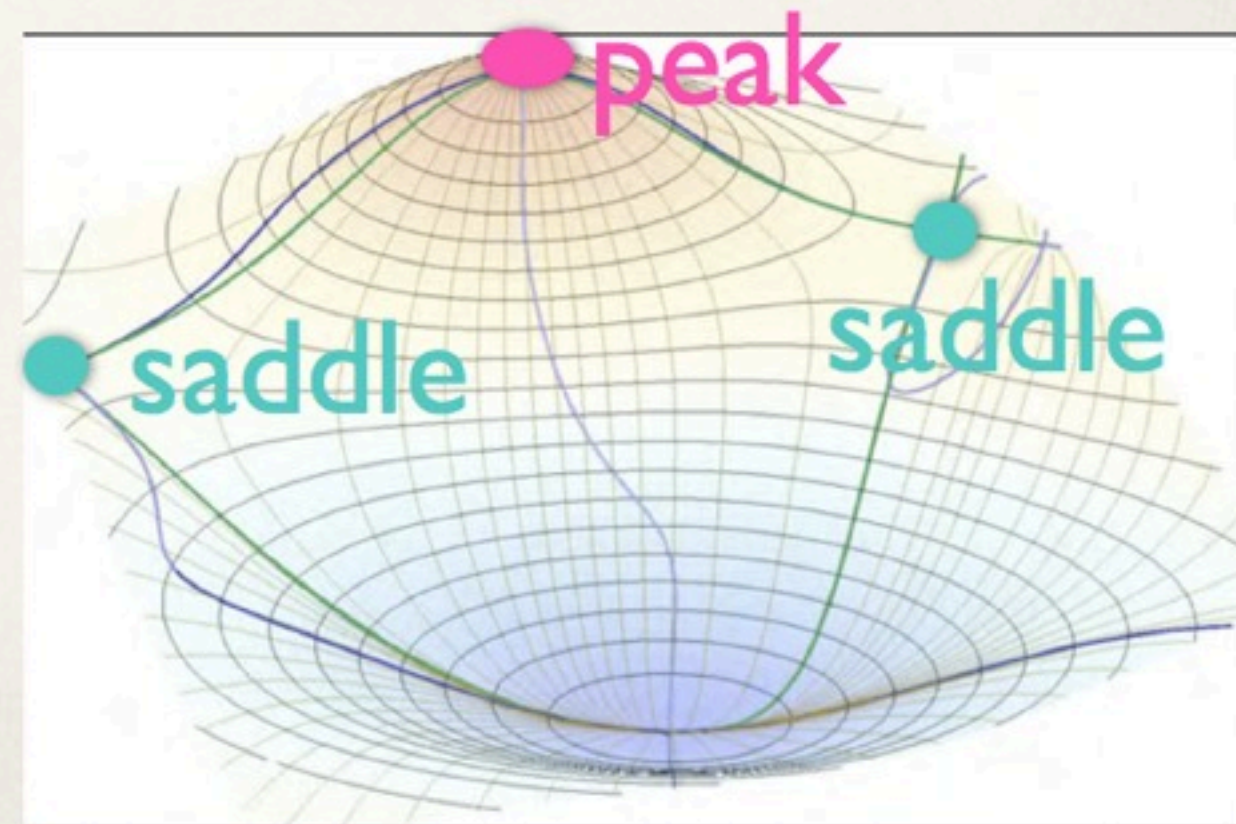
Can we understand where spin and vorticity alignments come from?

-usual tidal torque theory

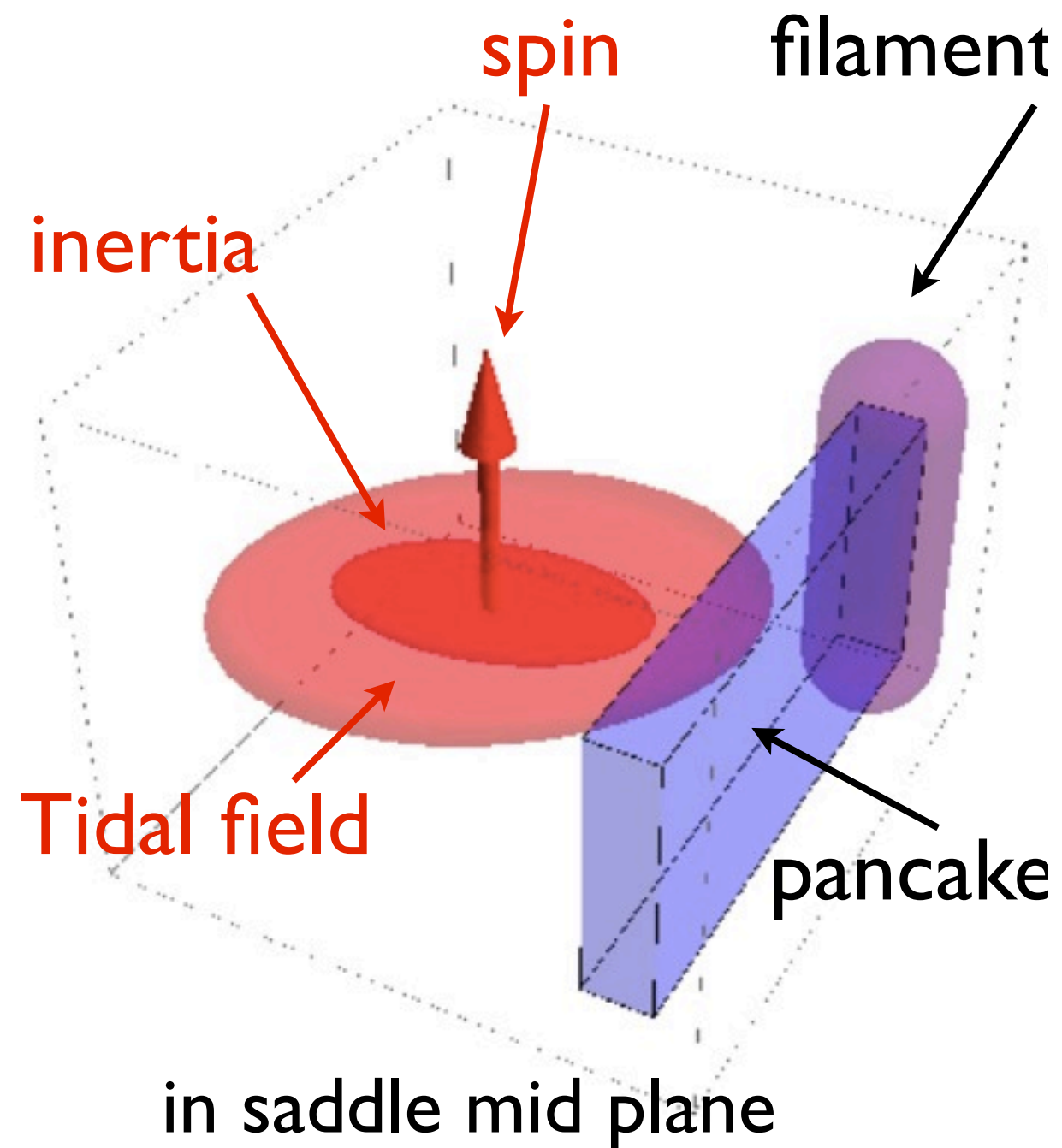
$$L_k = \varepsilon_{ijk} I_{li} T_{lj}$$



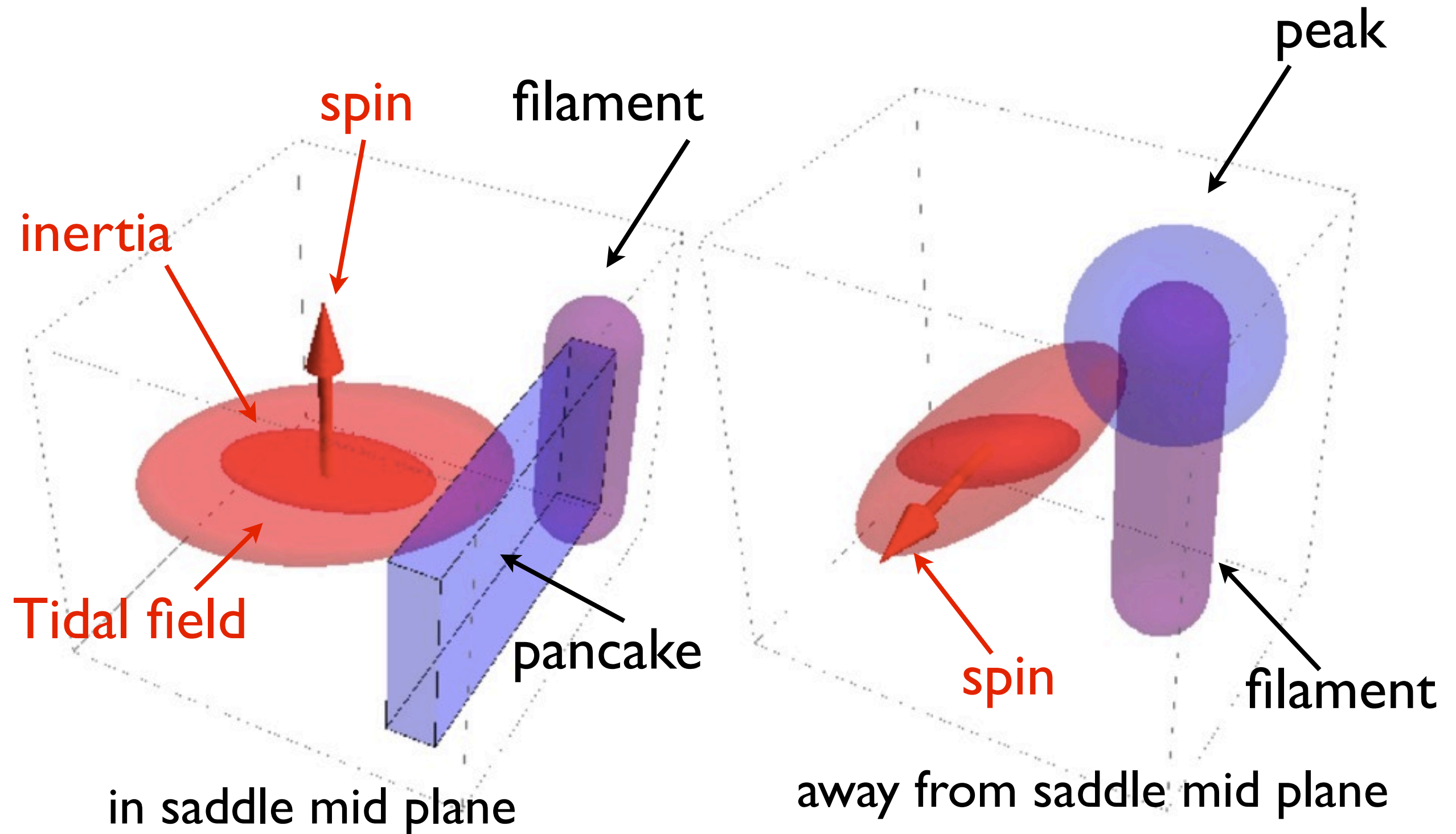
-anisotropy of the cosmic web:
surrounding of a saddle point
with typical geometry



Tidal/Inertia mis-alignment



Tidal/Inertia mis-alignment



Spin structure near Saddle

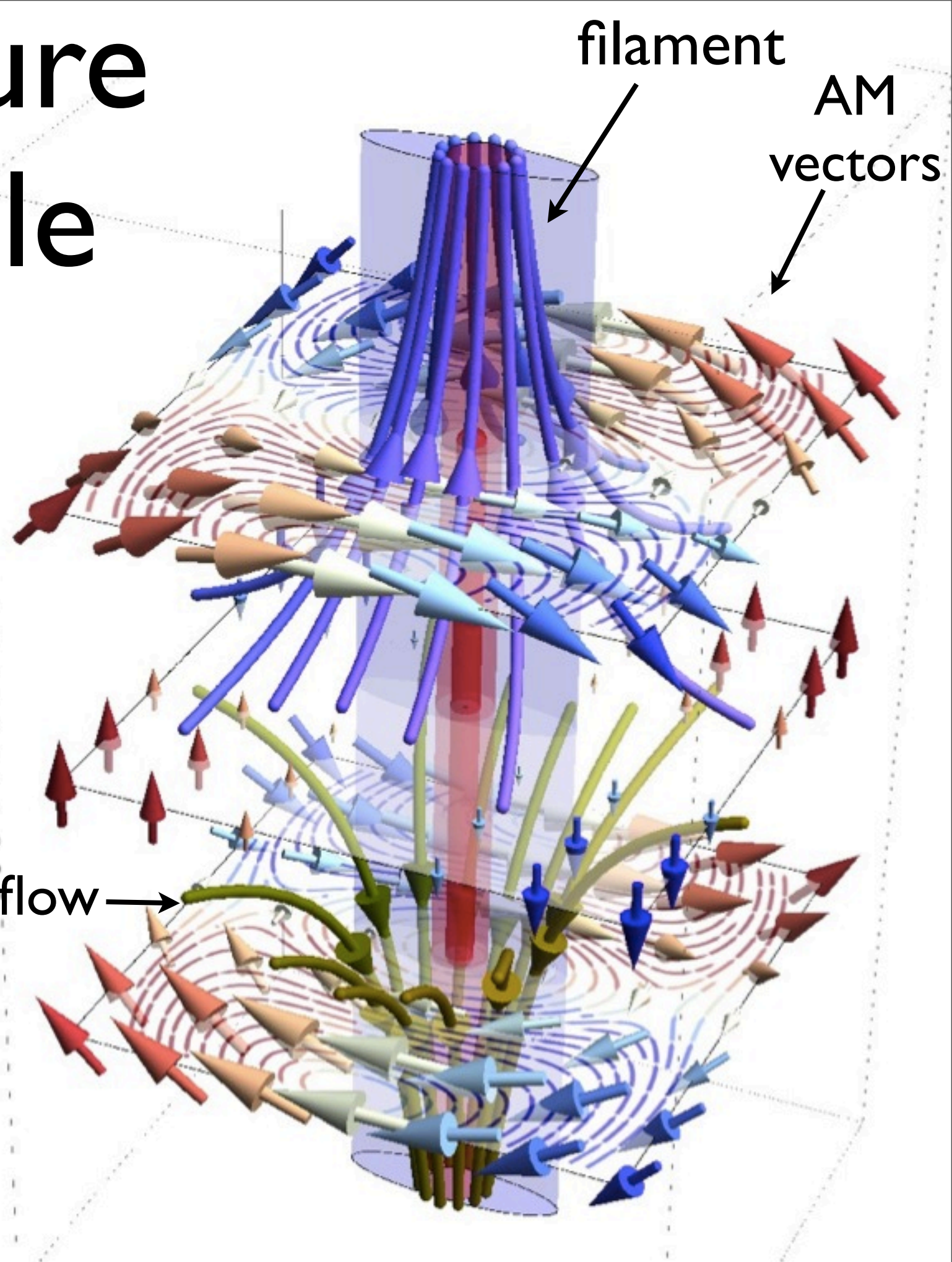
$$L_k = \varepsilon_{ijk} I_{li} T_{lj}$$

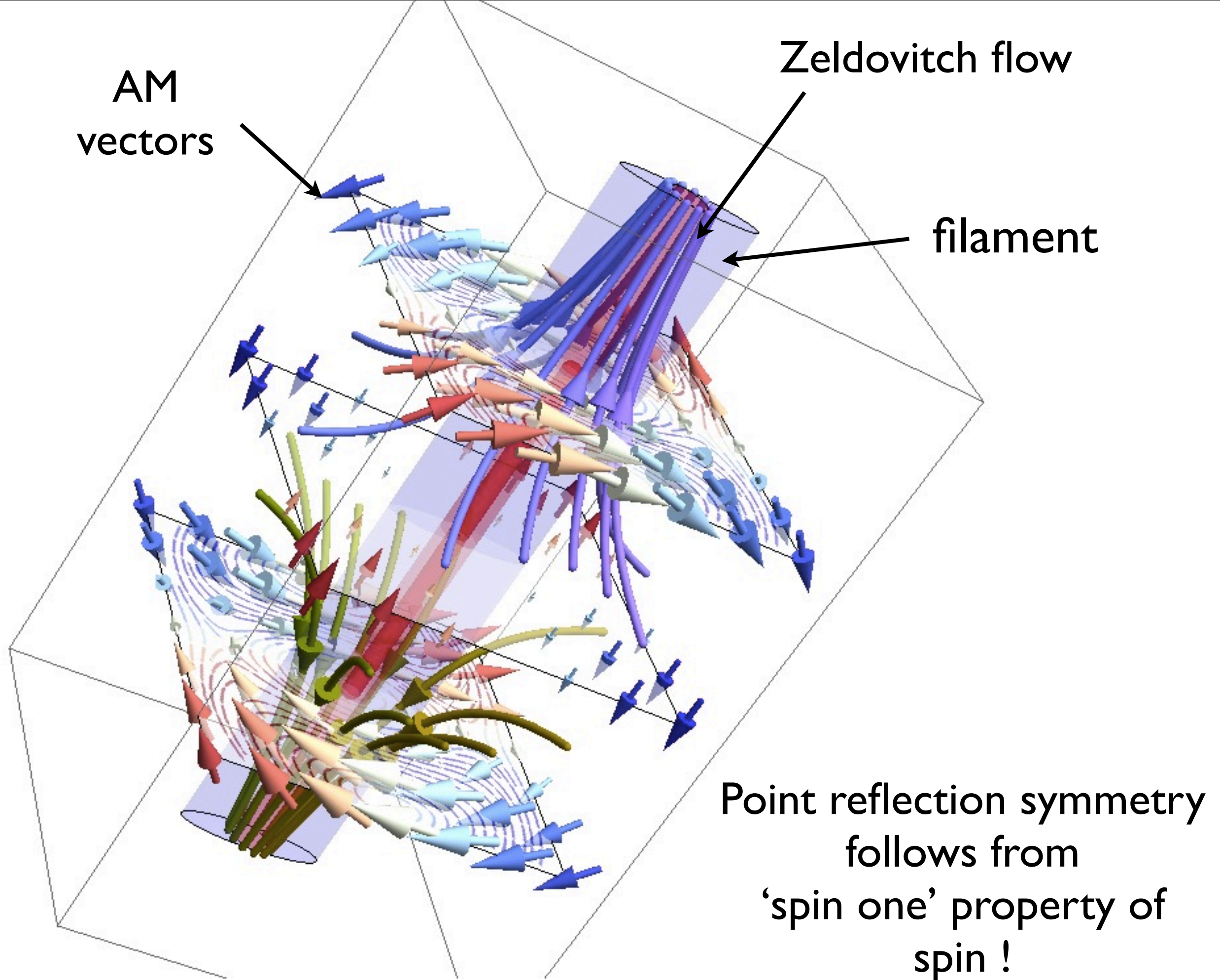
$$\approx \varepsilon_{ijk} H_{li} T_{lj}$$

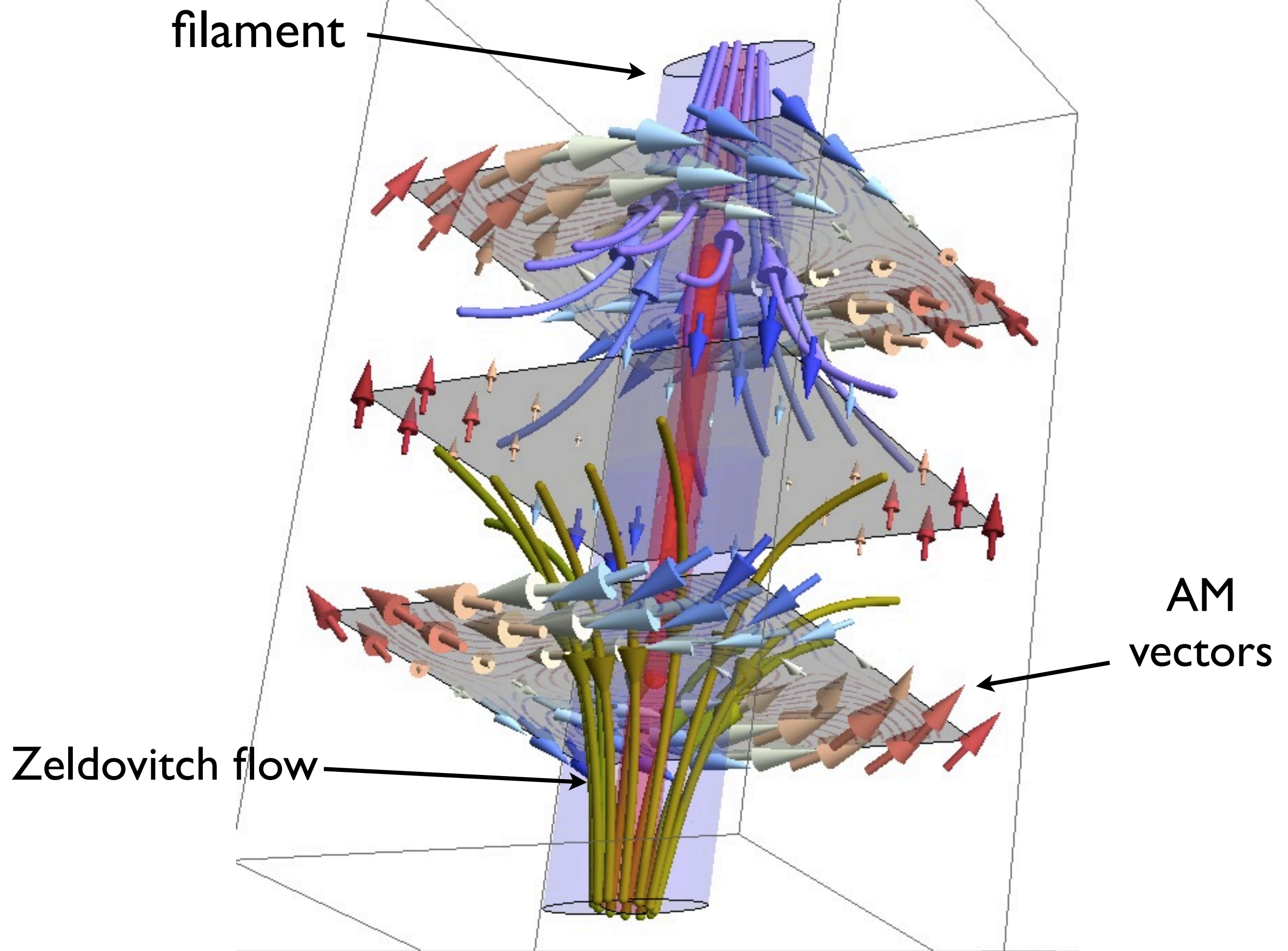
Hessian

Tidal

Zeldovitch flow







TTT@ saddle?

the Gaussain joint PDF of the derivatives of the field, $\mathbf{X} = \{x_{ij}, x_{ijk}, x_{ijkl}\}$ and $\mathbf{Y} = \{y_{ij}, y_{ijk}, y_{ijkl}\}$ in two given locations (\mathbf{r}_x and \mathbf{r}_y separated by a distance $r = |\mathbf{r}_x - \mathbf{r}_y|$) obeys

$$\text{PDF}(\mathbf{X}, \mathbf{Y}) = \frac{1}{\det|2\pi\mathbf{C}|^{1/2}} \times$$

$$\exp\left(-\frac{1}{2} \begin{bmatrix} \mathbf{X} \\ \mathbf{Y} \end{bmatrix}^T \cdot \begin{bmatrix} \mathbf{C}_0 & \mathbf{C}_\gamma \\ \mathbf{C}_\gamma^T & \mathbf{C}_0 \end{bmatrix}^{-1} \cdot \begin{bmatrix} \mathbf{X} \\ \mathbf{Y} \end{bmatrix}\right), \quad (\text{A2})$$

subject to the "saddle" **constraints** (2D)

height $x_{0,2} + x_{2,0} = \nu, \quad x_{1,2} + x_{3,0} = 0, \quad x_{0,3} + x_{2,1} = 0, \quad \text{zero gradient}$

$$\kappa \cos(2\theta) = \frac{1}{2} (x_{4,0} - x_{0,4}), \quad \kappa \sin(2\theta) = -x_{1,3} - x_{3,1}.$$

parametrized curvature

TTT@ saddle?

the Gaussain joint PDF of the derivatives of the field, $\mathbf{X} = \{x_{ij}, x_{ijk}, x_{ijkl}\}$ and $\mathbf{Y} = \{y_{ij}, y_{ijk}, y_{ijkl}\}$ in two given locations (\mathbf{r}_x and \mathbf{r}_y separated by a distance $r = |\mathbf{r}_x - \mathbf{r}_y|$) obeys

PDF(\mathbf{X}, \mathbf{Y})

$$\exp \left(-\frac{1}{2} \begin{bmatrix} \mathbf{X} \\ \mathbf{Y} \end{bmatrix}^T \right).$$

$$\begin{aligned} x_{0,0,2} + x_{0,2,0} + x_{2,0,0} &= \nu, \quad x_{1,0,2} + x_{1,2,0} + x_{3,0,0} = 0, \\ x_{0,1,2} + x_{0,3,0} + x_{2,1,0} &= 0, \quad x_{0,0,3} + x_{0,2,1} + x_{2,0,1} = 0, \\ \kappa_{1,1} &= \frac{1}{3} (x_{2,0,2} - x_{0,0,4} - 2x_{0,2,2} - x_{0,4,0} + x_{2,2,0} + 2x_{4,0,0}), \\ \kappa_{1,2} &= x_{1,1,2} + x_{1,3,0} + x_{3,1,0}, \quad \kappa_{1,3} = x_{1,0,3} + x_{1,2,1} + x_{3,0,1}, \\ \kappa_{2,2} &= \frac{1}{3} (x_{0,2,2} - x_{0,0,4} + 2x_{0,4,0} - 2x_{2,0,2} + x_{2,2,0} - x_{4,0,0}), \\ \kappa_{2,3} &= x_{0,1,3} + x_{0,3,1} + x_{2,1,1}. \end{aligned} \quad \text{3D} \quad (\text{B4})$$

subject to the "saddle" constraints (2D)

height $x_{0,2} + x_{2,0} = \nu, \quad x_{1,2} + x_{3,0} = 0, \quad x_{0,3} + x_{2,1} = 0, \quad \text{zero gradient}$

$$\kappa \cos(2\theta) = \frac{1}{2} (x_{4,0} - x_{0,4}), \quad \kappa \sin(2\theta) = -x_{1,3} - x_{3,1}.$$

parametrized curvature

Define the spin at point \mathbf{r}_y along the z direction as the anti-symmetric contraction of the de-traced tidal field and hessian:

(2D)

$$L(\mathbf{r}_y) = \varepsilon_{ij} \bar{y}_{il} \bar{y}_{jmm} = (y_{2,0} - y_{0,2}) (y_{1,3} + y_{3,1}) + \frac{y_{1,1}}{2} (y_{0,4} - y_{4,0}) - \frac{y_{1,1}}{2} (y_{4,0} - y_{0,4}) . \quad (\text{A3})$$

It is then fairly straightforward to compute the corresponding constrained expectation, $\langle L | \text{pk} \rangle$, for L as

$$L_z(r, \theta, \kappa, \nu) = \int L(\mathbf{Y}) \text{PDF}(\mathbf{X}, \mathbf{Y} | \text{pk}) d\mathbf{X} d\mathbf{Y} . \quad (\text{A4})$$



e.g. for $n=-2$

Incredibly simple prediction !

$$L_z = \kappa \frac{r^4 \sin(2\theta)}{144} e^{-\frac{r^2}{2}} \left(\sqrt{6} \kappa (r^2 - 4) \cos(2\theta) + 6 \right) .$$

asymmetry

2D Theory of Tidal Torque @ saddle?

$$L_z \propto r^4 \sin 2\theta \text{ at small radius}$$

$$L_z \propto \sin 2\theta \exp(-r^2) \text{ at large radius}$$

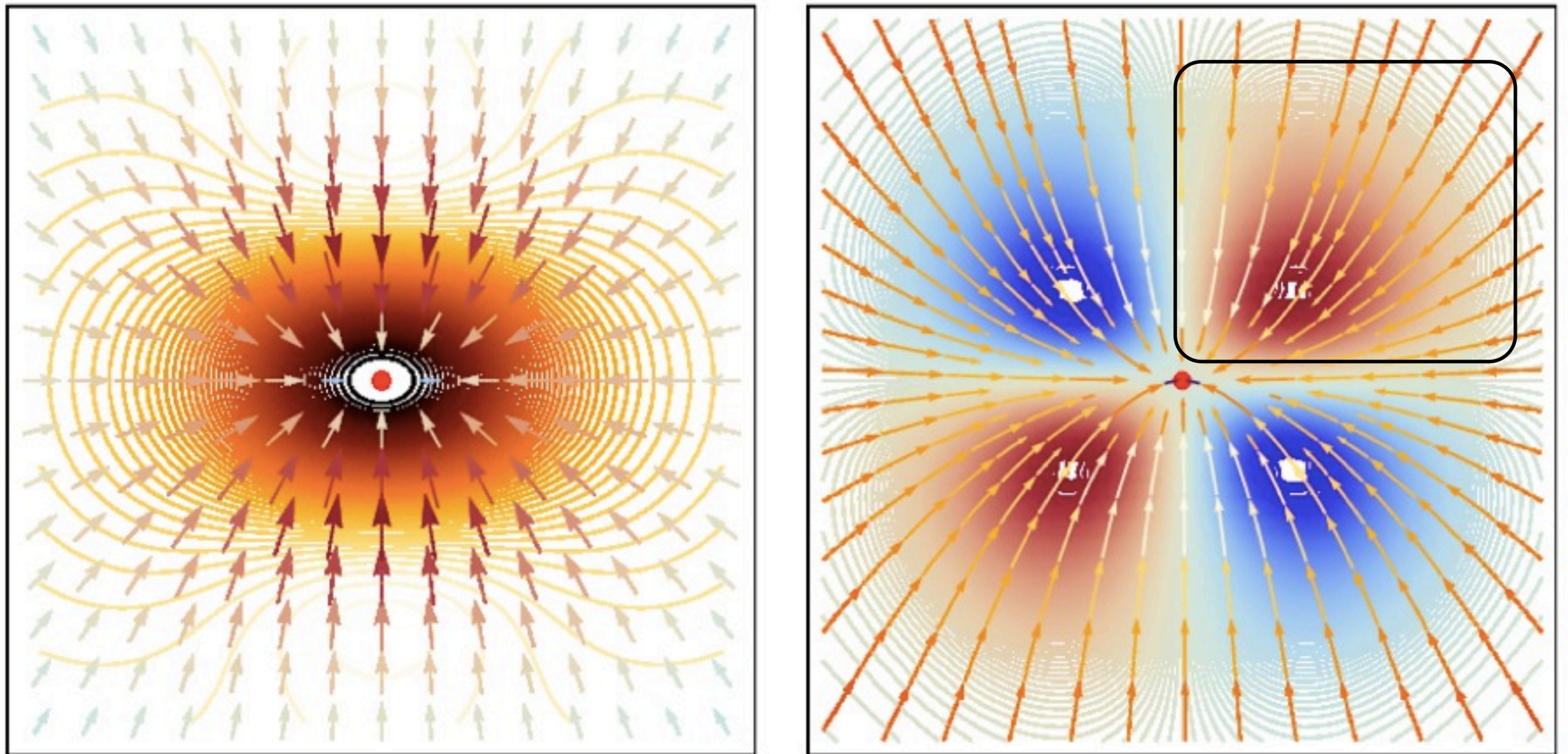


Figure 4. left: cross section of 2D Lagrangian patch near a saddle point; right: corresponding momentum (colour coded) and transverse velocity flow.

Link with Eulerian vorticity?

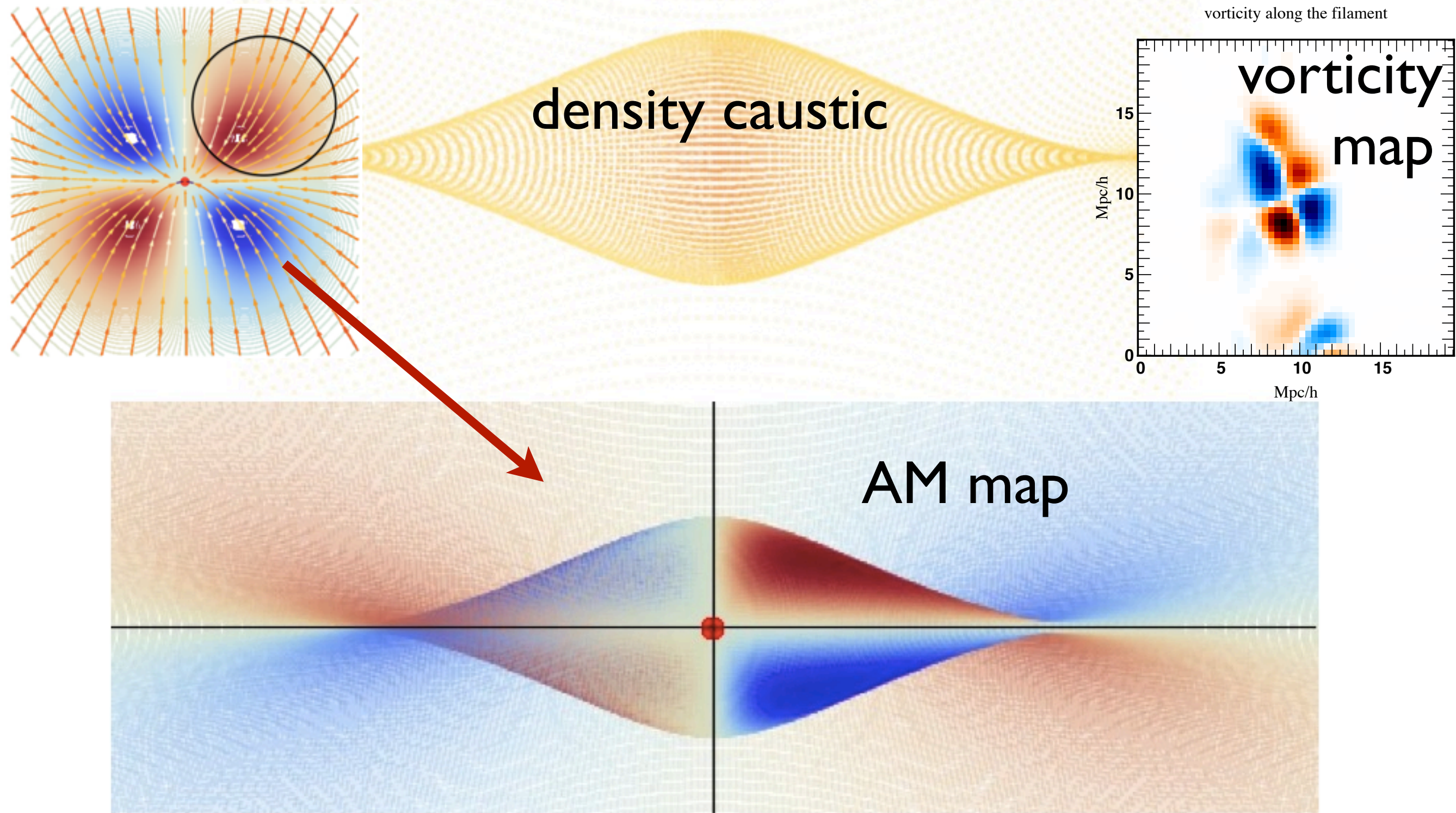


Figure 5. top: Density caustic; Bottom: Zeldovitch mapping of the spin distribution

Explain transition mass?

$$\mathcal{M}_{\text{tr}}(L_s) = \Sigma_0 \int d\nu d\kappa \mathcal{A}(\nu, \kappa) \mathcal{P}(\nu, \kappa | \text{pk}),$$

where the area of a quadrant is

$$\mathcal{A} = \frac{1}{4} \pi \frac{\lambda_2}{\lambda_1} (2r_\star)^2,$$

The pdf of the shape is

$$P(\nu, \kappa) = \frac{\kappa}{\sqrt{2\pi}} e^{-4\kappa^2 - \frac{1}{2}\nu^2} \text{erfc} \left(\frac{2\kappa - \gamma\nu}{\sqrt{2 - 2\gamma^2}} \right).$$

height

asymmetry

shape parameter

so the scaling of the transition mass

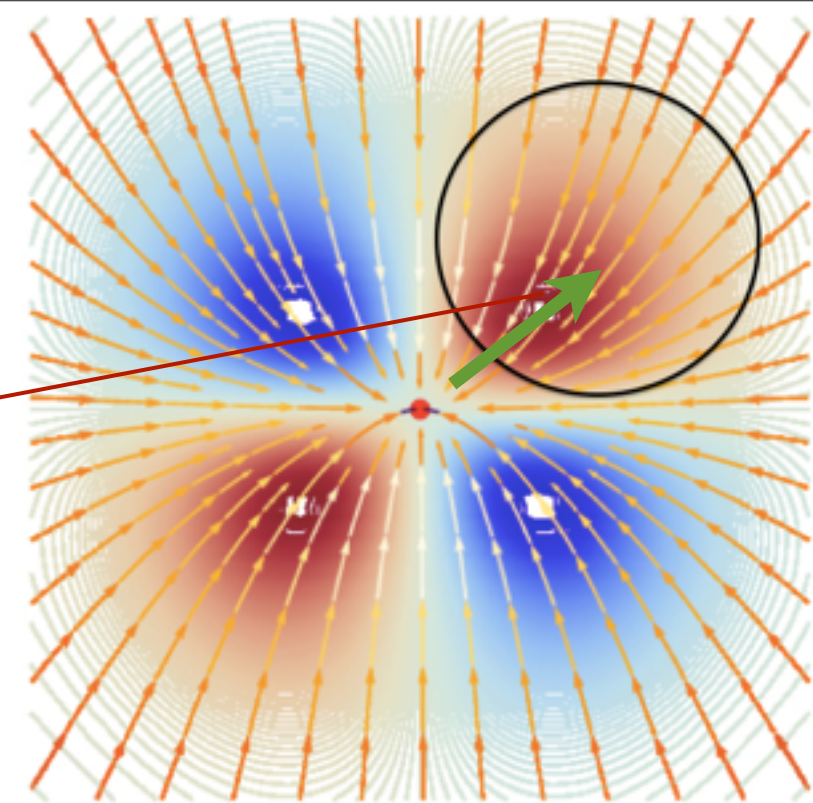
$$\mathcal{M}_{\text{tr}}(z) = \Delta N \frac{\lambda_{2,\star}}{\lambda_{1,\star}} \left(\frac{r_\star}{L_s} \right)^2 M_s(z) \equiv \alpha M_s,$$

where $\Delta N = \mathcal{P}(\nu_\star, \kappa_\star | \text{pk}) \Delta\nu \Delta\kappa$, and $M_s(z) \equiv \pi L_s^2(z) \Sigma_0$.

power index

$$\alpha(n) = \frac{n+5}{27} + \frac{1}{11}.$$

$$\alpha(-5) \sim 1/11.$$



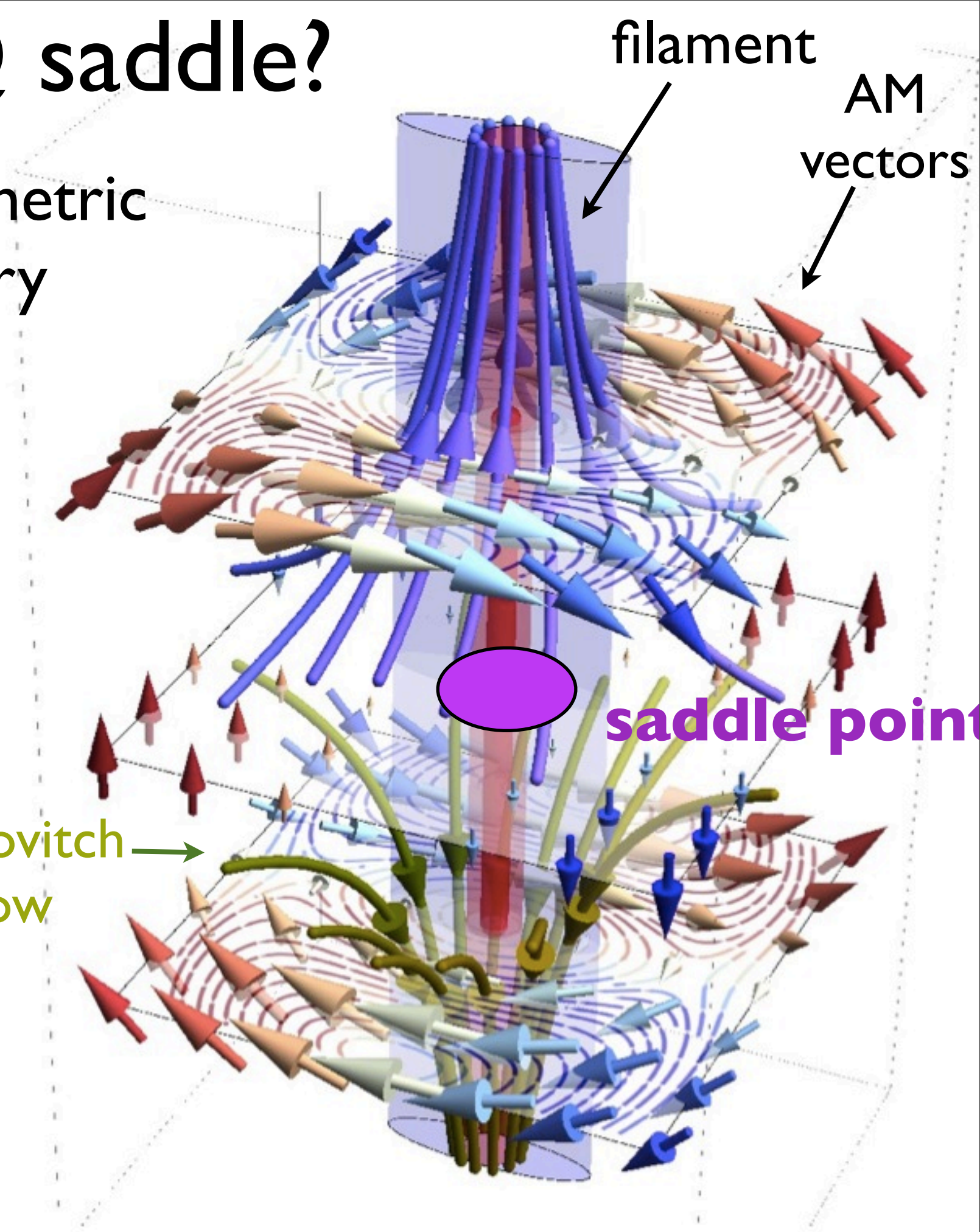
2D

3D TTT@ saddle?

- point reflection symmetric
- vanish if no asymmetry

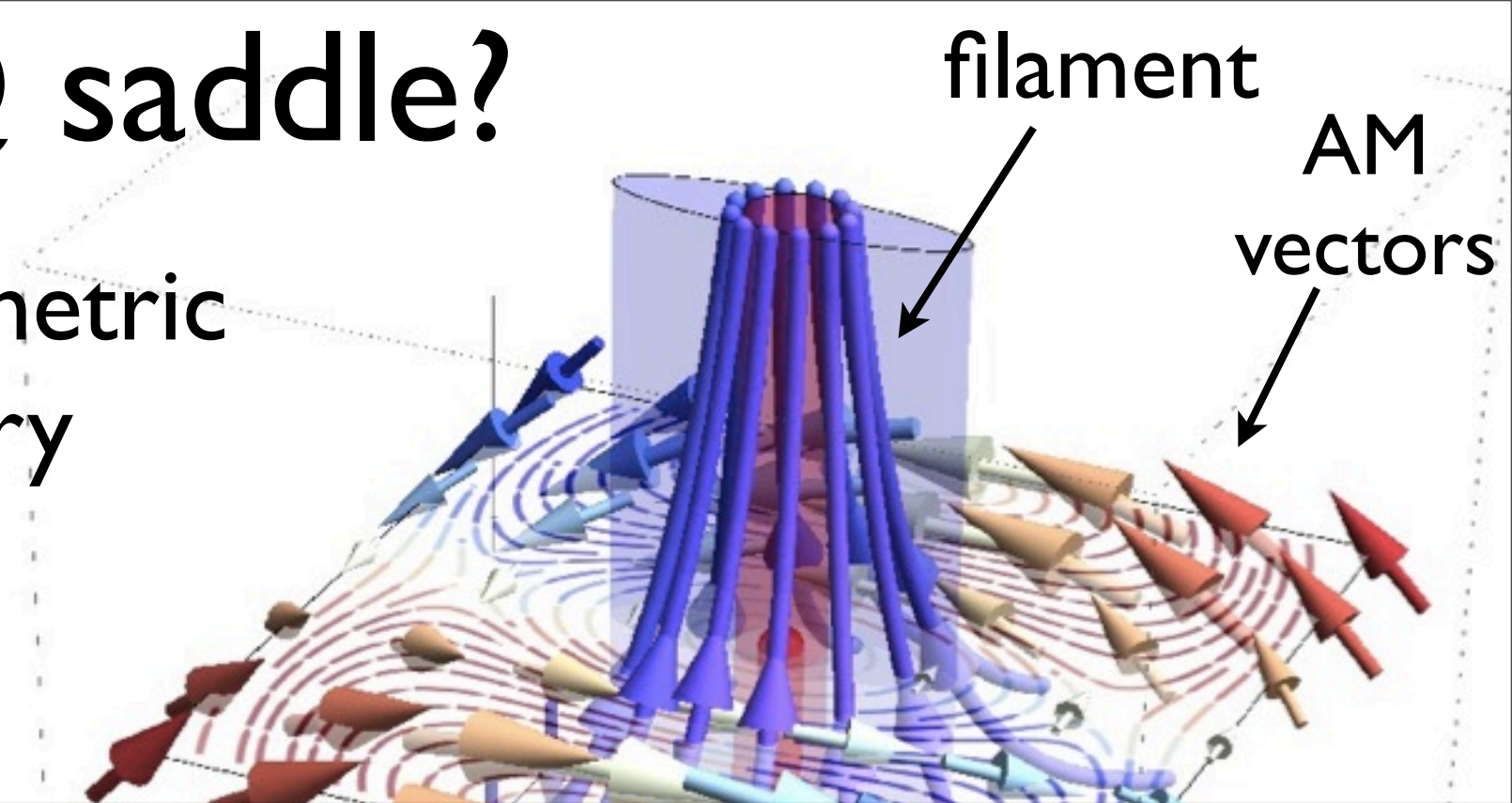
perp.
↑
spin //
to filament
↓
perp.

Zeldovitch
flow →

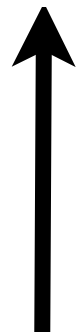


3D TTT@ saddle?

- point reflection symmetric
- vanish if no asymmetry



perp.



spin //

to filament



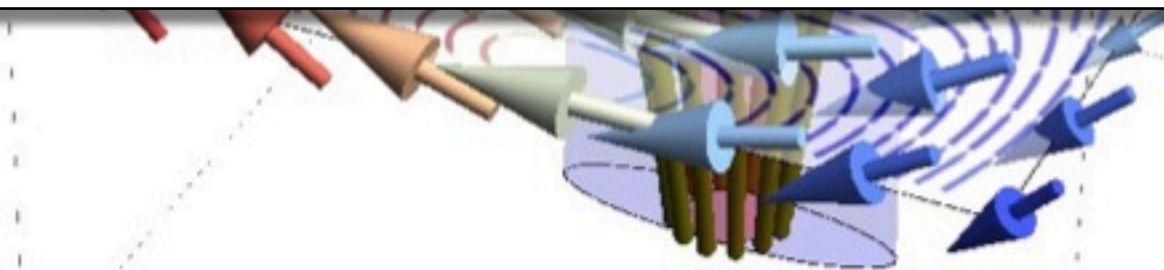
perp.

$$L_x = \frac{1}{210} (\kappa_1 + 2\kappa_2) r z \sin(\theta) (7\nu - 10\sqrt{3}\kappa_1),$$

$$L_y = \frac{1}{210} (2\kappa_1 + \kappa_2) r z \cos(\theta) (10\sqrt{3}\kappa_2 - 7\nu),$$

$$L_z = \frac{1}{420} (\kappa_1 - \kappa_2) r^2 \sin(2\theta) (10\sqrt{3}(\kappa_1 + \kappa_2) + \nu(7 - 4z^2))$$

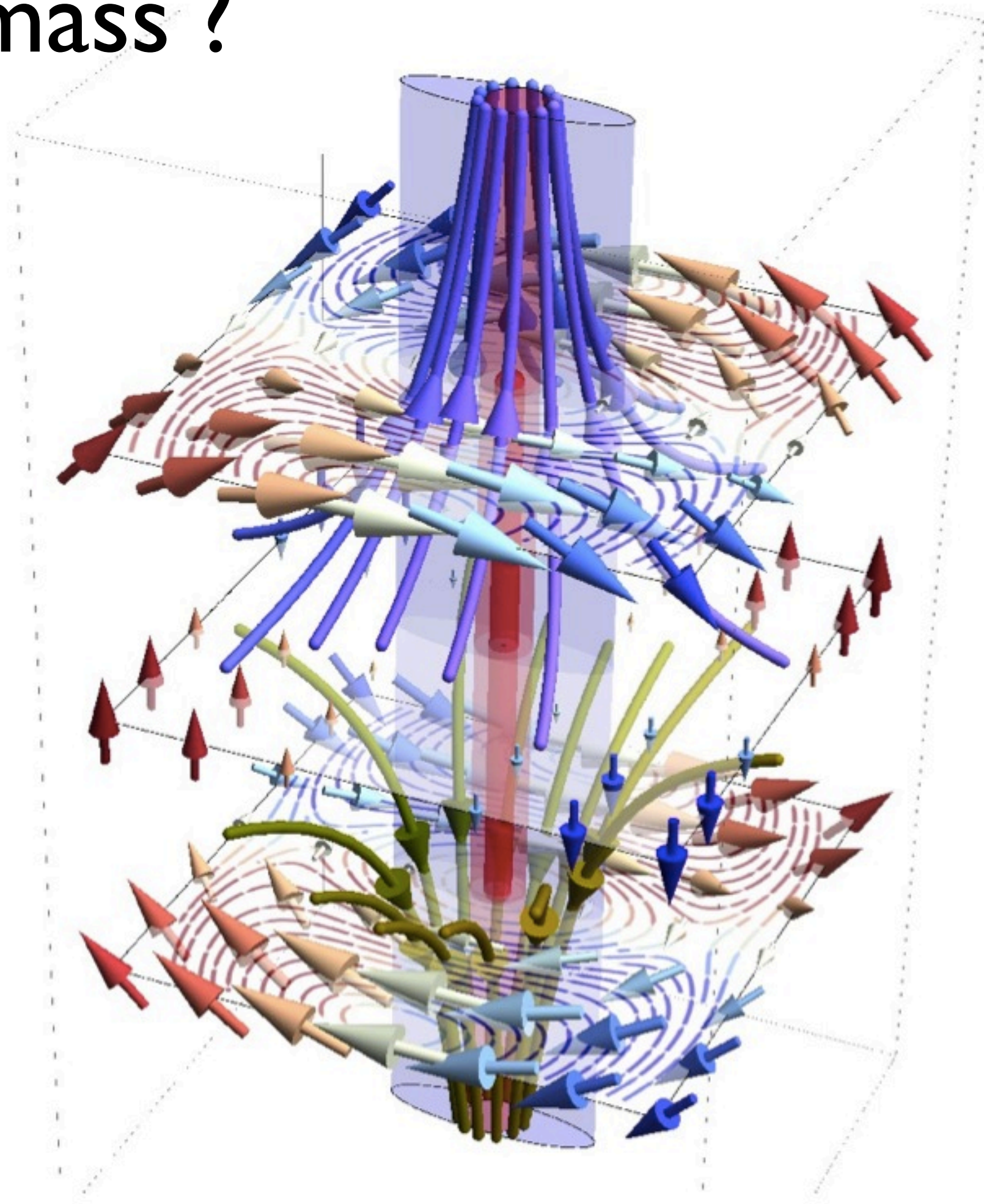
We check that it is indeed point reflection symmetric with respect to the saddle point, and it vanishes when $\kappa_1 = \kappa_2 =$



3D Transition mass ?

Lagrangian theory
capture spin flip !

Transition mass
associated
with **size**
of quadrant



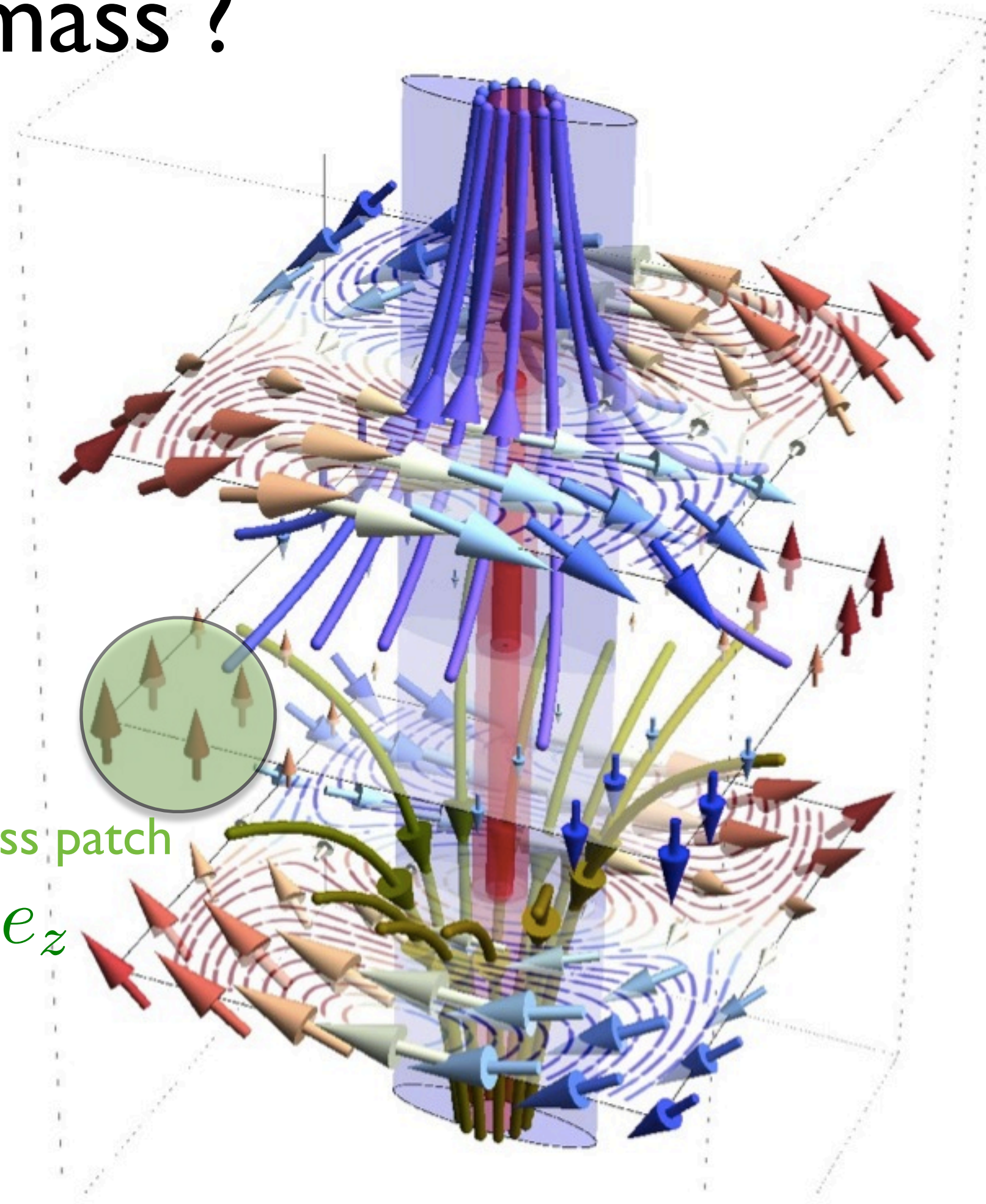
3D Transition mass ?

Lagrangian theory
capture spin flip !

Transition mass
associated
with **size**
of quadrant

Low mass patch

$$L \propto e_z$$



3D Transition mass ?

Lagrangian theory
capture spin flip !

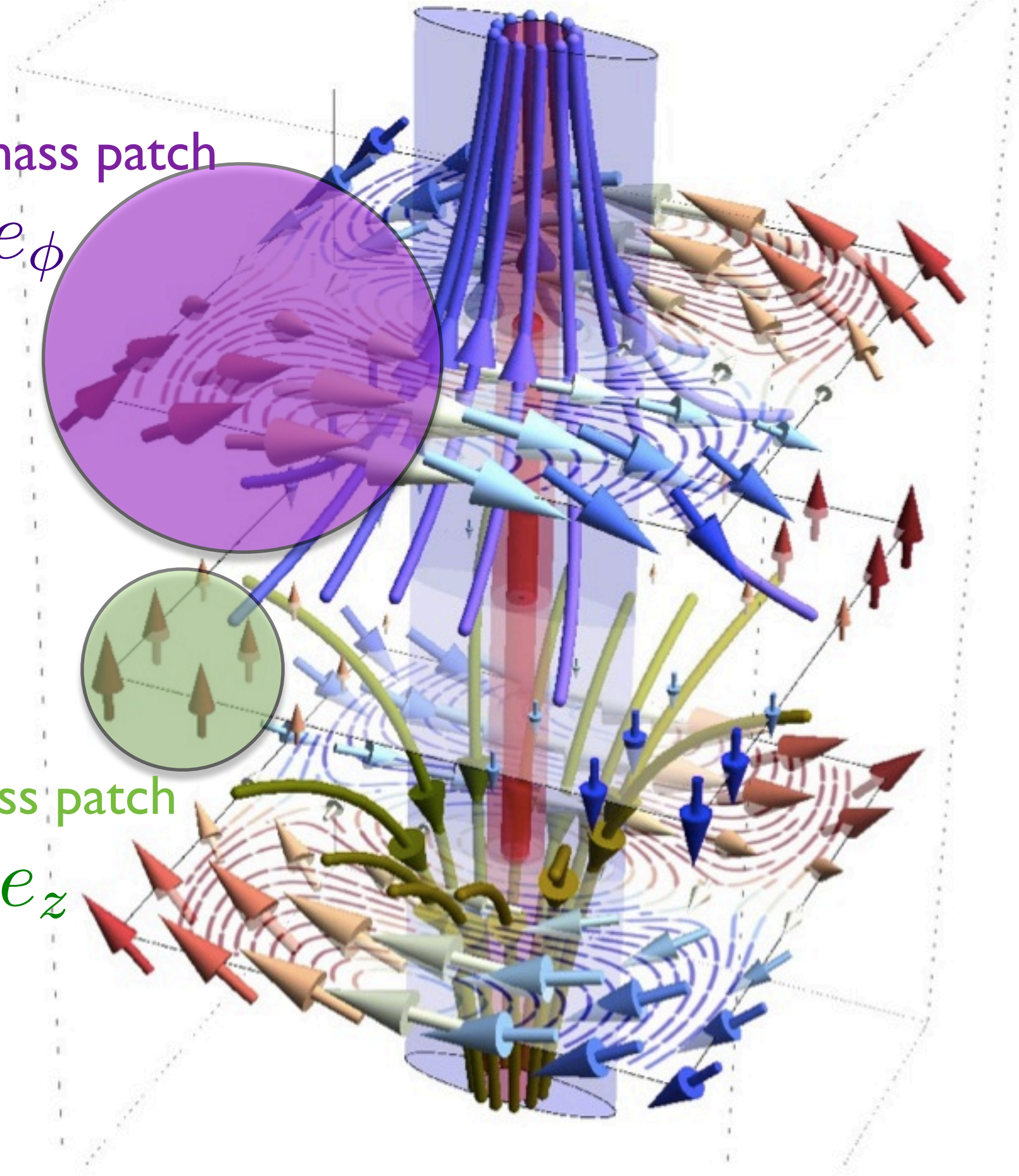
Transition mass
associated
with **size**
of quadrant

High mass patch

$$L \propto e_{\phi}$$

Low mass patch

$$L \propto e_z$$



Geometry of the saddle provides a **natural ‘metric’** (local frame as defined by Hessian @ saddle) relative to which **dynamical evolution** of DH is predicted.

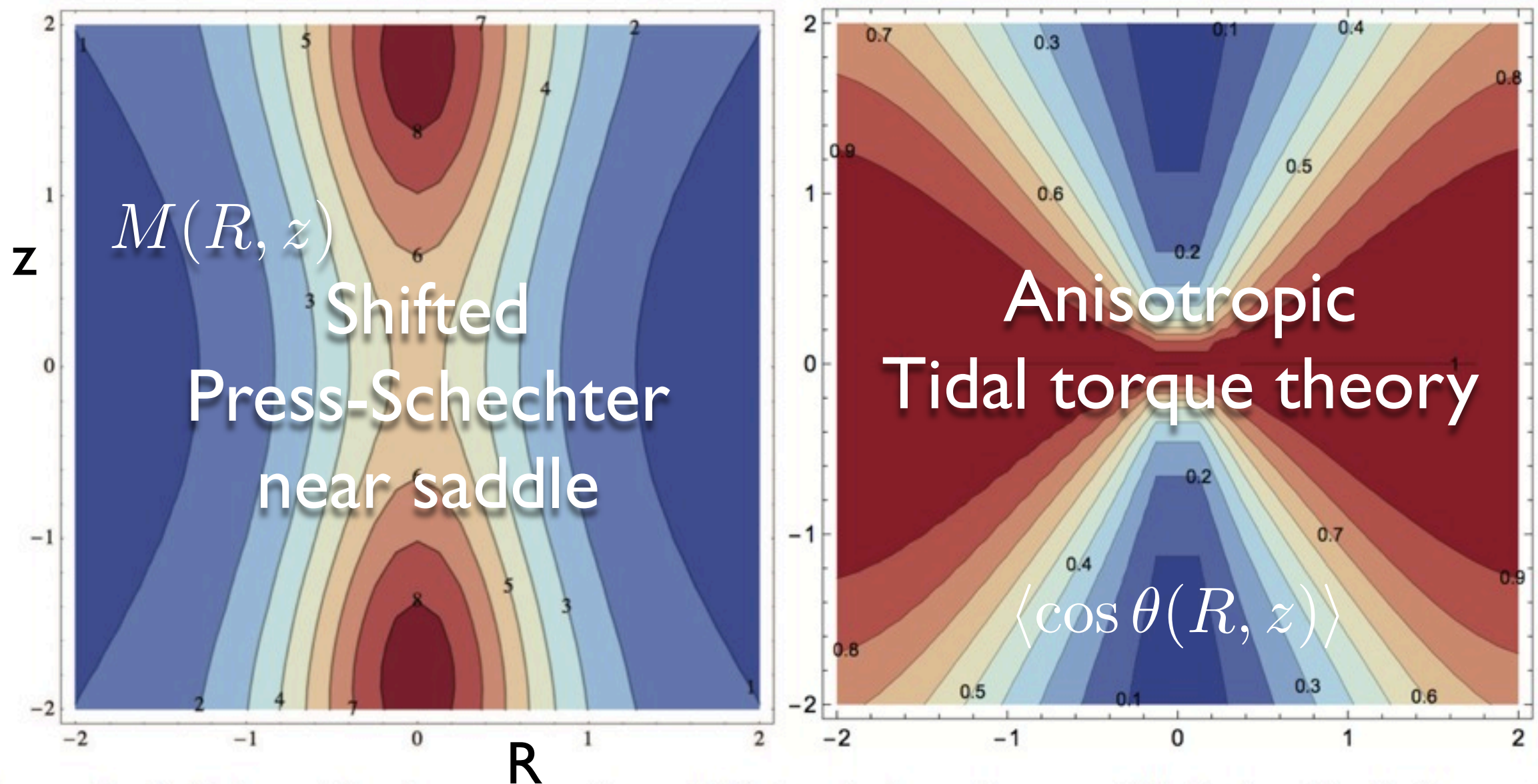


Figure 5. *Left:* logarithmic cross section of $M_p(r, z)$ along the most likely (vertical) filament (in units of $10^{12} M_\odot$). *Right:* corresponding cross section of $\langle \cos \hat{\theta} \rangle(r, z)$. The mass of halos increases towards the nodes, while the spin flips.

Geometry of the saddle provides a **natural ‘metric’** (local frame as defined by Hessian @ saddle) relative to which **dynamical evolution** of DH is predicted.

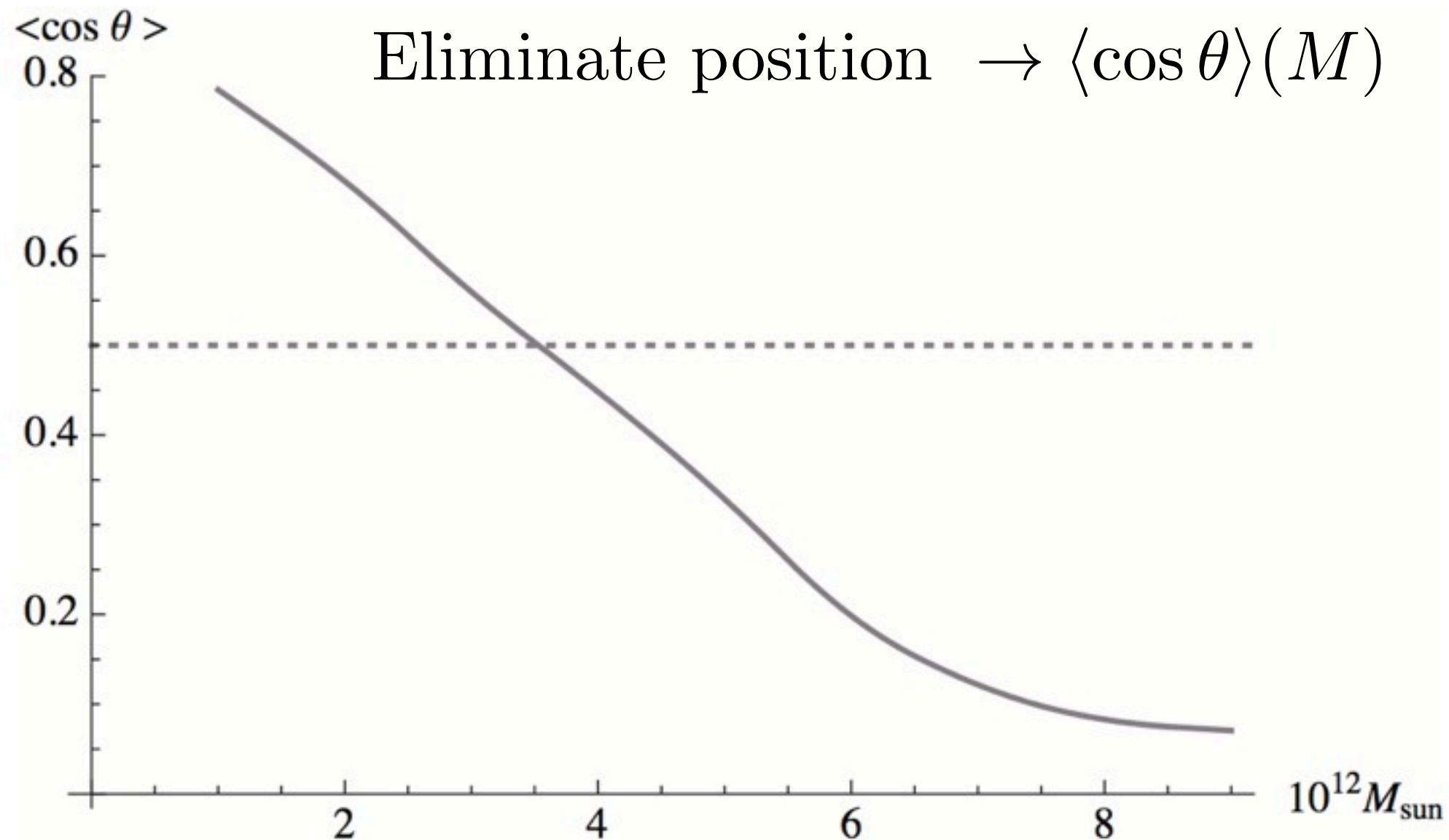
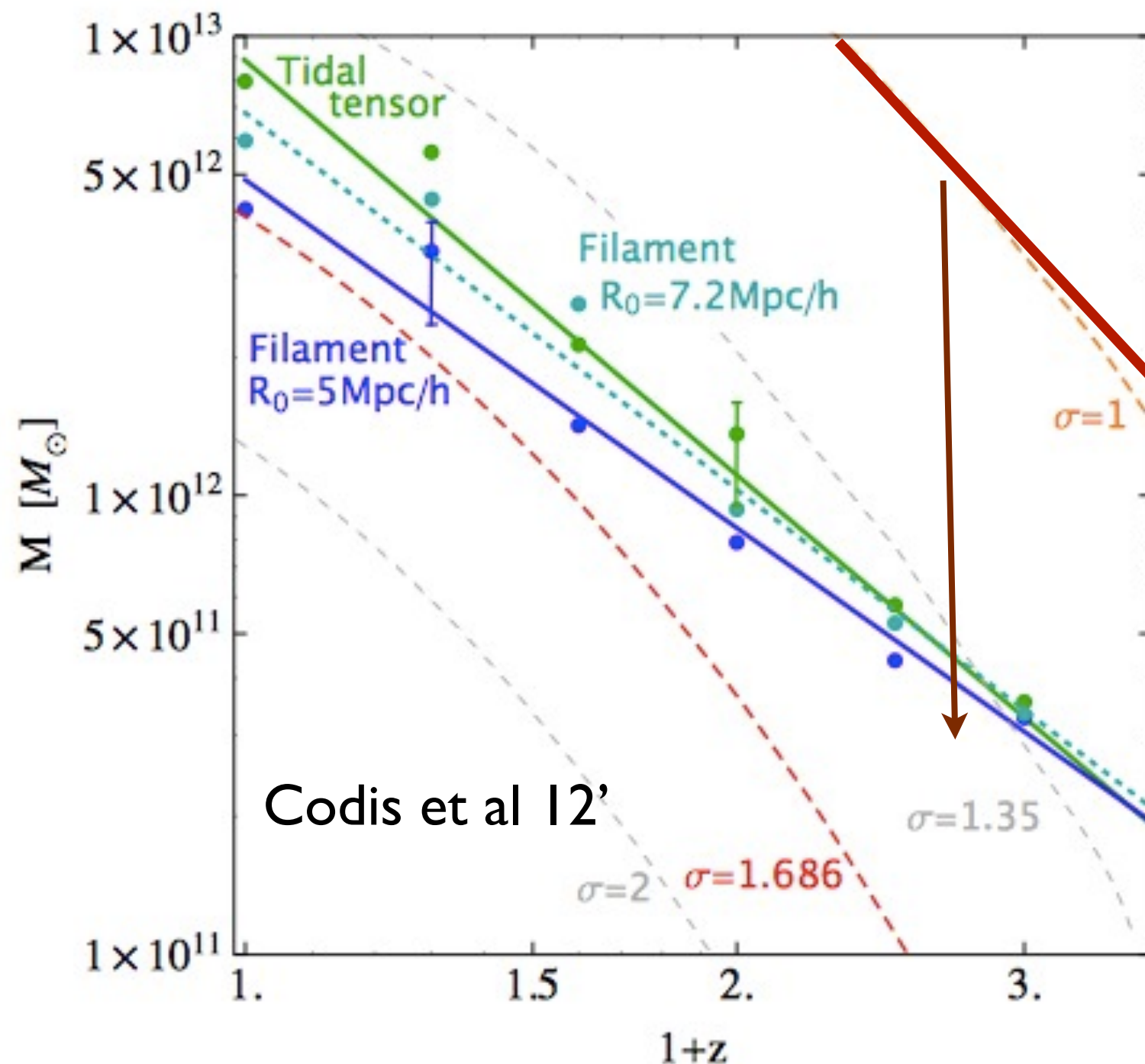


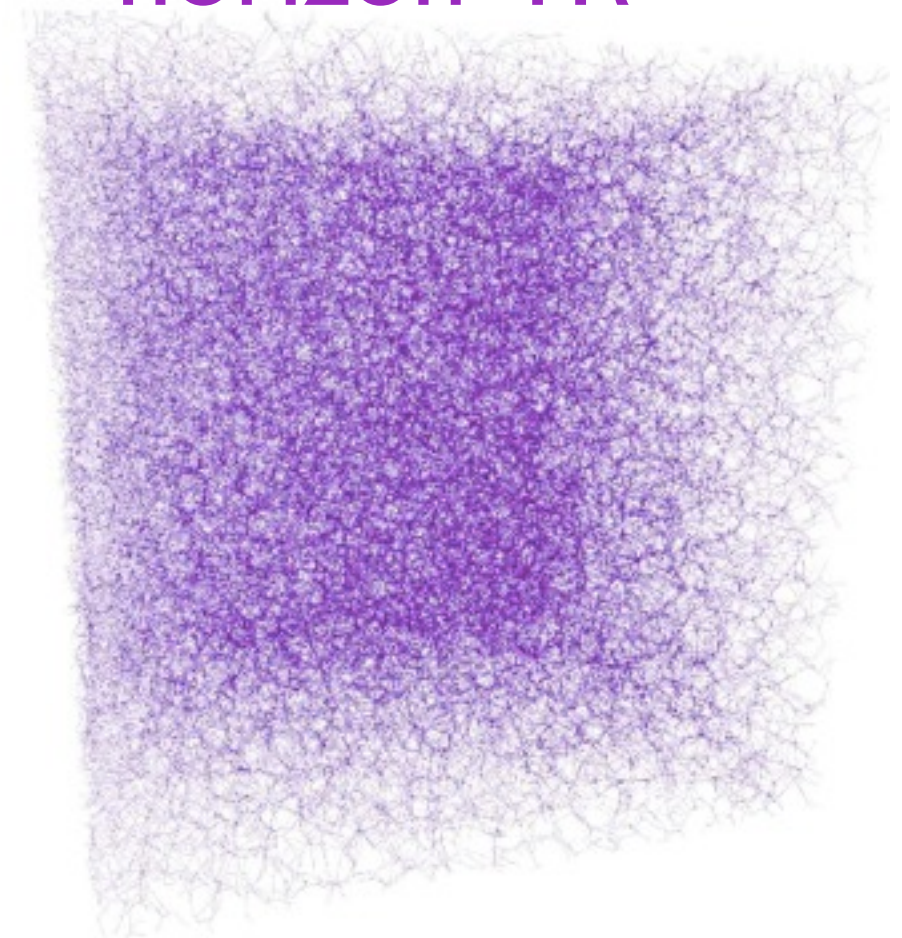
Figure 6. Mean alignment between spin and filament as a function of mass for a filament smoothing scale of 5 Mpc/ h . The spin flip transition mass is around $4 \cdot 10^{12} M_{\odot}$.
towards the nodes, while the spin flips.

Explain transition mass?

Transition mass versus redshift



horizon 4π

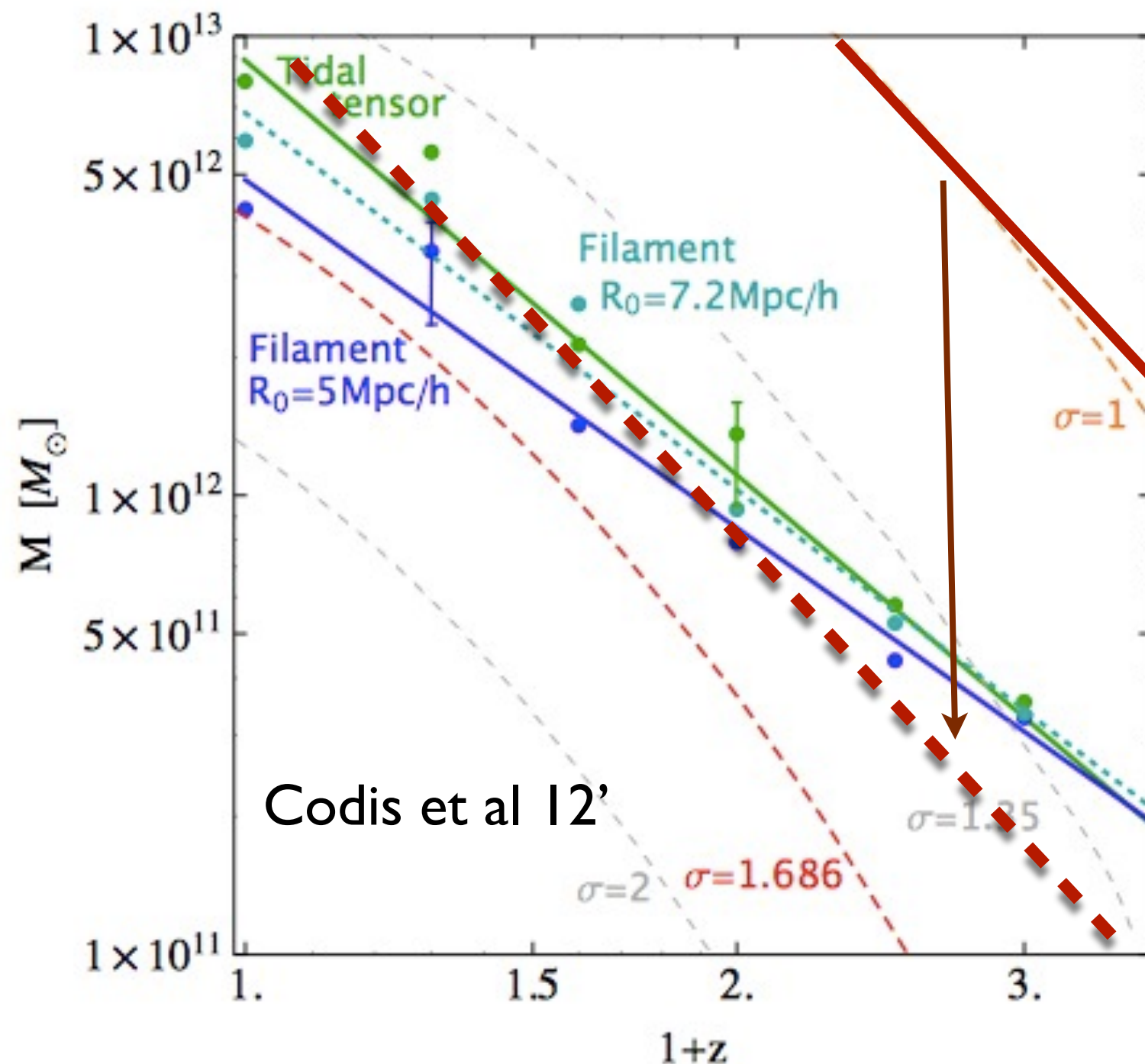


skeleton of LSS

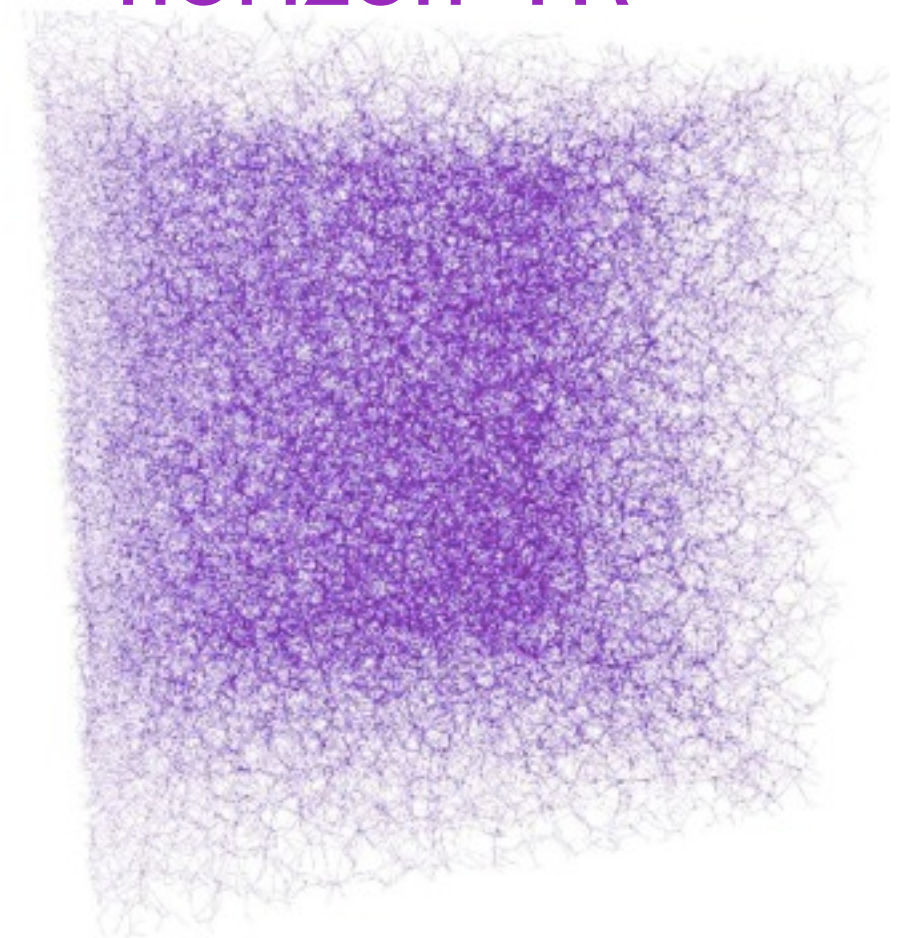
Only 2 ingredients: a) spin is spin one b) filaments flattened

Explain transition mass?

Transition mass versus redshift



horizon 4π



skeleton of LSS

Only 2 ingredients: a) spin is spin one b) filaments flattened

Does it work with log-Gaussian Random Fields?

2D

point reflection symmetry
for realistic sets of saddles
from log GRF

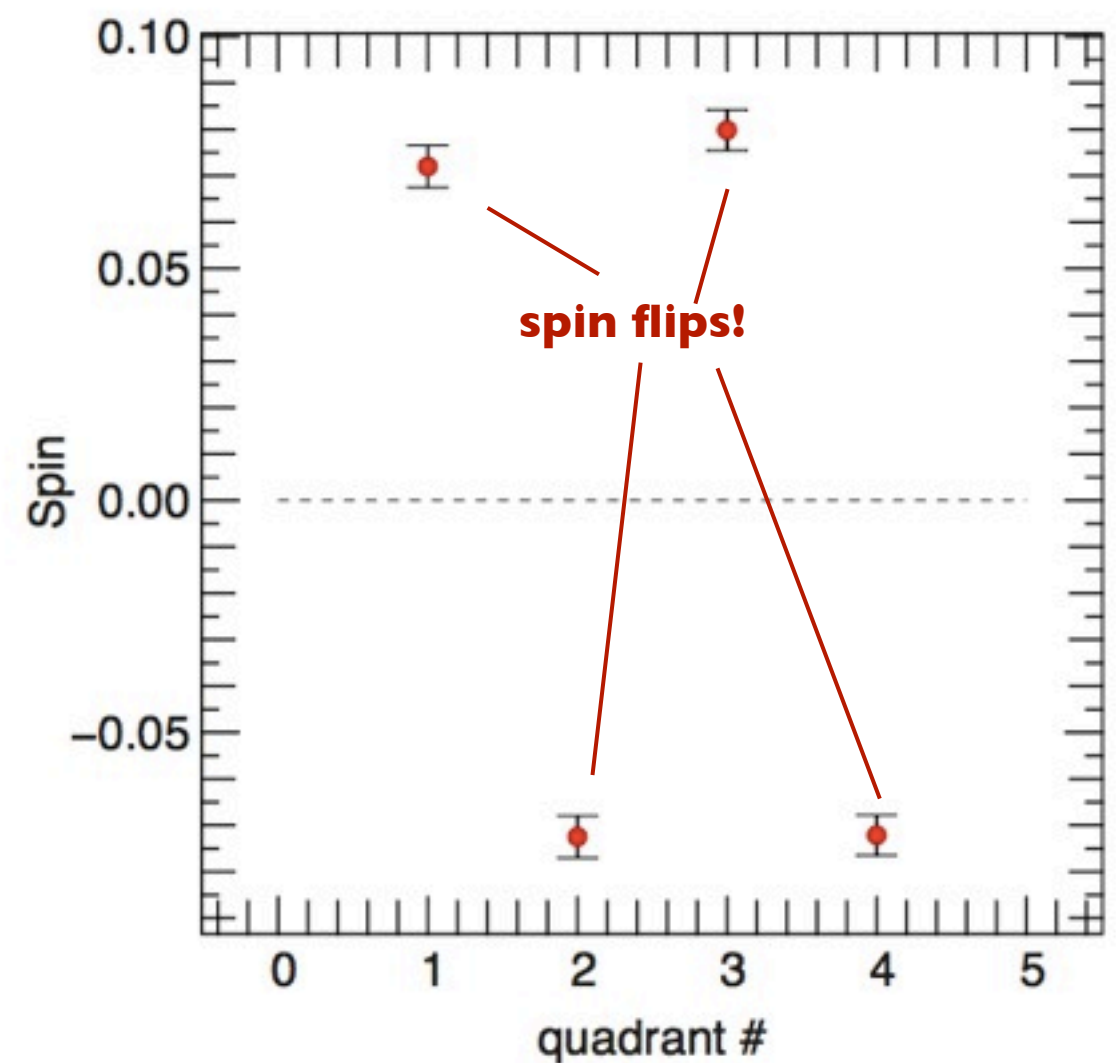
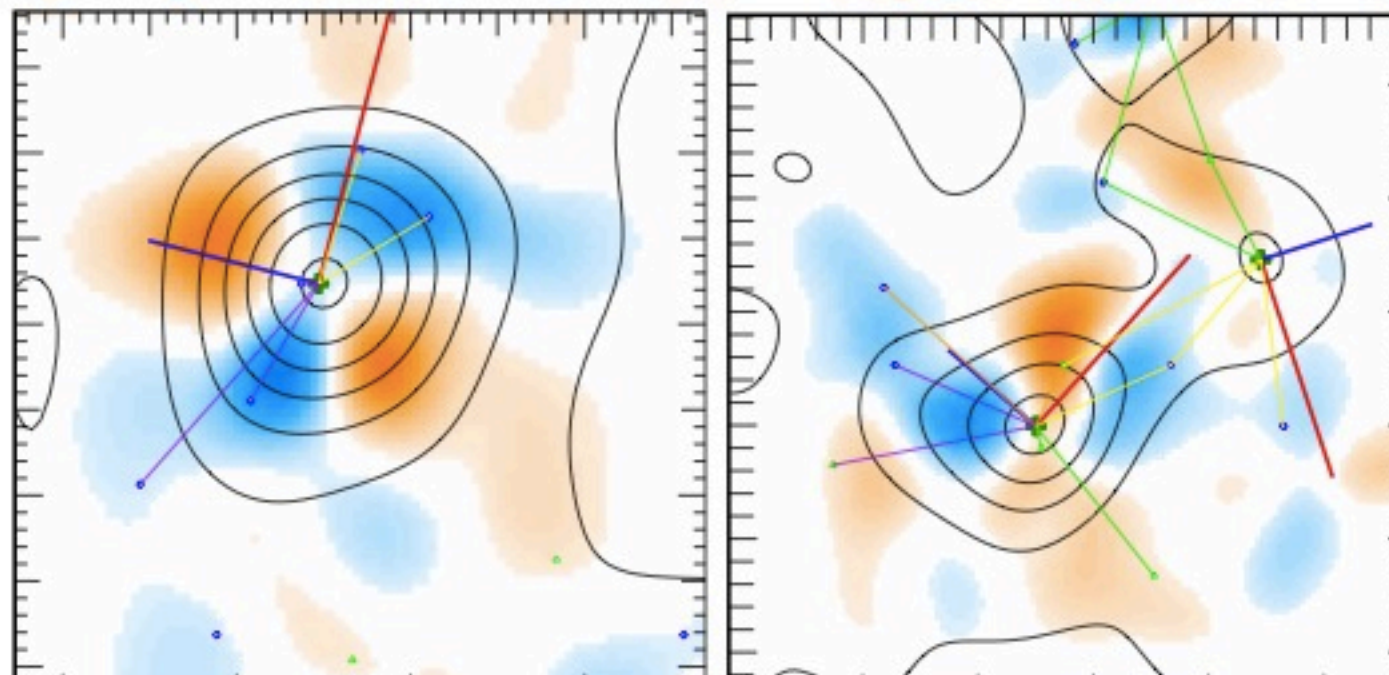


Figure 11. Alignment of 'spin' along e_z in 2D as a function of quadrant rank, clockwise. As expected, from one quadrant to the next, the spin is flipping sign.

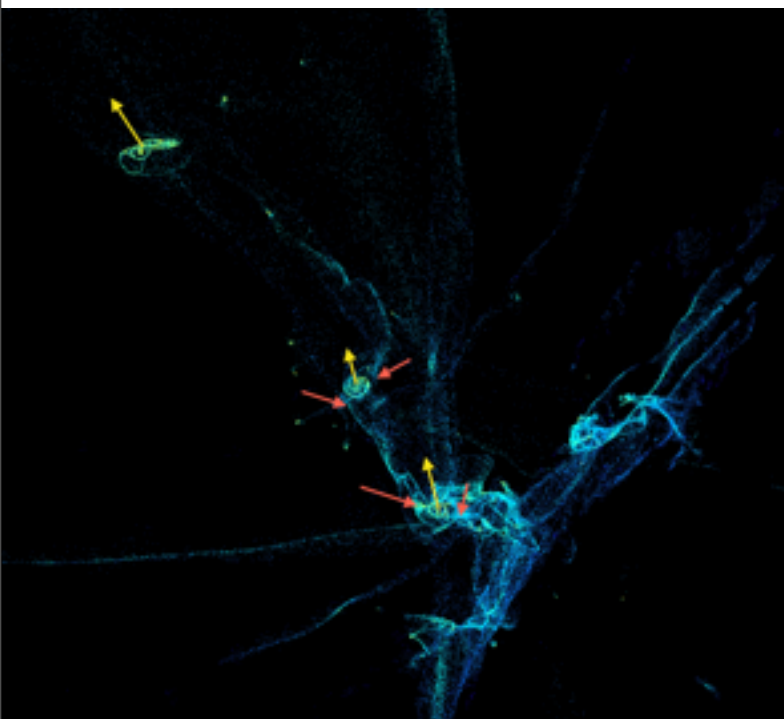
Filaments

=

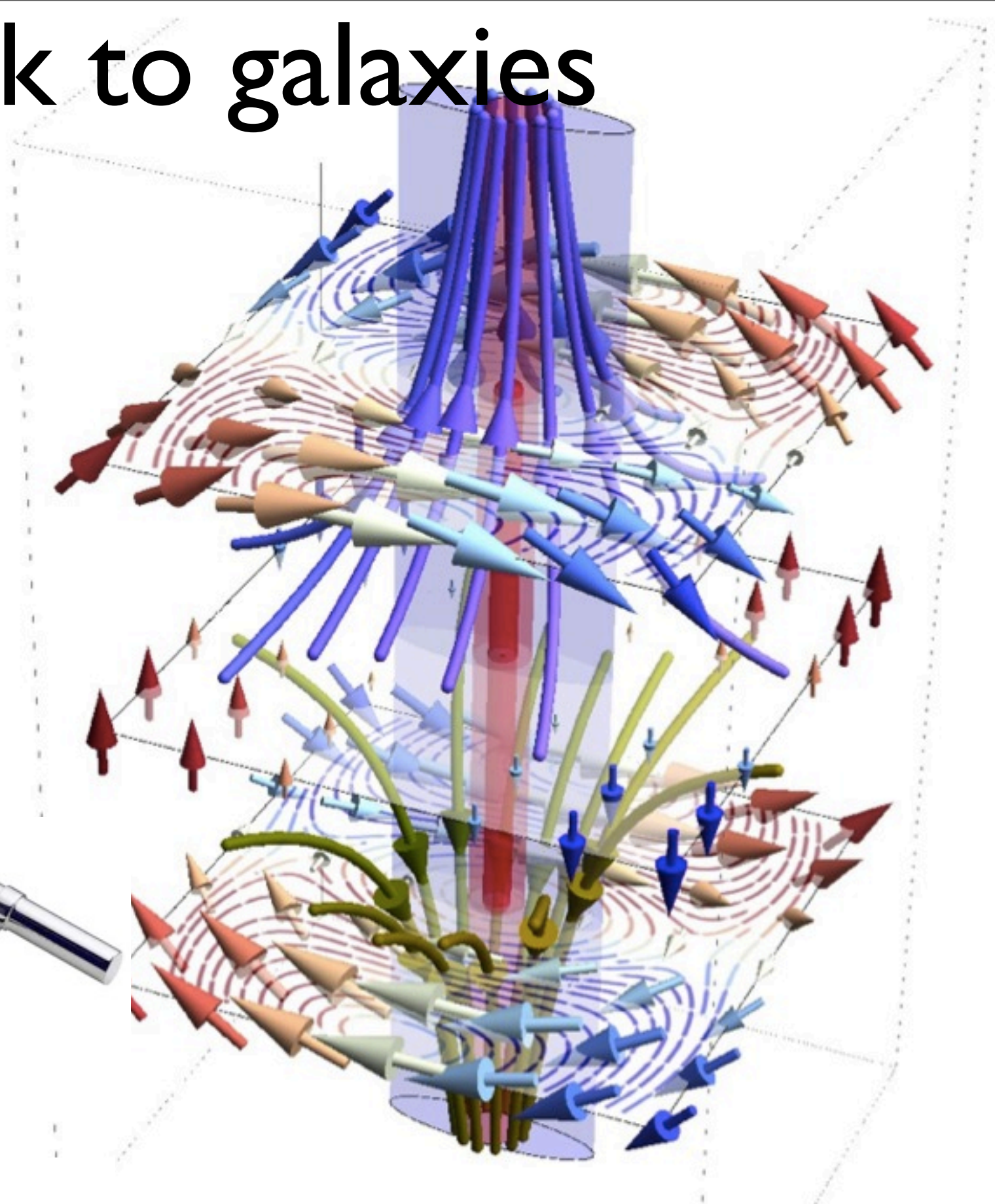
metric for spin build up
of galaxies

AM built up via cold flows
occurs near **extrema of
helicity** (v.L) either side of
saddles!

$$\varepsilon_{ijk} \delta_{li} \psi_{lj} \psi_k$$



Back to galaxies



define epoch of maximal AM accretion

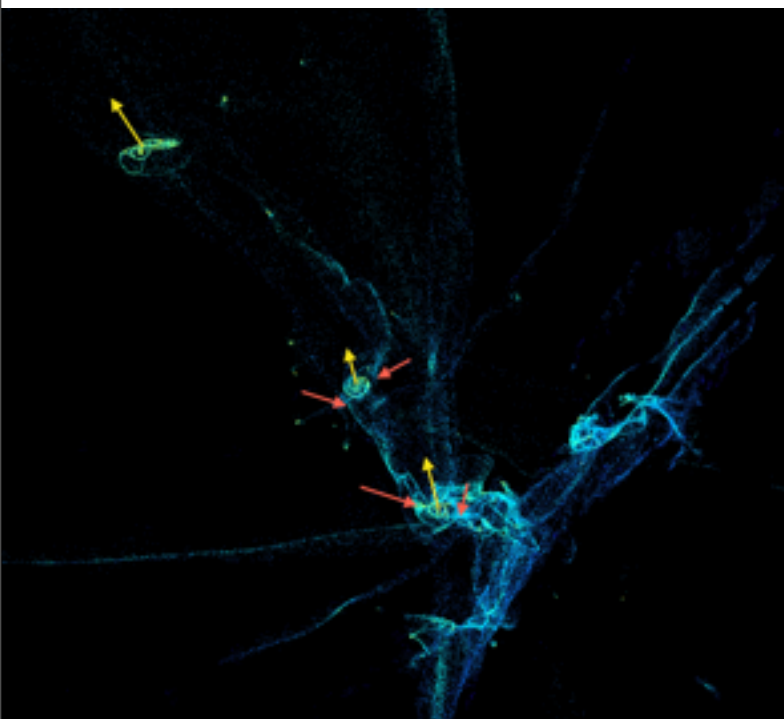
Filaments

=

metric for spin build up
of galaxies

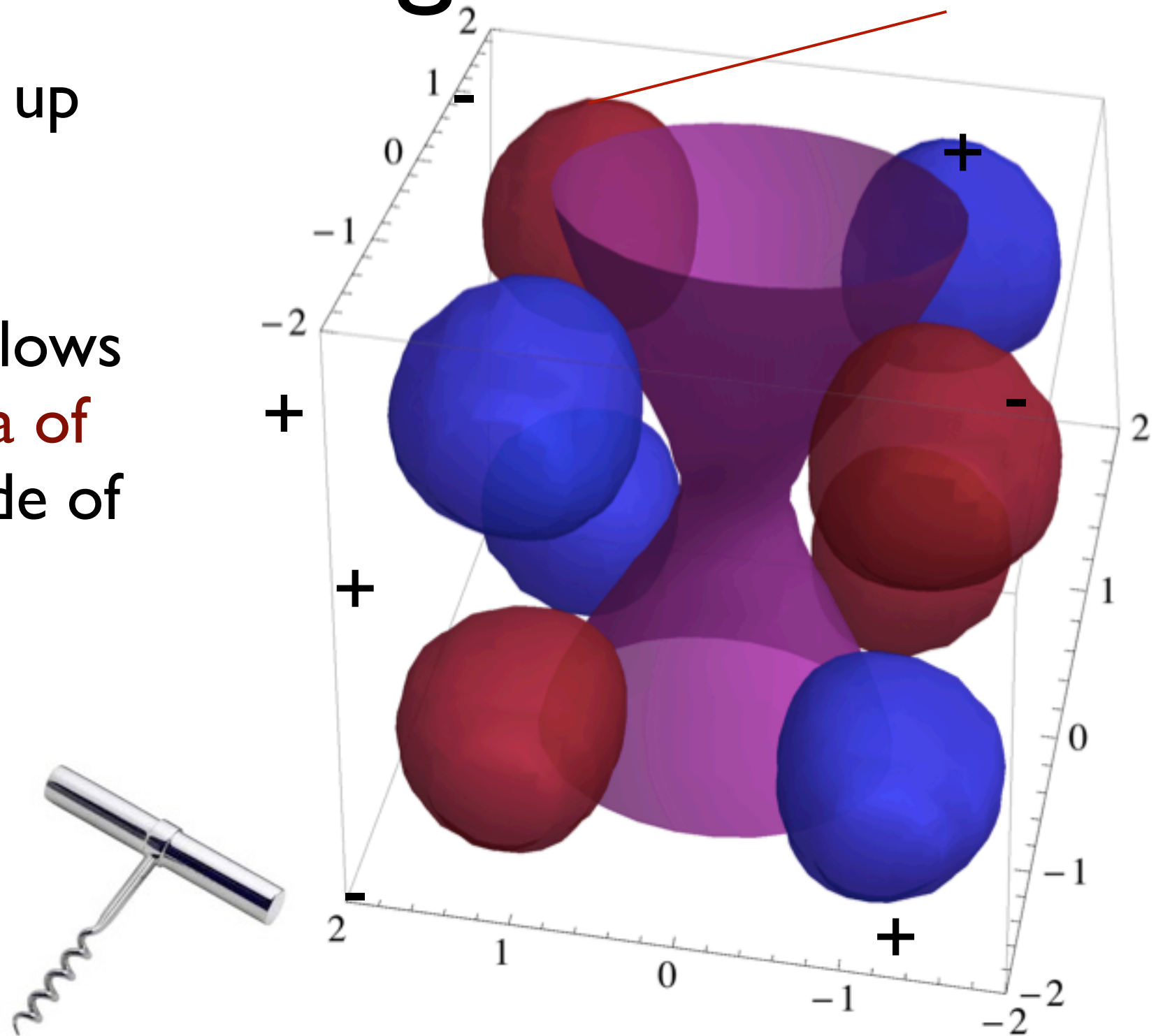
AM built up via cold flows
occurs near **extrema of
helicity** (v.L) either side of
saddles!

$$\varepsilon_{ijk} \delta_{li} \psi_{lj} \psi_k$$



Back to galaxies

**predicted Lagrangian
extrema of 'helicity'**



define epoch of maximal AM accretion

Take home message...

- Morphology (= AM stratification) driven by LSS via *cold flows* in cosmic web: it explains Es & Sps where, how & why from **ICs**
- Signature in correlation between *morphology* and *internal kinematic structure* of cosmic web.
- Process driven by simple dynamics: $t_{\text{dyn}} \sim 1/\sqrt{\rho}$, shock
 - requires updating TTT to **saddles**: simple theory :-)

Where galaxies form does matter, and can be traced back to ICs
Flattened filaments generate point-reflection-symmetric AM/vorticity distribution:
they induce the observed spin transition mass & helicity of cold flows

For more details: Pichon et al. 2011 Codis et al 2012, Tillson et al 2012,
Laigle et al 2014 Dubois et al 2014 Welker et al 2014, Pichon et al. 2014

What about galaxies ??

- Horizon-AGN simulation Jade (CINES)
(PI Y. Dubois, Co-I J. Devriendt & C. Pichon)
 - $L_{\text{box}}=100 \text{ Mpc}/h$
 - 1024^3 DM particles $M_{\text{DM,res}}=8 \times 10^7 M_{\text{sun}}$
 - Finest cell resolution $dx=1 \text{ kpc}$
 - Gas cooling & UV background heating
 - Low efficiency star formation
 - Stellar winds + SNII + SNIa
 - O, Fe, C, N, Si, Mg, H
 - AGN feedback radio/quasar
- Outputs
(backed up and analyzed on BEYOND)
 - Simulation outputs
 - Lightcones ($1^\circ \times 1^\circ$) performed on-the-fly
 - Dark Matter (position, velocity)
 - Gas (position, density, velocity, pressure, chemistry)
 - Stars (position, mass, velocity, age, chemistry)
 - Black holes (position, mass, velocity, accretion rate)
- $z=1.5$ using 3 Mhours on 4096 cores

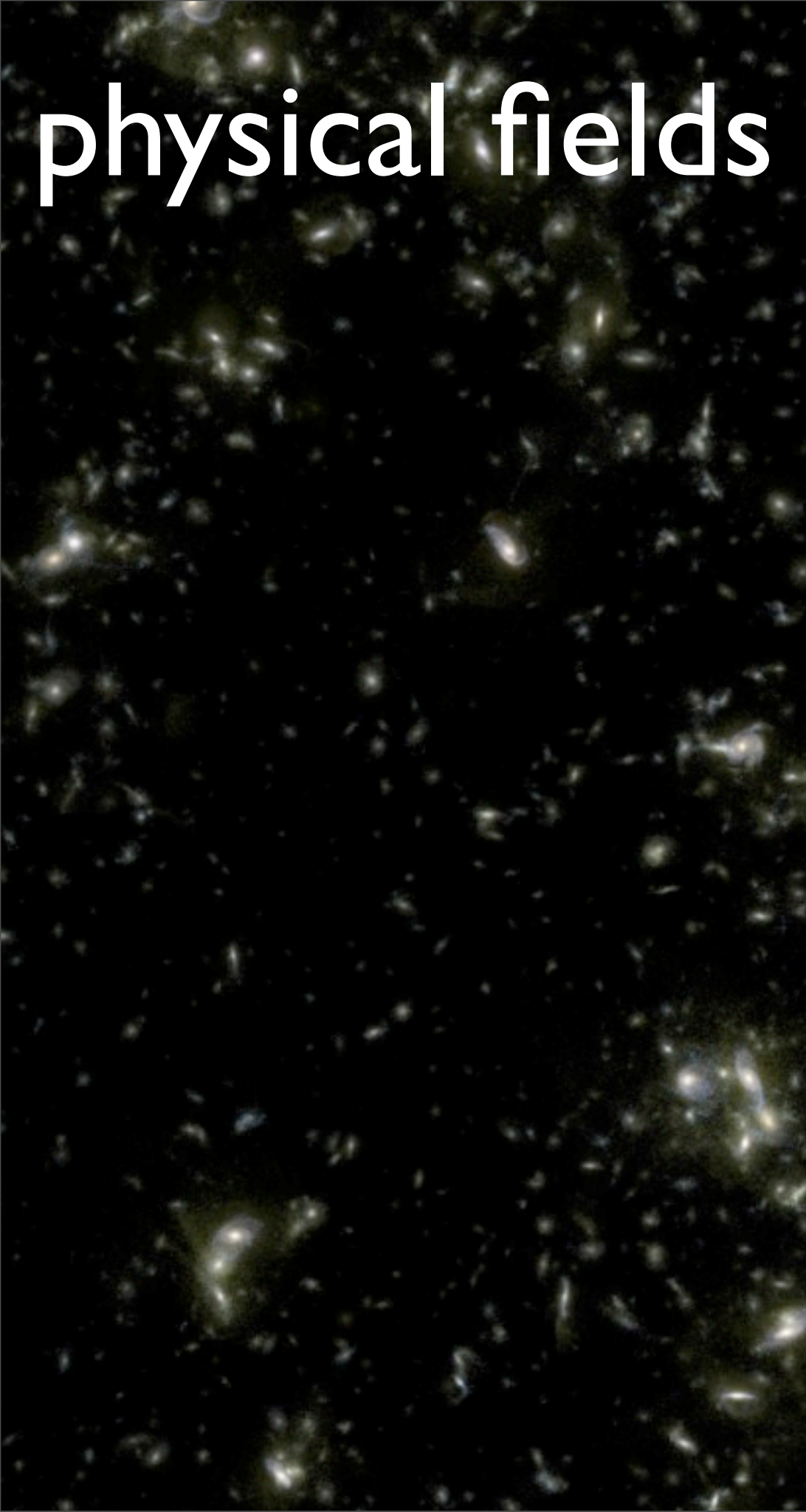
PART IV

horizon-AGN.projet-horizon.fr

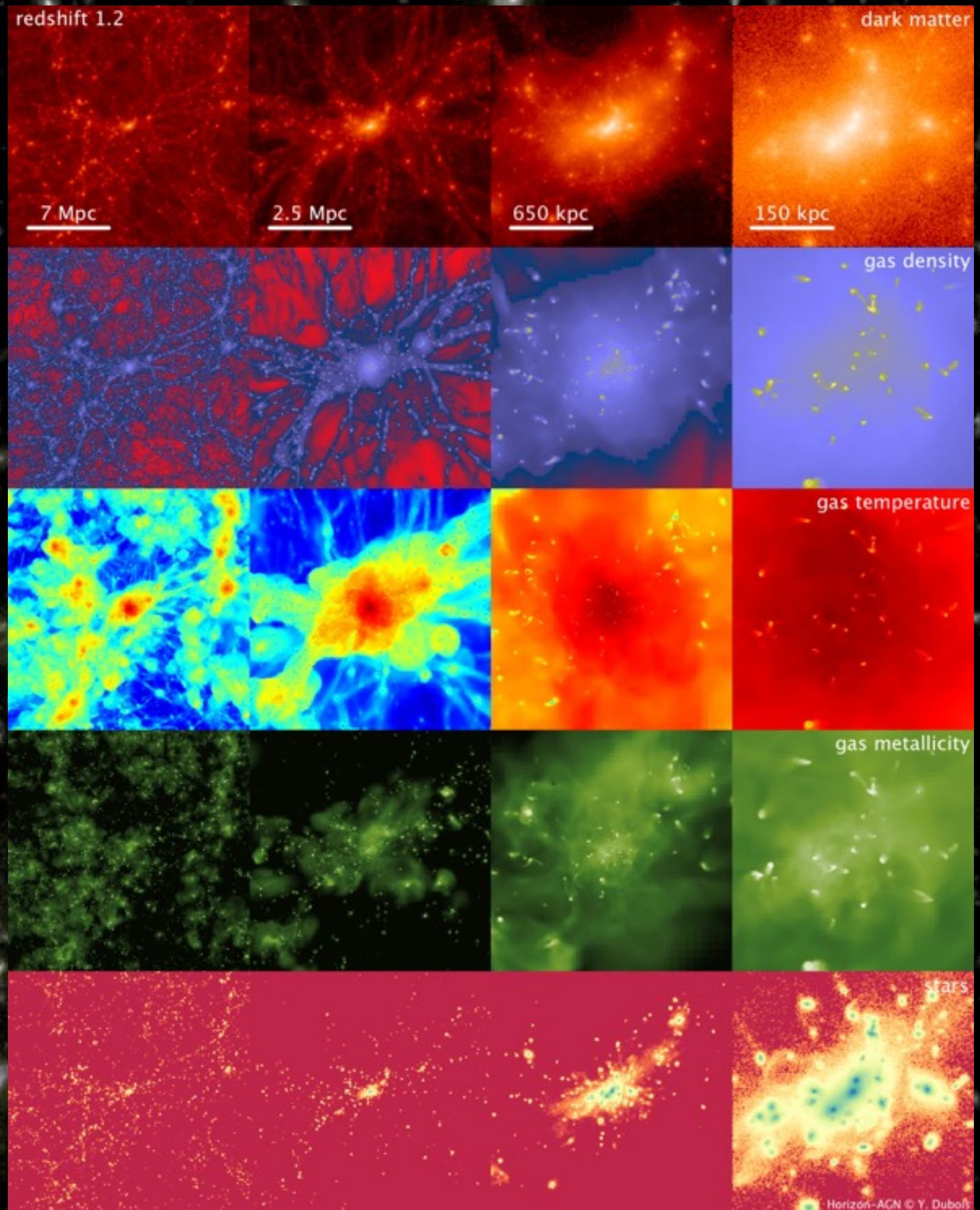
$z=1.2$

Part V Outline

- Can morphology/physics trace spin flip?
- Are transition masses consistent?
- The fate of forming galaxies
- The fate of merging galaxies



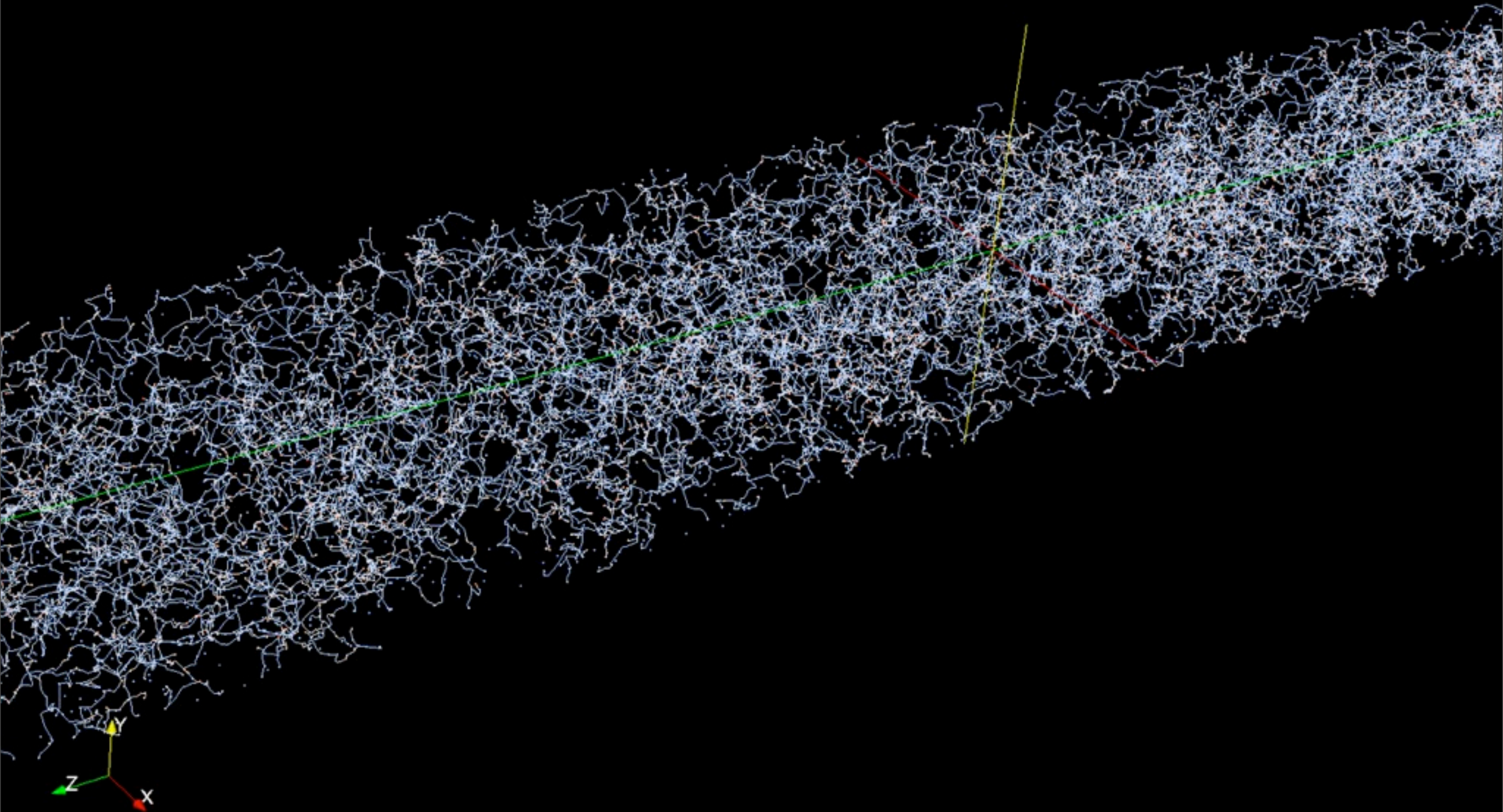
physical fields



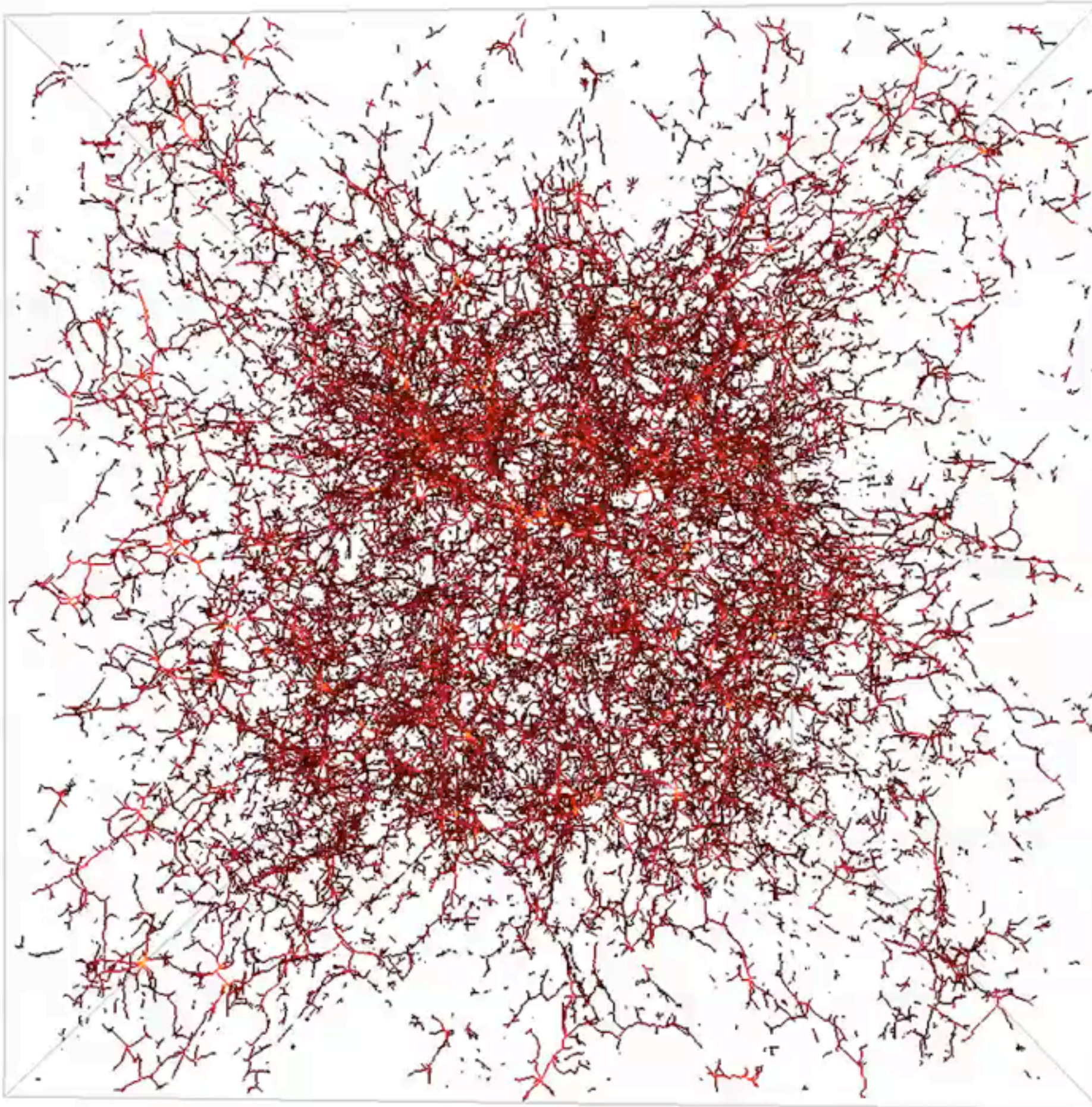
A visualization of the cosmic web, showing a dense network of purple and blue filaments and clusters against a dark green background. The structure represents the distribution of baryonic matter in the universe.

Baryonic light cone

Skeleton from BH catalogue

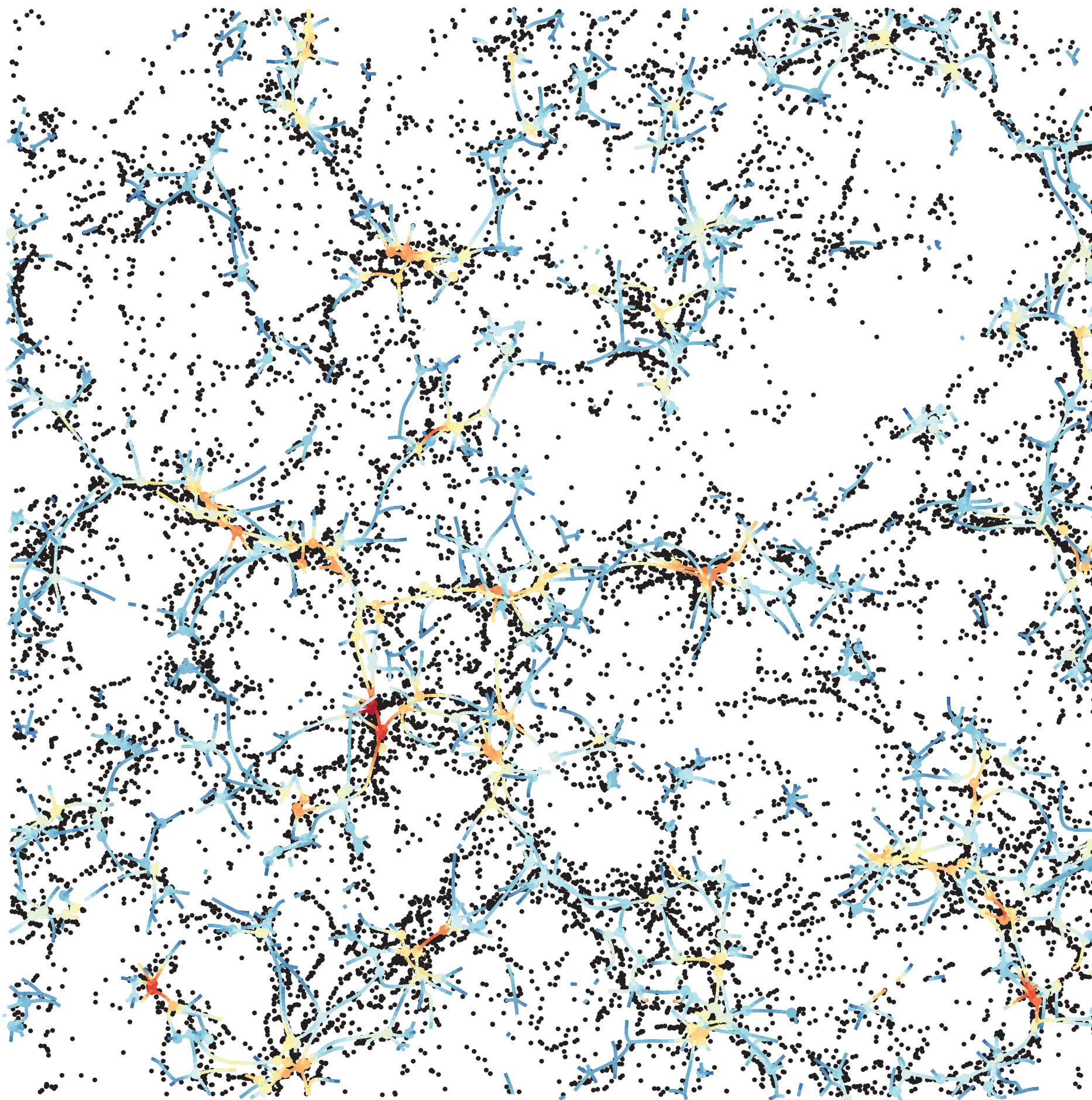


Skeleton of box



results on spin
alignment independent
of code used
(persistence/watershed)

Galaxies versus dense filaments

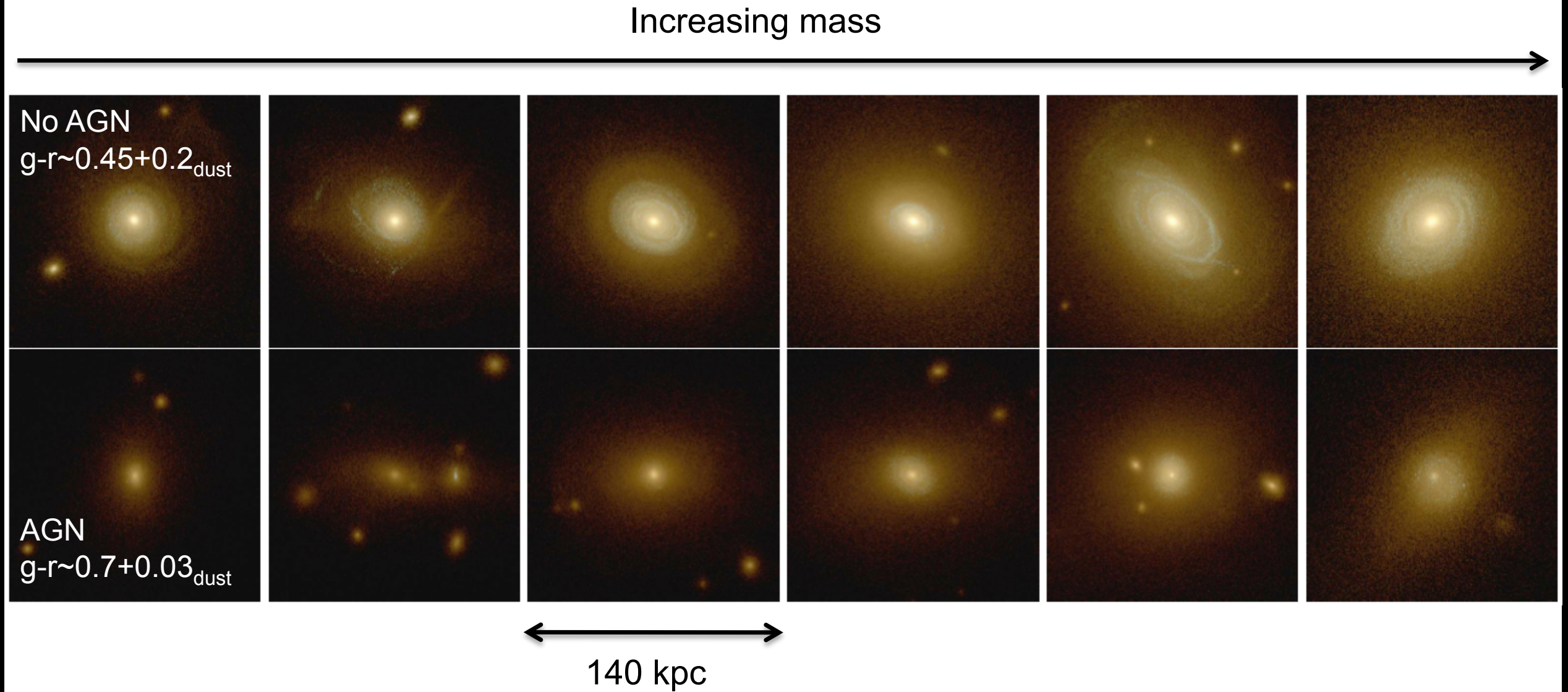


Galaxies are strongly
clustered
near filaments

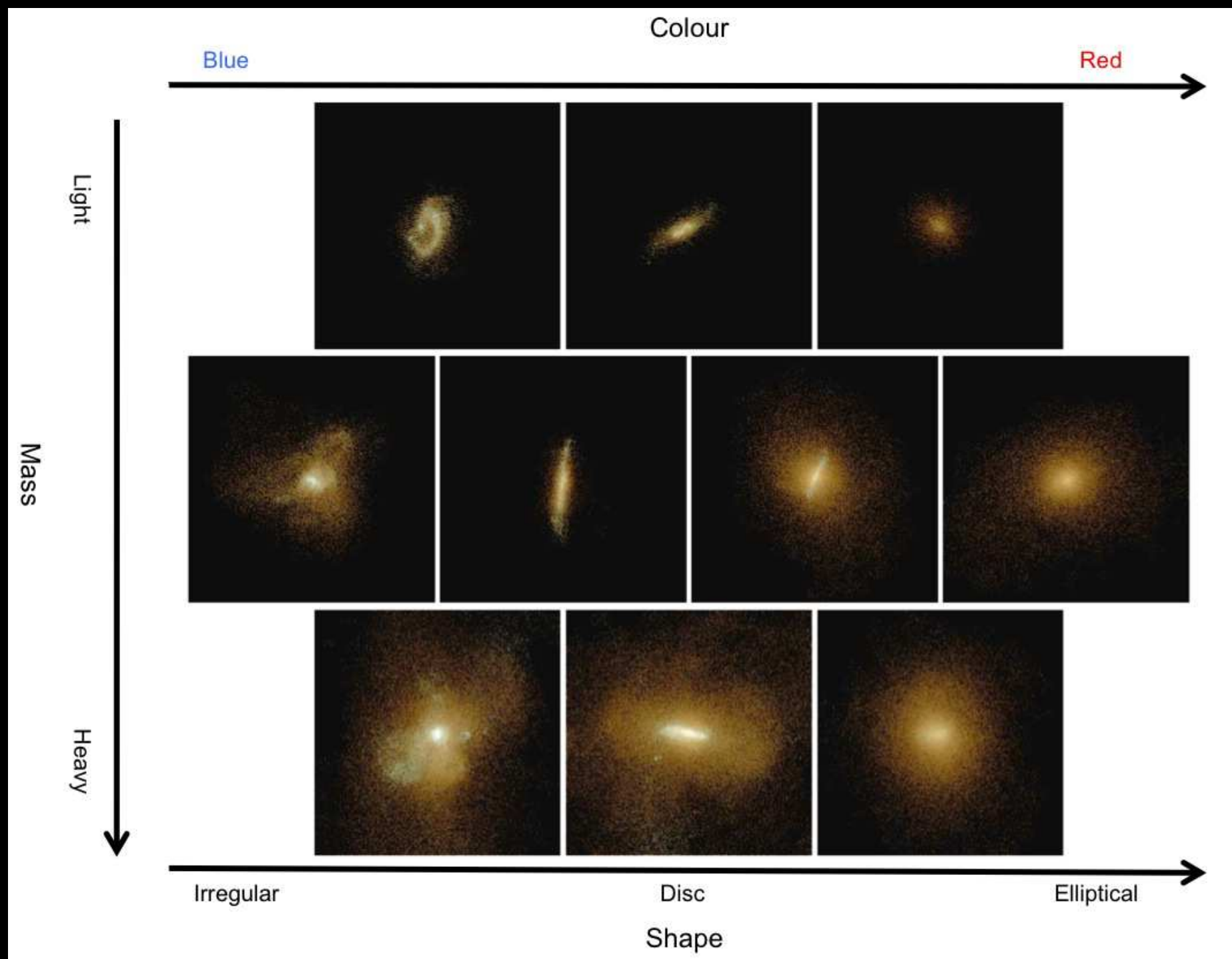
can morphology
trace spin flip ?

- thanks to AGN feedback we have
morphological diversity

Can we get massive galaxies that look like ellipticals ?



Dubois, Gavazzi, Peirani, Silk, 2013



Filament-galactic spin & mass

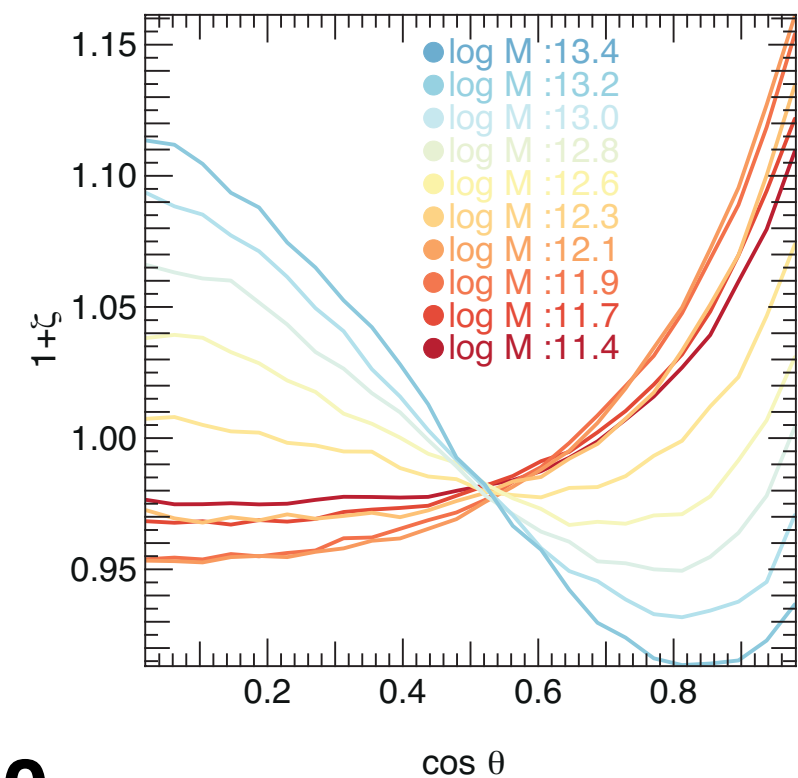
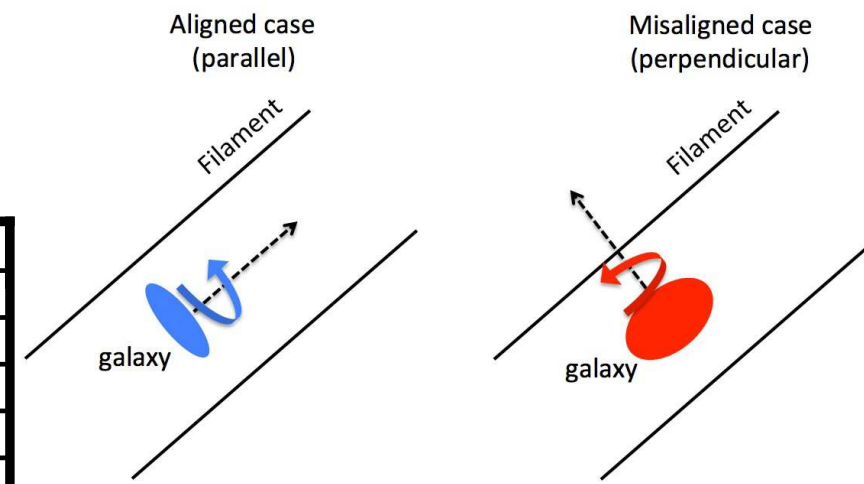
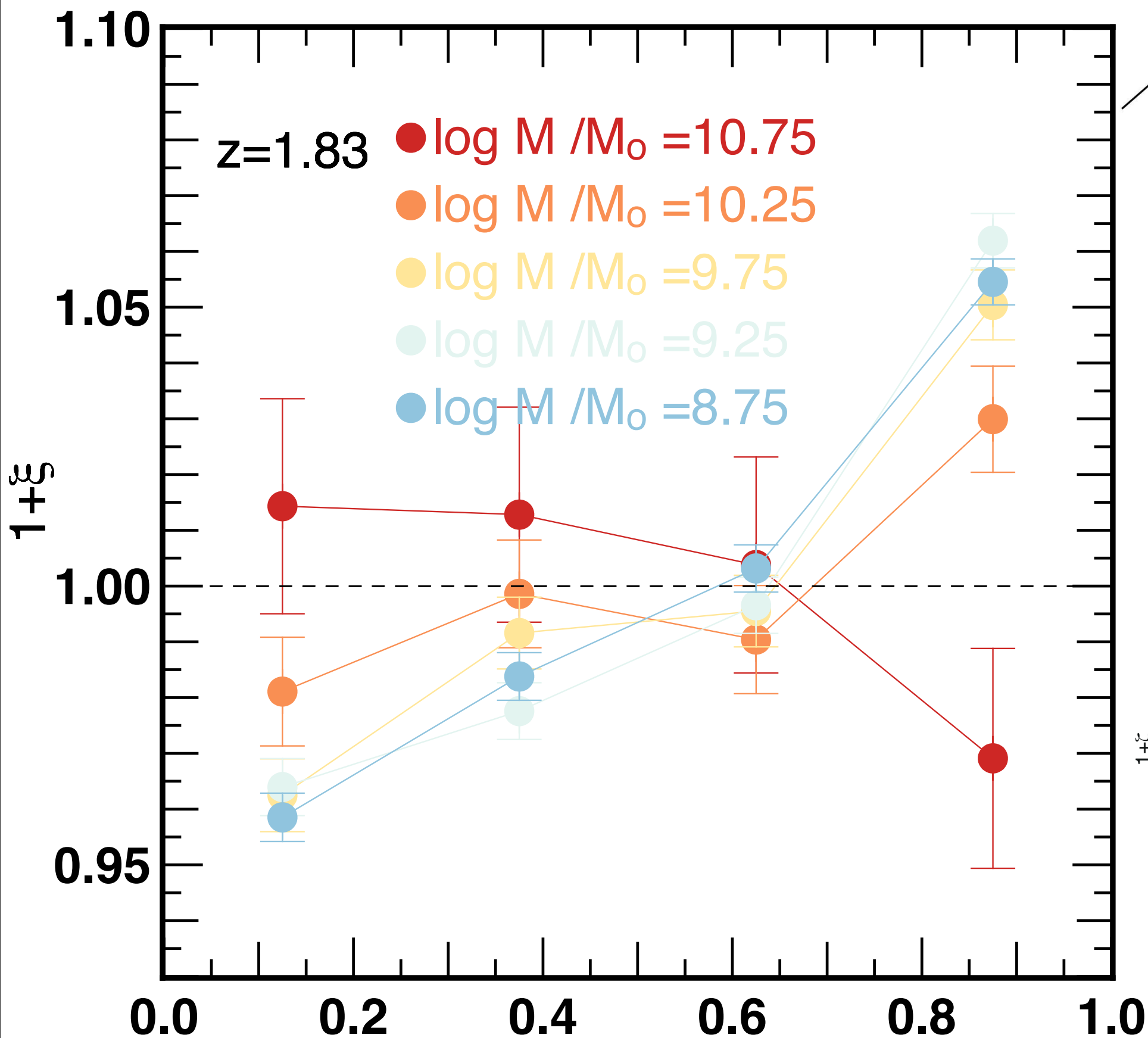
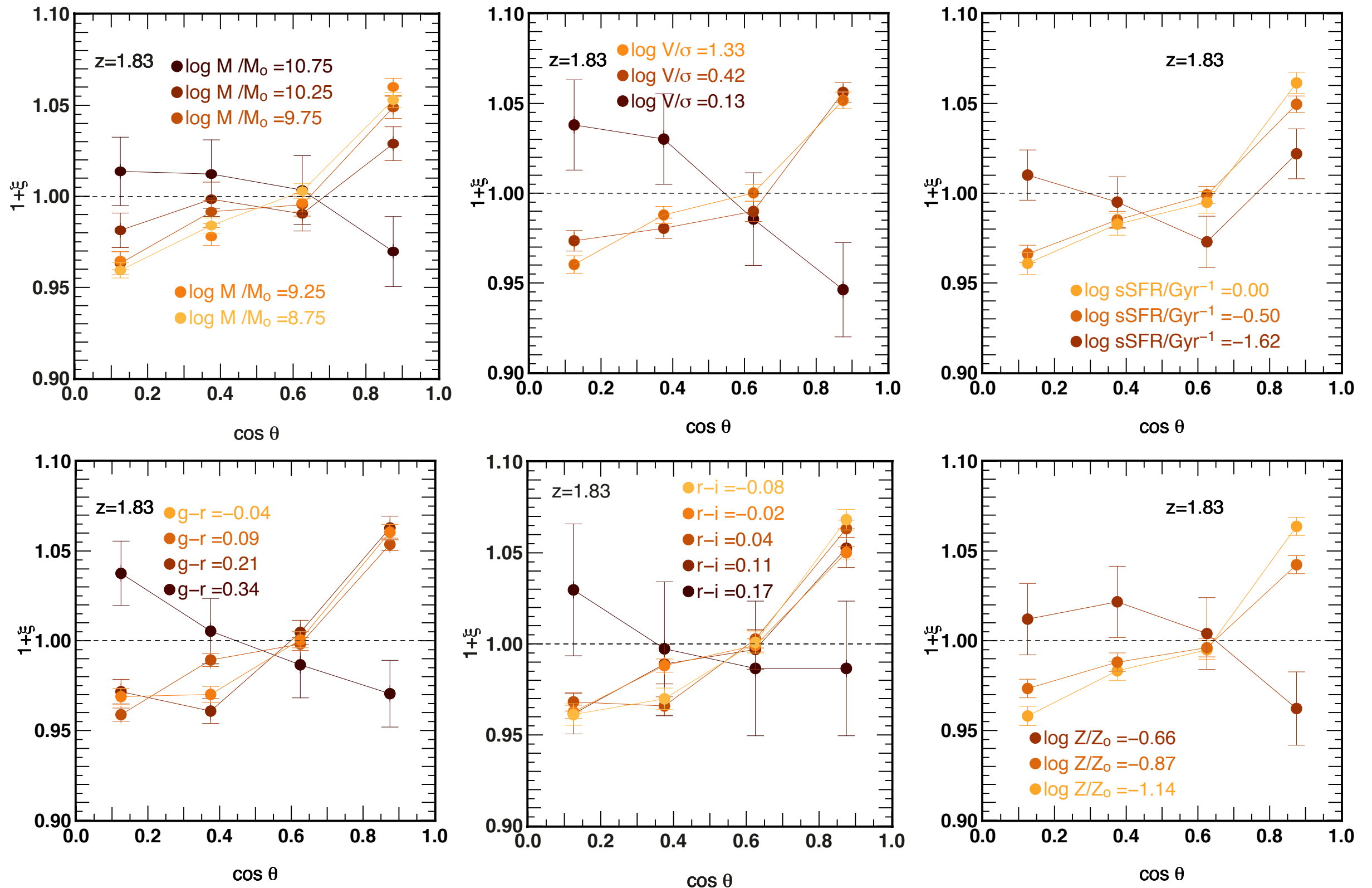


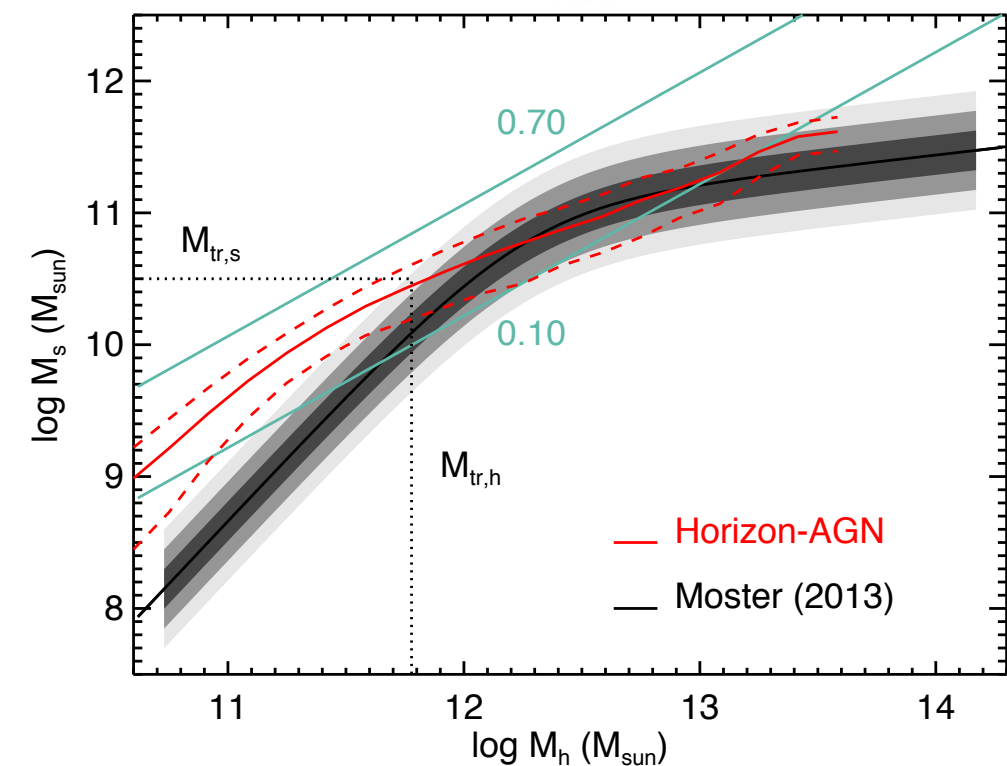
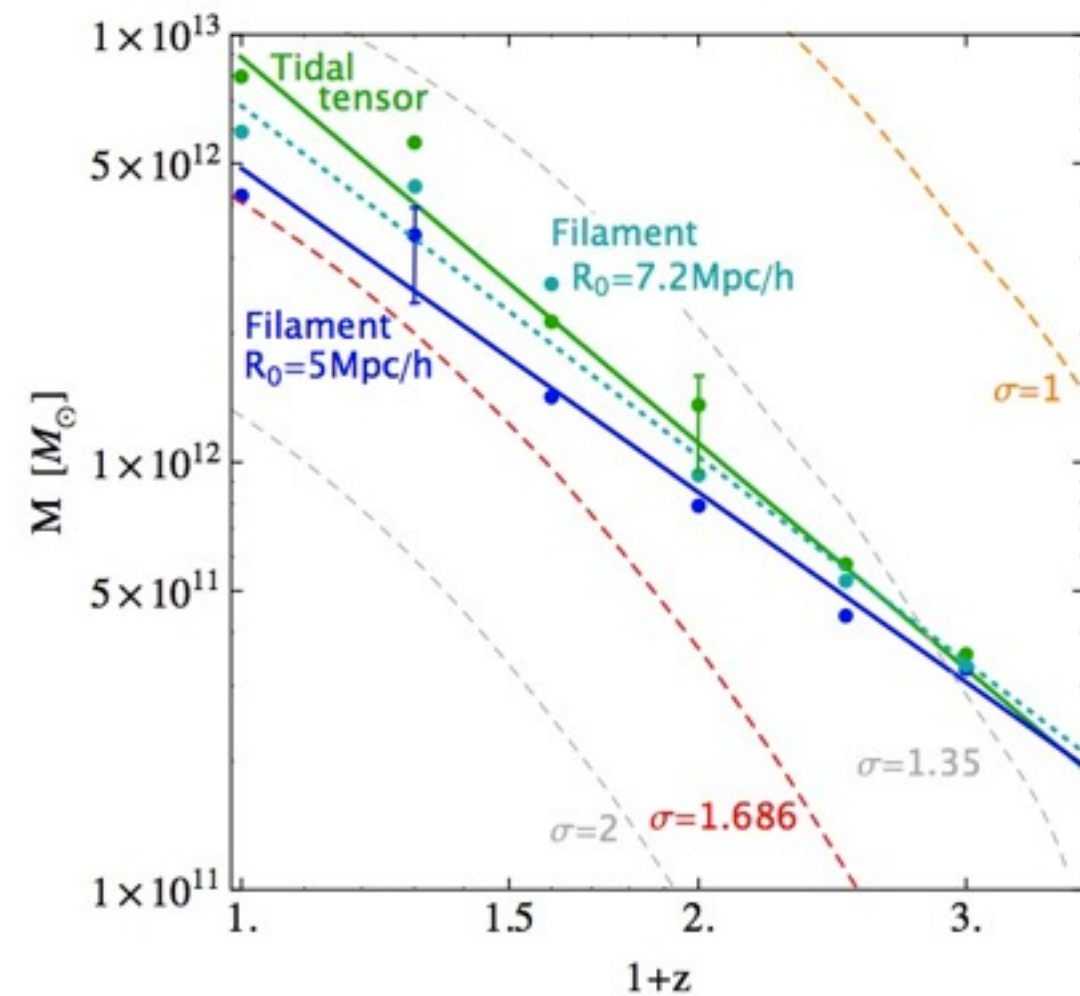
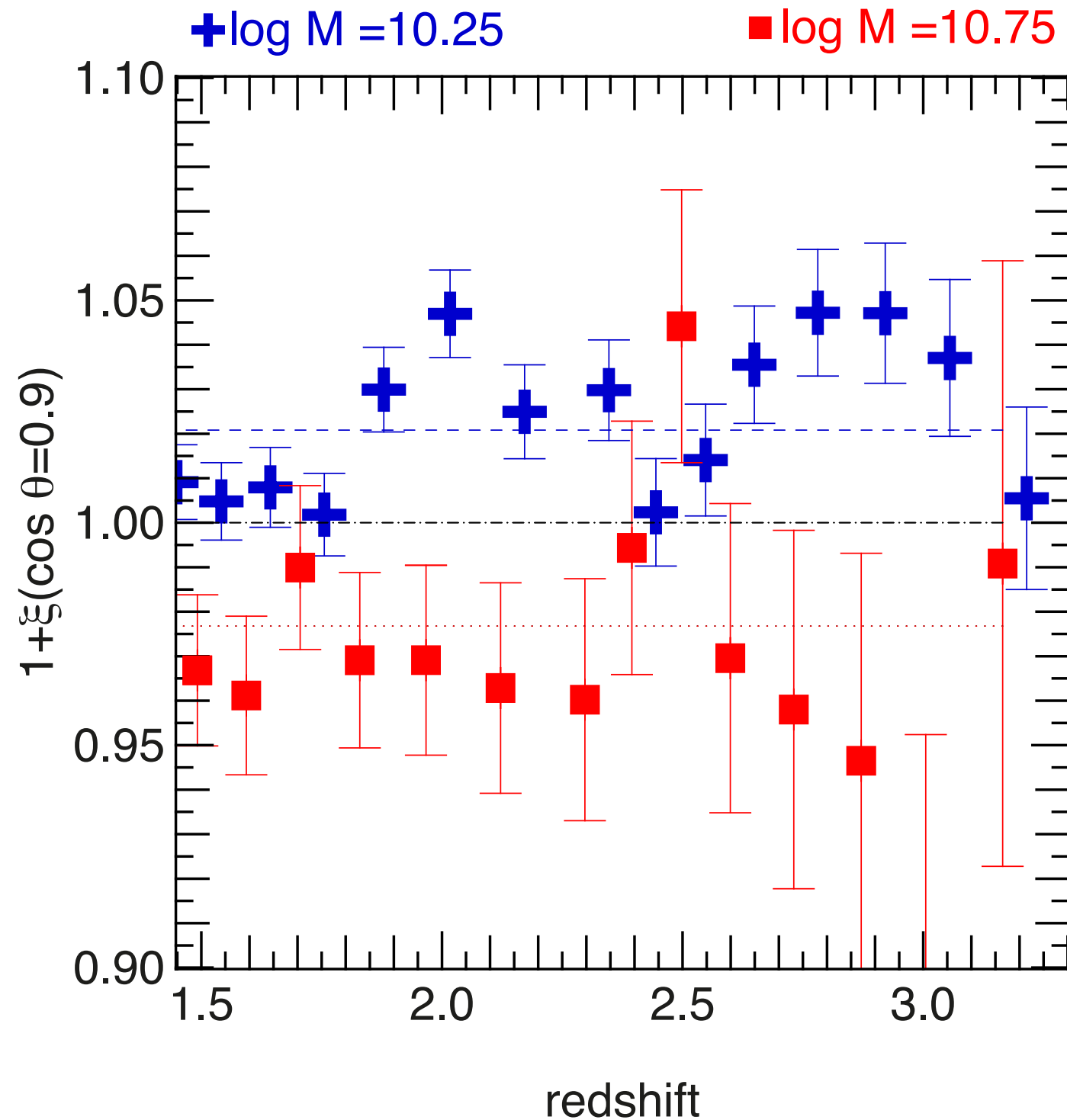
Figure 10. The probability distribution of the cosine of the angle between the spin of dark halos and the direction of the closest filament.

Can morphological/physical properties of galaxies trace spin flip?



is morphometric
transition mass
consistent
with DM ?

Mass transition as a function of redshift

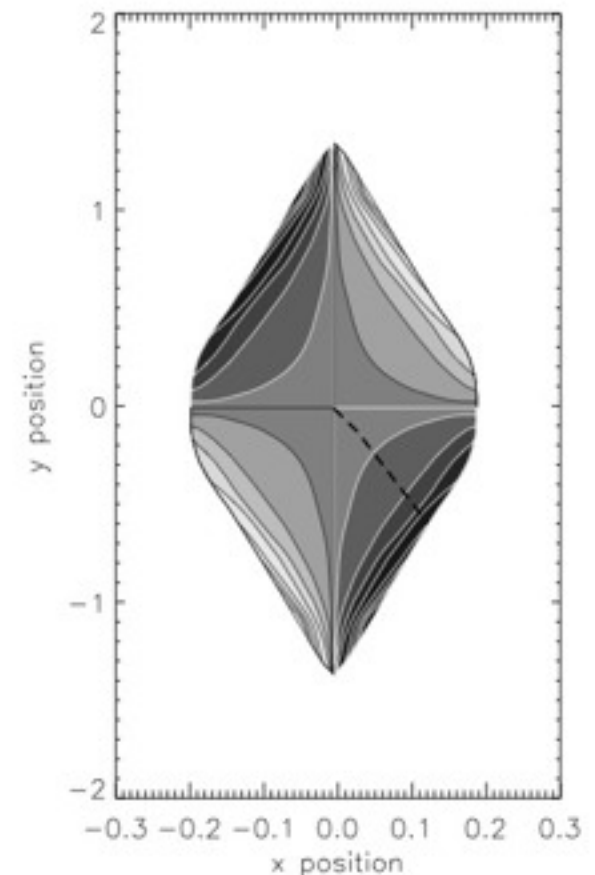
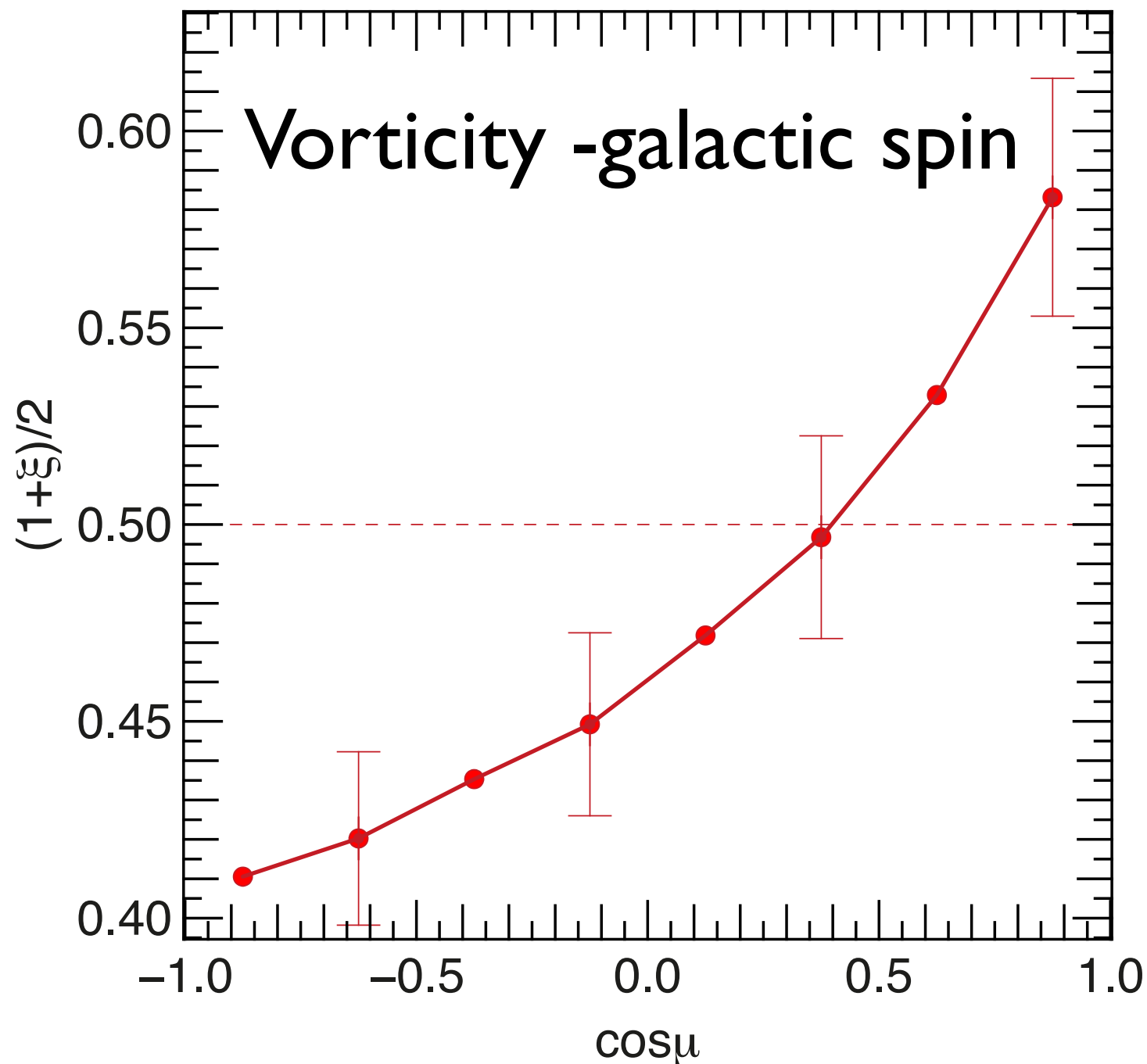


Stellar mass transition $\sim 2 \cdot 10^{10} M_\odot$ is consistent with earlier findings of a dark matter mass transition for the orientation of the spin of dark halos of $6 \cdot 10^{11} M_\odot$ at that redshift.

Final point 1/2:
low mass galaxies

What is the *physical* origin of low mass galaxies spin-filament alignment?

Vorticity arising from kin. structure of filament!



Final point 2/2:
high mass galaxies

Are the imprints of LSS noticeable on galaxy properties ?

Horizon-AGN
>150 000 galaxies

