## The impact of Galaxy Evolution on Star Formation

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# (ONE OF) THE PROBLEM OF COSMOLOGY SIMULATIONS

Sub-grid recipes Galactic disk barely resolved (100 pc) Movie by O. Agertz

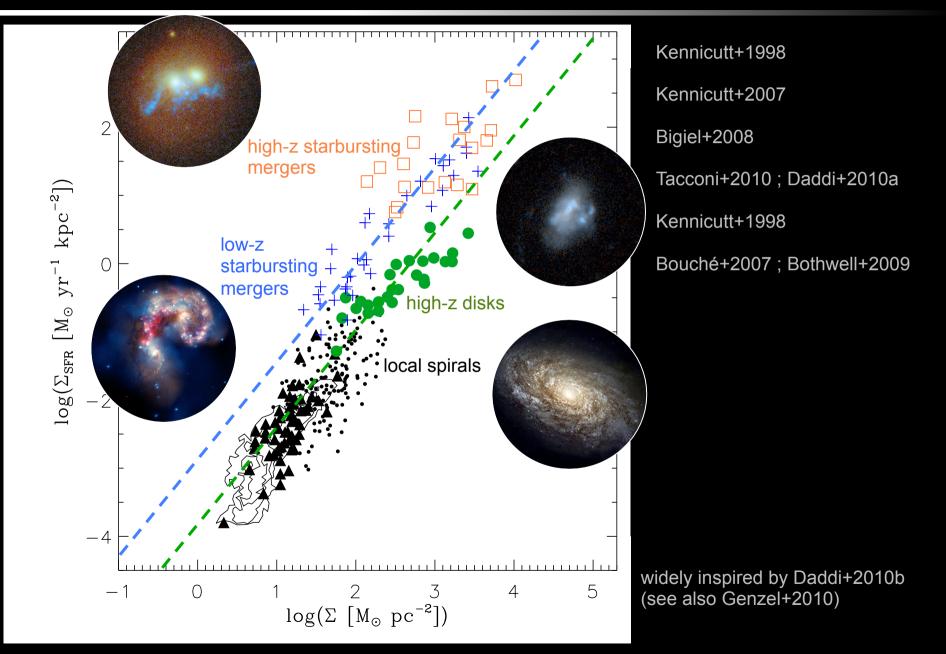
# (ONE OF) THE PROBLEM OF SMALL SCALE SIMULATIONS



Movie by M. Bate

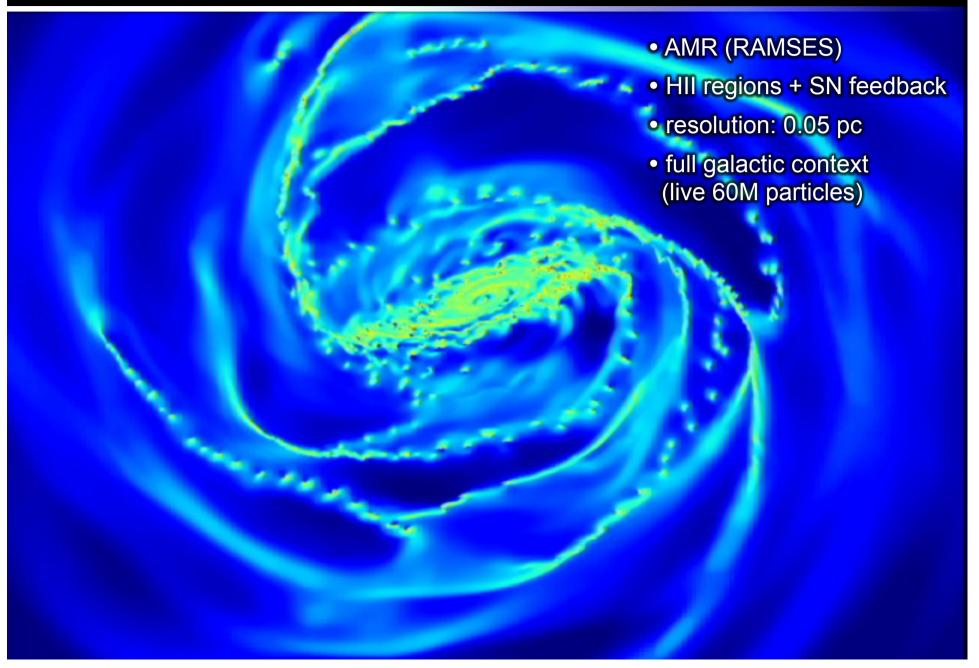
Arbitrary or idealised initial / boundary conditions (geometry, turbulence spectrum ...)

## UNIVERSALITY OF STAR FORMATION?



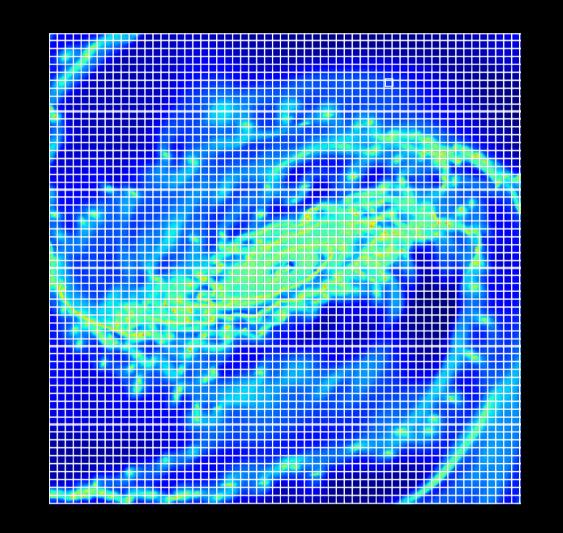
# THE MILKY WAY

#### Renaud et al. (2013)



#### SF AT 100 PC SCALE

Kraljic, Renaud et al. (2014)

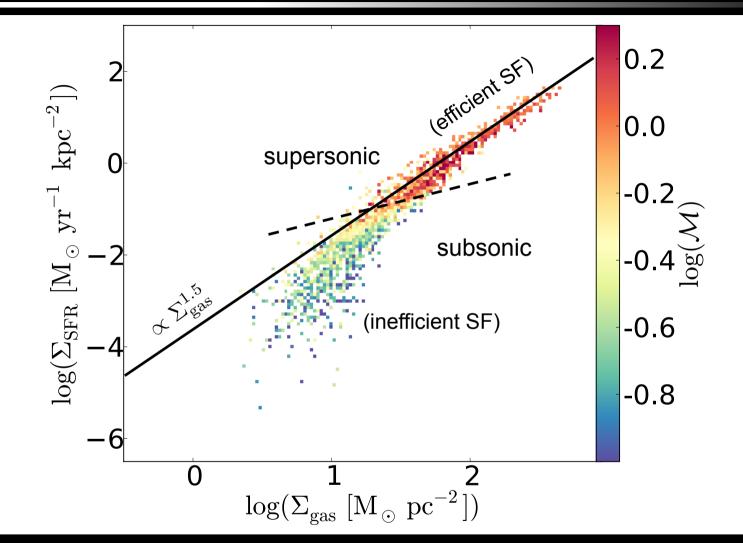


#### = 100 pc \* 100 pc

≥gas SFR

## SCHMIDT-KENNICUTT RELATION

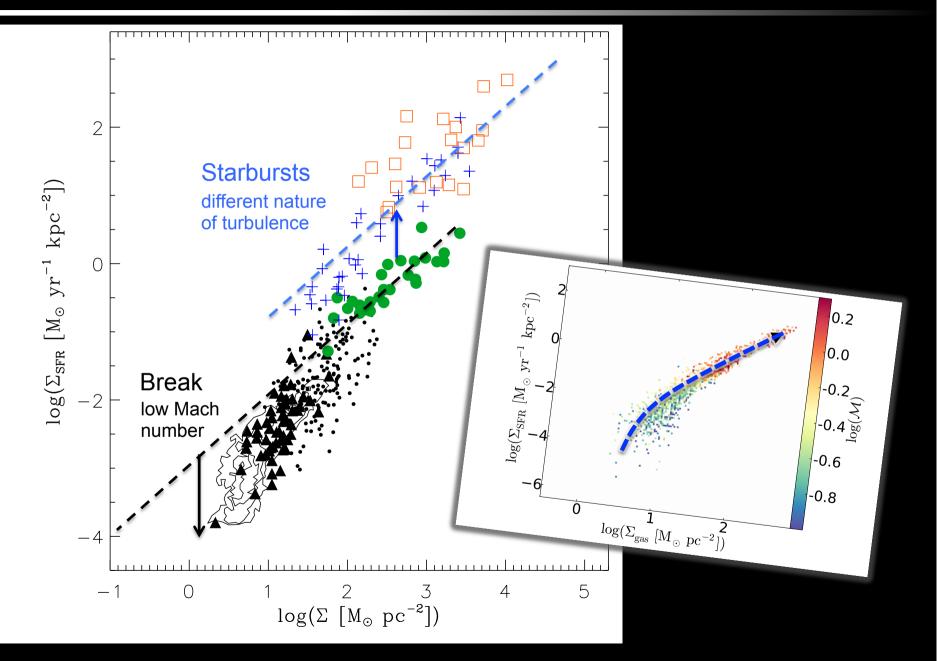
#### Kraljic, Renaud et al. (2014)



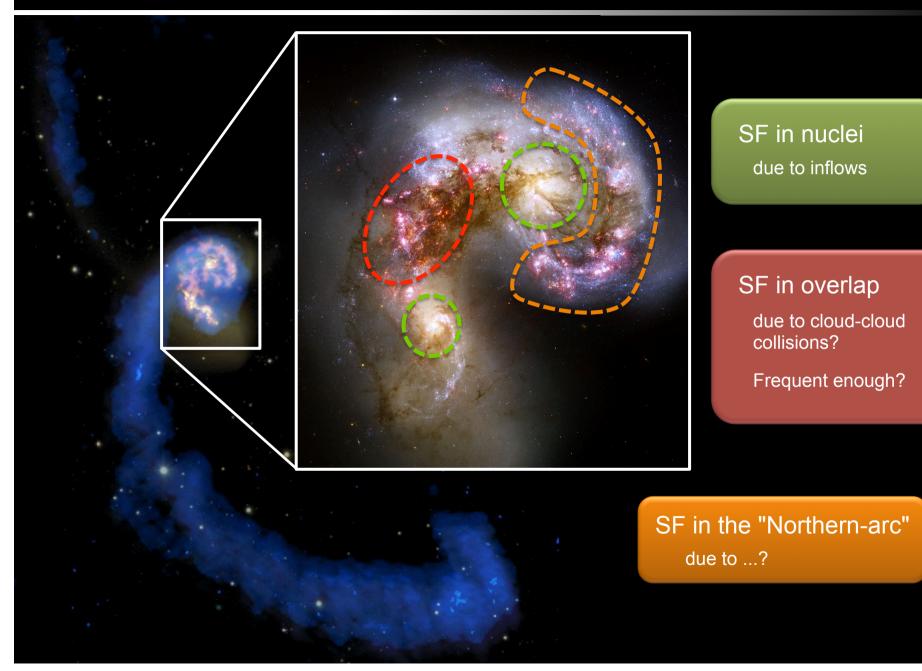
(See also theoretical prediction in Renaud, Kraljic & Bournaud 2012)

Variations of turbulence  $\rightarrow$  spread in SF efficiencies

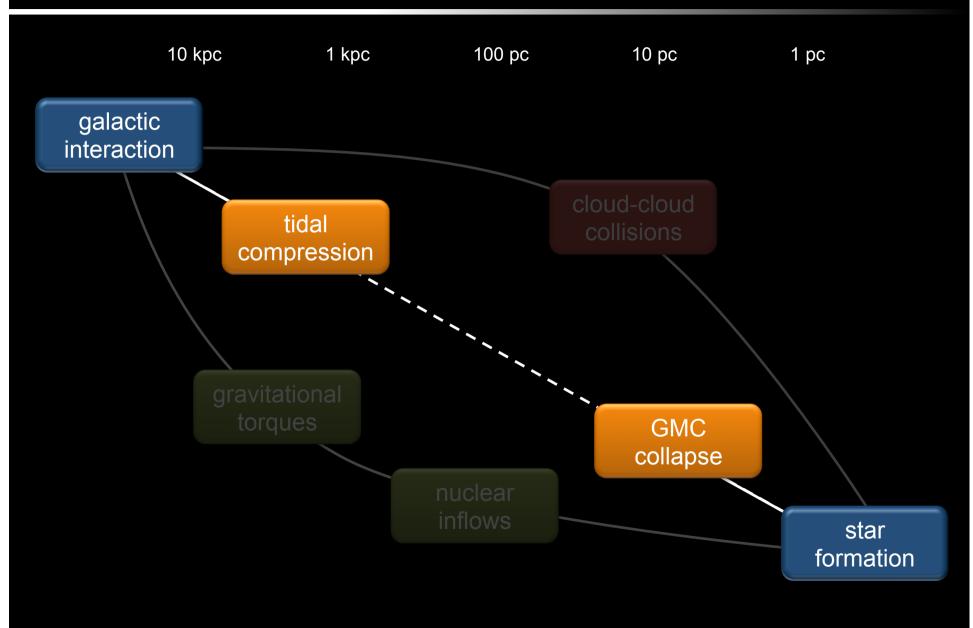
## WHAT ABOUT STARBURSTS?



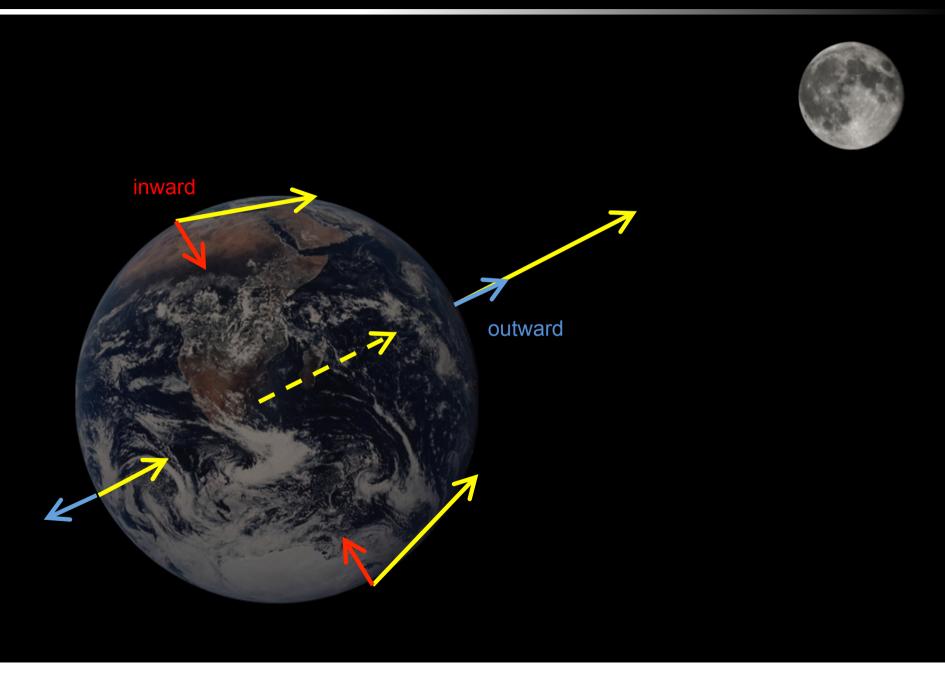
### EXTENDED STARBURSTS



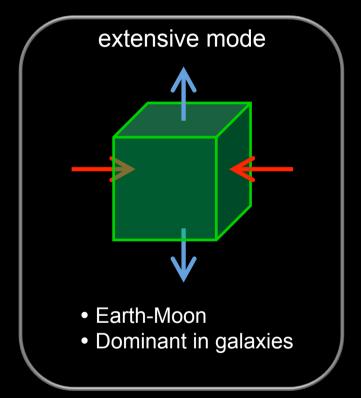
## TRIGGERED, ENHANCED STAR FORMATION

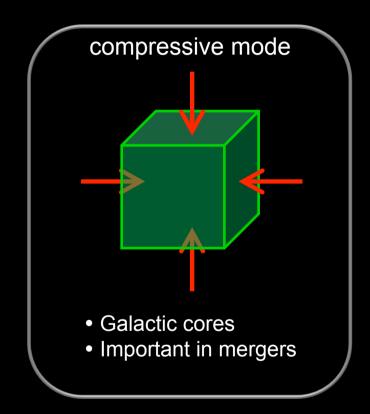


# A QUICK REMINDER ABOUT TIDES



# TIDAL MODES





### COMPRESSIVE TIDES

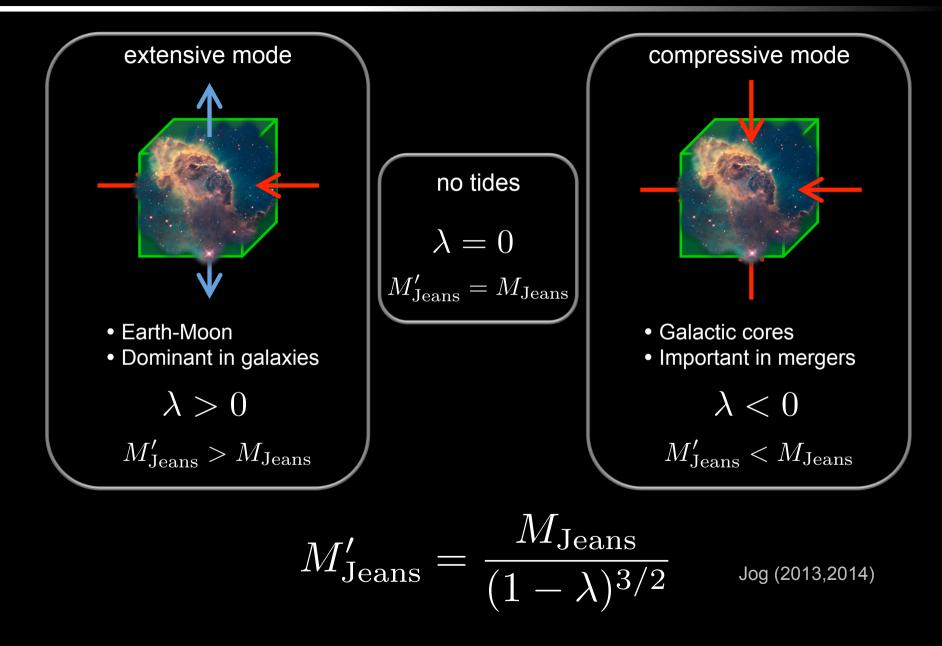
.



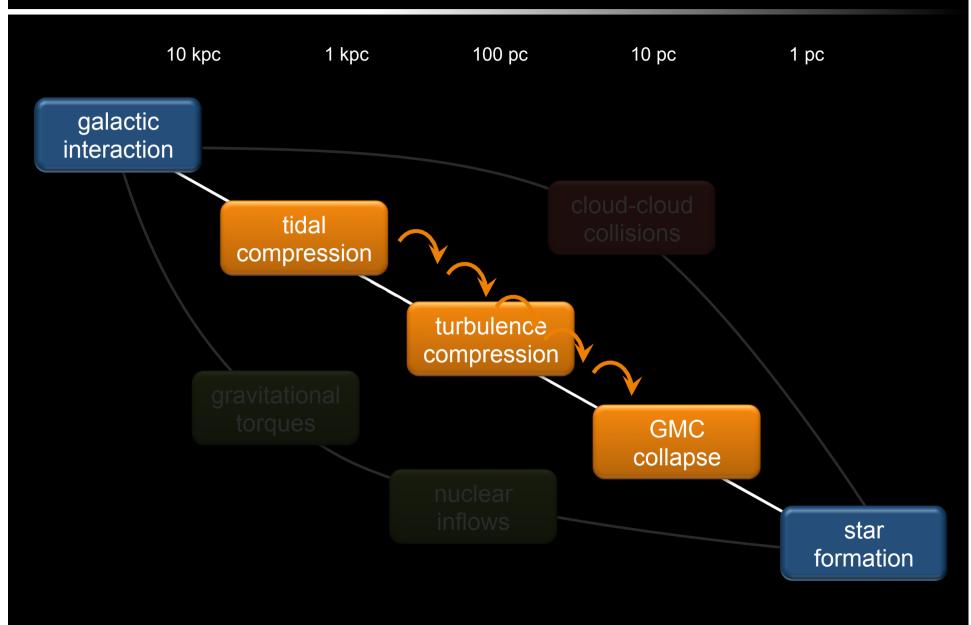
- triggered by the collisions
- Over large volumes
- Valid for all mergers

#### TIDES AND ISM STABILITY

#### Renaud et al. (2008, 2009)

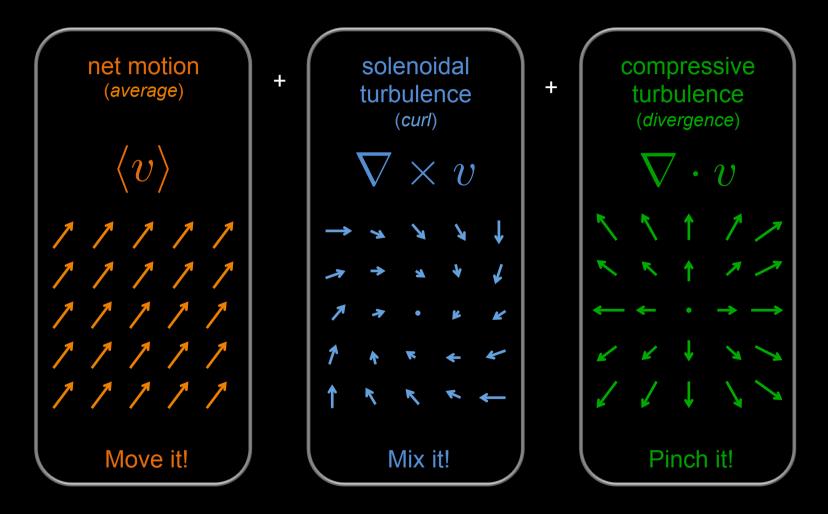


## TRIGGERED, ENHANCED STAR FORMATION



## A QUICK REMINDER ABOUT TURBULENCE

Local velocity field =



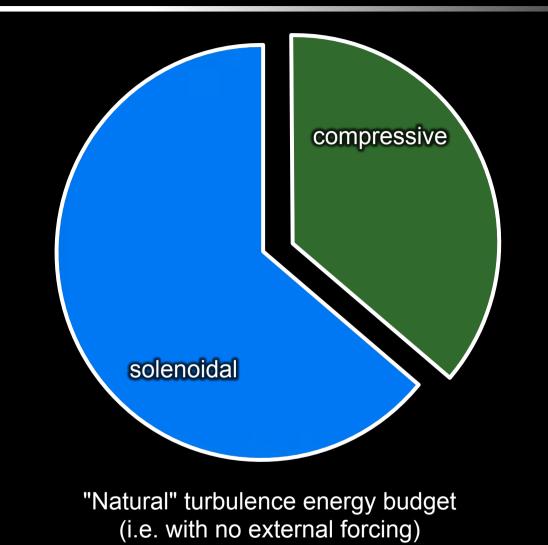
### SOLENOIDAL AND COMPRESSIVE TURBULENCE

Candle smoke Photo taken in the Inter-Sofa Medium (ISM) of my living room

> compressive turbulence

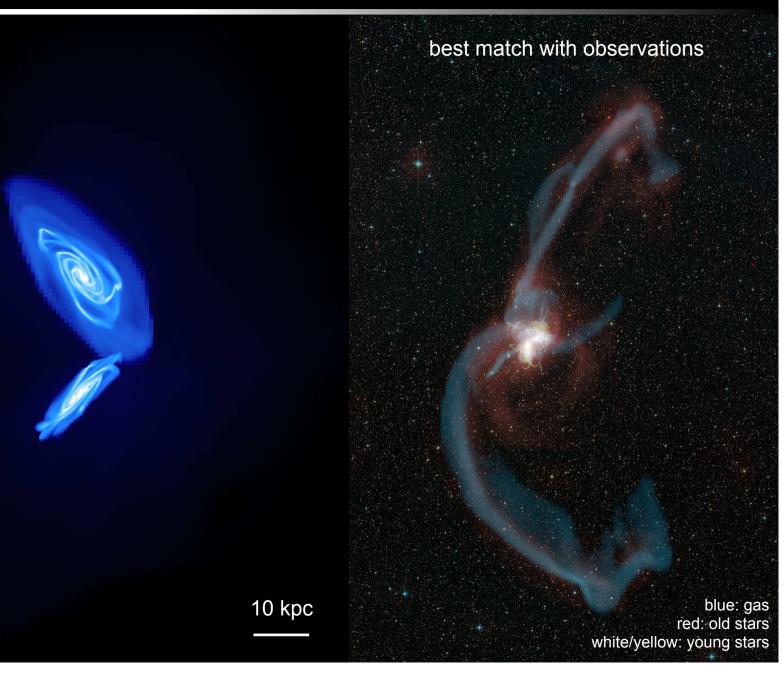
solenoidal turbulence

### SOLENOIDAL AND COMPRESSIVE TURBULENCE



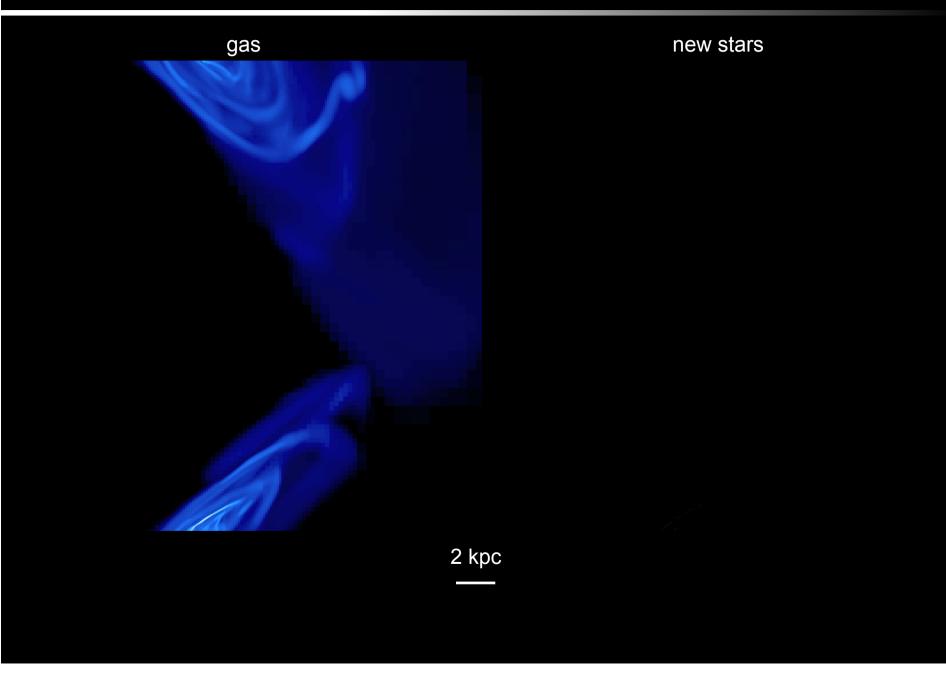
Federrath et al. (2010)

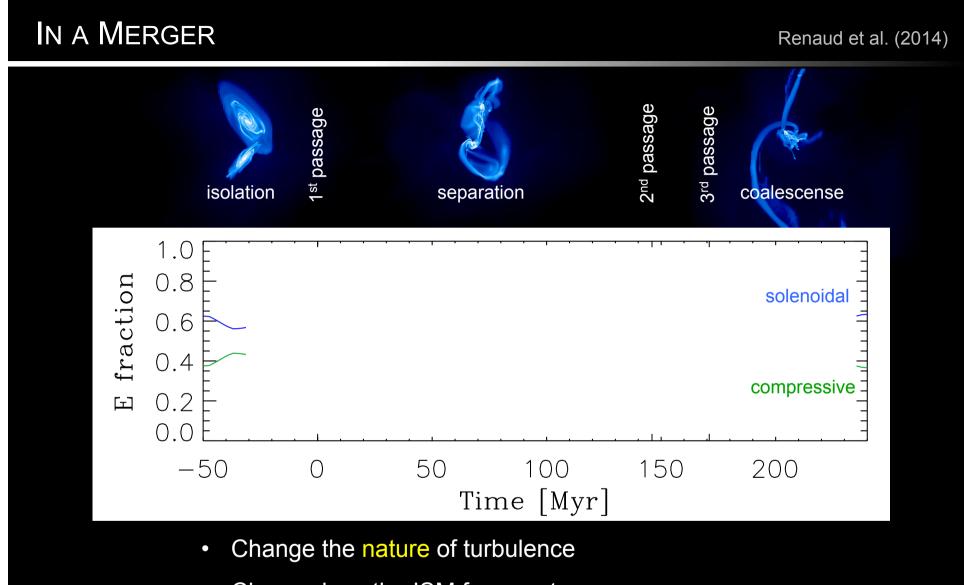
# YET ANOTHER SIMULATION OF THE ANTENNAE Renaud, Bournaud & Duc (2014)



# THE ANTENNAE

Renaud, Bournaud & Duc (2014)



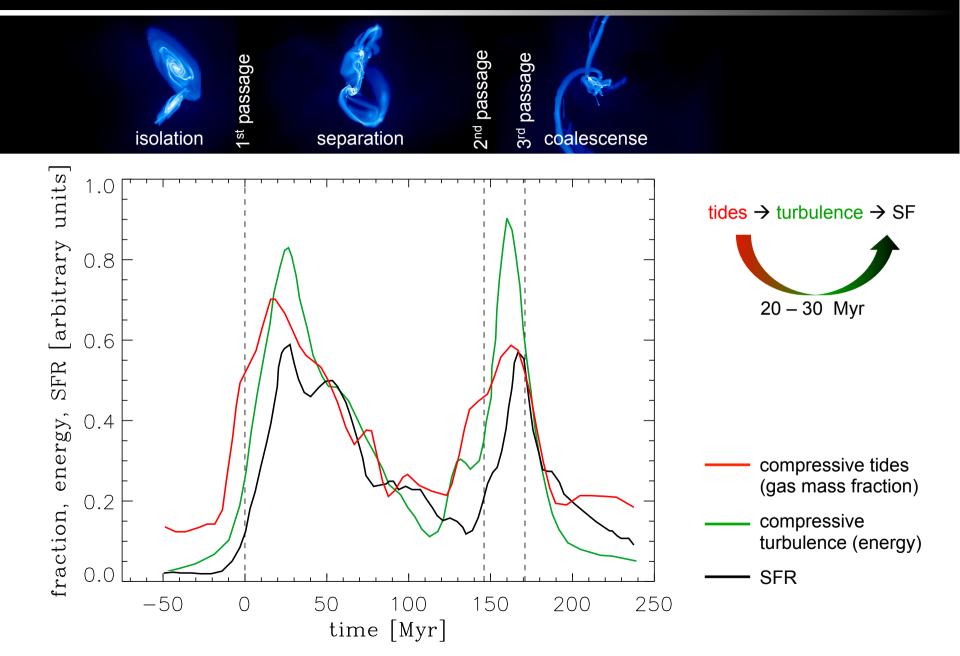


- Change how the ISM fragments
- Change the IMF ( $\rightarrow$  bottom-heavy)

Chabrier, Hennebelle & Charlot (2014)

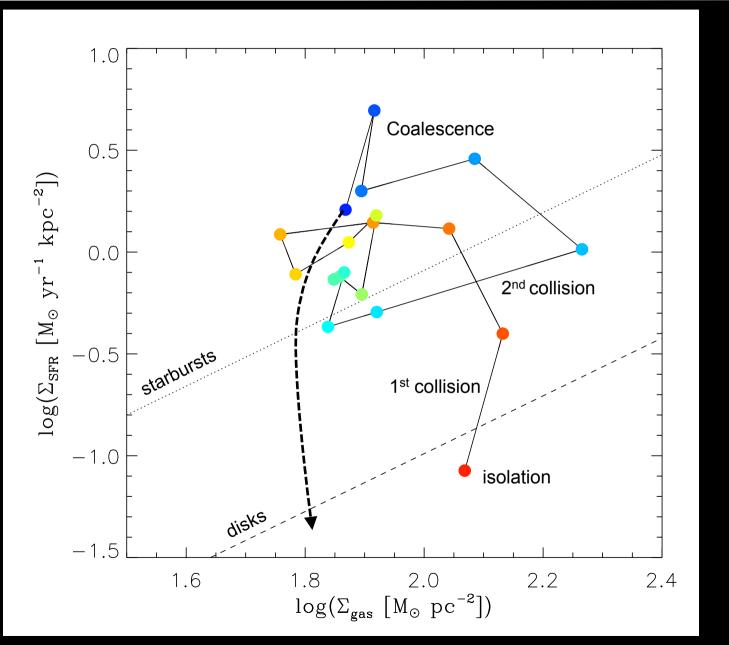
### TIME EVOLUTION

#### Renaud et al. (2014)



## DISK VS MERGER

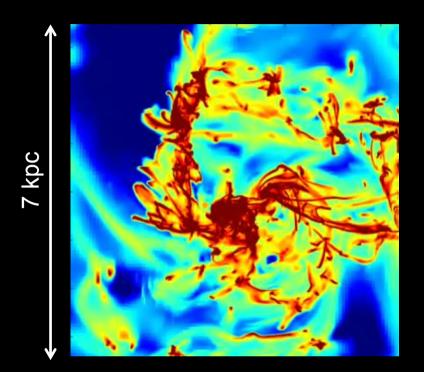
#### Renaud et al. (2014)

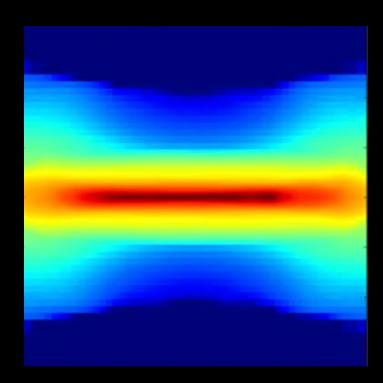




## CLUMPY DISK

#### Bournaud, Perret, Renaud et al. (2014)





Milky Way progenitor at *z*~2 gas fraction ~50%



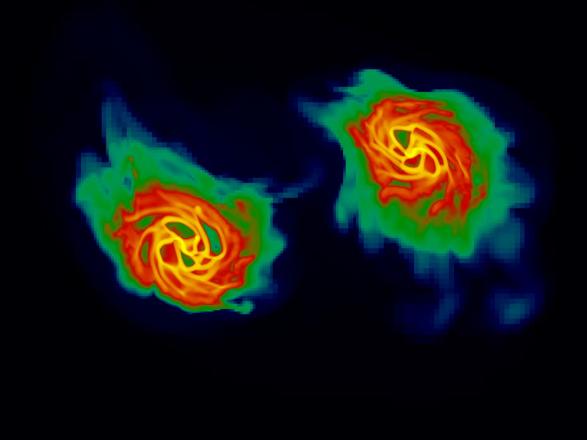
Formation of massive gas clumps (10<sup>9</sup>  $M_{\odot}$ ) High SFR (~50  $M_{\odot}$ /yr)

# HIGH REDSHIFT MERGERS

83 Myrs

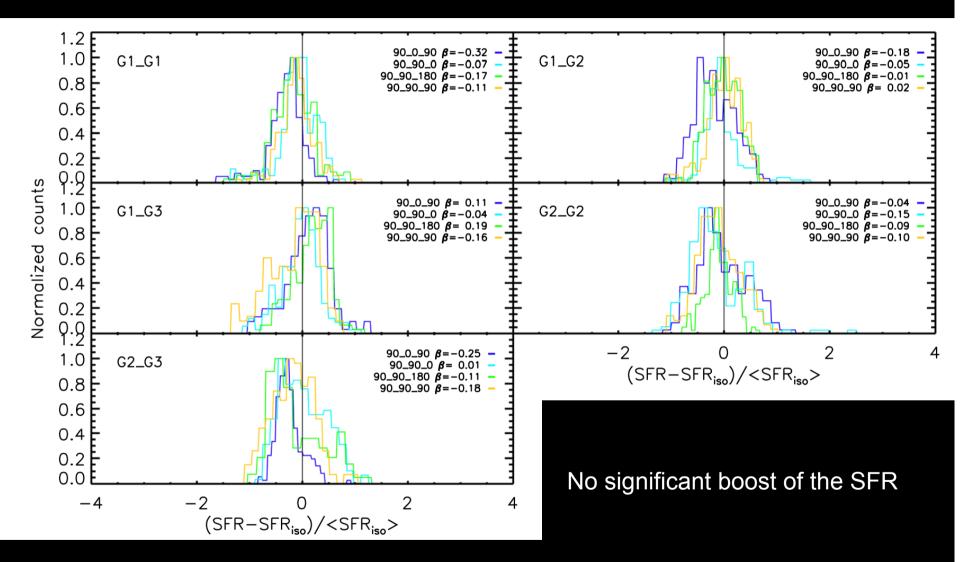
Perret, Renaud et al. (2014)

Merger of 2 gas-rich disks



### HIGH REDSHIFT MERGERS

Perret, Renaud et al. (2014)



Saturation due to the regulation by stellar feedback?

#### CONCLUSIONS

- Galaxy evolution leads to a diversity of environments, clouds, SF
- Coupling between large scales and small scales (turbulence, feedback...)
- SF closely related to supersonic turbulence.

- Galaxy mergers change the nature of turbulence
- (and probably the IMF)
- which explains properties of starbursts (e.g. off-nuclear SF)