



February 15, 2014

First KIAS-NCTS Joint Workshop on Particle
Physics, String Theory and Cosmology @ High I



DARK MATTER IN MODELS WITH DISCRETE GAUGE SYMMETRY

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EVIDENCE OF BSM

- Celestial evidence of DM (familiar ones):

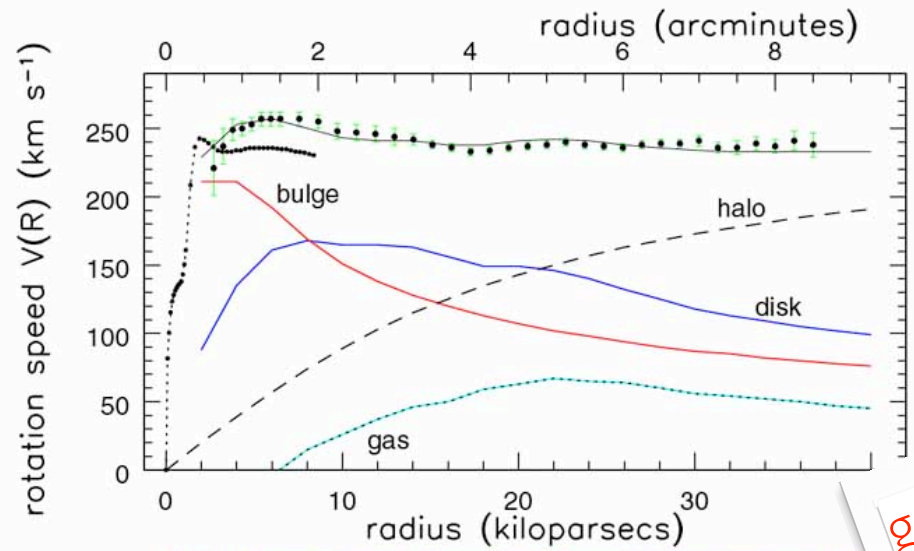
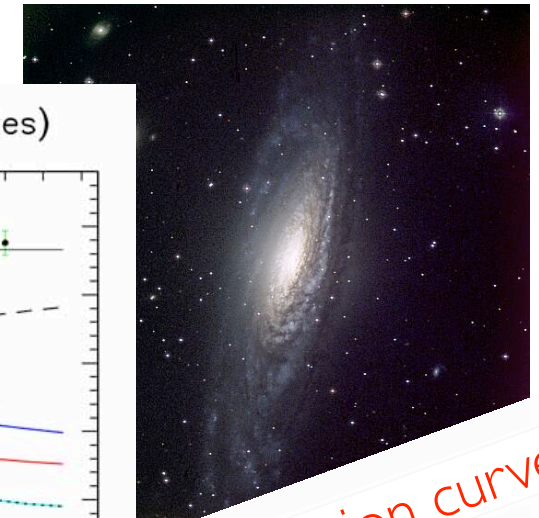
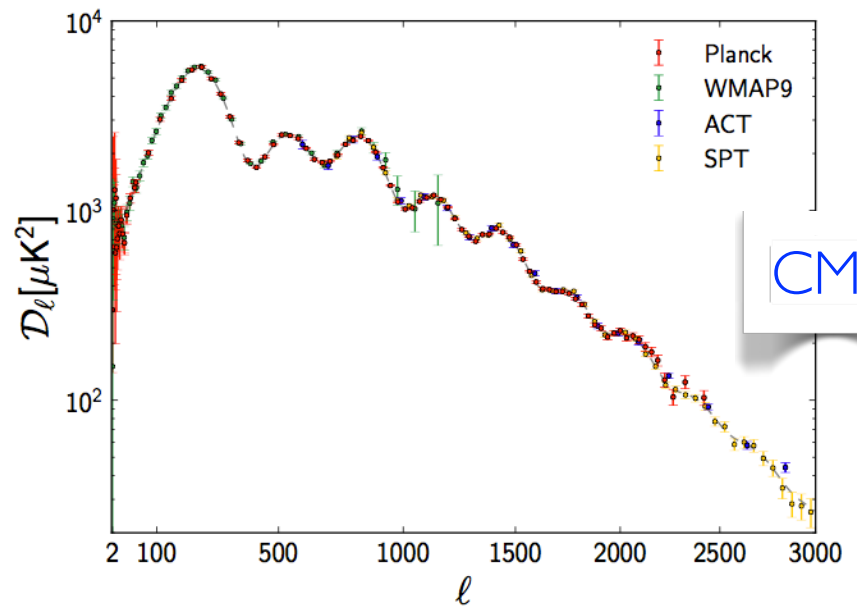
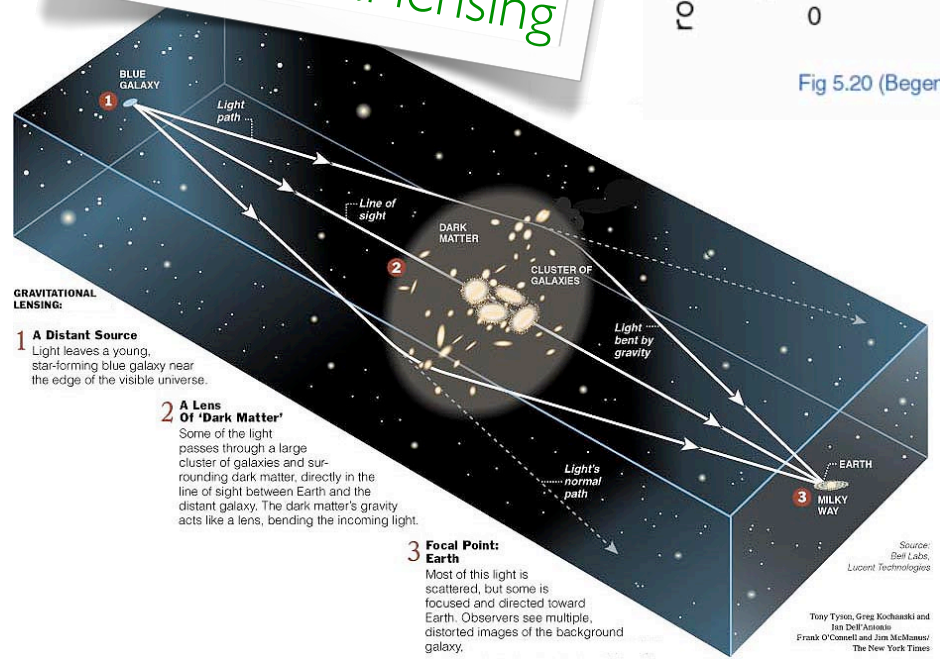


Fig 5.20 (Begeman, Sofue) 'Galaxies in the Universe' Sparke/Gallagher CUP 20

galaxy rotation curve

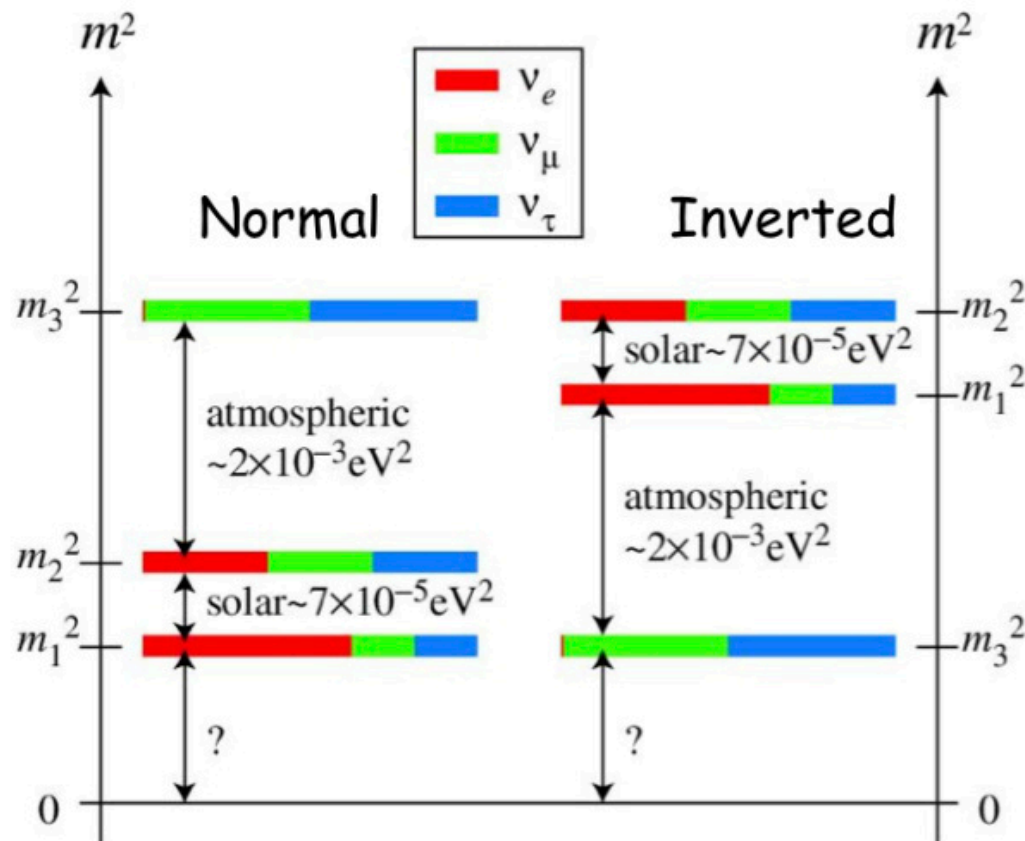
galaxy clusters and gravitational lensing



CMB

EVIDENCE OF BSM

- Celestial / terrestrial evidence: neutrino mass



by Steve King 2011

ONE PROPERTY IN COMMON

- Both DM and neutrinos are difficult to see (catch in detectors)

WHAT IS ESSENTIAL?

What is essential is invisible to the eye.

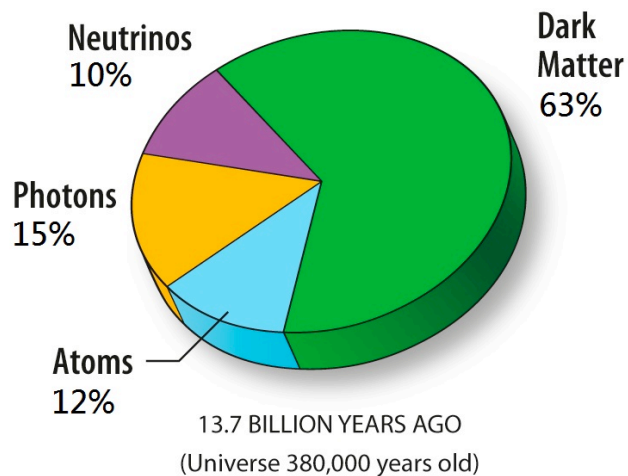
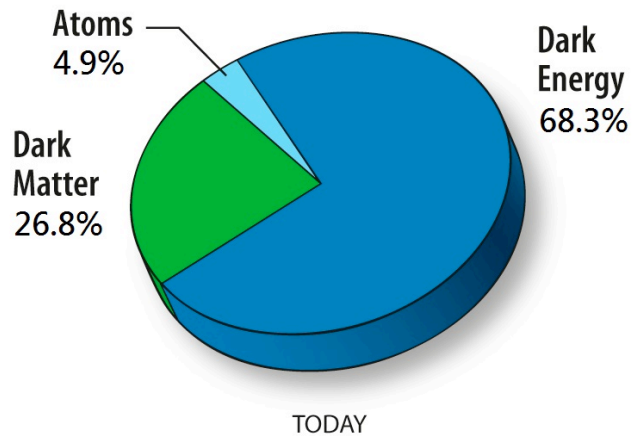
"The Little Prince" - Antoine de Saint-Exupéry



INTRODUCING THE ELEPHANT

- When there's an elephant in the room, introduce them.

--- Randy Pausch, *The Last Lecture*



EXAMINING ELEPHANT BY THE BLIND



EVERYONE HAS HIS OWN STORY

scalar

- Dark matter is fermionic or a combination of them.

vectorial

OUR STORY

- Model with Z_2 symmetry emerged from $U(1)$
 - ▣▣▣▣➔ scalar DM
- Model with Z_2 symmetry emerged from $SU(2)$
 - ▣▣▣▣➔ non-Abelian vector DM
- Summary

MODEL I

HUMBLE CRITERIA OF MODEL

- Stabilized DM candidate
- Simplest gauge group extension
- Minimal new particle contents
- Generating neutrino mass

THE MODEL

- Extend SM gauge group by extra $U(1)_\zeta$
- Add 3 RH neutrinos and 2 complex scalar fields, S and D
- Anomaly cancellation demands $U(1)_\zeta$ nothing but $U(1)_{B-L}$
 $\Rightarrow G_{SM} \times U(1)_{B-L}$
- Quantum number assignment:

	f_{SM}	ν_{kR}	H	S	D
$SU(2), U(1)_Y$	g_{SM}^f	1, 0	2, 1/2	1, 0	1, 0
$U(1)_\zeta [Z_2]$	$\zeta_f [-]$	-1 [-]	0 [+]	2 [+]	1 [-]

Table 1: Charge assignments of the fermions and scalars in the model. f_{SM} (g_{SM}^f) denotes SM fermions (their assignments) and H the usual complex doublet. For quarks and leptons, $\zeta_f = 1/3$ and -1 , respectively.

DISCRETE GAUGE SYMMETRY

Krauss, Wilczek 1989

Nakayama, Takahashi, Yanagida 2011

- New scalar Lagrangian:

$$\mathcal{L} = (\mathcal{D}^\mu D)^\dagger \mathcal{D}_\mu D + (\mathcal{D}^\mu S)^\dagger \mathcal{D}_\mu S - \mathcal{V}$$

$$\mathcal{D}_\mu = \partial_\mu + ig_\zeta \zeta Z'_\mu$$

zero VEV to maintain Z_2
and DM longevity

induces nonzero $\langle S \rangle$
and breaks $U(1)_{B-L}$

$$\begin{aligned} \mathcal{V} = & \mu_D^2 |D|^2 - \mu_S^2 |S|^2 + \mu_{DS} (D^2 S^\dagger + \text{h.c.}) \\ & + 2\lambda_{DS} |D|^2 |S|^2 + 2(\lambda_{DH} |D|^2 + \lambda_{HS} |S|^2) H^\dagger H \\ & + \lambda_D |D|^4 + \lambda_S |S|^4 + (\lambda_H H^\dagger H - \mu_H^2) H^\dagger H \end{aligned}$$

all λ 's > 0 for
vacuum stability

essential to the breaking
 $U(1)_{B-L} \rightarrow Z_2$

S-H mixing, assumed
to be negligible

SCALAR DM

- Scalar VEV's:

$$\langle H \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_H \end{pmatrix}, \quad \langle S \rangle = \frac{v_S}{\sqrt{2}}$$

- Mass eigenstates of $D = (D_R + i D_I)/\sqrt{2}$ (both dubbed **darkons**):

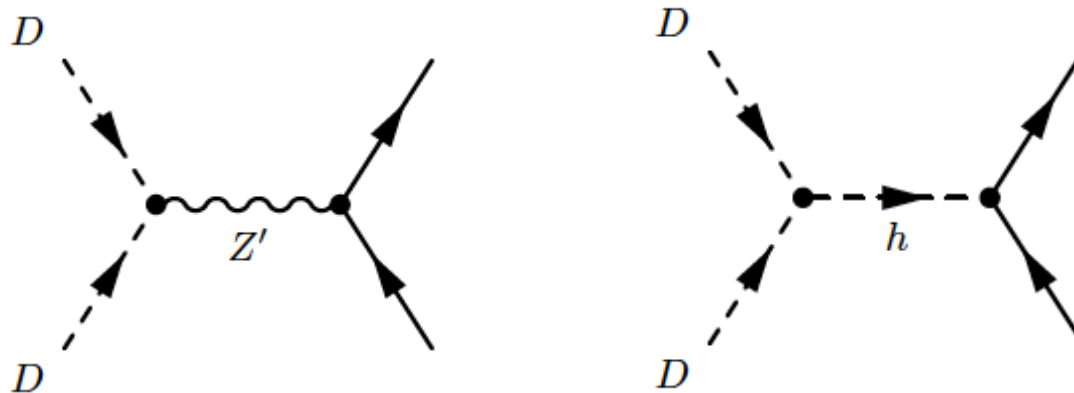
$$m_{D_R, D_I}^2 = \mu_D^2 + \lambda_{DH} v_H^2 + \lambda_{DS} v_S^2 \pm \sqrt{2} \mu_{DS} v_S > 0$$

- Consider **nearly-degenerate case** ($\mu_{DS} > 0$ and ~ 0)

▣▣▣ D_I as **WIMP DM** (\sim simplest darkon)

Silveria, Zee 1985

▣▣▣ **coannihilation with D_R possible**



NEUTRINO MASS

- Neutrino sector contains both **Dirac and Majorana** terms:

$$i\lambda_{kl} \bar{\nu}_{kR} H^T \tau_2 L_{lL} - \frac{1}{2} \lambda'_{kl} \bar{\nu}_{kR} (\nu_{lR})^c S^\dagger + \text{h.c.}$$

$$\mathcal{M}_D = \frac{1}{\sqrt{2}} \lambda v_H, \quad \mathcal{M}_{\nu_R} = \frac{1}{\sqrt{2}} \lambda' v_S \sim \text{multi-TeV}$$

- Majorana neutrino mass is generated through usual **Type-I seesaw**.



A WORD ABOUT HIGGSSES

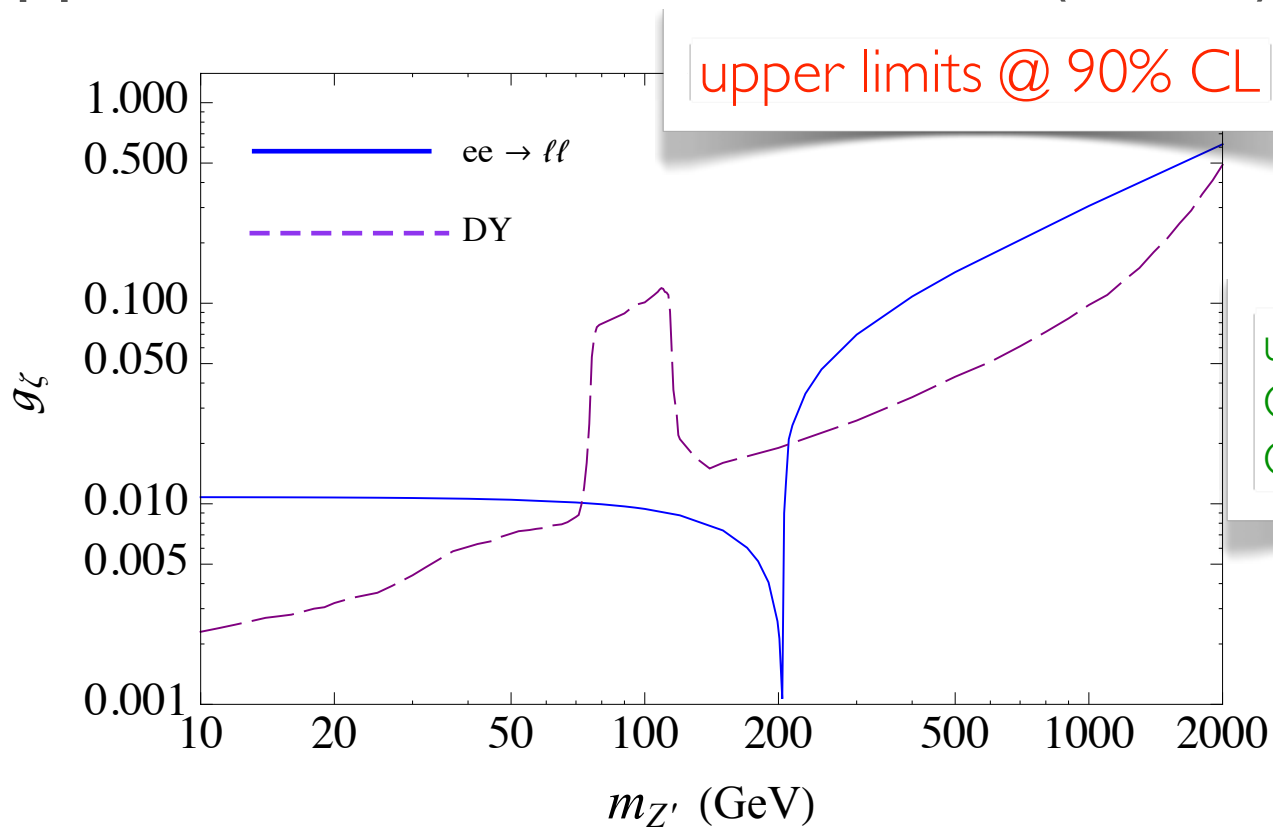
- Physical h and s are almost purely from H and S , respectively, under our assumption of **negligible mixing**
- Mass eigenvalues

$$M_{h,s}^2 \simeq 2\lambda_{H,S} v_{H,S}^2$$

- M_h fixed at 125 GeV
- M_s is multi-TeV in view of RH neutrino's Majorana mass, provided all λ 's are of $O(1)$

CONSTRAINTS ON GAUGE COUPLING

- Z' mass induced purely by S: $m_{Z'} = 2g_\zeta v_S$
 ▶ **no Z - Z' mixing at tree level**
- $e^+e^- \rightarrow Z' \rightarrow \ell^+\ell^-$ @ LEP-II: $\sigma + A_{\text{FB}}$
- $pp \rightarrow Z' \rightarrow \ell^+\ell^-X$ @ LHC 7 TeV (4.5/fb): σ



using method adapted from
CWC, Christensen, Ding, Han 2012
CWC, Lin, Tandean 2012

RELIC DENSITY OF DM

- Assume mass degeneracy between darkons
- Employ approximate Boltzmann equation solution

$$\Omega_D h_0^2 = \frac{1.07 \times 10^9}{\sqrt{g_*} m_{\text{Pl}} J \text{ GeV}} \quad J = \int_{x_f}^{\infty} dx \frac{\langle \sigma_{\text{eff}} v_{\text{rel}} \rangle}{x^2} = J_{Z'} + J_h$$

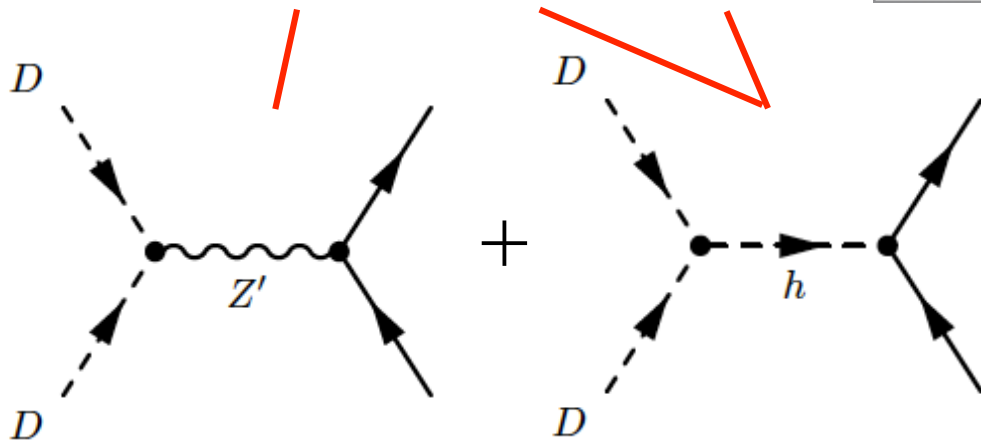
Hubble constant in units of 100/km/s/Mpc

$$x_f = \ln \left[0.038 g_{\text{eff}} m_D m_{\text{Pl}} \langle \sigma_{\text{eff}} v_{\text{rel}} \rangle (g_* x_f)^{-1/2} \right]$$

darkon's effective # of dof's in coannihilation

of relativistic dof's below freeze-out temp T_f

$$\sigma_{\text{eff}} \simeq \frac{1}{4} (2\sigma_{IR} + \sigma_{II} + \sigma_{RR})$$



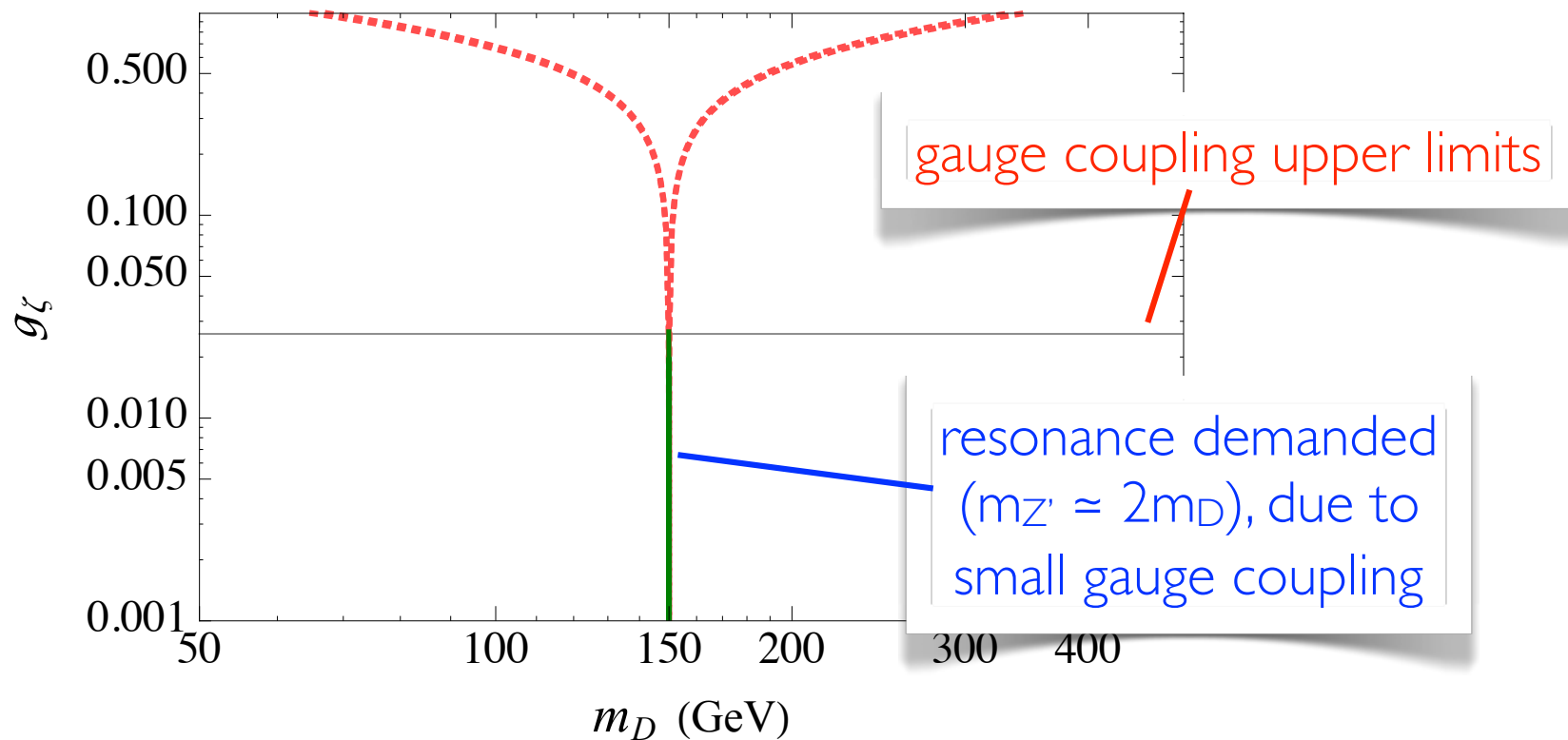
RELIC DENSITY CONSTRAINT

- Employ 90%-CL range:

$$0.092 \leq \Omega_D h_0^2 \leq 0.118$$

WMAP 2011

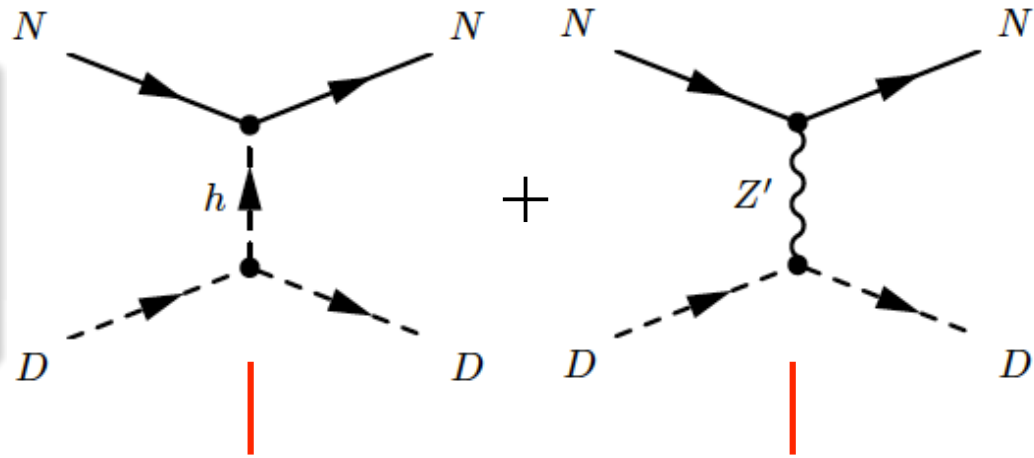
- Reference value $m_{Z'} = 300$ GeV as an example:



DM-NUCLEON SCATTERING

- Scattering cross section for direct detection

resonance relation,
 $m_{Z'} \approx 2m_D$, is assumed
 (rather strong assumption)



$$\sigma_{DN} \simeq \frac{\lambda_{DH}^2 g_{NNh}^2 \mu_{DN}^2 v_H^2}{\pi m_D^2 m_h^4} + \frac{g_\zeta^4 \mu_{DN}^2}{\pi m_{Z'}^4}$$

$$\mu_{DN} = \frac{m_D m_N}{m_D + m_N}$$

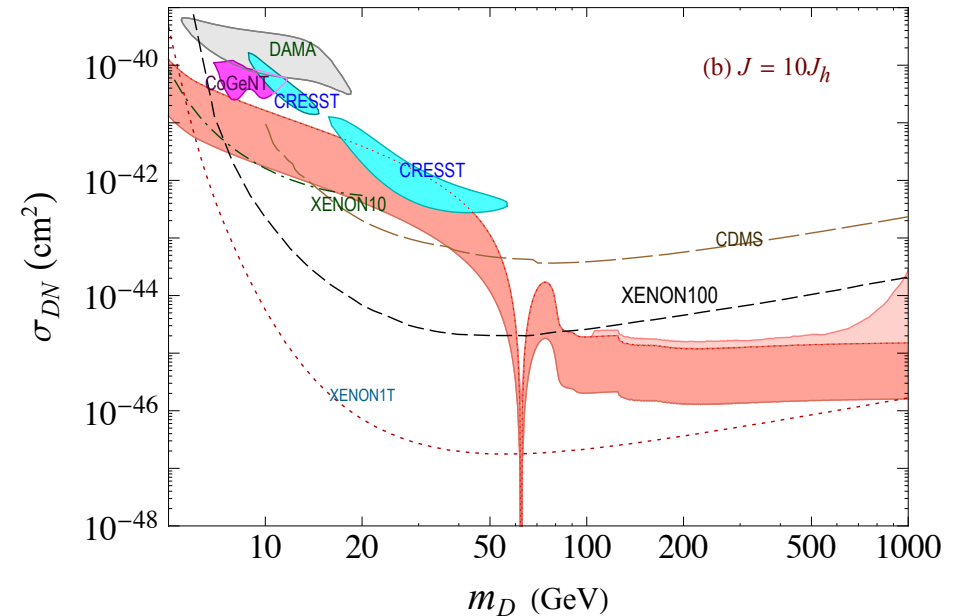
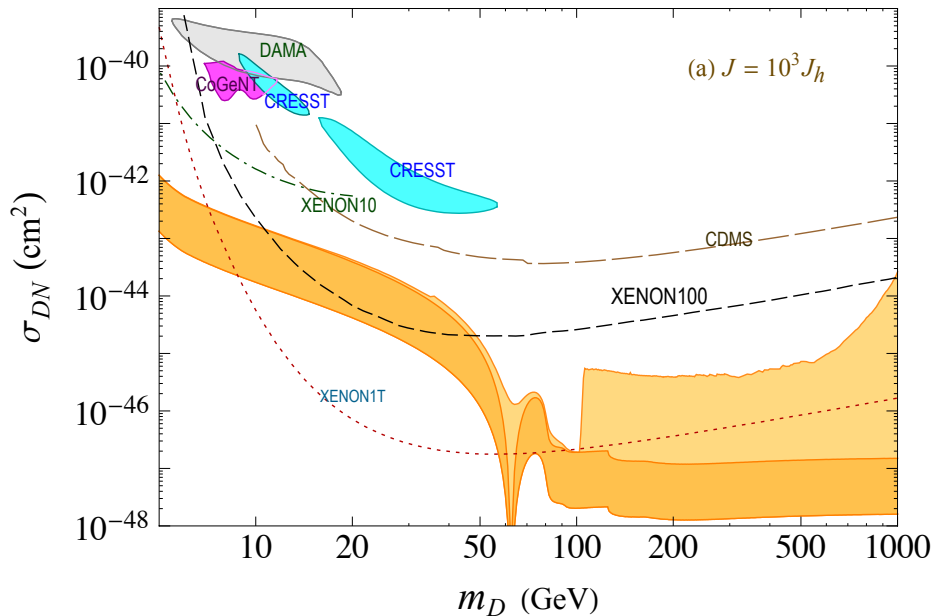
Higgs-nucleon effective coupling
 $0.0011 \leq g_{NNh} \leq 0.0032$

Cheng, CWC 2012

DIRECT SEARCH CONSTRAINT

$$J_{Z'} = 999J_h$$

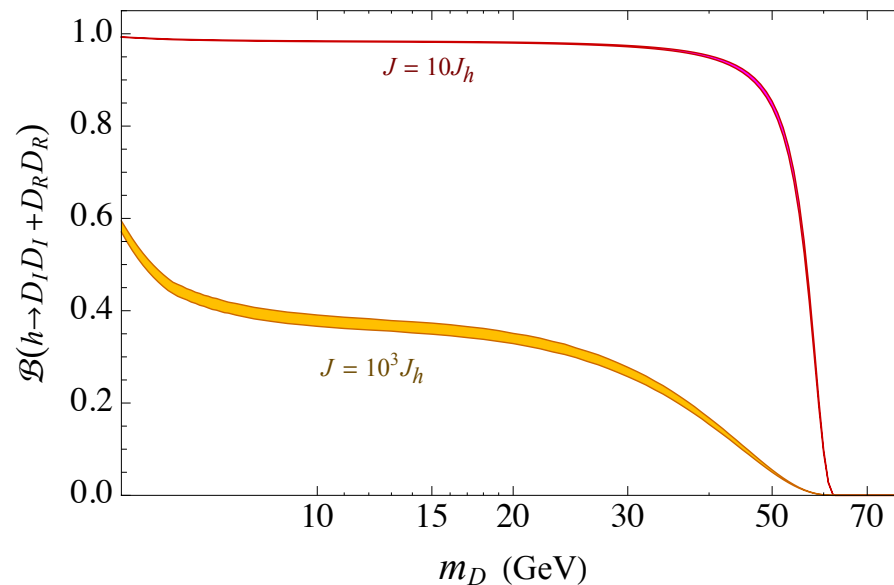
$$J_{Z'} = 9J_h$$



- darker region: purely Higgs; lighter region: Higgs + Z'
- Only some space below 50 GeV ruled out by data
- Higgs dominant in small m_D region
- Z' dominant for $m_D \gtrsim m_h$, also allowed by data
- Wait for XENON1T to probe

INVISIBLE HIGGS DECAYS

- Invisible $h \rightarrow DD$ decays possible if $m_D < m_h/2$
- hDD coupling $\propto \lambda_{DH} v_H$
- $\text{BR}(h \rightarrow \text{inv}) \approx 0.2$ from LHC data Giardino, Kannike, Raidal, Strumia 2012



cf. simplest darkon model is ruled out in this region

- $\text{BR}(h \rightarrow DD)$ large even if h subdominant in DD annihilation
 ⇒ hDD coupling more constrained for $m_D < m_h/2$

MODEL II

RESONANCE RELATION

- Both direct DM search and Higgs data favor Z' -dominated DM annihilation.
- All assumptions in previous model are acceptable except for one, namely, **resonance mass relation between darkons and Z'** .
- Is it possible to obtain resonance mass relation more naturally, while keeping gauge structure and representations as simple as possible?

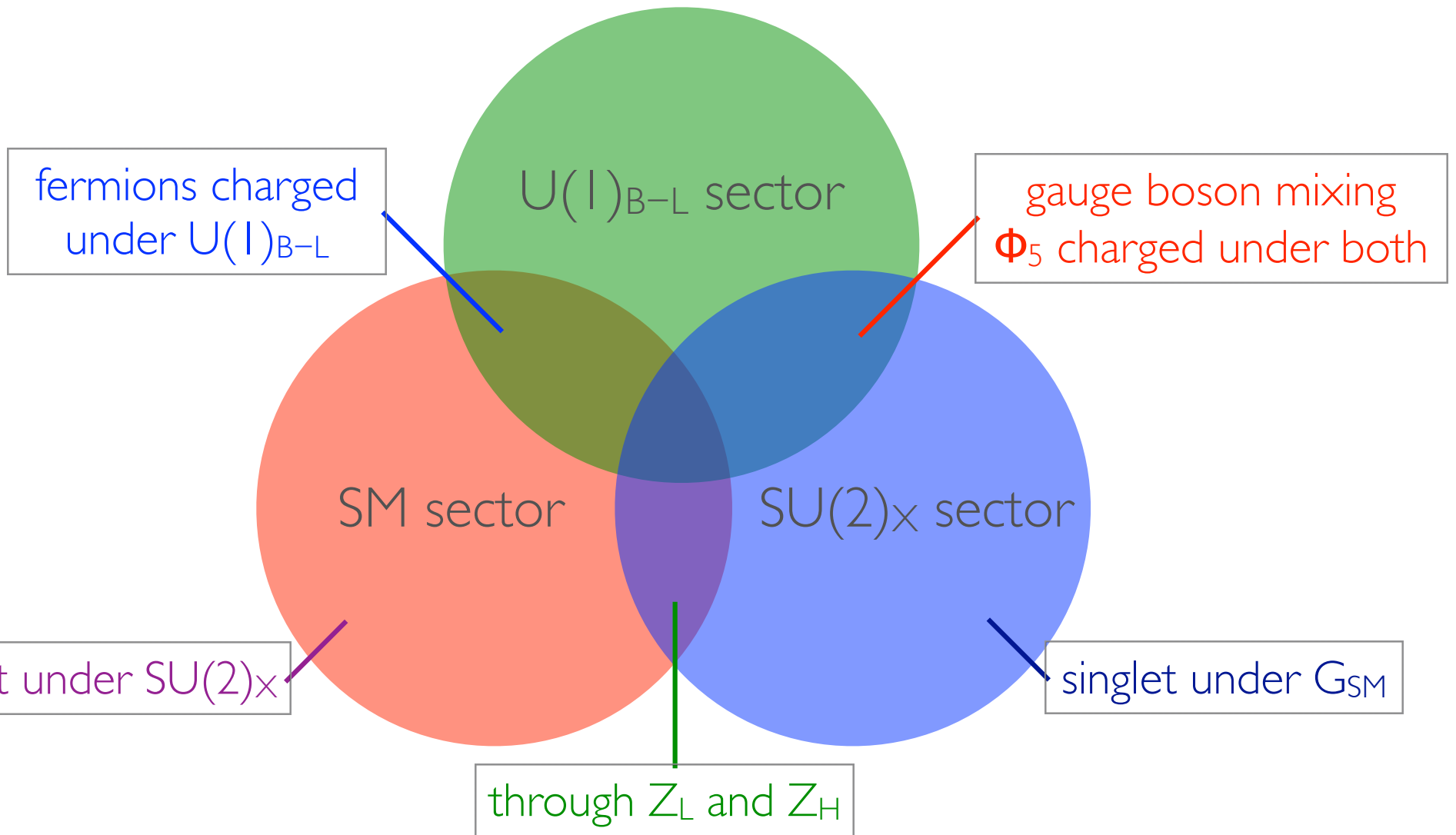
IMPROVED MODEL

- Extend SM gauge group by $SU(2)_X \times U(1)_{B-L}$
- Add 3 RH neutrinos, 1 5-plet scalar fields (Φ_5), 1 singlet scalar (S), 3 $SU(2)$ gauge bosons (X_μ, X^*_μ, C_μ), and 1 $U(1)$ gauge boson (E_μ)
- Anomaly cancellation satisfied
 - ▣ $G_{SM} \times SU(2)_X \times U(1)_{B-L}$
 - ↳ Z_2^X : even/odd T_{3X} component
- Quantum number assignment:

	f_{SM}	ν_R	H	S	ϕ_2	ϕ_1	ϕ_0	ϕ_{-1}	ϕ_{-2}	X	X^\dagger	C_3	E
$SU(2)_X [U(1)_{B-L}]$	1 [$B-L$]	1 [-1]	1 [0]	1 [2]	5 [2]	5 [2]	5 [2]	5 [2]	5 [2]	3 [0]	3 [0]	3 [0]	1 [0]
T_{3X}	0	0	0	0	2	1	0	-1	-2	1	-1	0	0
Z_2^X	+	+	+	+	+	-	+	-	+	-	-	+	+

TABLE I: The charge assignments under $SU(2)_X \times U(1)_{B-L}$ and Z_2^X parity of the fermions, scalars and new gauge bosons in the model, with f_{SM} referring to SM fermions, $X = (C_1 - iC_2)/\sqrt{2}$, and T_{3X} denoting the eigenvalue of the third generator of $SU(2)_X$.

STRUCTURE OF MODEL




SYMMETRY BREAKING


- VEV's

$$\langle H \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_H \end{pmatrix}, \quad \langle S \rangle = \frac{v_S}{\sqrt{2}}, \quad \langle \Phi_5 \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} v_\Phi \\ 0 \\ 0 \\ 0 \\ 0 \end{pmatrix}$$

$v_S \gg v_\Phi > 0$



$U(1)_{B-L} \longrightarrow Z_2^{B-L}$



$SU(2)_X \longrightarrow Z_2^X$

- Since $\langle \Phi_5 \rangle \neq 0$ occurs via its $T_{3X} = 2$ component, Z_2^X symmetry emerges naturally as subgroup of $SU(2)_X$
 - ▣► stabilizing X and X^\dagger as DM candidates
 - ▣► take Z_2^X -odd scalars to be more massive
- Z_2^{B-L} is remnant of $U(1)_{B-L}$ after $\langle S \rangle \neq 0$ as before, but does not play any role in stabilizing X

NEW GAUGE BOSONS

- X and X[†]:

$$X = \frac{1}{\sqrt{2}} (C_1 - iC_2) , \quad X^\dagger = \frac{1}{\sqrt{2}} (C_1 + iC_2)$$

- W boson-like, but electrically neutral
- sub-TeV mass $m_X^2 = g_X^2 v_\Phi^2$

- Z_L and Z_H:

$$\begin{pmatrix} Z_L \\ Z_H \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} C_3 \\ E \end{pmatrix}$$

$$|\theta| \simeq \frac{g_X}{g_{B-L}} R_v \simeq R_v$$

$$R_v = \frac{v_\Phi^2}{v_S^2} \ll 1$$

$$\begin{aligned} v_S &\gg v_\Phi > 0 \\ g_X &\simeq g_{B-L} \end{aligned}$$



$$m_{Z_L}^2 \simeq 4m_X^2 (1 - R_v)$$

$$m_{Z_H}^2 \simeq 4m_X^2 \frac{g_{B-L}^2}{g_X^2 R_v} (1 + R_v)$$

resonance relation
cf. ρ parameter in
SM

NEUTRINO MASS

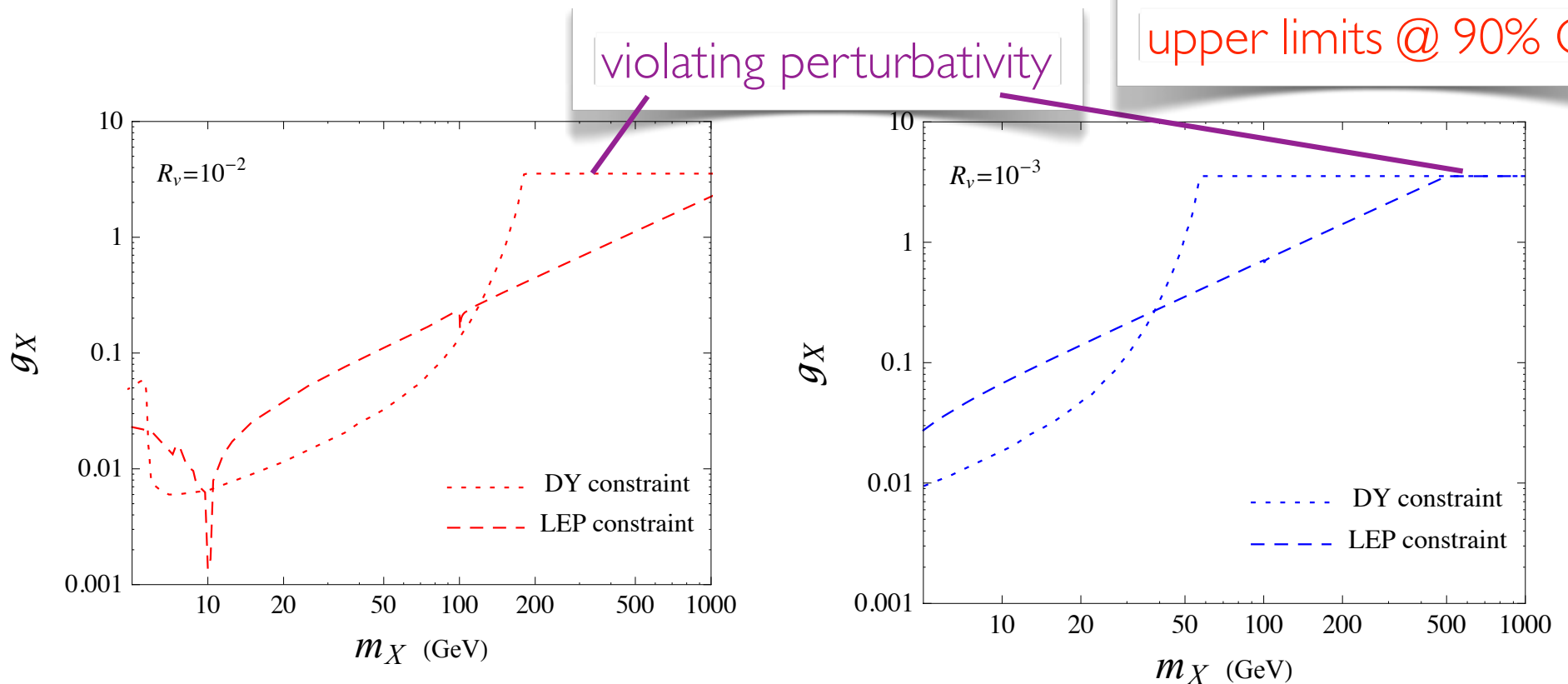
- The story of neutrino mass generation is same as before

$Z_{L,H}$ COUPLING TO FERMIONS

- At tree level:
 - Z_L -f-fbar coupling $\propto g_{B-L} \sin\theta \sim g_{B-L} \theta$
 - Z_H -f-fbar coupling $\propto g_{B-L} \cos\theta \sim g_{B-L}$
- Although Z_H is much heavier than Z_L , it turns out that the coupling between the former and SM fermions makes its contributions to the $e^+e^- \rightarrow Z_{L,H} \rightarrow \ell^+\ell^-$ and $pp \rightarrow Z_{L,H} \rightarrow \ell^+\ell^- X$ more dominant.

CONSTRAINTS ON GAUGE COUPLING

- Take $g_X = g_{B-L}$ for definiteness and simplicity
- $e^+e^- \rightarrow Z_{L,H} \rightarrow \ell^+\ell^-$ @ LEP-II: $\sigma + A_{FB}$
- $pp \rightarrow Z_{L,H} \rightarrow \ell^+\ell^-X$ @ LHC 7 TeV (4.5/fb): σ



RELIC DENSITY OF DM

- Pair annihilation is dominated by Z_L due to **resonance** and **lighter mass** than Z_H
- Employ approximate Boltzmann equation solution

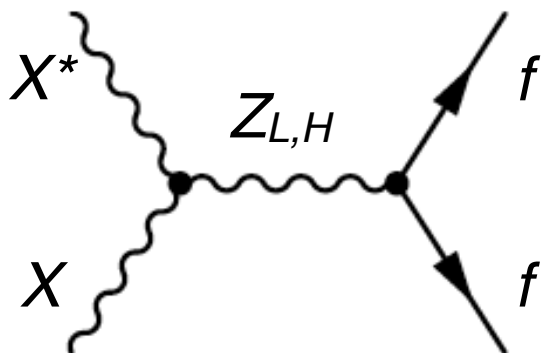
$$\Omega_D h_0^2 = \frac{1.07 \times 10^9}{\sqrt{g_*} m_{\text{Pl}} J \text{ GeV}} \quad J = \int_{x_f}^{\infty} dx \frac{\langle \sigma_{\text{eff}} v_{\text{rel}} \rangle}{x^2}$$

Hubble constant in units of 100/km/s/Mpc

$$x_f = \ln \left[0.038 g_{\text{eff}} m_X m_{\text{Pl}} \langle \sigma_{\text{eff}} v_{\text{rel}} \rangle (g_* x_f)^{-1/2} \right]$$

$g=3$ for spin-1 X

of relativistic dof's below freeze-out temp T_f



$$\sigma \propto g_X^2 g_{B-L}^2 \cos^2 \theta \sin^2 \theta \sim g_X^2 g_{B-L}^2 \theta^2$$

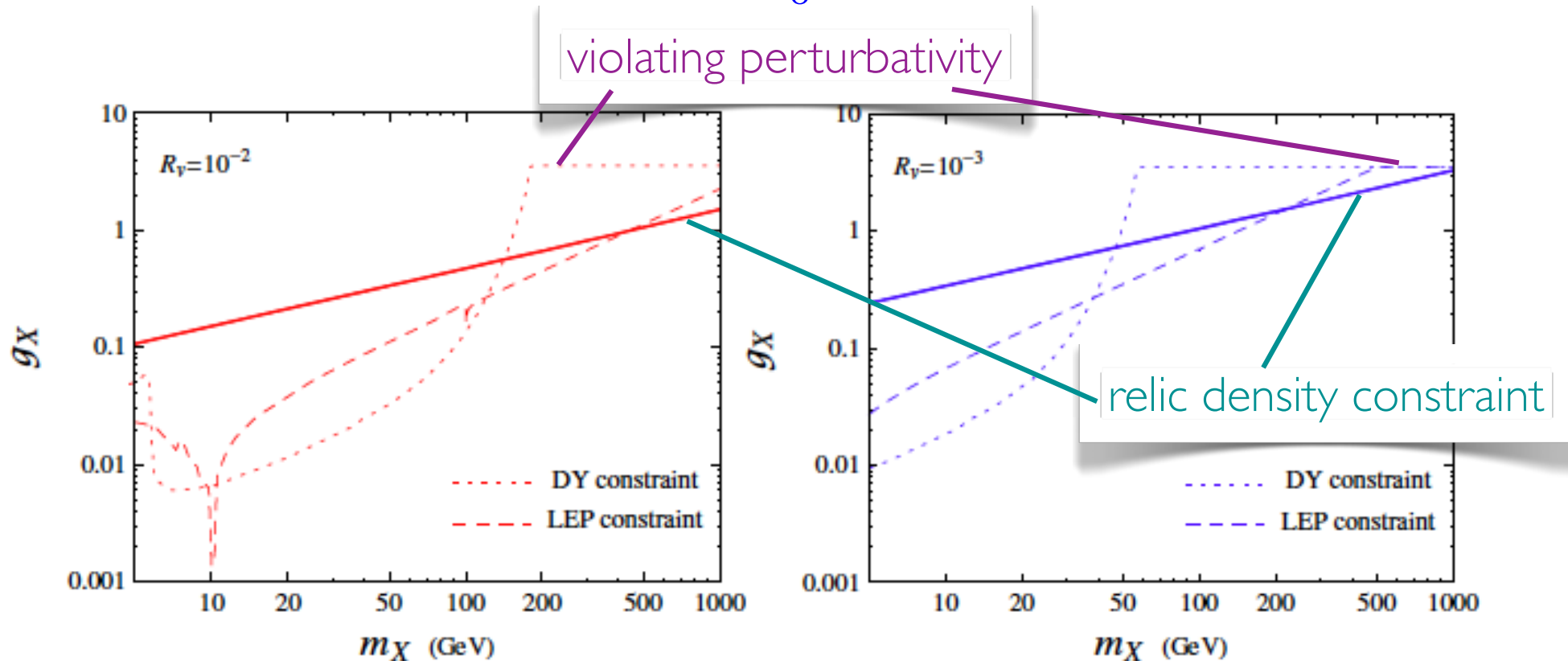
larger gauge couplings demanded

RELIC DENSITY CONSTRAINT

- Employ 90%-CL range:

$$0.1159 \leq \Omega_D h_0^2 \leq 0.1215$$

Planck 2013

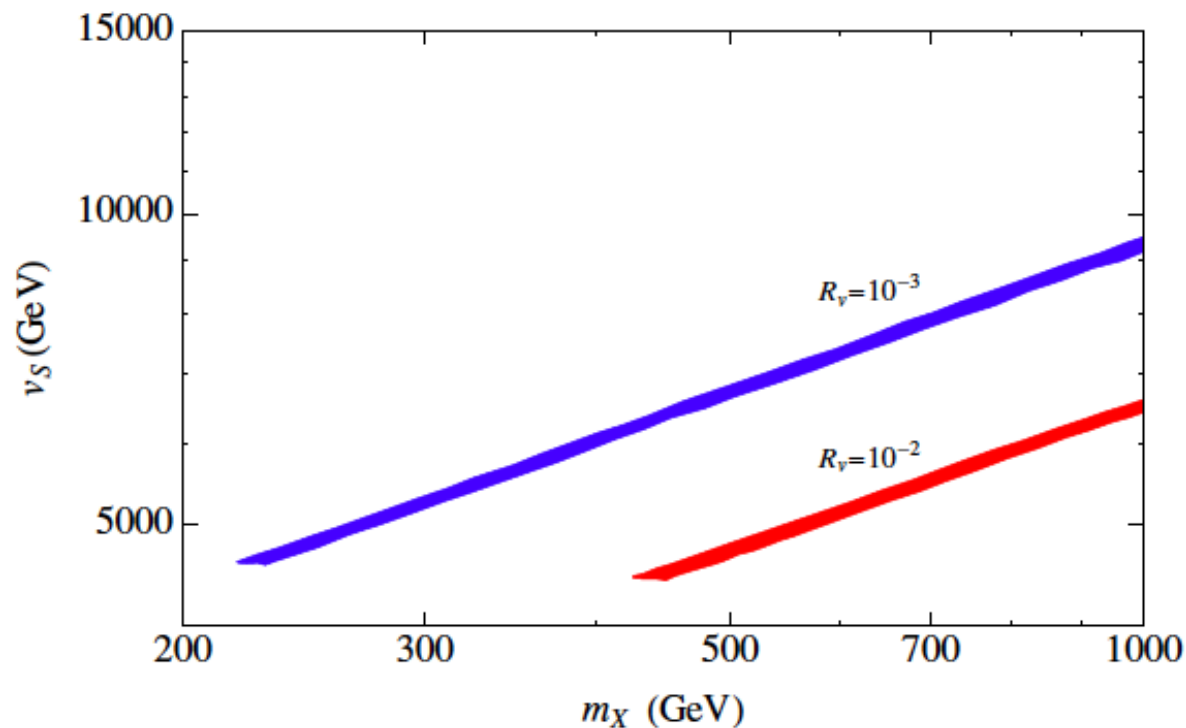


- $m_X > 420$ (220) GeV is allowed for $R_v = 10^{-2}$ (10^{-3})
- $O(1)$ gauge coupling constant is required

A WORD ABOUT v_S

- A linear relation between DM mass and the S VEV:

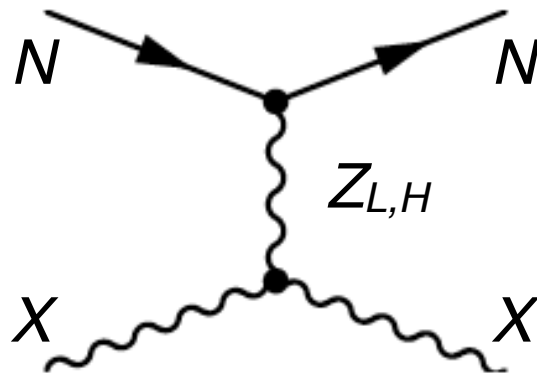
$$m_X = g_X v_\Phi = g_X \sqrt{R_\nu} v_S$$



- v_S should be above $\sim 5-10$ TeV, preferring TeV-scale type-I seesaw

DM-NUCLEON SCATTERING

- Scattering cross section for direct detection in non-relativistic limit

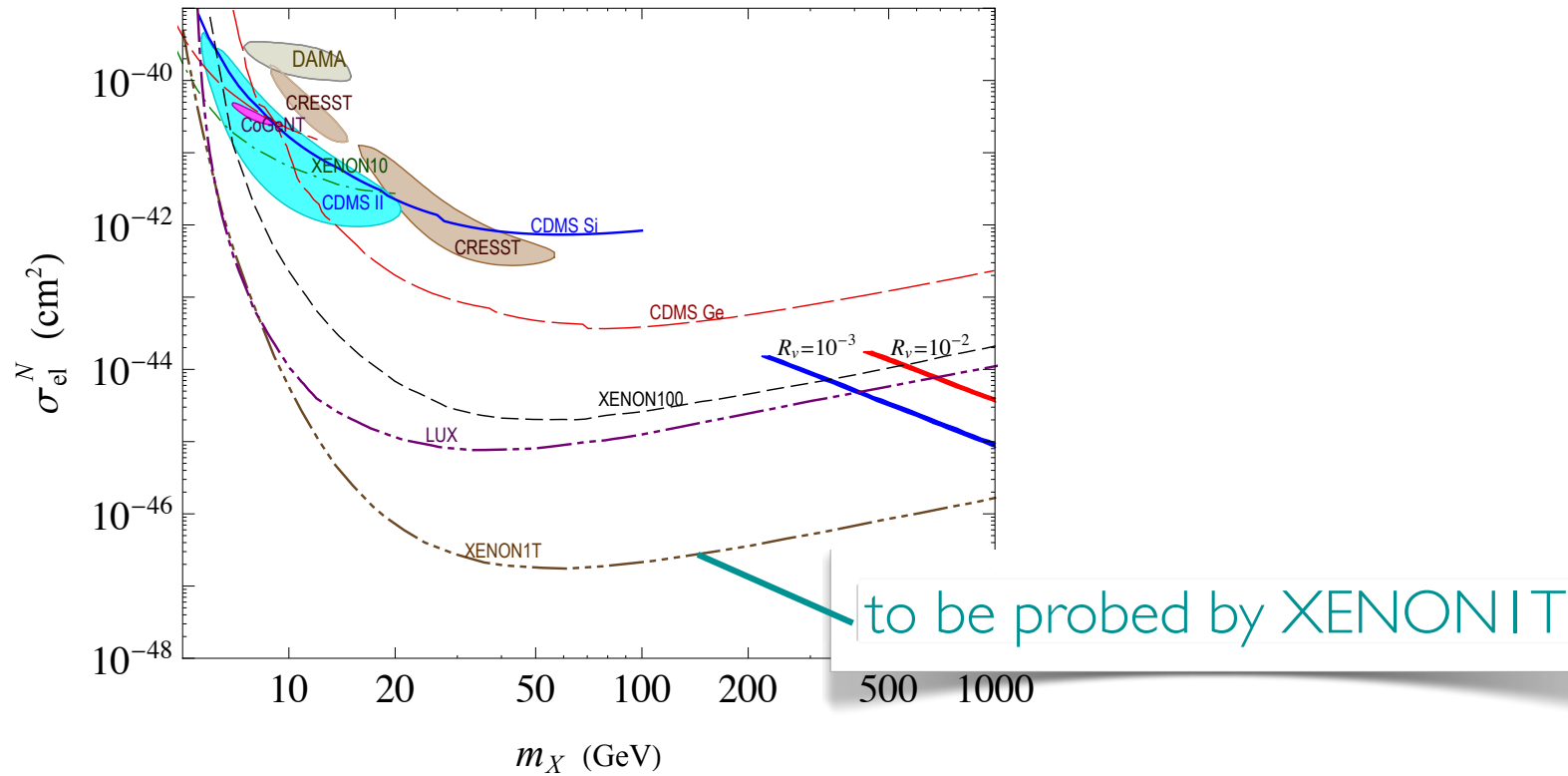


$$\mu_{XN} = \frac{m_X m_N}{m_X + m_N}$$

$$\sigma_{XN}^{\text{el}} = \frac{g_X^2 g_{B-L}^2 \cos^2 \theta \sin^2 \theta \mu_{XN}^2}{\pi m_{Z_L}^4} \simeq \frac{g_X^4 R_\nu^2 \mu_{XN}^2}{16\pi m_X^4}$$

- Still dominated by Z_L due to lighter mass

DIRECT SEARCH CONSTRAINT



- A significant portion of parameter space at high m_X is allowed by direct search
- Smaller R_ν is not helpful as it is ruled out by gauge coupling constraints

SUMMARY

- Z_2 symmetry for stabilizing DM candidates (darkons) emerges naturally as remnant of extra gauge groups, rather than *ad hoc*.
- Models easily accommodate SM-like Higgs, but also leave ample room for non-SM-like Higgs (including invisible Higgs decays).
- Incorporated minimal mechanism for (1) stabilizing DM using DGS and (2) generating neutrino mass through breaking of same group.
- Z' as well as Higgs contribute to DM interactions with SM particles.
 - consider Z' dominant in DM relic density determination
 - coannihilation is taken into account in U(1) model
 - natural resonance effect in non-Abelian extensions
- Checked constraints of new gauge coupling, DM relic density, and DM direct detection. WIMP of $O(100)$ GeV is favored.
- Computed invisible Higgs decays.
- Comments on collider pheno are given in paper.

SUMMARY

- We had a very successful workshop.
- I learned physics and ski.
- Many thanks to Eung Jin and organizing staff!

THANK YOU AND
LOOK FORWARD TO
SEEING YOU ALL IN “T-HIGH I”