

New routes of reactions by a long-lived negatively charged massive particle during big bang nucleosynthesis

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collaborators

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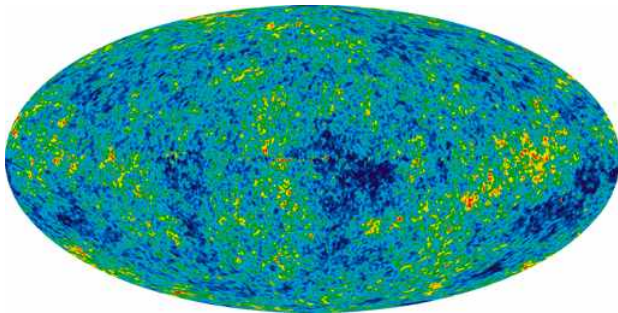
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Introduction

1. Light element abundances

- Standard big bang nucleosynthesis (BBN) parameter: baryon-to-photon ratio η
- Observation of CMB $\rightarrow$ constraint on η

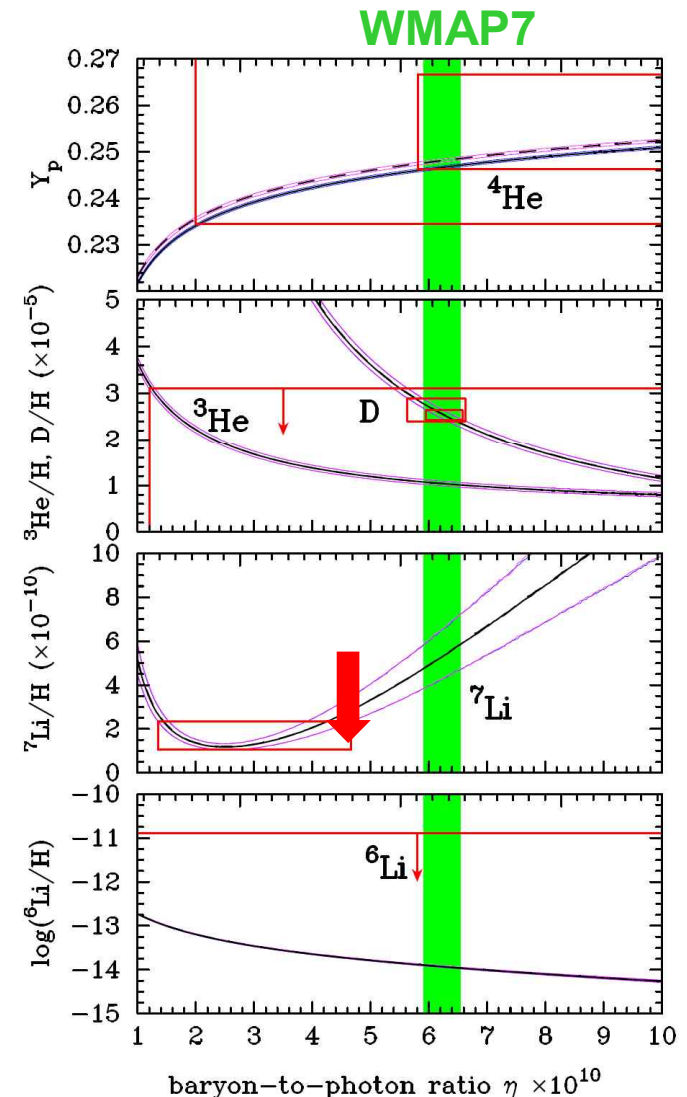


WMAP Science Team

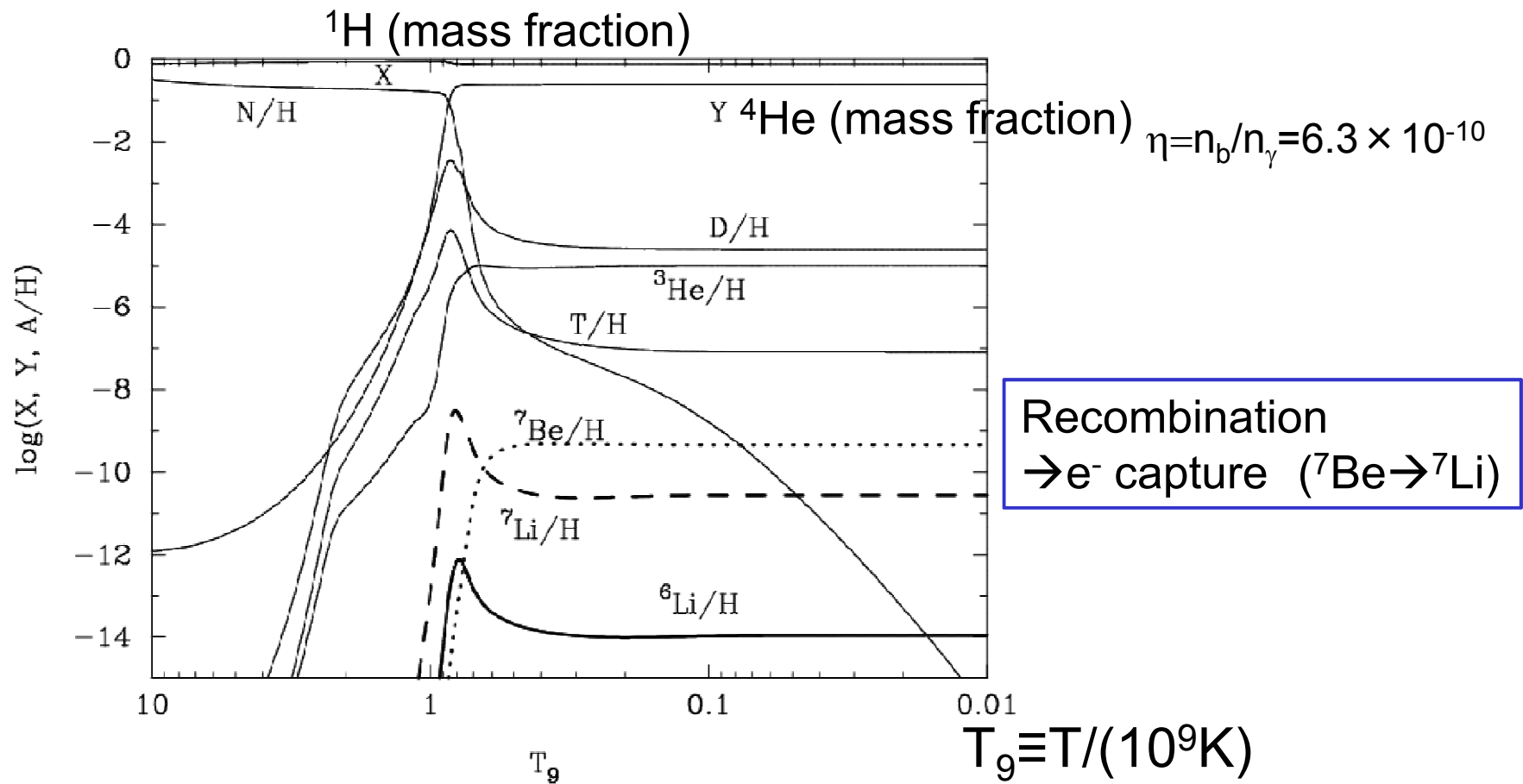
- Observation of metal-poor stars
 - ✓ ${}^7\text{Li}$ abundance is smaller than theory by a factor of ~ 3

Signature of new physics?

- Primordial abundances of ${}^9\text{Be}$, ${}^{10,11}\text{B}$, ... are not detected yet.



2. Standard Big Bang nucleosynthesis (SBBN)



3. Li problem

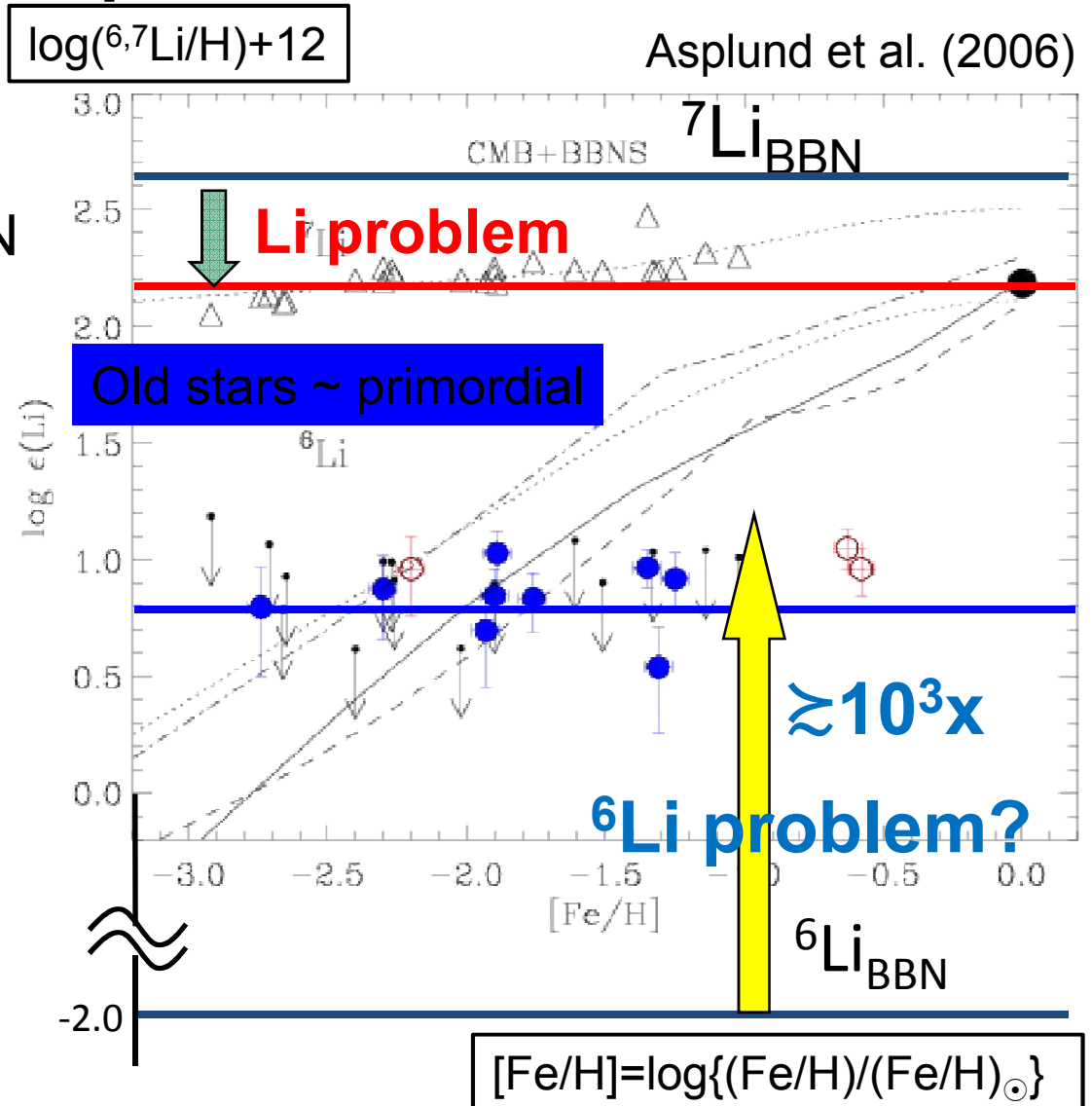
- ^7Li abundance observed in metal-poor stars are a factor ~ 3 smaller than SBBN

$$^7\text{Li}/\text{H} = (1.1 - 1.5) \times 10^{-10}$$

- Some of the stars have large abundances of ^6Li

$$^6\text{Li}/\text{H} \approx 6 \times 10^{-12}$$

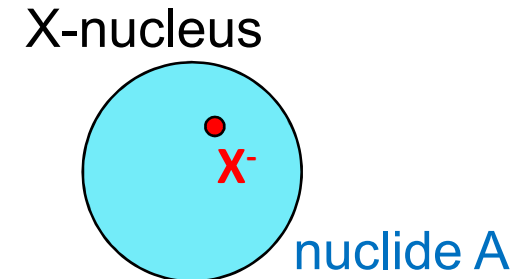
- But, Lind et al. perform 3D NLTE analysis for atmosphere, and found **no detection**



News: Lind et al. Astron. Astrophys. (2013)

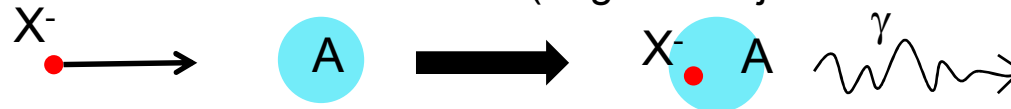
4. Effect of long-lived negatively charged massive particle (CHAMP) on BBN

- CHAMP X^- recombines with nuclide A, and X-nuclide (A_X) forms (Cahn & Glashow 1981)

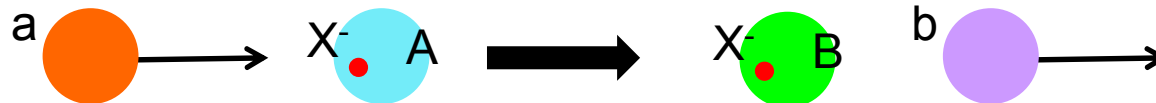


- exotic atoms form $\rightarrow$ exotic nucleosynthesis
(Pospelov 2007, Kohri & Takayama 2007, Kawasaki et al. 2007-, Hamaguchi et al. 2007, Jedamzik 2008-, MK et al. 2007-)

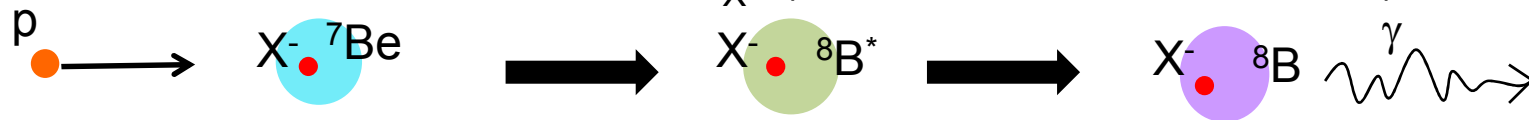
- ✓ Recombination of X^- and nuclide (e. g. De Rujula et al. 1990, Dimopoulos et al 1990)



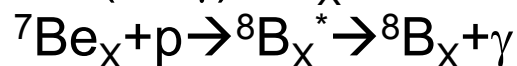
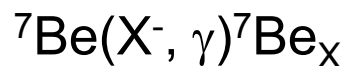
- ✓ Nuclear reaction of X-nuclide (e.g. Kohri & Takayama 2007)



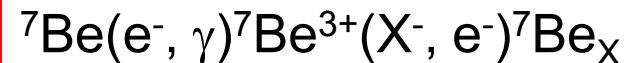
- ✓ resonant reaction via exotic atoms of ${}^8\text{B}_X^*$ (Bird et al. 2007; MK et al. 2007)



- Important reactions related to ${}^7\text{Be}$ abundance



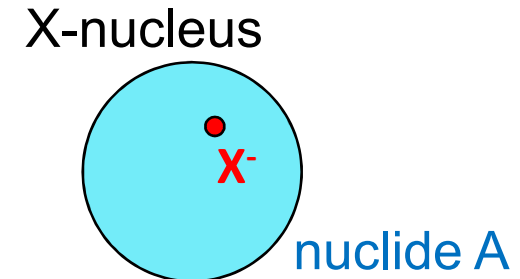
(1) Another route for ${}^7\text{Be}_X$ formation



(MK, Kim, Cheoun, Kajino, Kino, PRD 88, 063514, 2013)

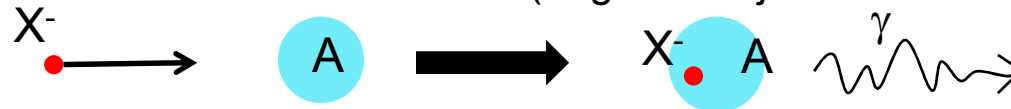
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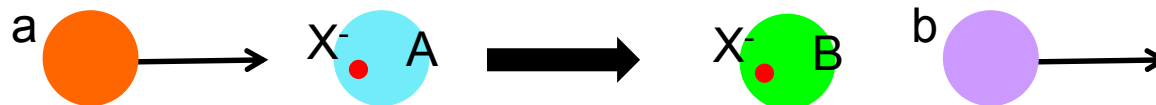


- exotic atoms form $\rightarrow$ exotic nucleosynthesis
(Pospelov 2007, Kohri & Takayama 2007, Kawasaki et al. 2007-, Hamaguchi et al. 2007, Jedamzik 2008-, MK et al. 2007-)

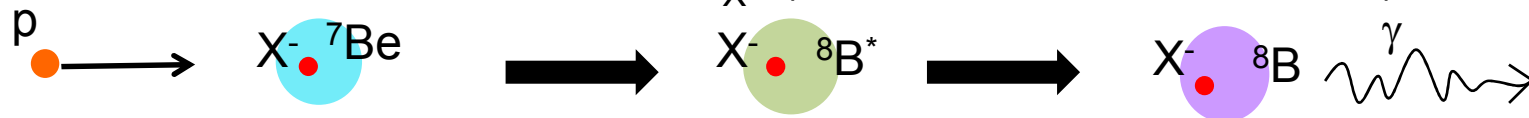
- ✓ Recombination of X^- and nuclide (e. g. De Rujula et al. 1990, Dimopoulos et al 1990)



- ✓ Nuclear reaction of X-nuclide (e.g. Kohri & Takayama 2007)



- ✓ resonant reaction via exotic atoms of ${}^8\text{B}_X^*$ (Bird et al. 2007; MK et al. 2007)



(2) extensive study of XBBN

Goal

- Precise calculation of recombination rates
- Estimation of dependences of BBN on the mass of X^- (m_X), and the nuclear charge distribution

(1) Recombination rate via ${}^7\text{Be}(e^-, \gamma){}^7\text{Be}^{3+}(\text{X}^-, e^-){}^7\text{Be}_x$

$$\Gamma_{\text{rec}} = \frac{n_{\text{Be}^{3+}}}{n_{\text{Be}^{4+}}} \left[\Gamma_{\text{Be}^{3+} \rightarrow \text{Be}_x^*} \frac{\Gamma_{\text{Be}_x^*, \text{tr}}}{\Gamma_{\text{Be}_x^*, \text{de}} + \Gamma_{\text{Be}_x^*, \text{tr}}} \right]$$

① number ratio of Be^{3+} and Be^{4+} is given by equilibrium value (Saha eq.)

② Reaction rate of ${}^7\text{Be}^{3+}$ via ${}^7\text{Be}^{3+}(\text{X}^-, e^-){}^7\text{Be}_x^*$

$$\sigma(E) = \sigma(I({}^7\text{Be}^{3+})) \left[\frac{E}{I({}^7\text{Be}^{3+})} \right]^{-1/2} H(I({}^7\text{Be}^{3+}) - E)$$

$$\sigma(I({}^7\text{Be}^{3+})) = 10 / (Z_{\text{Be}} \alpha m_e)^2 = 1.75 \times 10^7 \text{ b.}$$

MK, Kim, Cheoun, Kajino, Kino,
PRD 88, 063514, (2013)

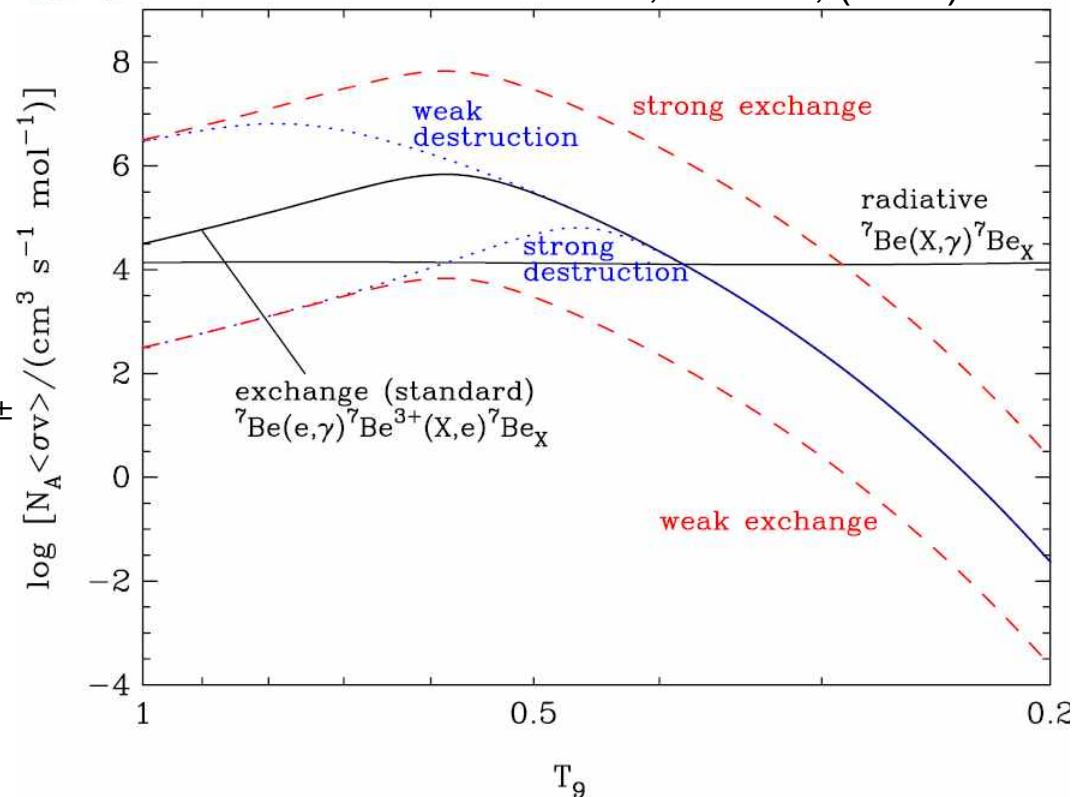
③ Probability that ${}^7\text{Be}_x^*$ is converted to the GS ${}^7\text{Be}_x$

➤ The dominant destruction process:
Collisional ionization: $A_x^* + e^\pm \rightarrow A + X + e^\pm$

Cross section is assumed as

$$\sigma(E) = \sigma_{\text{de}} H(E - E_{\text{th}})$$

$$\sigma_{\text{de}} \sim 10^{-2} \times \pi [2n^2 / (Z_1 Z_2 \alpha \mu)]^2$$



Abundance

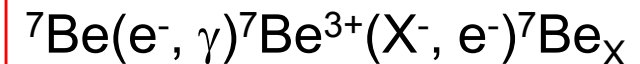
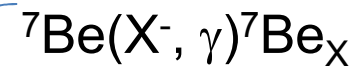
Results

MK, Kim, Cheoun, Kajino, Kino,
PRD 88, 063514, (2013)

abundances evolution

$n_x = 0.05 n_b$, $m_x \gg 1$ GeV, $\tau_x \gg 200$ s,
 $\eta = 6.19 \times 10^{-10}$ (WMAP 9yr)

${}^7\text{Be}$ recombination

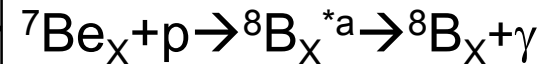


— $\sigma = 0$

— $\sigma_s = 17.5$ Mb (standard)

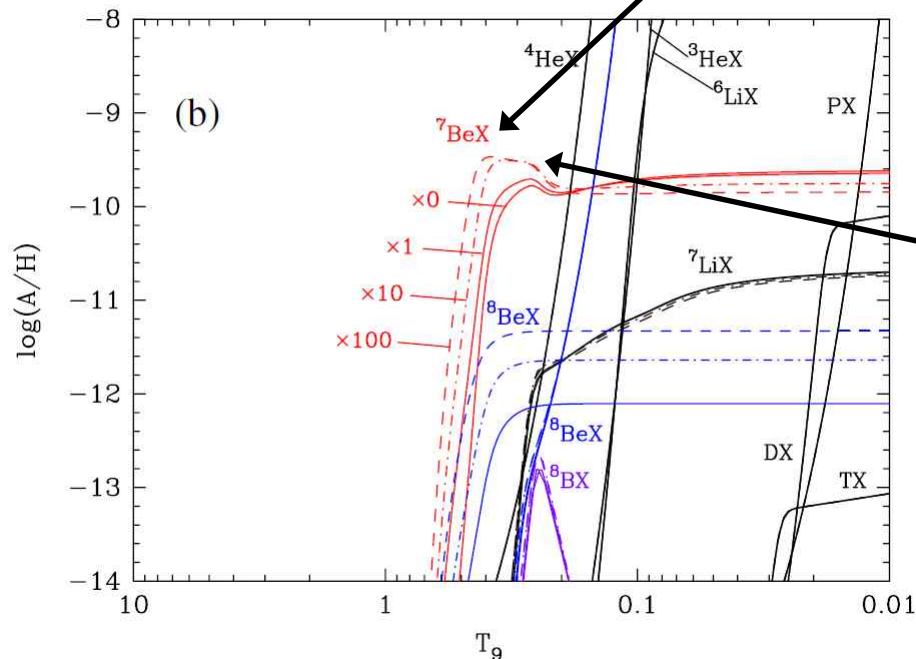
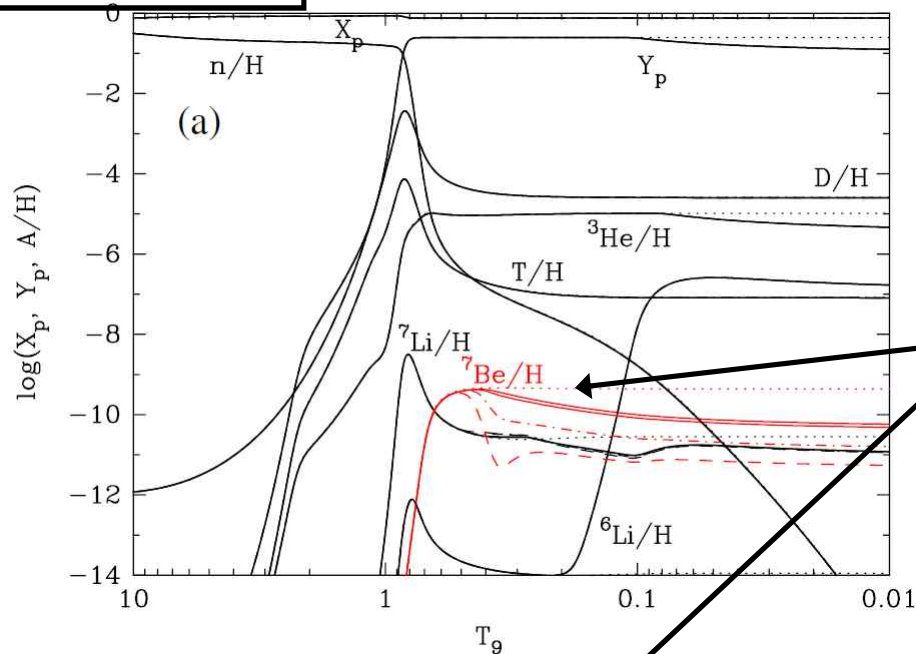
- - - $\sigma = 10\sigma_s$

- - - $\sigma = 100\sigma_s$



➤ amount of ${}^7\text{Be}$ destruction depends
on the rate of ${}^7\text{Be}^{3+}(X^-, e^-) {}^7\text{Be}_X^*$

Temperature $T_9 = T/(10^9 \text{ K})$



(2) Extensive study on XBBN Model

1. Binding energy of X-nuclides

[assumption]

➤ X^- : spin 0, charge -e, mass m_X (parameter)

➤ **Nuclear charge density**

1) Woods-Saxon

$$\rho(r') = \frac{ZeC}{1 + \exp[(r' - R)/a]}$$

2) Gaussian

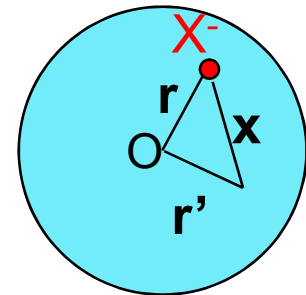
$$\rho(r') = Ze(\pi b)^{-3/2} \exp(-r'^2 / b^2) \quad b = \sqrt{2/3} \langle r_c^2 \rangle^{1/2}$$

3) homogeneous

$$\rho(r') = \frac{3Ze}{4\pi r_0^3} H(r_0 - r') \quad r_0 = \sqrt{5/3} \langle r_c^2 \rangle^{1/2}$$

Root mean square charge radius

X-nucleus



Nucleus A

Schrödinger equation

$$\left[-\frac{\hbar^2}{2\mu} \nabla^2 + V(r) - E \right] \psi(\mathbf{r}) = 0$$

$$V(r) = \int_0^\infty \frac{-e\rho(r')}{x} d^3r'$$

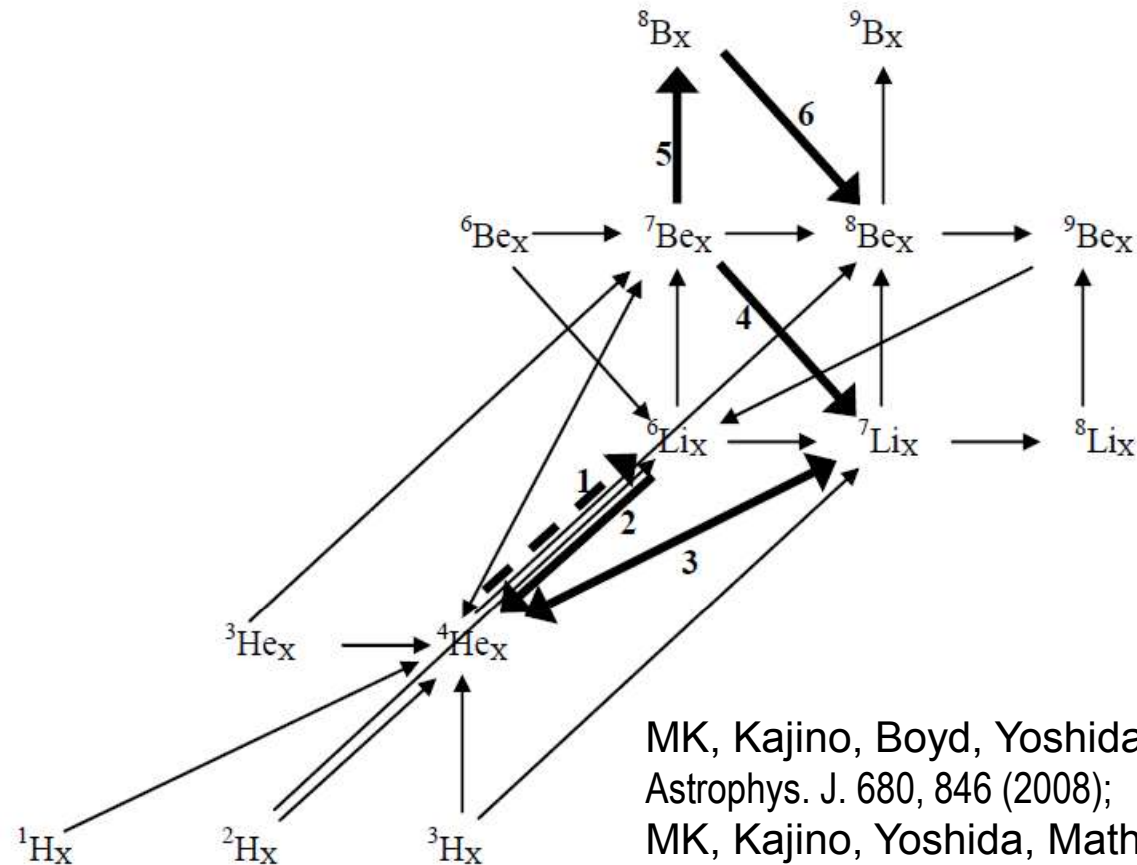
✓ Binding energies and wave functions are derived with

1) variational calculation (Gaussian expansion method, c.f. Hiyama et al. 2003)

2) numerical integration (RADCAP, code by Bertulani 2003)

2. Reaction network

➤ Up to C_X



MK, Kajino, Boyd, Yoshida, Mathews,
Astrophys. J. 680, 846 (2008);
MK, Kajino, Yoshida, Mathews,
PRD 81, 083521 (2010)

Important reactions

1. ${}^4\text{He}_X(d, X^-){}^6\text{Li}$
2. ${}^6\text{Li}_X(p, {}^3\text{He } \alpha)X^-$
3. ${}^4\text{He}_X(t, \gamma){}^7\text{Li}_X$ & ${}^7\text{Li}_X(p, 2\alpha)X^-$
4. ${}^7\text{Be}_X(, X^0){}^7\text{Li}$
5. ${}^7\text{Be}_X(p, \gamma){}^8\text{B}_X$
6. ${}^8\text{B}_X(, e^+ \nu_e){}^8\text{Be}_X$

- ✓ X^- recombination: 16
- ✓ X nuclear reaction: 59
(including β -decay: 2)
- ✓ X^- charge transfer: 3
- ✓ X -decay: 11

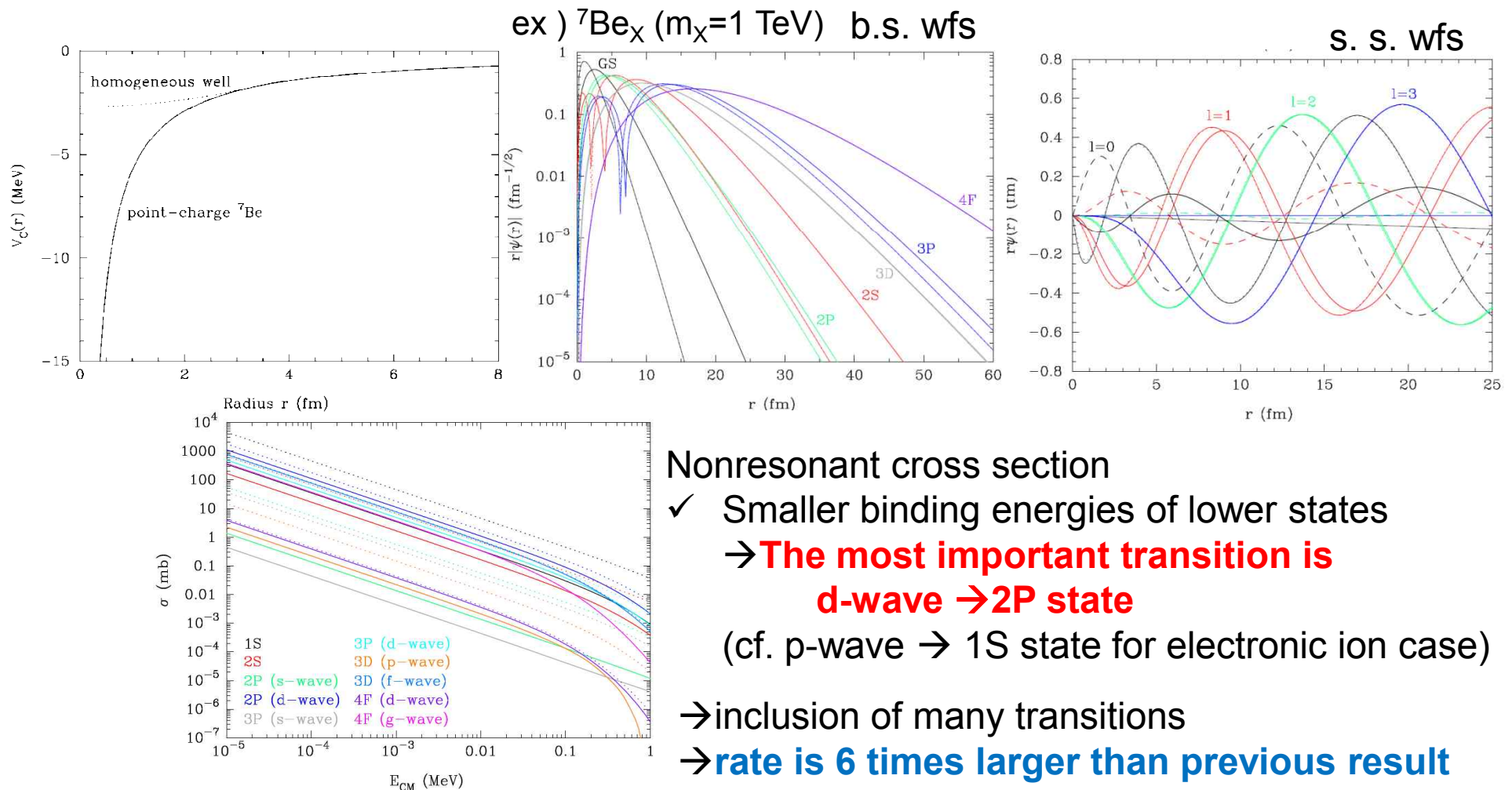
3. Recombination cross section

➤ Resonant and nonresonant cross sections for (${}^7\text{Be}$, ${}^7\text{Li}$, ${}^9\text{Be}$, ${}^4\text{He}$)

Finite size of nuclear charge

→ binding energies of tightly bound states are smaller than those of point-charges

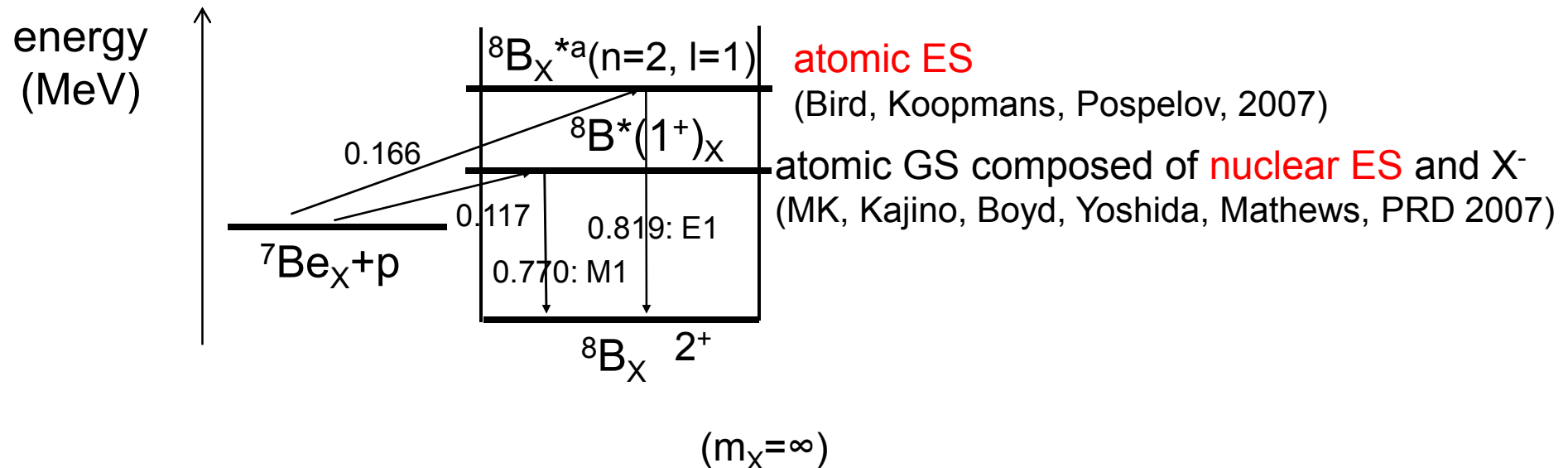
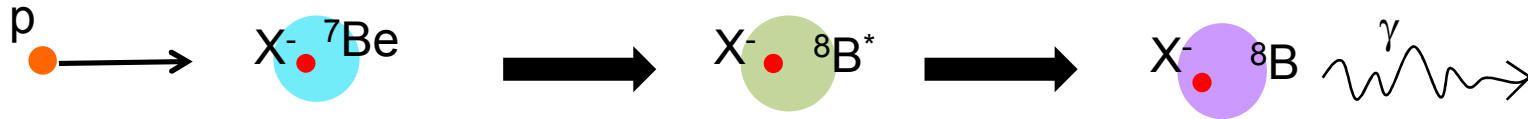
→ wave functions and recombination cross sections are also different



4. Nuclear reaction rate

- Binding energies of X-nuclei → reaction Q-values, detailed balance
- Reaction rates of X-nuclei are estimated with those of normal nuclei correcting for charge number and reduced mass
- We adopt cross sections calculated with a quantum three-body model
(Hamaguchi et al. 2007, Kamimura et al. 2008)

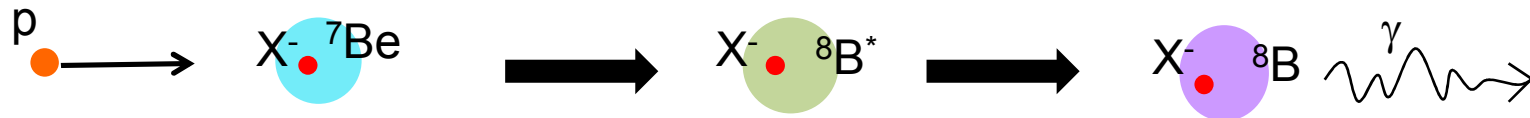
- Important resonant reaction:



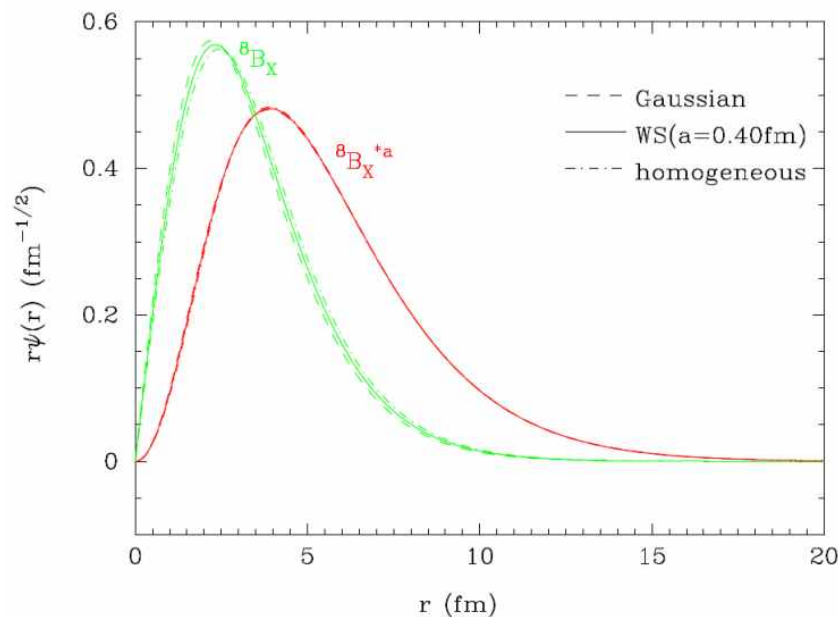
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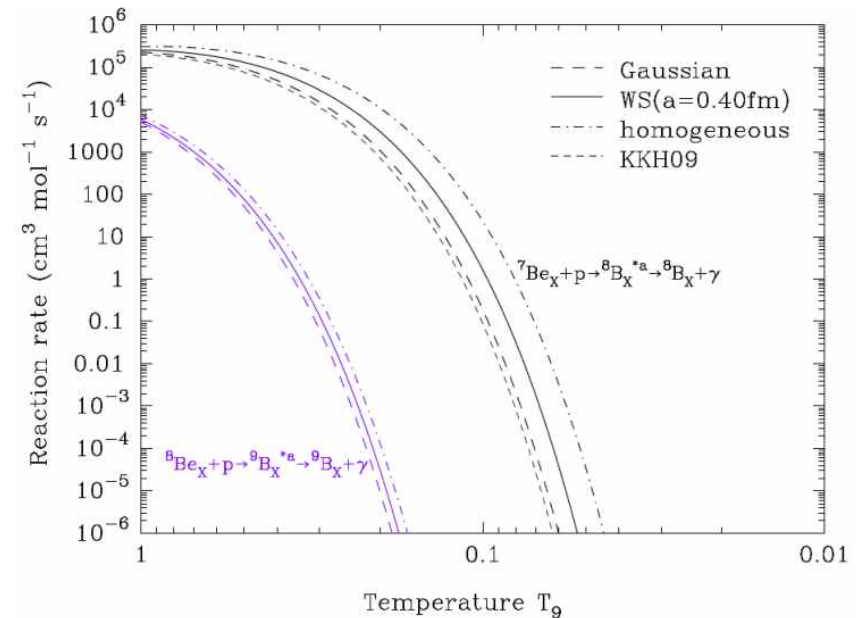
- Important resonant reaction:



ex) $m_X = \infty$ b.s. wfs ($^8\text{B}_X$)



Thermal reaction rates



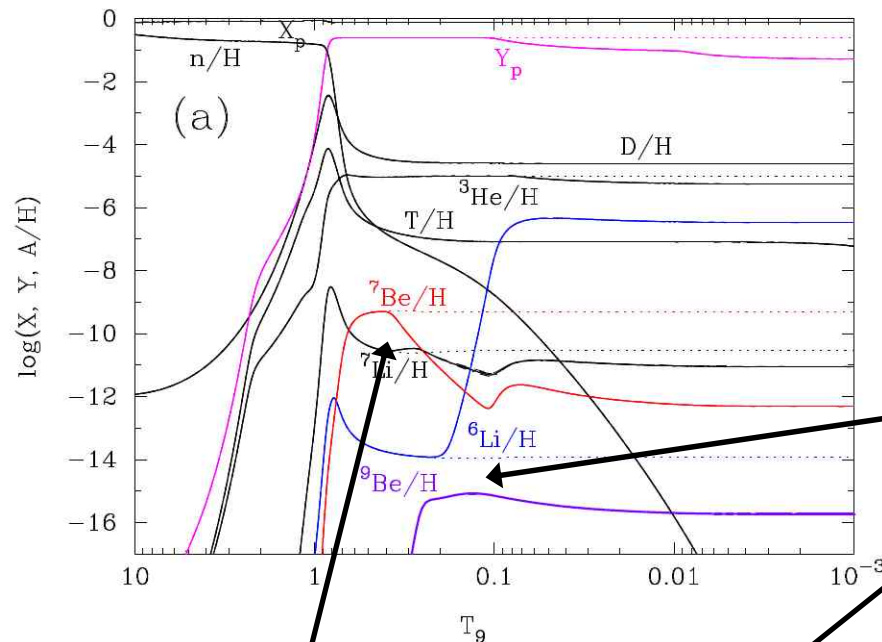
Abundance

Results

abundances evolution

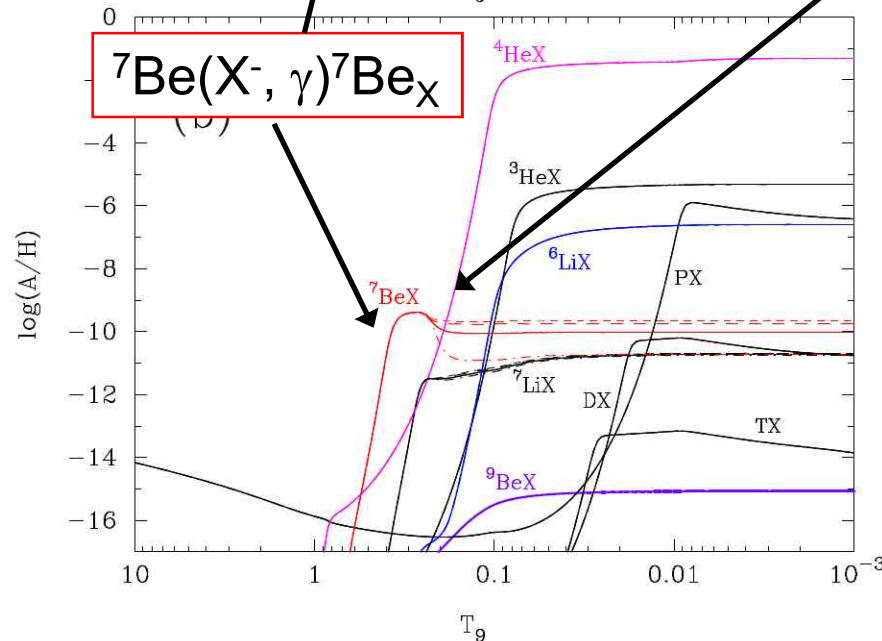
case 1 ($m_X = 1$ TeV)

$n_X = 0.05 n_b$, $\tau_X \gg 200$ s, $\eta = 6.19 \times 10^{-10}$
(WMAP 9yr)



${}^7\text{Li}_X(d, X^-){}^9\text{Be}$ (this study)

${}^7\text{Be}_X + p \rightarrow {}^8\text{B}_X^{*a} \rightarrow {}^8\text{B}_X + \gamma$ (Bird et al. 2008)



${}^7\text{Be}(X^-, \gamma){}^7\text{Be}_X$

--- Gaussian
— Woods-Saxon
- - - homogeneous

➤ the resonance height of ${}^8\text{B}_X^*$ is sensitive to the nuclear charge distribution

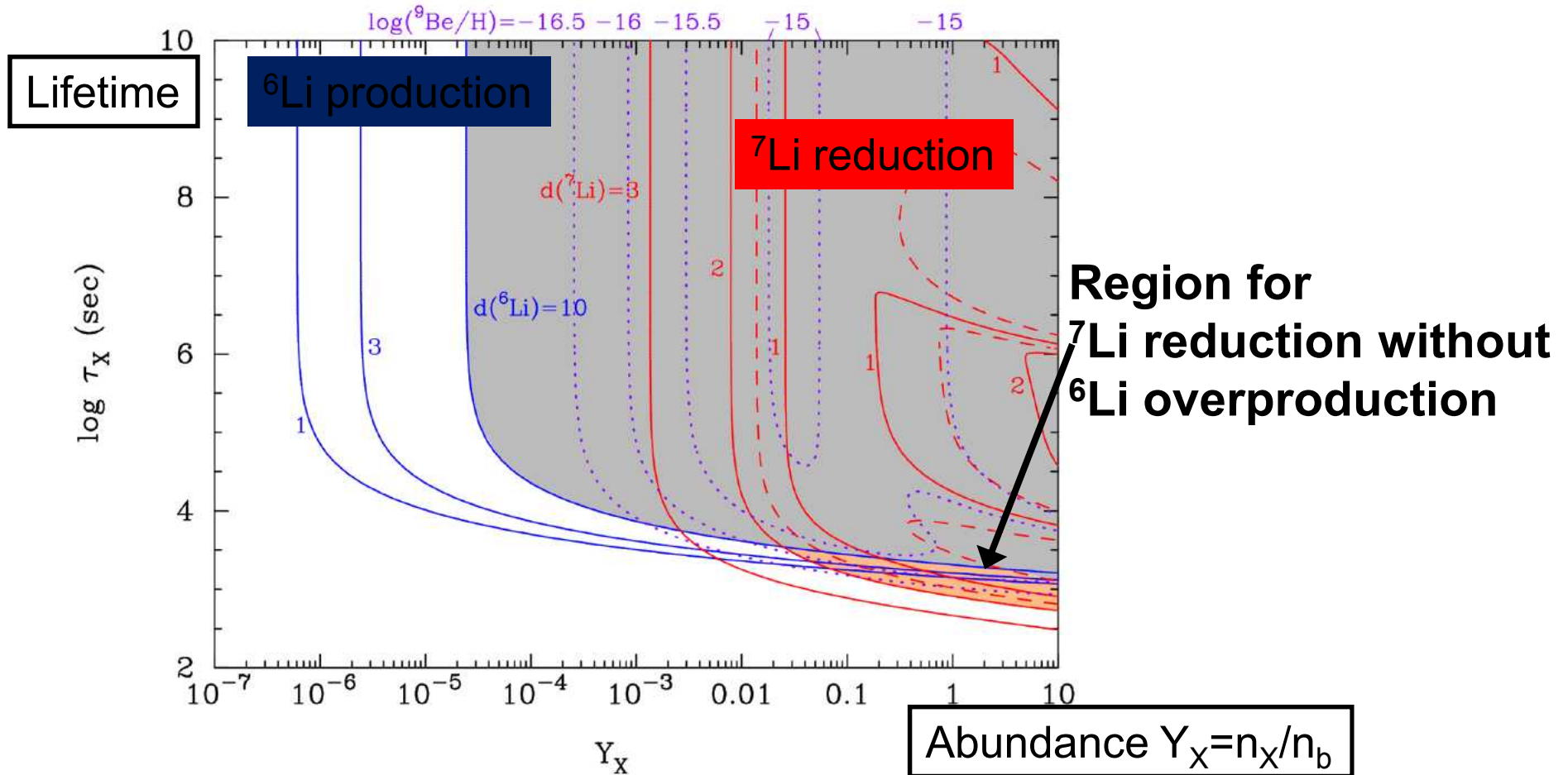


➤ amount of ${}^7\text{Be}$ destruction significantly depends on the charge distribution

Temperature $T_9 = T/(10^9 \text{ K})$

Parameter search: case 1 ($m_X=1$ TeV)

Contours of calculated Li abundance relative to the observed value: $d(^A\text{Li}) = A_{\text{Li}}^{\text{Calc}} / A_{\text{Li}}^{\text{Obs}}$

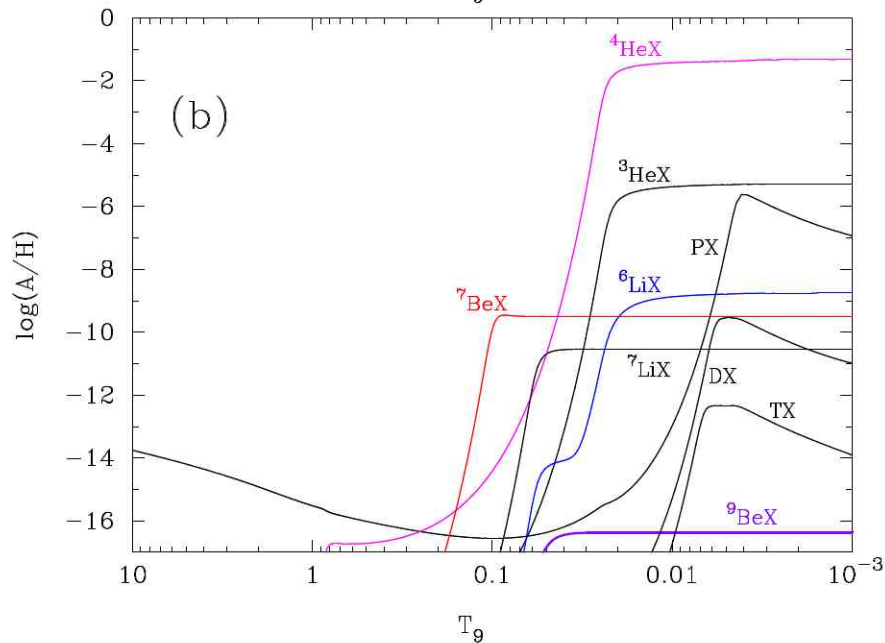
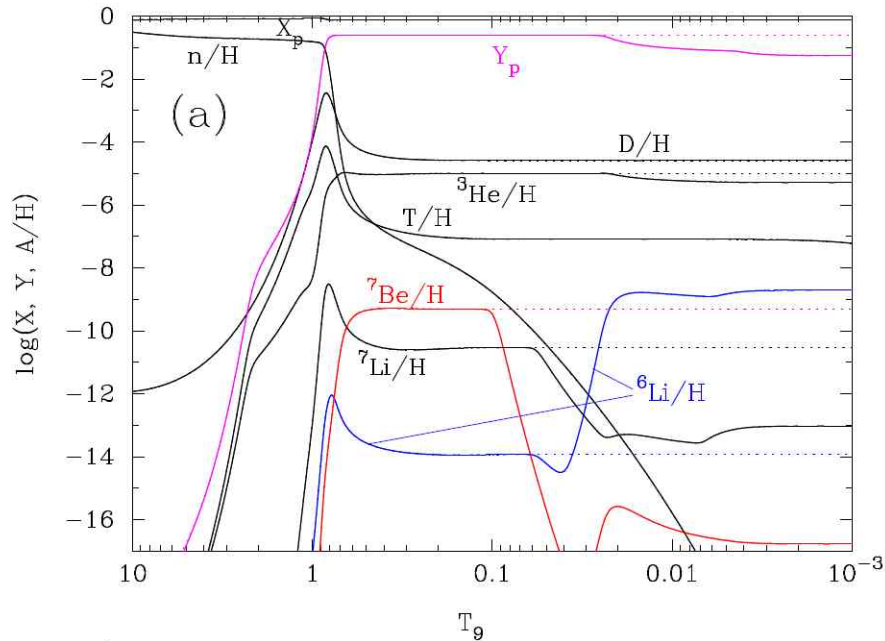


- Realistic parameter region is shifted to **$Y_X \gtrsim 0.02$ and $\tau_X \approx (0.6-3) \times 10^3 \text{ s}$** from the previous one: $Y_X \gtrsim 1$ and $\tau_X \approx (1-2) \times 10^3 \text{ s}$ (MK et al. 2010)

Summary

- We calculated light-element nucleosynthesis during BBN with negatively-charged X^- particles dynamically.
- **New route of ${}^7\text{Be}_x$ formation** (${}^7\text{Be}$ exchange between ${}^7\text{Be}^{3+}$ and X^-)
- Rates of recombination of X^- and ${}^7\text{Be}$, ${}^7\text{Li}$, ${}^9\text{Be}$, ${}^4\text{He}$ are derived.
- **New ${}^9\text{Be}$ production reaction** [${}^7\text{Li}_x(d, X^-){}^9\text{Be}$].
- **Parameter region of ${}^7\text{Li}$ reduction is moved**
 - ✓ $Y_X \gtrsim 0.02$ and $\tau_X \approx (0.6-3) \times 10^3 \text{s}$ (for $m_X = 1 \text{ TeV}$)
→ required abundance of X^- particle is smaller than the previous estimate by more than a factor of 10
- **Resulting ${}^7\text{Li}$ abundance depends significantly on assumed nuclear charge distribution**
 - ✓ Energy levels of X -nuclides are affected by charge distribution
→ resonant reaction rates are also affected

Abundance



abundances evolution

case 2 ($m_\chi = 1$ GeV)

$n_\chi = 0.05 n_b$, $\tau_\chi \gg 200$ s, $\eta = 6.19 \times 10^{-10}$
(WMAP 9yr)

➤ m_χ is small $\rightarrow$ binding energies are small
 $\rightarrow A_\chi$ forms at low temperature

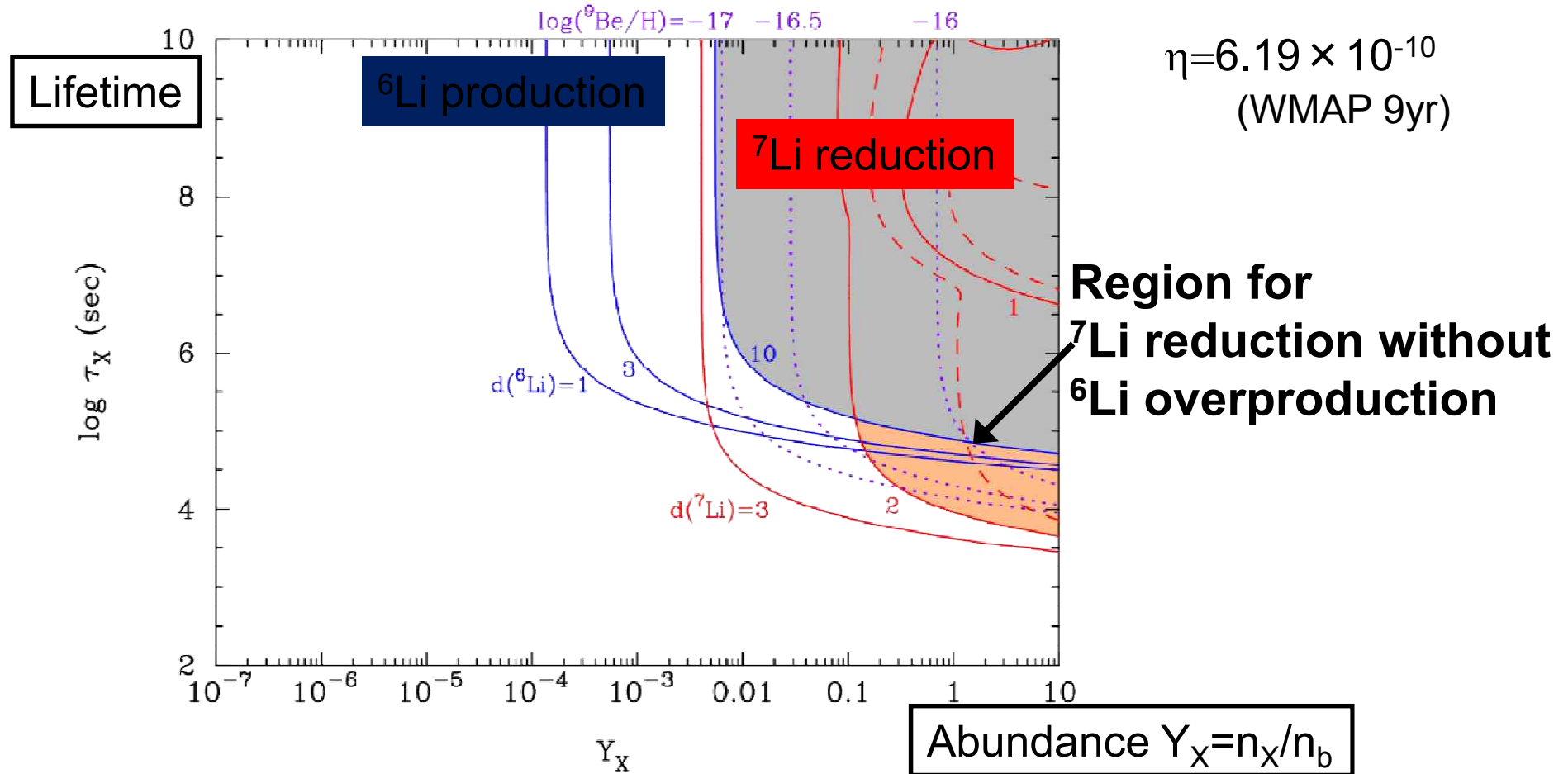


➤ ${}^7\text{Be}$ destruction and ${}^6\text{Li}$ production
are less efficient

temperature $T_9 = T/(10^9 \text{ K})$

Parameter search: case 2 ($m_X=1$ GeV)

Contours of calculated Li abundance relative to the observed value: $d(A\text{Li})=A\text{Li}^{\text{Calc}}/A\text{Li}^{\text{Obs}}$



➤ Parameter region for smaller mass cases locates at longer lifetime region:

$Y_X \gtrsim 0.1$ and $\tau_X \approx 5 \times 10^3 - 2 \times 10^5$ s (for $m_X=1$ GeV)

Processes affecting elemental abundances

Existence of particle [z~10 ⁹] Early stars [z~O(10)]	Model	⁷ Li problem solved ?	Signatures on other nuclides ?
	sub-SIMP X⁰	✓	⁶Li, ⁹Be
	SIMP X⁰	no	⁹Be and/or ¹⁰B
	CHAMP X ⁻ *	✓	⁶ Li, ⁹Be
	Early cosmic ray	no	⁶ Li, ⁹ Be & ^{10,11} B

Model

1. Recombination rate via ${}^7\text{Be}(e^-, \gamma){}^7\text{Be}^{3+}(\text{X}^-, e^-){}^7\text{Be}_\text{X}$

$$\Gamma_{\text{rec}} = \frac{n_{\text{Be}^{3+}}}{n_{\text{Be}^{4+}}} \left[\Gamma_{\text{Be}^{3+} \rightarrow \text{Be}_\text{X}^*} \frac{\Gamma_{\text{Be}_\text{X}^*, \text{tr}}}{\Gamma_{\text{Be}_\text{X}^*, \text{de}} + \Gamma_{\text{Be}_\text{X}^*, \text{tr}}} \right]$$

① number ratio of Be^{3+} and Be^{4+}

② Reaction rate of ${}^7\text{Be}^{3+}$ for the charge exchange reaction ${}^7\text{Be}^{3+}(\text{X}^-, e^-){}^7\text{Be}_\text{X}^*$

③ Probability that ${}^7\text{Be}_\text{X}^*$ excited states (ESs) transit to the ${}^7\text{Be}_\text{X}$ ground state (GS)
 $\Gamma_{\text{Be}_\text{X}^*, \text{tr}}$ rate for $\text{ES} \rightarrow \text{GS}$ $\Gamma_{\text{Be}_\text{X}^*, \text{de}}$ destruction rate

2. number ratio $\text{Be}^{3+}/\text{Be}^{4+}$

1) Hydrogen like-ion

$$\text{BE}(n) = (Z_1 Z_2 \alpha)^2 \mu / (2n^2); \quad \langle r \rangle \sim n^2 / (Z_1 Z_2 \alpha \mu) = Z_1 Z_2 \alpha / [2\text{BE}(n)]$$

$${}^7\text{Be}^{3+} \text{ GS: } \text{BE}(1) = 218 \text{ eV}, \quad \langle r \rangle = 1.98 \times 10^{-9} \text{ cm}$$

$${}^7\text{Be}_\text{X} \text{ (} n \gg 1 \text{) states: } \text{BE}(n) = 2.78 \text{ MeV}/n^2, \quad \langle r \rangle \sim 1.04 n^2 \times 10^{-13} \text{ cm}$$

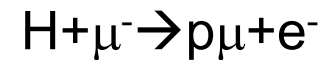
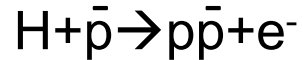
$\rightarrow {}^7\text{Be}_\text{X} \text{ (} n=113 \text{) ESs has nearly the same binding energy as the } {}^7\text{Be}^{3+} \text{ GS}$

2) ${}^7\text{Be}^{3+}$ can be considered as an isolated ionic state $\leftarrow \langle r \rangle < l_{\text{ave}}(e^\pm)$

3) The equilibrium value $\frac{n_{\text{Be}^{3+}}}{n_{\text{Be}^{4+}}} = \left(\frac{2\pi}{m_e T} \right)^{3/2} \exp \left[\frac{I({}^7\text{Be}^{3+})}{T} \right] n_e \sim 2e^{-m_e/T}$

3. Reaction rate for the reaction ${}^7\text{Be}^{3+}(\text{X}^-, \text{e}^-) {}^7\text{Be}_\text{X}^*$

- cross section is assumed by analogies of protonium formation and muonic hydrogen formation



$$\sigma(E) = \sigma(I({}^7\text{Be}^{3+})) \left[\frac{E}{I({}^7\text{Be}^{3+})} \right]^{-1/2} H(I({}^7\text{Be}^{3+}) - E)$$

$$\sigma(I({}^7\text{Be}^{3+})) = 10 / (Z_{{}^7\text{Be}} \alpha m_e)^2 = 1.75 \times 10^7 \text{ b.}$$

- binding energies of final states are similar to that of the initial state
 $\rightarrow {}^7\text{Be}^{3+}(\text{n})$ is converted to ${}^7\text{Be}_\text{X}^*$ (113n)

4. Probability that ${}^7\text{Be}_\text{X}^*$ is converted to ${}^7\text{Be}_\text{X}$

I) transition

1) Spontaneous emission: $A_\text{X}^* \rightarrow A_\text{X}^{*'} + \gamma$

2) Stimulate emission: $A_\text{X}^* + \gamma \rightarrow A_\text{X}^{*'} + 2\gamma$

3) Photo-absorption: $A_\text{X}^{*'} + \gamma \rightarrow A_\text{X}^*$

II) destruction

1) Collisional ionization: $A_\text{X}^* + \text{e}^\pm \rightarrow \text{A} + \text{X}^- + \text{e}^\pm$

2) Charge exchange: $A_\text{X}^* + \text{e}^- \rightarrow \text{A}^{(Z-1)+} + \text{X}^-$

3) Photo-ionization: $A_\text{X}^* + \gamma \rightarrow \text{A} + \text{X}^-$

$$\sigma_{\text{des}} \sim 10^{-2} \times \pi [2n^2 / (Z_1 Z_2 \alpha \mu)]^2$$

$$\Gamma_{\text{Be}_\text{X}^*, \text{tr}} \gtrsim \Gamma_{\text{Be}_\text{X}^*, \text{de}} \rightarrow \text{GS } {}^7\text{Be}^{3+} \text{ is only available path to GS } {}^7\text{Be}_\text{X}$$

5. Recombination rate via ${}^7\text{Be}(e^-, \gamma){}^7\text{Be}^{3+}(X^-, e^-){}^7\text{Be}_X$

$$\Gamma_{\text{rec}} = \frac{n_{\text{Be}^{3+}}}{n_{\text{Be}^{4+}}} \left[\Gamma_{\text{Be}^{3+} \rightarrow \text{Be}_X^*} \frac{\Gamma_{\text{Be}_X^*, \text{tr}}}{\Gamma_{\text{Be}_X^*, \text{de}} + \Gamma_{\text{Be}_X^*, \text{tr}}} \right]$$

① number ratio of Be^{3+} and Be^{4+} is given by equilibrium value (Saha eq.)

② cross section for charge exchange reaction ${}^7\text{Be}^{3+}(X^-, e^-){}^7\text{Be}_X^*$ is assumed as

$$\sigma(E) = \sigma(I({}^7\text{Be}^{3+})) \left[\frac{E}{I({}^7\text{Be}^{3+})} \right]^{-1/2} H(I({}^7\text{Be}^{3+}) - E)$$

$$\sigma(I({}^7\text{Be}^{3+})) = 10 / (Z_{\text{Be}} \alpha m_e)^2 = 1.75 \times 10^7 \text{ b.}$$

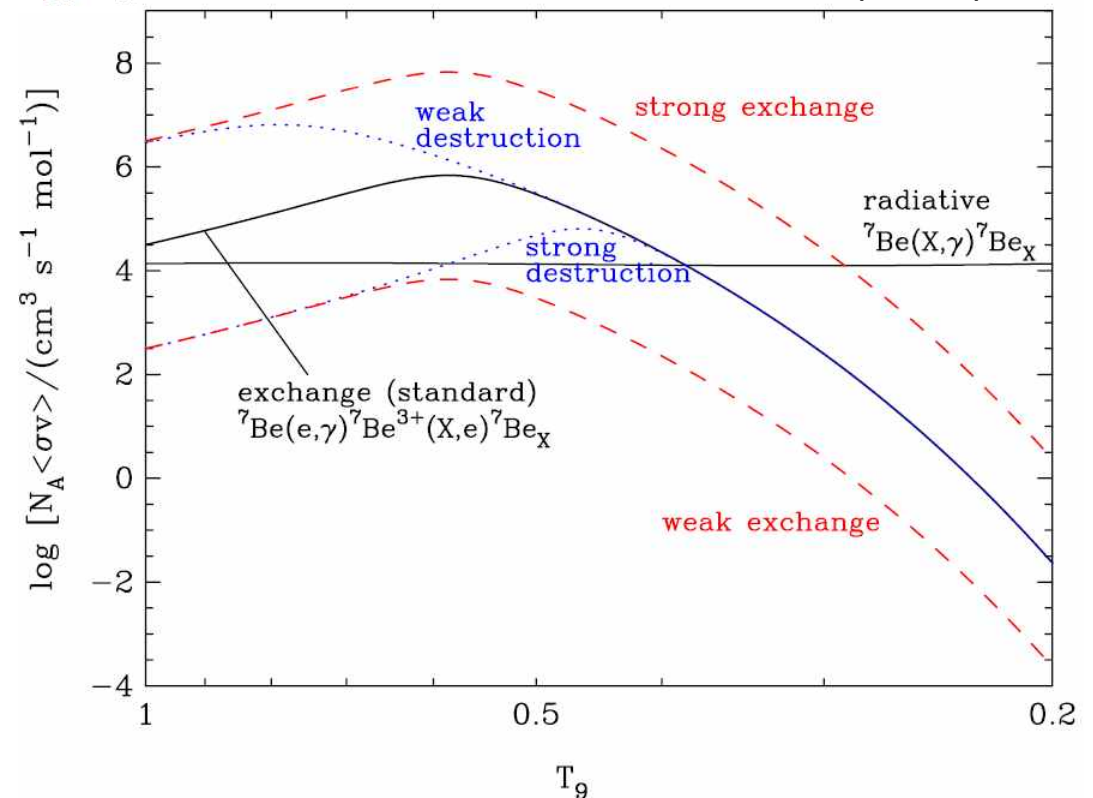
MK, Kim, Cheoun, Kajino, Kino,
PRD 88, 063514, (2013)

③ Probability of the GS ${}^7\text{Be}_X$ formation

➤ The dominant destruction process:
ionization via $A_X^* + e^\pm$ collision

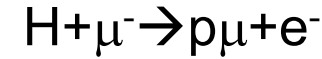
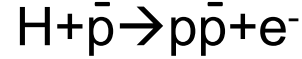
Cross section is assumed as

$$\sigma(E) = \sigma_{\text{de}} H(E - E_{\text{th}})$$



3. Reaction rate for the reaction ${}^7\text{Be}^{3+}(\text{X}^-, \text{e}^-) {}^7\text{Be}_\text{X}^*$

- cross section is assumed by analogies of protonium formation and muonic hydrogen formation



$$\sigma(E) = \sigma(I({}^7\text{Be}^{3+})) \left[\frac{E}{I({}^7\text{Be}^{3+})} \right]^{-1/2} H(I({}^7\text{Be}^{3+}) - E)$$

$$\sigma(I({}^7\text{Be}^{3+})) = 10 / (Z_{{}^7\text{Be}} \alpha m_e)^2 = 1.75 \times 10^7 \text{ b.}$$

- binding energies of final states are similar to that of the initial state

$\rightarrow {}^7\text{Be}^{3+}(\text{n})$ is converted to ${}^7\text{Be}_\text{X}^*$ (113n)

$$\begin{aligned} \Gamma_{\text{Be}^{3+} \rightarrow \text{Be}_\text{X}^*} &= n_\text{X} \langle \sigma v \rangle = n_\text{b} Y_\text{X} \langle \sigma v \rangle \\ &= 3.96 Y_\text{X} \times 10^4 \text{ s}^{-1} \left(\frac{h}{0.700} \right)^2 \left(\frac{\Omega_\text{b}}{0.0463} \right) \\ &\quad \times \left(\frac{T_9}{0.4} \right)^{3/2} \left[\frac{\sigma(I({}^7\text{Be}^{3+}))}{1.75 \times 10^7 \text{ b}} \right], \end{aligned}$$

Sakimoto, J. Phys. B (2001)

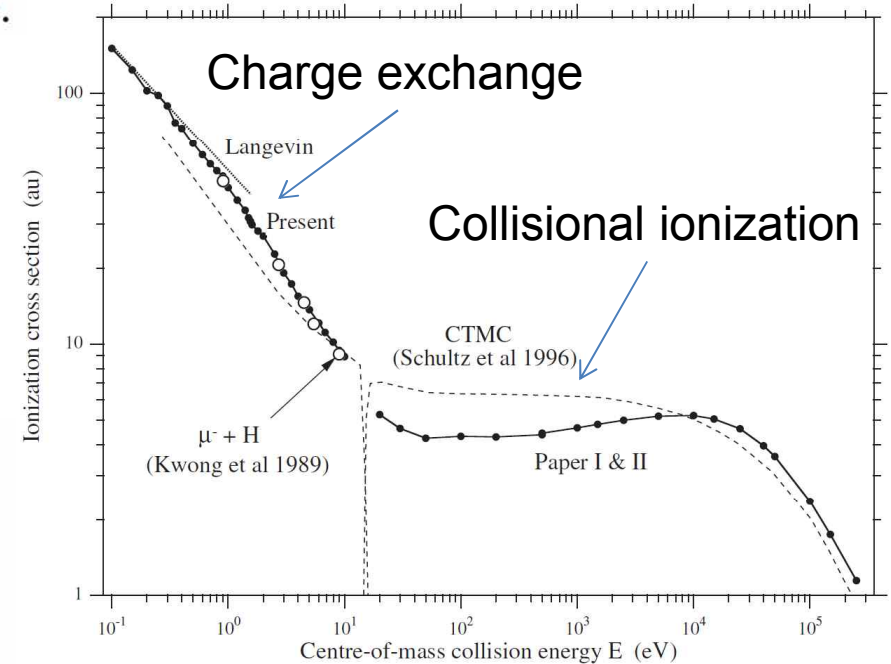


Figure 9. Ionization cross sections in the energy range $E = 10^{-1} - 2.5 \times 10^5$ eV calculated by the present method of direct numerical solution. The values at $E < I$ are the present results for the protonium formation (1), and the ones at $E > I$ are the results of paper I ($E \leq 500$ eV) and paper II ($E \geq 500$ eV) for the break-up ionization (2). The CTMC results are given by Schultz *et al* (1996). The $\mu^- + \text{H}$ results are the capture cross section σ_cap calculated by the semiclassical method (Kwong *et al* 1989, Cohen 1998).

4. Probability that ${}^7\text{Be}_x^*$ is converted to ${}^7\text{Be}_x$

I) transition

1) Spontaneous emission: $A_x^* \rightarrow A_x^{*'} + \gamma$

2) Stimulate emission: $A_x^* + \gamma \rightarrow A_x^{*'} + 2\gamma$

3) Photo-absorption: $A_x^{*'} + \gamma \rightarrow A_x^*$

$$\begin{aligned}\Gamma_{u,st}^\gamma &= \sum_l B_{ul} B_{\nu_{ul}}(T) = \sum_l \frac{2Z_A^2 \alpha}{m_A} \frac{g_l}{g_u} f_{lu} \frac{E_{ul}^2}{\exp(E_{ul}/T) - 1} \\ &\sim \frac{2N_l Z_A^2 \alpha}{m_A} T \overline{E_{ul}} \\ &= 9.21 \times 10^{13} \text{ s}^{-1} \left(\frac{N_l}{226}\right) \left(\frac{\overline{E_{ul}}}{218 \text{ eV}}\right) \left(\frac{T_9}{0.4}\right) \left(\frac{Z_A}{4}\right)^2 \left(\frac{m_A}{6.53 \text{ GeV}}\right)^{-1} \\ \overline{E_{ul}} &\equiv \sum_l [(g_l/g_u) f_{lu} E_{ul}] / N_l\end{aligned}$$

II) destruction

1) Collisional ionization: $A_x^* + e^\pm \rightarrow A + X^- + e^\pm$

2) Charge exchange: $A_x^* + e^- \rightarrow A^{(Z-1)+} + X^-$

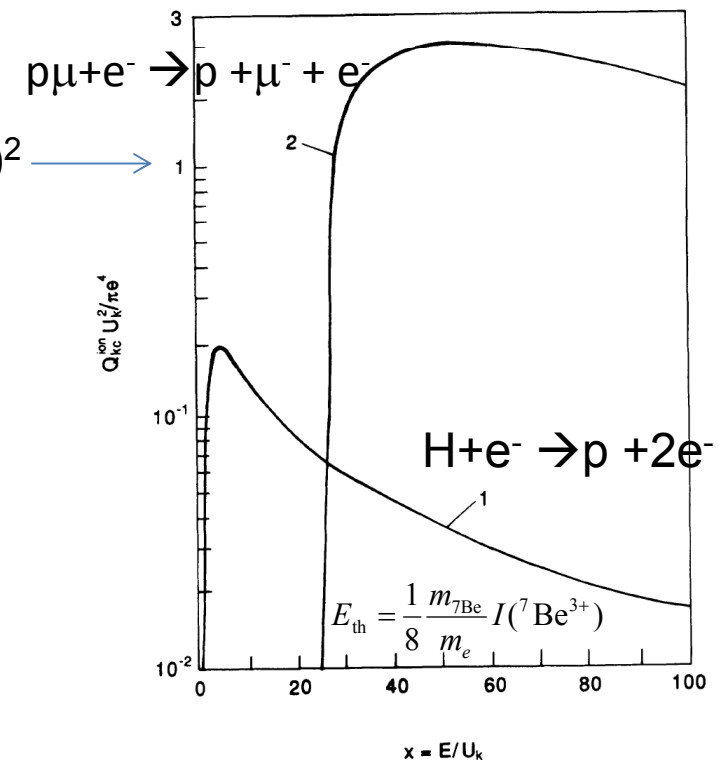
3) Photo-ionization: $A_x^* + \gamma \rightarrow A + X^-$

$$\begin{aligned}4 \times (\text{Atomic radius})^2 &\longrightarrow \\ = \pi [2n^2 / (Z_1 Z_2 \alpha \mu)]^2\end{aligned}$$

$$\begin{aligned}\Gamma_{\text{Be}_x^*, \text{de}} &= (n_{e^-} + n_{e^+}) \langle \sigma v \rangle \\ &= 2.79 \times 10^{10} \text{ s}^{-1} \left(\frac{T_9}{0.4}\right)^2 \frac{e^{-m_e/T}}{2.97 \times 10^{-7}} \\ &\quad \times \frac{(1 + E_{\text{th}}/T)}{11.1} \frac{e^{-E_{\text{th}}/T}}{4.13 \times 10^{-5}} \left(\frac{n}{113}\right)^4, \quad (29)\end{aligned}$$

GS ${}^7\text{Be}^{3+}$ is only available path

Kunc, Phys. Rev. A (1994)



4. Probability that ${}^7\text{Be}_x^*$ is converted to ${}^7\text{Be}_x$

I) transition

- 1) Spontaneous emission: $A_X^* \rightarrow A_X^{*'} + \gamma$
- 2) Stimulate emission: $A_X^* + \gamma \rightarrow A_X^{*'} + 2\gamma$
- 3) Photo-absorption: $A_X^{*'} + \gamma \rightarrow A_X^*$

$$\begin{aligned}\Gamma_{u,st}^\gamma &= \sum_l B_{ul} B_{\nu_{ul}}(T) = \sum_l \frac{2Z_A^2 \alpha}{m_A} \frac{g_l}{g_u} f_{lu} \frac{E_{ul}^2}{\exp(E_{ul}/T) - 1} \\ &\sim \frac{2N_l Z_A^2 \alpha}{m_A} T \overline{E_{ul}} \\ &= 9.21 \times 10^{13} \text{ s}^{-1} \left(\frac{N_l}{226} \right) \left(\frac{\overline{E_{ul}}}{218 \text{ eV}} \right) \left(\frac{T_9}{0.4} \right) \left(\frac{Z_A}{4} \right)^2 \left(\frac{m_A}{6.53 \text{ GeV}} \right)^{-1} \\ \overline{E_{ul}} &\equiv \sum_l [(g_l/g_u) f_{lu} E_{ul}] / N_l\end{aligned}$$

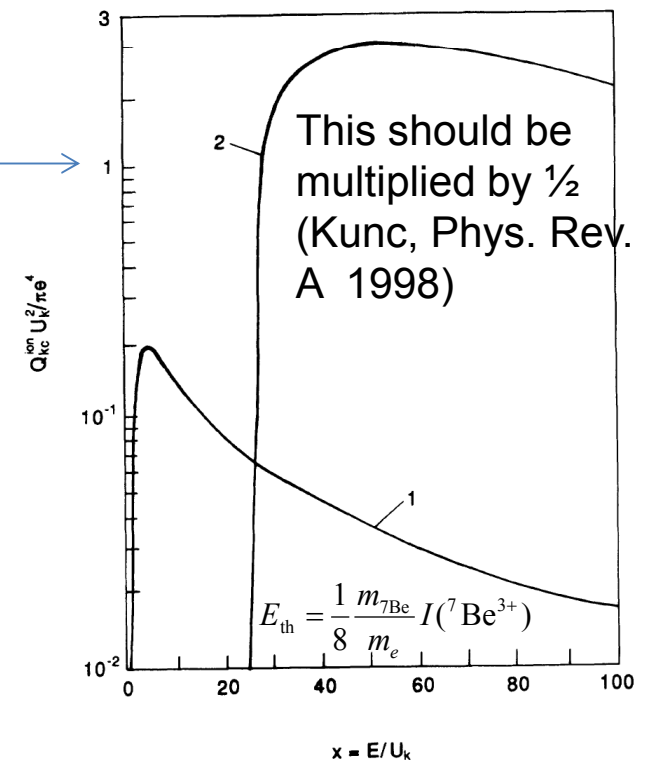
II) destruction

- 1) Collisional ionization: $A_X^* + e^\pm \rightarrow A + X + e^\pm$
- 2) Charge exchange: $A_X^* + e^- \rightarrow A^{(Z-1)+} + X^-$
- 3) Photo-ionization: $A_X^* + \gamma \rightarrow A + X^-$

$$\begin{aligned}\Gamma_{\text{Be}_X^*, \text{de}} &= (n_{e^-} + n_{e^+}) \langle \sigma v \rangle \\ &= 2.79 \times 10^{10} \text{ s}^{-1} \left(\frac{T_9}{0.4} \right)^2 \frac{e^{-m_e/T}}{2.97 \times 10^{-7}} \\ &\quad \times \frac{(1 + E_{\text{th}}/T)}{11.1} \frac{e^{-E_{\text{th}}/T}}{4.13 \times 10^{-5}} \left(\frac{n}{113} \right)^4, \quad (29)\end{aligned}$$

$$4 \times (\text{Atomic radius})^2 \rightarrow \pi [2n^2 / (Z_1 Z_2 \alpha \mu)]^2$$

Kunc, Phys. Rev. A (1994)



GS ${}^7\text{Be}^{3+}$ is only available path

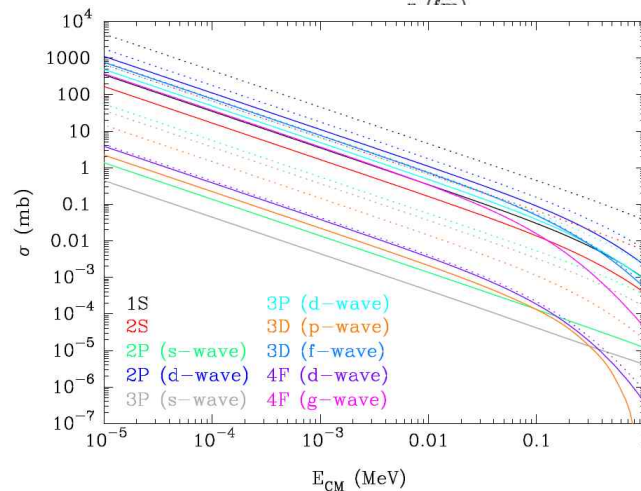
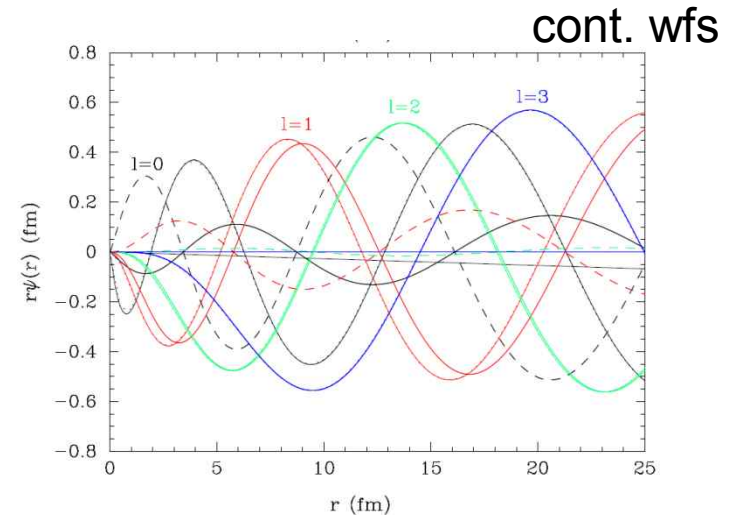
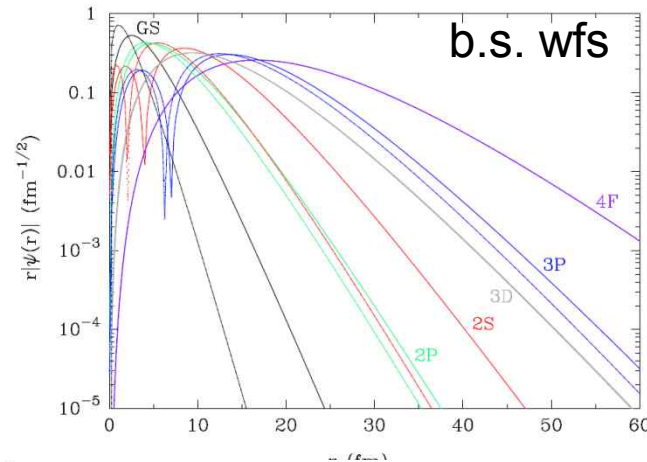
← ESs ${}^7\text{Be}^{3+*}$ are converted to highly ESs ${}^7\text{Be}_x^*$ ($n \gtrsim 226$)

which are predominantly destroyed and do not transit to ${}^7\text{Be}_x$

3. Recombination cross sections

➤ Resonant and nonresonant cross sections for (${}^7\text{Be}$, ${}^7\text{Li}$, ${}^9\text{Be}$, ${}^4\text{He}$)

ex)
 ${}^7\text{Be}_X$ ($m_X=1\text{ TeV}$)



Nonresonant cross section

✓ Finite size of nuclear charge
→ difference from point charge

➤ cross sections for other nuclei are approximately given by those for the point-charge case into ground states.

$$\sigma \approx \frac{2^9 \pi^2 e_1^2}{3 \exp(4)} \frac{E_{\text{bin}}}{\mu^3 v^2}$$

Transition from bound states of the first nuclear excited state of ${}^7\text{Be}^*$ and heavy exotic particle X^- to those of the ground state ${}^7\text{Be}$ and X^-

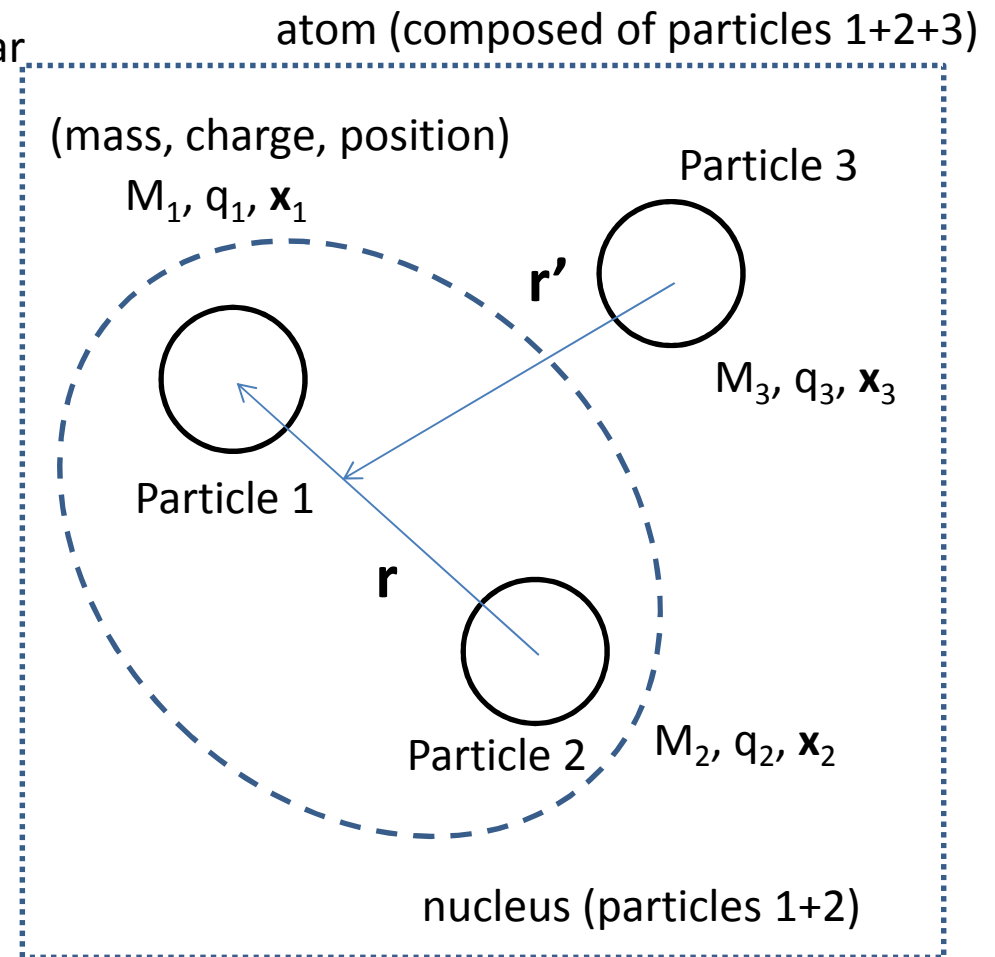
→ ${}^7\text{Be} = \alpha + {}^3\text{He}$: two-body bound state
 α , ${}^3\text{He}$, X^- are called 1, 2, 3, respectively.
 Consider the situation show in the figure.

➤ Assume that the atom can be described as the sum of two wave functions for
 1) nucleus (1+2 two-body system), and
 2) atom ([1+2]+3 two-body system)
 ...Eq. (A. 2)

➤ The electric dipole (E1) moment is written with Jacobi coordinates

$$\mathbf{d} = \sum_{i=1}^3 q_i \mathbf{x}_i = \frac{(q_1 + q_2)M_3 - q_3(M_1 + M_2)}{M_1 + M_2 + M_3} \mathbf{r}' + \frac{M_2 q_1 - M_1 q_3}{M_1 + M_2} \mathbf{r}$$

➤ Matrix element for the E1 transition is derived as integration over $\mathbf{r}$ and $\mathbf{r}'$ of E1 moment multiplied by wave functions of initial and final states. The E1 operator is the sum of a term of $\mathbf{r}$ and $\mathbf{r}'$. By orthogonalities of wave functions, matrix element is zero for a transition to a final state whose nuclear and atomic states are different from those of the initial state



➤ ex) initial state

$$\Psi_{j_1}^{n m_1}(\mathbf{r}) \Psi_{n_i l_i m_i}^a(\mathbf{r}')$$

final state

$$\Psi_{j_2}^{n m_2}(\mathbf{r}) \Psi_{n_f l_f m_f}^a(\mathbf{r}')$$

nuclear w.f.

atomic w.f.

Matrix element for the transition is

$$\begin{aligned} \langle f | O(E1, \mu) | i \rangle &= \int d\mathbf{r} \int d\mathbf{r}' \\ &\times \Psi_{j_2}^{n m_2*}(\mathbf{r}) \Psi_{n_f l_f m_f}^{a*}(\mathbf{r}') \left[\underline{q_{r'} r' Y_{1\mu}(\hat{r}')} + \underline{q_r r Y_{1\mu}(\hat{r})} \right] \Psi_{j_1}^{n m_1}(\mathbf{r}) \Psi_{n_i l_i m_i}^a(\mathbf{r}') \end{aligned}$$

r -independent → out of r-integration
r' -independent → out of r'-integration

$$\begin{aligned} \langle f | O(E1, \mu) | i \rangle &= \underbrace{\langle \Psi_{j_2}^{n m_2} | \Psi_{j_1}^{n m_1} \rangle}_{\text{Overlap of nuclear w.fs.}} \int d\mathbf{r}' \Psi_{n_f l_f m_f}^{a*}(\mathbf{r}') \left[q_{r'} r' Y_{1\mu}(\hat{r}') \right] \Psi_{n_i l_i m_i}^a(\mathbf{r}') \\ &\quad + \underbrace{\langle \Psi_{n_f l_f m_f}^a | \Psi_{n_i l_i m_i}^a \rangle}_{\text{Overlap of atomic w.fs.}} \int d\mathbf{r} \Psi_{j_2}^{n m_2*}(\mathbf{r}) \left[q_r r Y_{1\mu}(\hat{r}) \right] \Psi_{j_1}^{n m_1}(\mathbf{r}) \end{aligned}$$

Overlap of nuclear w.fs. = $\begin{cases} 1 & \text{(if the initial and final nuclear states are identical)} \\ 0 & \text{(otherwise)} \end{cases}$
 Overlap of atomic w.fs. = $\begin{cases} 1 & \text{(if the initial and final atomic states are identical)} \\ 0 & \text{(otherwise)} \end{cases}$

➤ Then, E1 matrix element is zero for a transition to a final state whose atomic and nuclear states are different from those of the initial state.

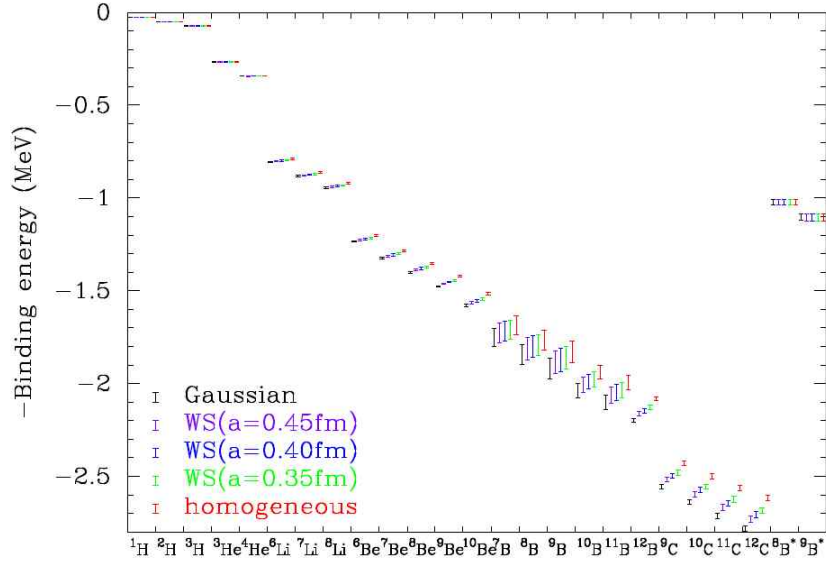


FIG. 1: Binding energies of nuclei and X^- particle with $m_X = 100$ TeV for different charge distributions of Gaussian (black lines), Woods-Saxon type with diffuseness parameters $a = 0.45$ fm (purple lines), 0.40 fm (blue lines), and 0.35 fm (green lines), and homogeneous well (red lines). Error bars indicate uncertainties determined from uncertainties in experimental **RMS** charge radii.

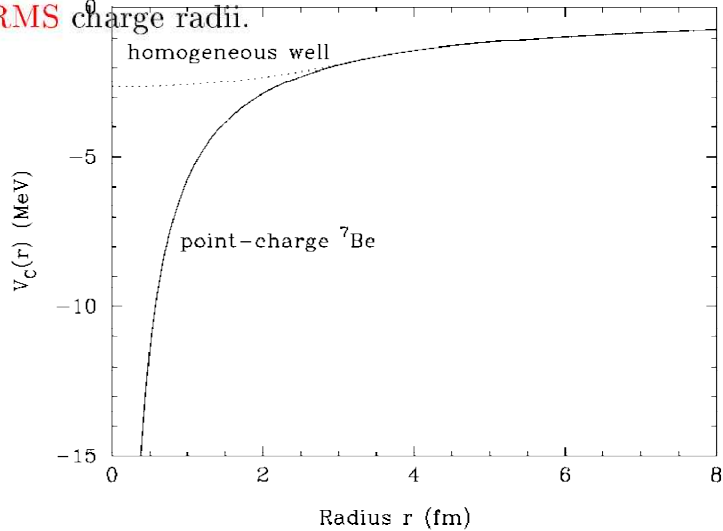
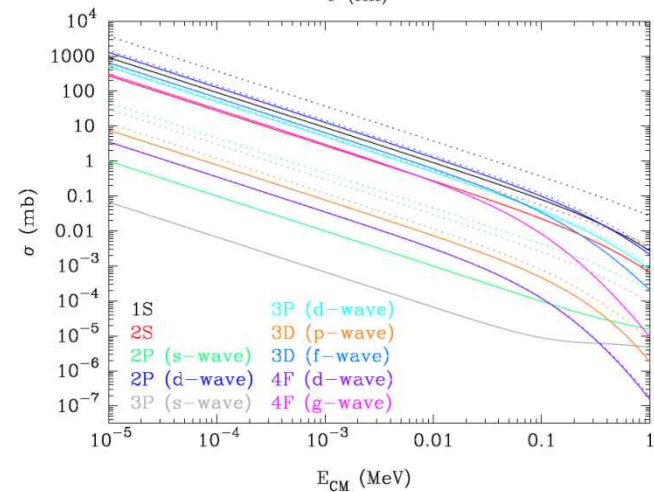
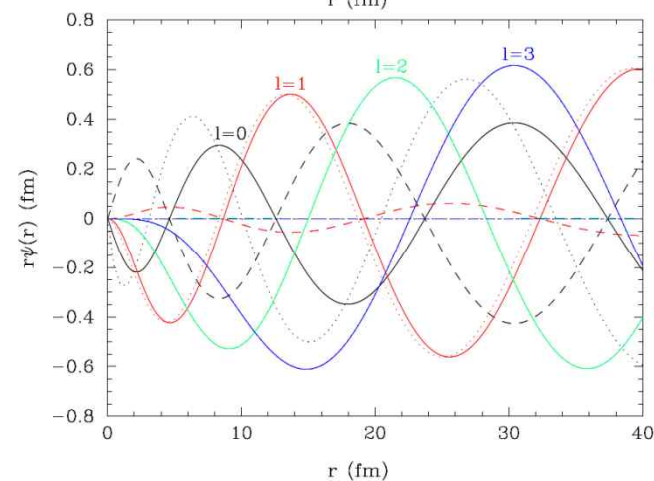
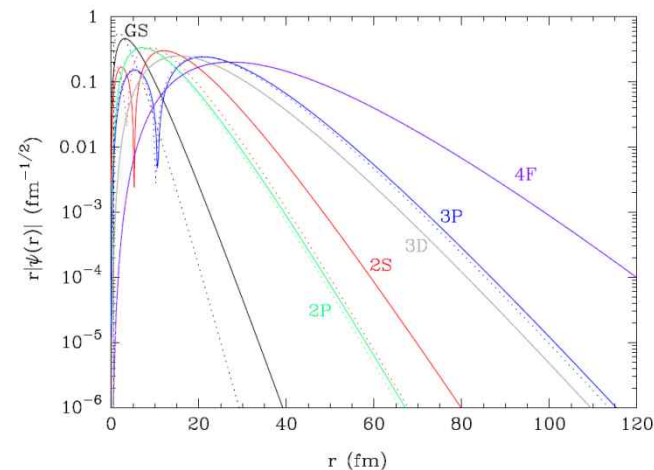
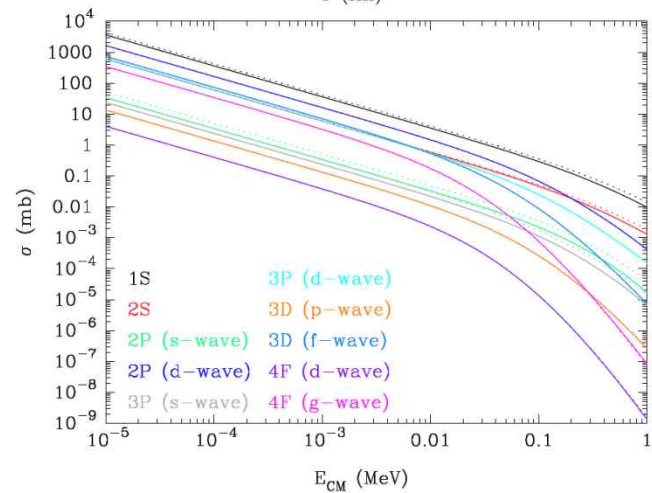
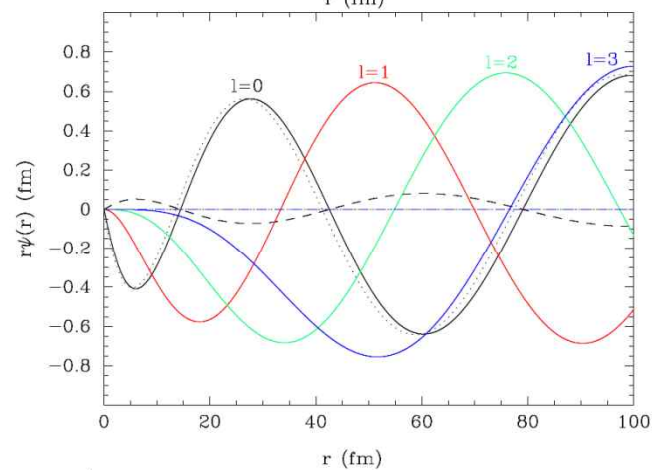
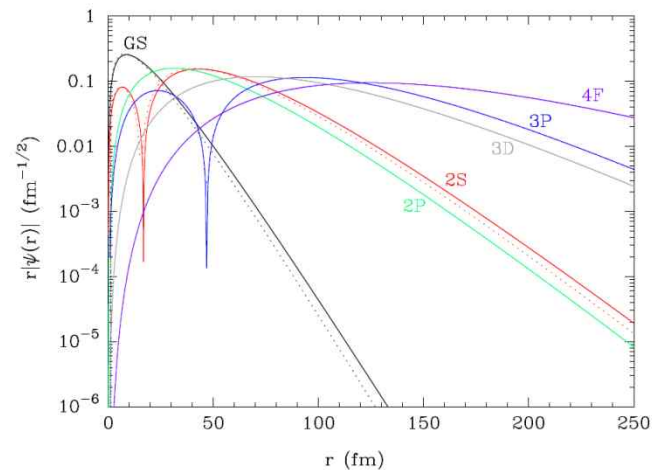


FIG. 2: Binding energies and resonance energies as a function of m_X . The upper black lines show resonance energies in the reactions ${}^7\text{Be}_X(p, \gamma){}^8\text{B}_X$ and ${}^8\text{Be}_X(p, \gamma){}^9\text{B}_X$. The lower lines show binding energies of ${}^7\text{Be}_X$ (black lines), ${}^8\text{Be}_X$ (purple lines), ${}^8\text{B}_X$ (green lines), ${}^9\text{B}_X$ (gray lines), and the first atomic excited states ${}^8\text{B}_X^{*a}$ (red lines) and ${}^9\text{B}_X^{*a}$ (blue lines). Results for different nuclear charge distributions of Gaussian (dashed lines), Woods-Saxon type with diffuseness parameter $a = 0.40$ fm (solid lines), and homogeneous well (dot-dashed lines) are drawn. Open circles show energy heights derived by the quantum many-body calculation for $m_X = \infty$ [39].

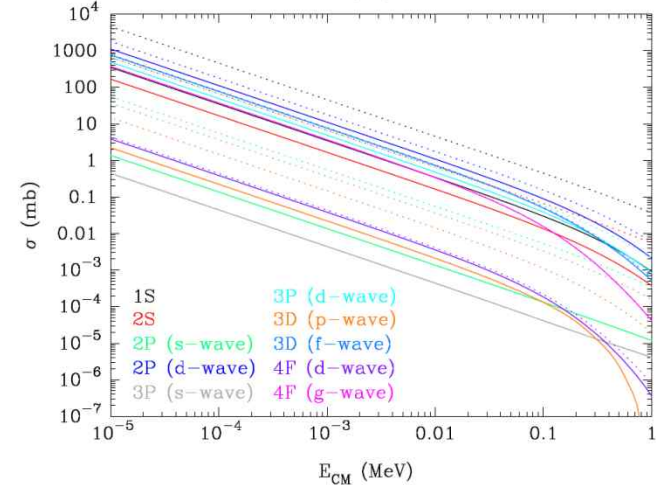
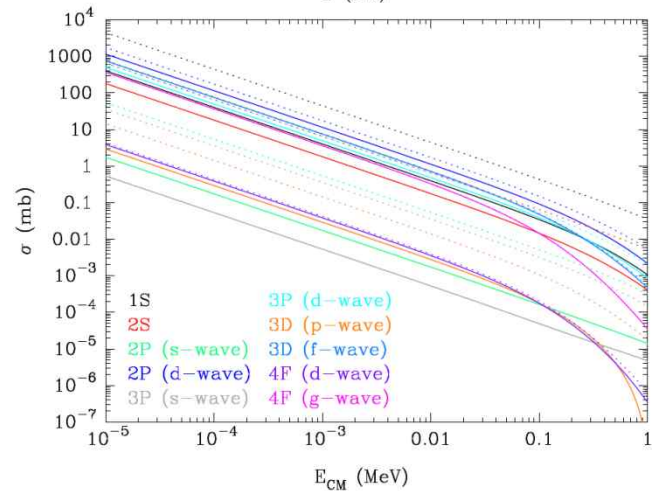
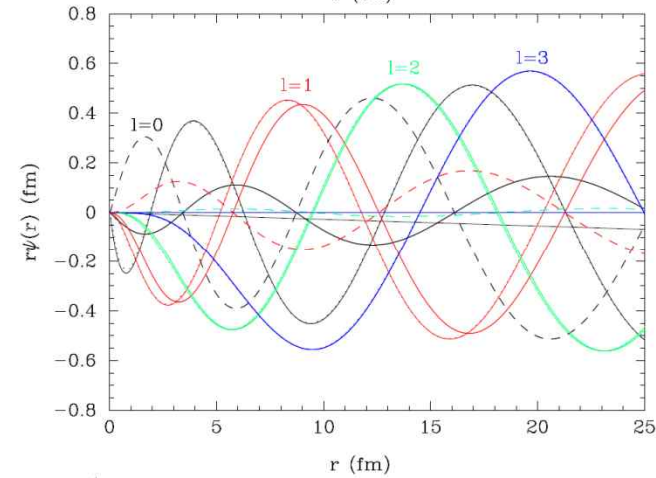
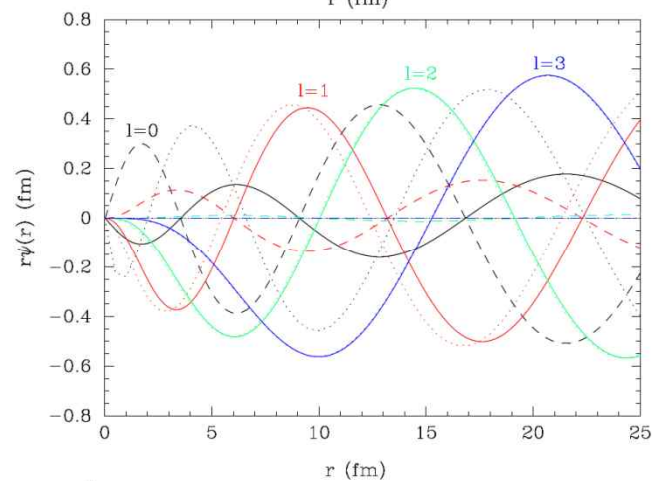
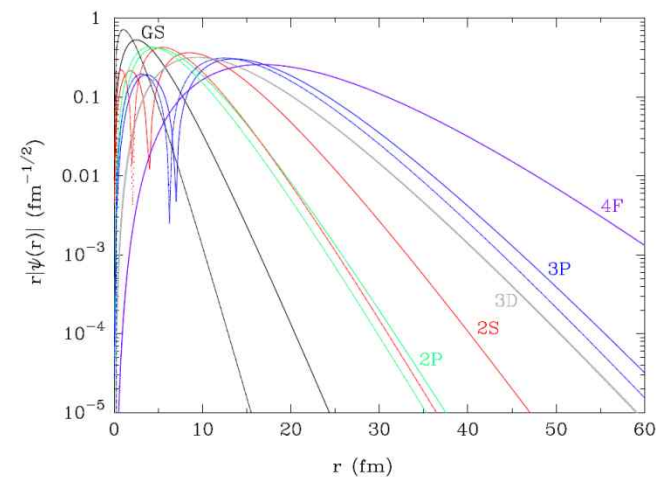
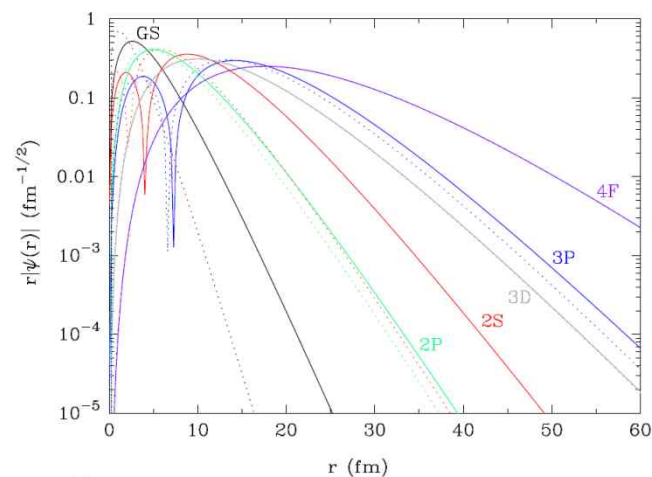
${}^7\text{Be}_x$ recombination

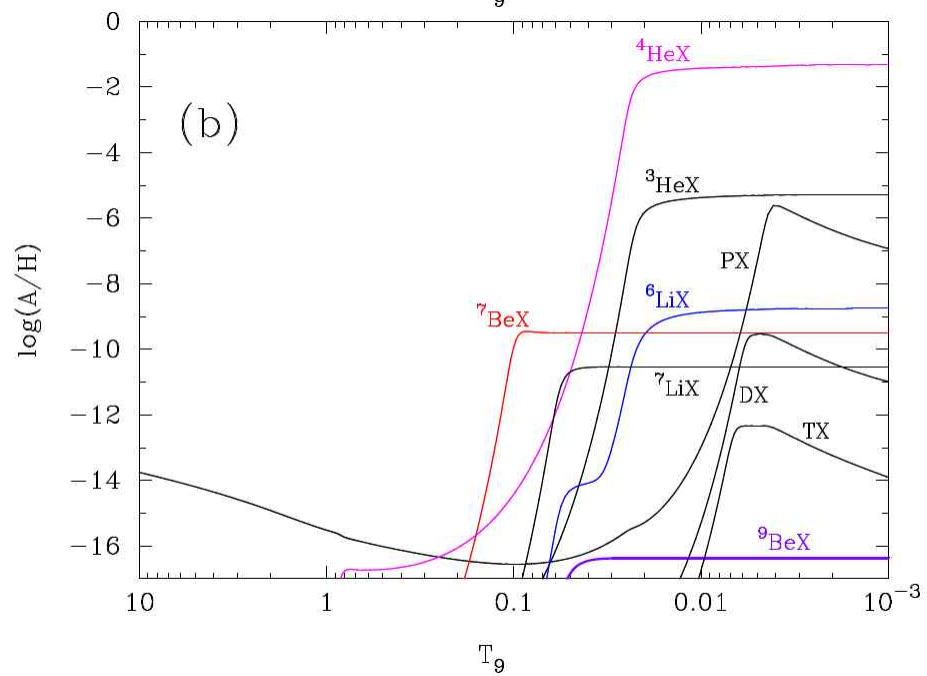
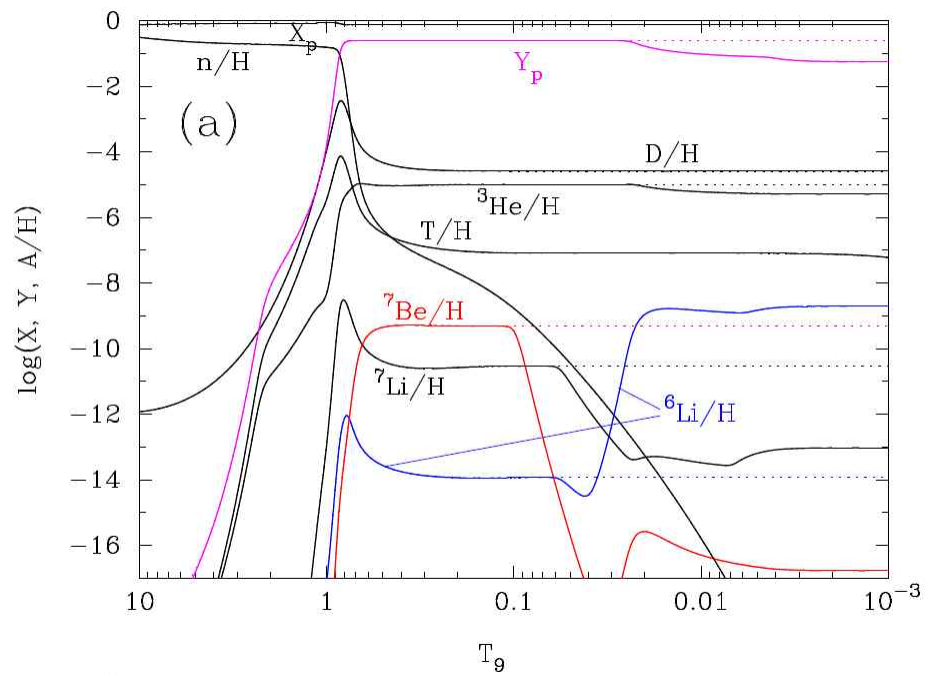
$m_X = 1 \text{ GeV}$,
10 GeV



${}^7\text{Be}_X$ recombination

$m_X = 100 \text{ GeV},$
 1000 GeV





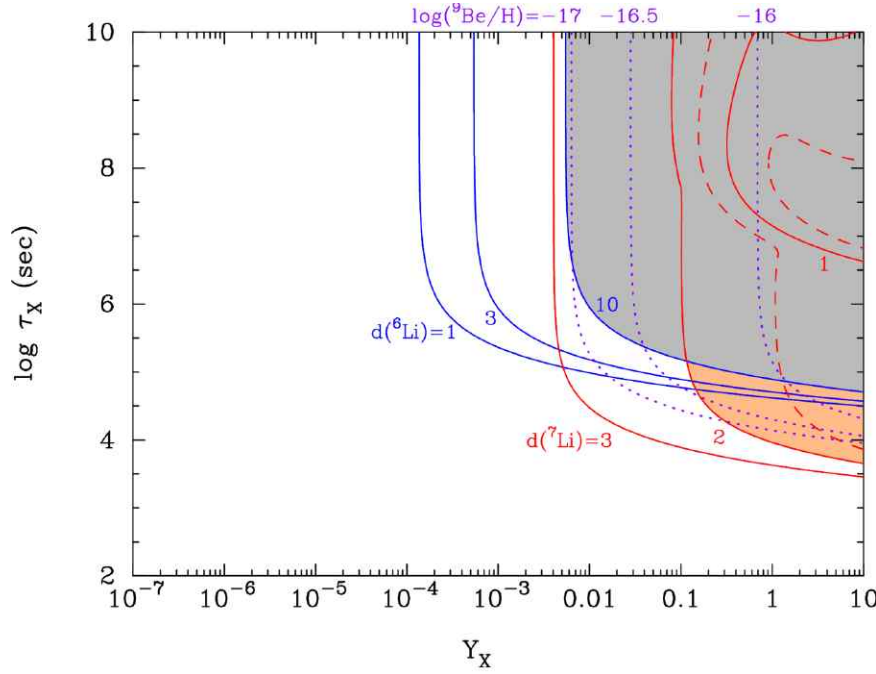


FIG. 23: Contours of constant lithium abundances relative to observed values in MPSs, i.e., $d(^6\text{Li}) = ^6\text{Li}^{\text{Cal}}/^6\text{Li}^{\text{Obs}}$ (blue lines) and $d(^7\text{Li}) = ^7\text{Li}^{\text{Cal}}/^7\text{Li}^{\text{Obs}}$ (red lines) for the case of $m_X = 1$ GeV. The adopted observational constraint on the ^7Li abundance is the center value of $\log(^7\text{Li}/\text{H}) = -12 + (2.199 \pm 0.086)$ derived in a 3D NLTE model [9], while that of ^6Li is taken from the two sigma upper limit of G64-12 (NLTE model with 5 parameters), $^6\text{Li}/\text{H} = (0.9 \pm 4.3) \times 10^{-12}$ [33]. Dashed lines around the line of $d(^7\text{Li}) = 1$ correspond to the 2 sigma uncertainty in the observational constraint. The gray region which locates right from contours of $d(^6\text{Li}) = 10$ or the 2 sigma lower limit, $d(^7\text{Li}) = 0.67$, are excluded by overproduction of ^6Li and underproduction of ^7Li , respectively. The orange region is interesting parameter region in which a significant ^7Li reduction realizes without an overproduction of ^6Li . Purple lines are contours of the abundance ratio $^9\text{Be}/\text{H}$ predicted when the unknown rate of the reaction $^7\text{Li}_X(d, X^-)^9\text{Be}$ is assumed as described in text.

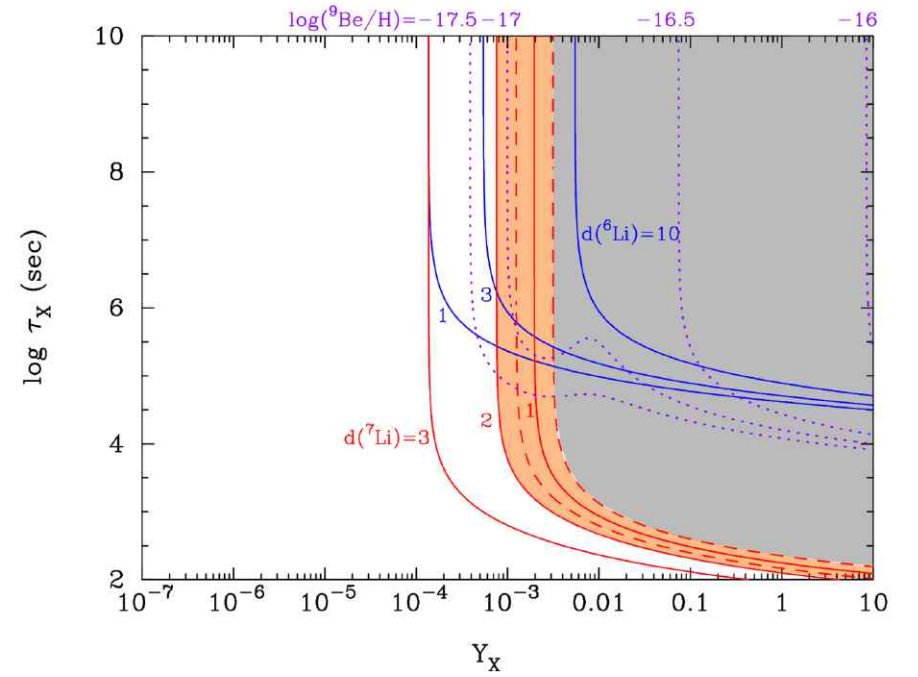
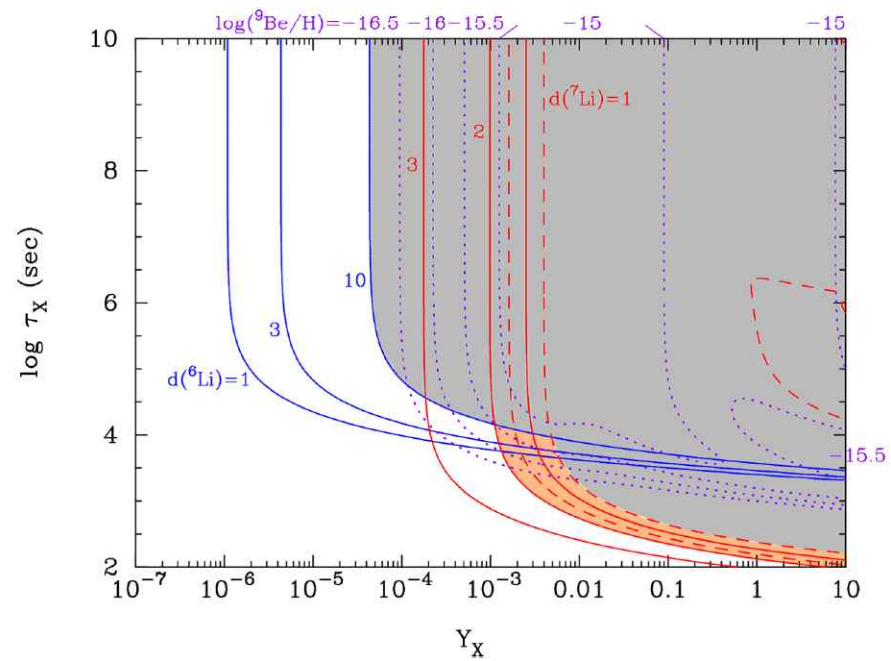
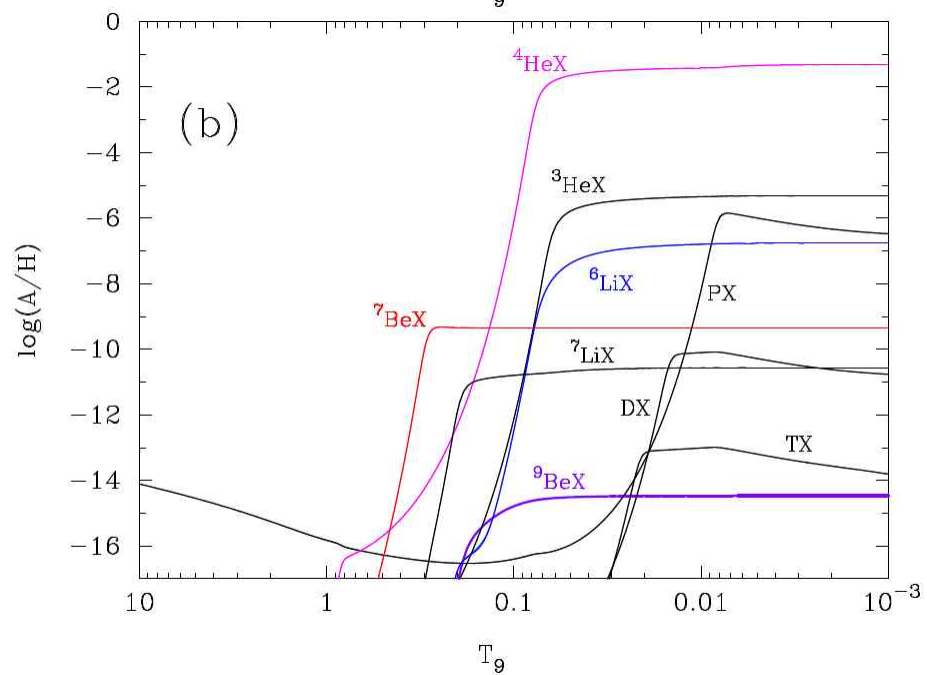
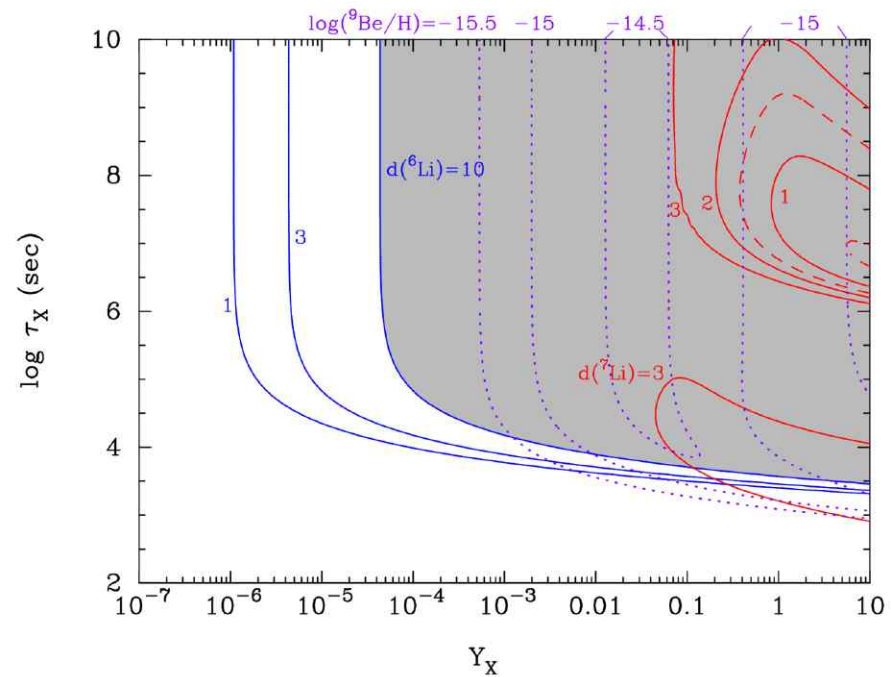
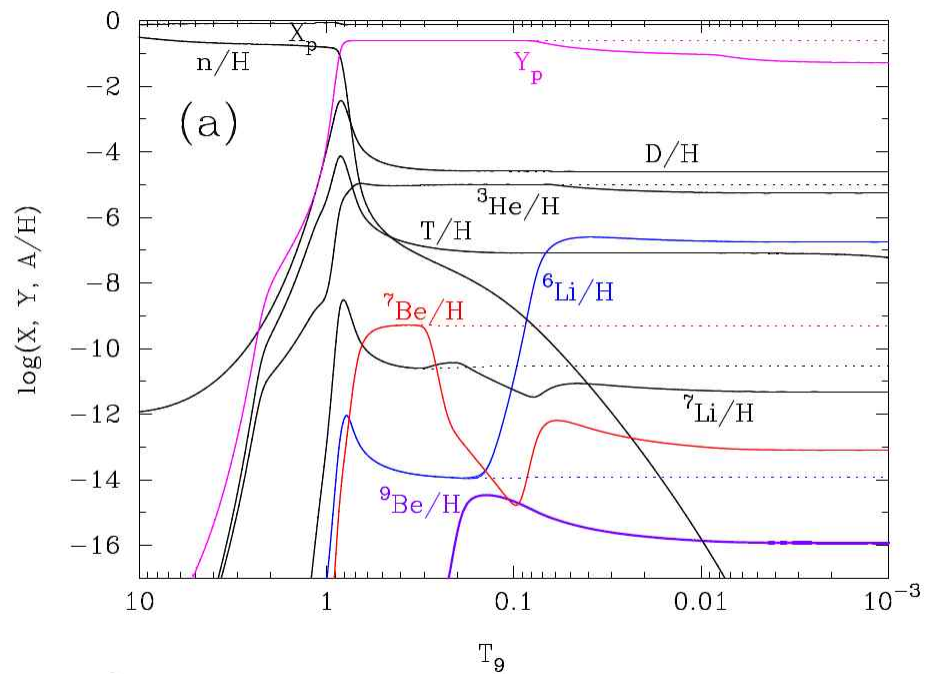
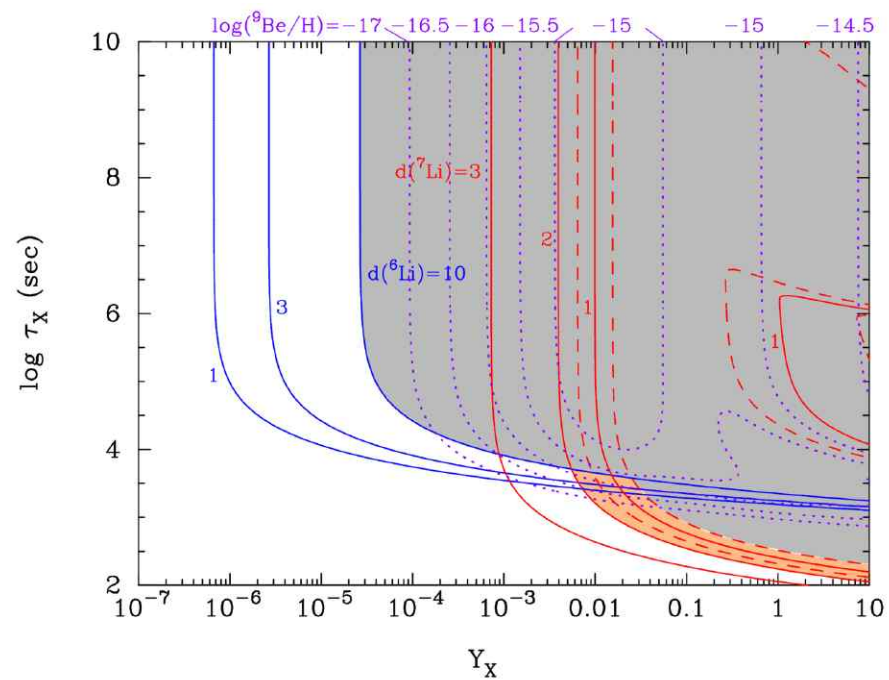
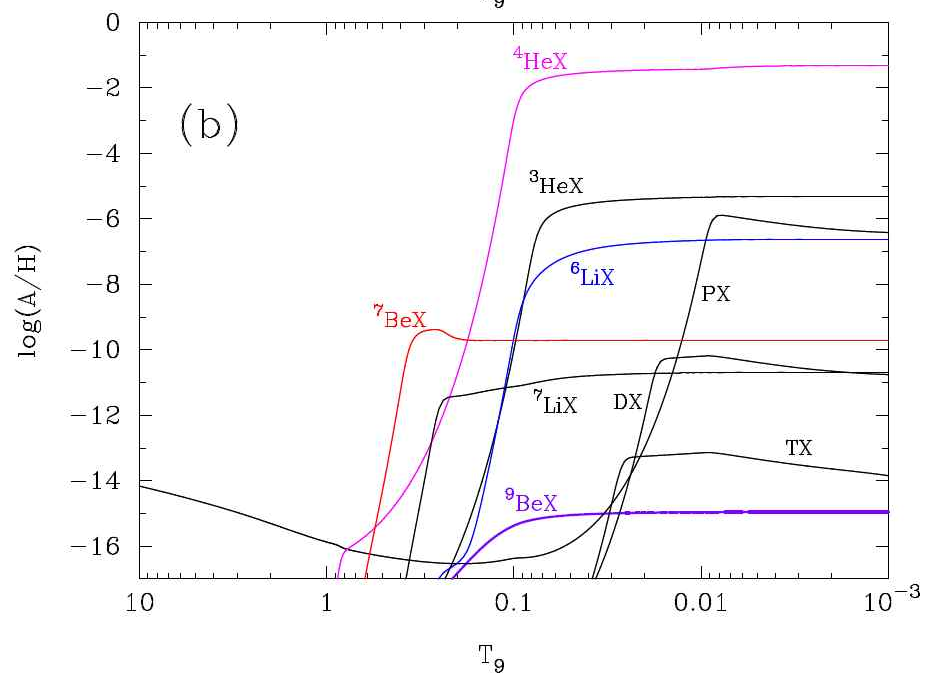
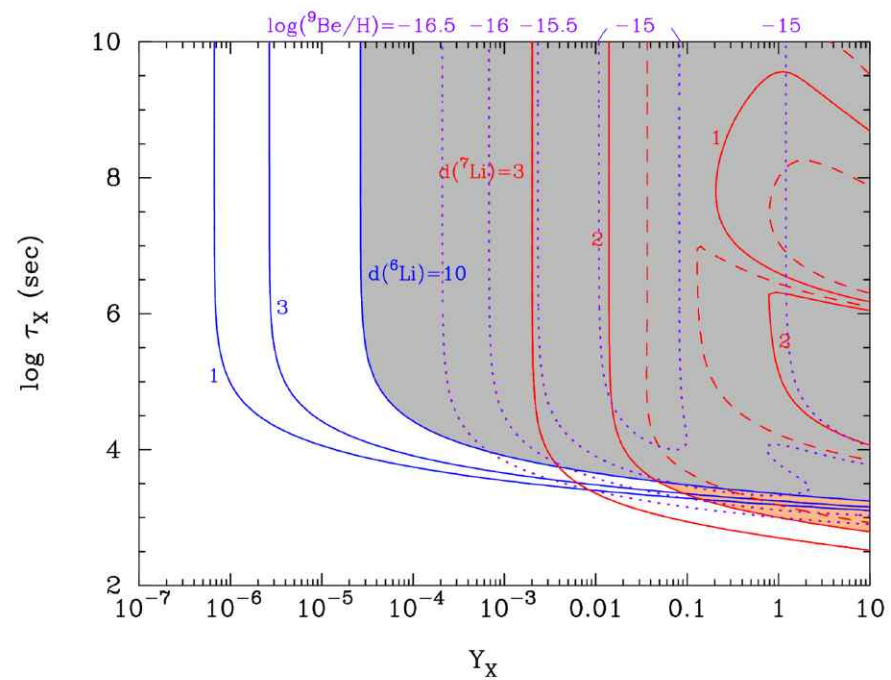
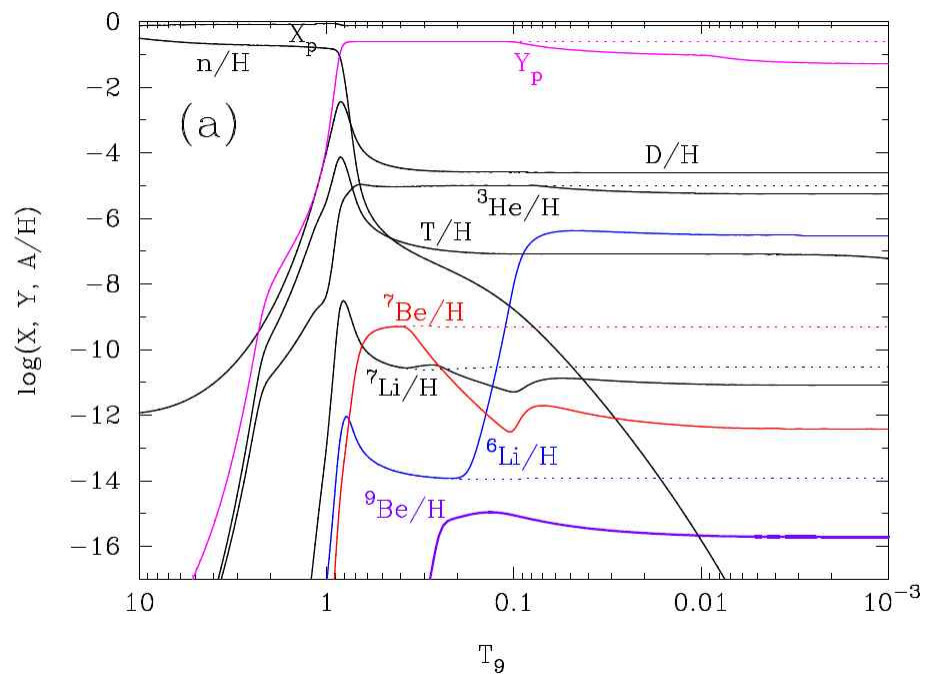
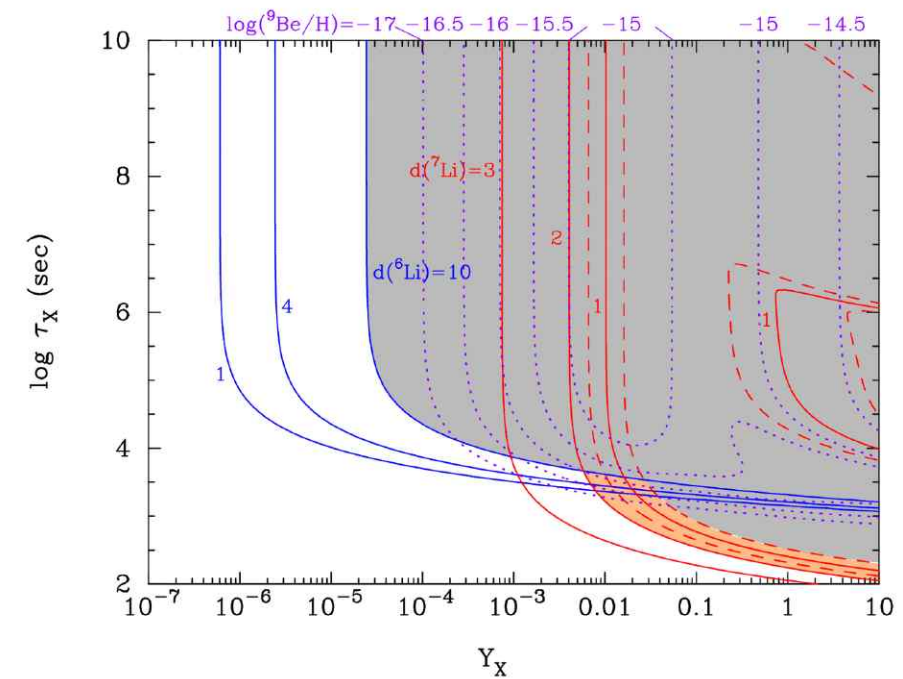
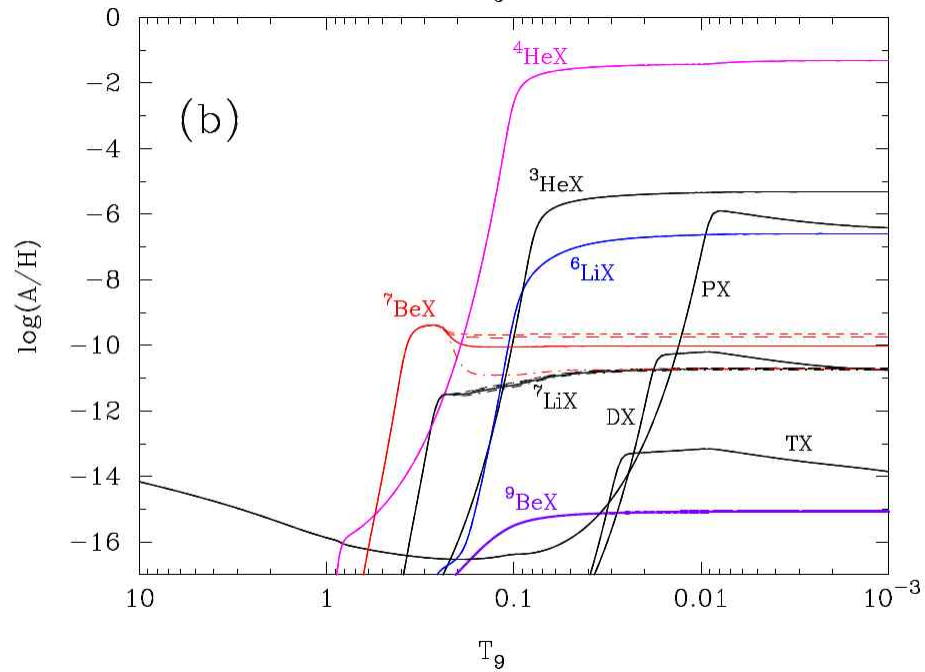
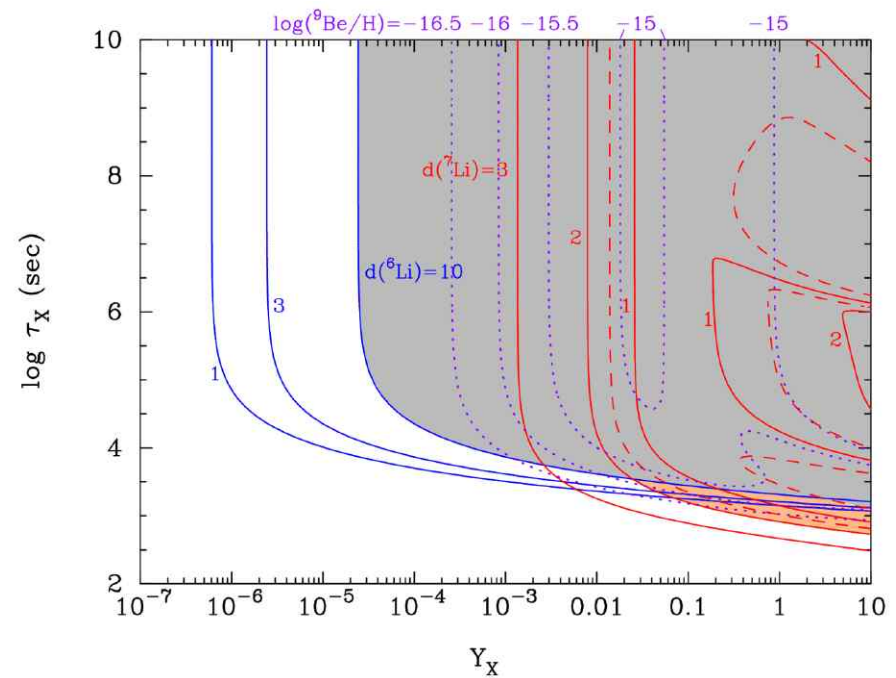
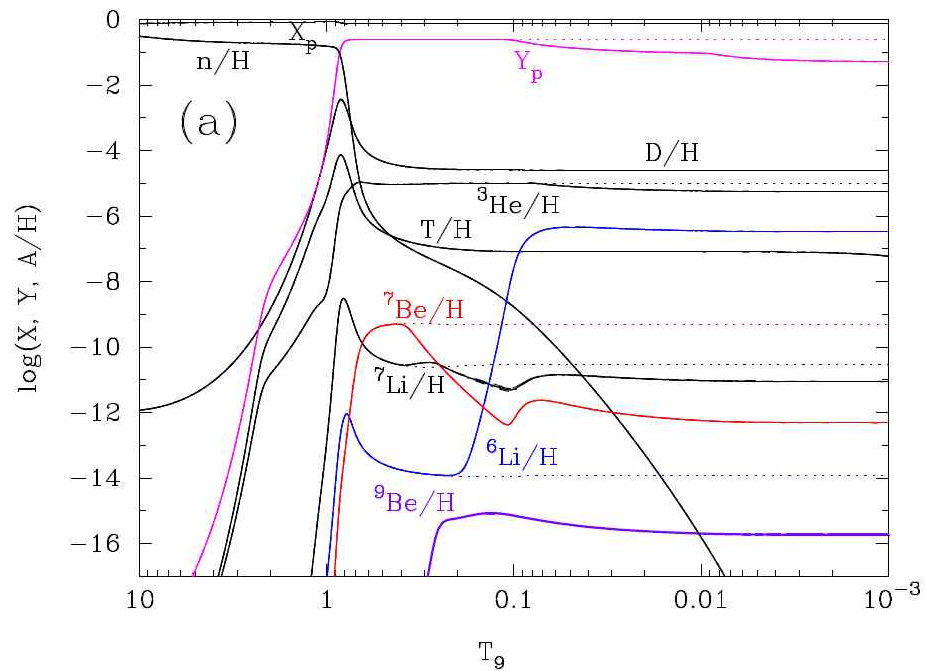


FIG. 24: Same as in Fig. 23 when the charged-current decay of $^7\text{Be}_X \rightarrow ^7\text{Li} + X^0$ is included.







Processes affecting elemental abundances

	Model	${}^7\text{Li}$ problem solved ?	Signatures on other nuclides ?
Existence of particle [$z \sim 10^9$]	sub-SIMP X^0	✓ [1]	${}^6\text{Li}$ [2], ${}^9\text{Be}$ [1]
	SIMP X^0 [3]	no	${}^9\text{Be}$ and/or ${}^{10}\text{B}$
	CHAMP $X^- *$	✓ [4,5,6]	${}^6\text{Li}$ [7], ${}^9\text{Be}$ [6]
Early stars [$z \sim \mathcal{O}(10)$]	Early cosmic ray	no	${}^6\text{Li}$ [8], ${}^9\text{Be}$ & ${}^{10,11}\text{B}$ [9, 10]

[1] Kawasaki, MK (2011)

[2] MK, Kawasaki (2012)

[3] MK, Kajino, Yoshida, Mathews (2009)

[4] Bird, Koopmans, Pospelov (2008) [stronger]

[5] MK, Kajino, Boyd, Yoshida, Mathews (2007)
[weaker]

[6] MK, Kim, Cheoun, Kajino, Mathews
(in preparation)

[7] Pospelov (2007)

[8] Rollinde, Vangioni, Olive (2006)

[9] Rollinde, Maurin, Vangioni, Olive, Inoue (2008);
MK (2008)

[10] MK, Kawasaki (2013)