QCD under a strong magnetic field

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KIAS , June 12, 2014 (Based on DKH 98, 2011, 2014)

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Motivations

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Magnetic field is relevant in QCD if strong enough:

 $|\textit{eB}| \gtrsim \Lambda_{\rm QCD}^2 \approx 10^{19}\,{\rm Gauss}\cdot\textit{e}.$

▶ Some neutron stars, called magnetars, have magnetic fields at the surface, $B \sim 10^{12-15}$ G (Magnetar SGR 1900+14):



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Motivations

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In the peripheral collisions of relativistic heavy ions huge magnetic fields are produced at the center:



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Contents

- Vector Condensation (DKH, work under progress)
- Neutron star cooling (DKH, 1998)
- Chiral Magnetic Effect (DKH, 2011)

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Vector condensation Neutron star Chiral Magnetic effect

Vector condensation

▶ The energy spectrum of (elementary) charged particle under the magnetic field $(\vec{B} = B\hat{z})$:

$$E(\vec{p}) = \pm \sqrt{p_z^2 + m^2 + n|qB|},$$

where $n = 2n_r + |m_L| + 1 - \text{sign}(qB)(m_L + 2s_z)$.

At the lowest Landau level the spin of the rho meson is along the B field direction and n = -1. If elementary,

 $m_\rho^2(B) = m_\rho^2 - |eB|.$

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Vector condensation Neutron star Chiral Magnetic effect

Vector condensation

Vector meson condensation: Vector order parameter develops under strong magnetic field (Chernodub 2011):

 $\langle \bar{u}\gamma_1 d \rangle = -i \langle \bar{u}\gamma_2 d \rangle = \rho(x_\perp).$



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Vector condensation

 Lattice calculation shows vector meson becomes lighter under the B field (Luschevskaya and Larina 2012):



Image: A matrix and a matrix

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Vector condensation

• Lattice calculation shows vector meson condensation at $B > B_c = 0.93 \text{GeV}^2/e$ (Barguta et al 1104.3767):



Vector condensation Neutron star Chiral Magnetic effect

Effective Lagrangian (DKH98 & 2014):

Quarks under strong B field occupy Landau levels:



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Vector condensation Neutron star Chiral Magnetic effect

Effective Lagrangian (DKH98 & 2014):

Quark propagator under B field is given as

$$S_{F}(x) = \sum_{n=0}^{\infty} (-1)^{n} \int_{k} e^{-ik \cdot x} e^{-k_{\perp}^{2}/|qB|} S_{n}(qB, k)$$
$$S_{n}(qB, k) = \frac{D_{n}(qB, k)}{\left[(1+i\epsilon)k_{0}\right]^{2} - k_{z}^{2} - 2|qB|n}$$

$$D_{n} = 2\tilde{k}_{\parallel} \left[P_{-}L_{n} \left(\frac{2\kappa_{\perp}}{|qB|} \right) - P_{+}L_{n-1} \left(\frac{2\kappa_{\perp}}{|qB|} \right) \right] + 4k_{\perp}L_{n-1}^{1} \left(\frac{2\kappa_{\perp}}{|qB|} \right)$$

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Vector condensation Neutron star Chiral Magnetic effect

Matching with QCD at Λ_L :

At low energy E < Λ_L we integrate out the modes in the higher Landau levels (n ≠ 0).

A new quark-gluon coupling:



$$\mathcal{L}_2 = c_2 rac{i g_s^2}{|qB|} ar{Q}_0 \, A \hspace{-0.5mm}/ \tilde{\gamma}_\mu \cdot \partial_\mu \, A \hspace{-0.5mm}/ \tilde{\gamma}_\mu Q_0 \, .$$

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Effective Lagrangian (DKH98):

Four-Fermi couplings for LLL quarks:



$$\mathcal{L}_{\text{eff}}^{1} \ni \frac{g_{1}^{s}}{4 |qB|} \left[\left(\bar{Q}_{0} Q_{0} \right)^{2} + \left(\bar{Q}_{0} i \gamma_{5} Q_{0} \right)^{2} \right]$$

Below Λ_L the quark-loop does not contribute to the beta-function of α_s: At one-loop

$$\frac{1}{\alpha_s(\mu)} = \frac{1}{\alpha_s(\Lambda_L)} + \frac{11}{2\pi} \ln\left(\frac{\mu}{\Lambda_L}\right) \,.$$

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Vector condensation Neutron star Chiral Magnetic effect

Effective Lagrangian (DKH98):

One-loop RGE for the four-quark interaction:

$$\mu \frac{d}{d\mu} g_1^s = -\frac{40}{9} \alpha_s^2 \; (\ln 2)^2$$

Solving RGE to get

 $g_1^s(\mu) = 1.1424 \left(lpha_s(\mu) - lpha_s(\Lambda_L) \right) + g_1^s(\Lambda_L)$.

If B ≥ 10²⁰ G, the four-quark interaction is stronger than gluon interaction. Therefore the chiral symmetry should break at a scale higher than the confinement scale for B ≥ 10²⁰ G.

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▶ If $B \ge 10^{20}$ G, the four-quark interaction is stronger than gluon interaction. Therefore the chiral symmetry should break at a scale higher than the confinement scale for $B \ge 10^{20}$ G.

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Vector condensation Neutron star Chiral Magnetic effect

Vector mesons in the Effective Lagrangian:

Running coupling under strong B field:



Vector condensation Neutron star Chiral Magnetic effect

Vector mesons in the Effective Lagrangian:

We need a stronger B field (B > m²_ρ/e) to condense vector mesons:

$$m_
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The critical B field occurs at (DKH 2014)

$$eB_c = m_
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▶ It agrees well with the lattice result by Barguta et al, $B_c = 0.93 \,\text{GeV}^2/e.$

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Vector condensation Neutron star Chiral Magnetic effect

QCD Vacuum Energy:

 The additional vacuum energy at one-loop is given by Schwinger as

$$\Delta \mathcal{E}_{\mathsf{vac}} = -rac{1}{16\pi^2} \int_0^\infty rac{ds}{s^3} e^{-M_\pi^2 s} \left[rac{eBs}{\mathsf{sinh}(eBs)} - 1
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 The chiral condensate becomes by the Gell-Mann-Oakes-Renner relation (Shushpanov+Smilga '97)

$$\langle \bar{q}q \rangle^B = \langle \bar{q}q \rangle^{B=0} \left(1 + \frac{|eB| \ln 2}{16\pi^2 F_\pi^2} + \cdots \right)$$

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Vector condensation Neutron star Chiral Magnetic effect

Neutron star cooling:

Cooling through axion bremsstrahlung (DKH98)



 Energy loss per unit volume per unit time at low temp. (Iwamoto, Ellis, Brinkman+Turner)

$$\mathcal{Q}_{a}^{1}\propto \left(\frac{f}{M_{\pi}}\right)^{4}m_{n}^{2.5}g_{an}^{2}T^{6.5}$$

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Neutron star cooling:

► The axion coupling $g_{an} \propto m_n / f_{PQ}$, $m_n \propto \langle \bar{q}q \rangle$, and since f_{PQ} does not get any corrections from the magnetic field, we have

$$g_{an}(B) = g_{an}(0) \left(1 + \frac{|eB| \ln 2}{16\pi^2 F_\pi^2} + \cdots\right).$$

 The correction then becomes (DKH '98), using the Goldberg-Treiman relation and GOR relation,

$$\frac{\mathcal{Q}_{a}^{1}(B)}{\mathcal{Q}_{a}^{1}(0)} \simeq \left(1 + \frac{g_{1}^{s}}{2\pi^{2}}\right)^{2} \left(1 + \frac{|eB|\ln 2}{16\pi^{2}F_{\pi}^{2}}\right)^{6.5}$$

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Neutron star cooling:

Similarly, the lowest-order energy emission rate per unit volume by the pion-axion conversion π[−] + p → n + a (Schramm) has corrections (DKH98):

$$\frac{\mathcal{Q}_{a}^{\pi^{-}}(B)}{\mathcal{Q}_{a}^{\pi^{-}}(0)} \simeq \left(1 + \frac{g_{1}^{s}}{2\pi^{2}}\right)^{2} \left(1 + \frac{|eB|\ln 2}{16\pi^{2}F_{\pi}^{2}}\right)^{-1}$$

Order of magnitude enhancement for B = 10²⁰ G. Magnetars cool quickly!

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Chiral Magnetic Effects (DKH2011)

An instanton number may be created in RHIC:

$$n_w = \frac{g_s^2}{32\pi^2} \int \mathrm{d}^4 x \ G^a \tilde{G}^a = n_L - n_R \,.$$

whose conjugate variable is μ_A .

 Chiral magnetic effect (Fukushima+Kharzeev+Warringa): Under a strong magnetic field

$$\vec{J}=rac{q^2}{2\pi^2}\mu_A\vec{B}$$
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- Off-center collision produces strong B field:



 Such CP-odd effect can be then observable (Fukushima+ Kharzeev+Warringa 2008):

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CME seen at RHIC? (STAR prl '09)



- May be (see talk by Bzdak at HIC10)
- May be not. (Muller+Schafer arXiv:1009.1053)

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Vector condensation Neutron star Chiral Magnetic effect

Quark matter under strong B field:

Under strong B field quark spectrum takes:

$$E_A = -\mu \pm \sqrt{k_z^2 + 2\left|qB\right|n}.$$

Quark propagator under B field is given as

$$S_{F}(x) = \sum_{n=0}^{\infty} (-1)^{n} \int_{k} e^{-ik \cdot x} e^{-k_{\perp}^{2}/|qB|} S_{n}(qB, k)$$

$$S_{n}(qB, k) = \frac{D_{n}(qB, k)}{[(1+i\epsilon)k_{0}+\mu]^{2}-k_{z}^{2}-2|qB|n}$$

$$D_{n} = 2\tilde{k}_{\parallel} \left[P_{-}L_{n} \left(\frac{2k_{\perp}^{2}}{|qB|} \right) - P_{+}L_{n-1} \left(\frac{2k_{\perp}^{2}}{|qB|} \right) \right] + 4k_{\perp}L_{n-1}^{1} \left(\frac{2k_{\perp}^{2}}{|qB|} \right) ,$$

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Vector condensation Neutron star Chiral Magnetic effect

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Anomalous current for LH fermions at one-loop

$$\Delta_{L}^{\alpha}(\mu_{L}) \equiv \left\langle \bar{\psi}_{L} \gamma^{\alpha} \psi_{L} \right\rangle = -\int \frac{\mathrm{d}^{4}k}{(2\pi)^{4}} \mathrm{Tr}\left(\gamma^{\alpha} \tilde{S}(k)_{L} \right) \,.$$

Matter-dependent part is finite and explicitly calculable:

We change variables:

$$k^{lpha} \longrightarrow k'^{lpha} = k^{lpha} + u^{lpha} \mu, \quad u^{lpha} = (1, \vec{0})$$

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Differentiating with respect to the chemical potential, we get

$$\frac{\partial}{\partial \mu}\tilde{S} = ie^{-\frac{k_{\perp}^{2}}{|qB|}} \sum_{n=0}^{\infty} (-1)^{n} D_{n} 2\pi i \,\delta\left(k_{\parallel}^{2} - \Lambda_{n}\right) \cdot \delta(k^{0} - \mu)$$

• Integrating over k_{\perp} , we get

$$\Delta_{\mathrm{mat}}^{\alpha} = |qB| \left[\Gamma_{L}^{\alpha\beta} I_{\beta}^{(0)} + 2g_{\mu}^{\alpha\beta} \sum_{n=1} I_{\beta}^{(n)} \right] \,,$$

where $\Gamma_L^{lphaeta}=\epsilon^{lphaeta12}\operatorname{sign}(qB)+g_{\scriptscriptstyle ||}^{lphaeta}$ and

$$I^{(n)\beta} = \int_0^{\mu} \mathrm{d}\mu' \int_{k_{\mathrm{H}}} k_{\mathrm{H}}^{\beta} \delta\left(k_{\mathrm{H}}^2 - \Lambda_n\right) \cdot \delta\left(k^0 - \mu'\right) = \frac{p_F^{(n)}}{4\pi^2} \delta^{\beta 0} \,.$$

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The Fermi momentum at the *n*-th Landau level:



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The density of states

$$n_L = \frac{|qB|}{4\pi} \cdot \sum_n \frac{p_F^{(n)}(\mu_L, B)}{\pi}.$$

As Δ^α(0, B) = 0, the anomalous electric (axial) vector currents become

$$J_{V}^{\alpha} \equiv q \left(\Delta_{L}^{\alpha} + \Delta_{R}^{\alpha} \right) = \delta^{\alpha 3} \frac{q^{2}B}{2\pi^{2}} \mu_{A} + \delta^{\alpha 0} q n,$$

$$J_{A}^{\alpha} \equiv q \left(\Delta_{L}^{\alpha} - \Delta_{R}^{\alpha} \right) = \delta^{\alpha 3} \frac{q^{2}B}{2\pi^{2}} \mu + \delta^{\alpha 0} q n_{A}.$$

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No more corrections

Full contributions to the anomalous current:

$$\langle J^{\alpha} \rangle = \left. \frac{\delta \Gamma_{\mathrm{mat}}(A, G; \mu)}{\delta A_{\alpha}} \right|_{A=0=G}$$

The full effective action is obtained by two steps:



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Anomalous Currents

Matter contribution to the anomalous current:

$$\langle J^{\alpha} \rangle_{\mathrm{mat}} = \frac{\delta \Gamma_{\mathrm{mat}}(\mathcal{A})}{\delta \mathcal{A}_{\alpha}} = \frac{\delta}{\delta \mathcal{A}_{\alpha}} \int_{0}^{\mu} \mathrm{d}\mu' \, \frac{\partial}{\partial \mu'} \Gamma(\mathcal{A};\mu') \,,$$

When derivative acts on the loop, we get vertex correction:



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Vector condensation Neutron star Chiral Magnetic effect

Anomalous Currents

- Because the electric charge is not renormalized (Ademollo-Gatto theorem), the vertex correction should vanish.
- Furthermore the density of states is also not subject to corrections due to interaction. (Luttinger theorem)
- ► The one-loop result is exact.

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- Magnetic field is relevant in QCD if $B \ge 10^{19} \, \mathrm{G}$.
- We derive an effective theory for LLL QCD, which has a new marginal four-quark interactions.
- Scale separation between chiral symmetry breaking and confinement.
- LLL quarks are one dimensional and does not contribute to running QCD coupling.
- Condensation along the vector channel occurs, when

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- The vector condensation is being calculated in the effective theory (DKH2014).
- ▶ We show that magnetars cool too quickly by emitting axions if B > 10²⁰ G.
- We calculate the spontaneous generation of anomalous current of dense quark matter under the magnetic field.
- The one-loop is shown to be exact:

$$J_V^{\alpha} = \delta^{\alpha 3} \frac{q^2 B}{2\pi^2} \mu_A + \delta^{\alpha 0} q n,$$

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