

# GROUP THEORY AND HYDRODYNAMICS

FORMALISM, GAUGE FIELDS & ANOMALIES

V. P. NAIR

CITY COLLEGE OF THE CUNY



*Recent Developments in Theoretical Physics*

KOREA INSTITUTE FOR ADVANCED STUDY

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- Lagrange's approach
  - Newton's equations for  $N$  point-particles  $\rightarrow$  coarse graining using a smooth density function  $\rightarrow$  fluid dynamics
- Point particle  $\equiv$  a unitary irreducible representation (UIR) of the Poincaré group
- Classical action which upon quantization gives a UIR of a group = A co-adjoint orbit action

Can we construct fluid dynamics as

Co-adjoint orbit action  $\rightarrow$  coarse graining  $\rightarrow$  fluid dynamics ?

- Advantages:
  - A single formalism where symmetries are foundational
  - Gauge fields  $\rightarrow$  Abelian and nonabelian Magnetohydrodynamics
  - Spin, magnetic moment effects
  - Gravity easily included (Mathisson-Papapetrou equation)
  - Anomalous symmetries (chiral magnetic effect, chiral vorticity effect, etc.)

- Consider nonrelativistic particles with internal “color” degrees of freedom. (Could be spin).
- Start with  $SU(2)$  color, denote the color charge as  $Q^a(t)$ . Equations of motion are (WONG)

$$\dot{Q}^a - f^{abc} A_i^c \dot{x}_i Q^b = 0$$

$$\dot{p}_i - F_{ij}^a \dot{x}_j Q^a = 0$$

- These can be obtained from the action

$$S = \int \left[ \frac{1}{2} m \dot{x}^2 + A_i^a Q^a \dot{x}_i \right] - i n \int \text{Tr}(\sigma_3 g^{-1} \dot{g}), \quad g \in SU(2)$$

with  $Q^a = \frac{n}{2} \text{Tr}(g \sigma_3 g^{-1} \sigma^a)$ .

- The crucial  $\text{Tr}(\sigma_3 g^{-1} \dot{g})$  term was introduced, in this and related contexts, by BALACHANDRAN & collaborators in the 70s and early 80s. (Also related to BOREL-WEIL-BOTT theory and to work by KOSTANT, SOURIAU, KIRILLOV + ....)

- Under  $g \rightarrow g \exp(i \frac{\sigma_3}{2} \varphi)$  it changes by  $\Delta S = n \Delta \varphi$ . For closed paths in  $SU(2)$ ,  $\Delta \varphi = 2\pi \times \text{integer} \Rightarrow n \in \mathbb{Z}$ .
- The co-adjoint term leads to something like a monopole field on  $S^2 = SU(2)/U(1)$ , which is the phase space for the “color” degrees of freedom.
- Quantizing  $g \implies$  one unitary irreducible representation of  $SU(2)$ , with  $j\text{-value} = n/2$ , leading to standard description of color by matrices.
- For group  $G$ , we have  $\text{rank}(G) \equiv r$  mutually commuting generators in  $\mathfrak{G}$ ,  $\text{rank}(G)$  coefficients, and  $g \in G$ . The action is given by

$$S = -i \sum_{s=1}^r w_s \int d\tau \text{Tr}(q_s g^{-1} \dot{g})$$

$q_s$  are diagonal generators of  $G$ , and  $\{w_s\}$  must be the weights for a UIR of  $G$ . Upon quantization, this gives the corresponding UIR.

- For relativistic point-particles, we must use this action with  $G =$  Poincaré group
- We consider Poincaré group = contraction of de Sitter group; this makes some traces easier to define.
- For de Sitter algebra, use standard Dirac  $\gamma$ -matrices with

$$J_{\mu\nu} = \frac{i}{4}[\gamma_\mu, \gamma_\nu], \quad P_\mu = \frac{\gamma_\mu}{r_0}, \quad \text{Poincaré} = r_0 \rightarrow \infty \text{ limit}$$

- A general element is given by

$$g = \exp(i\gamma_\alpha x^\alpha / r_0) \Lambda, \quad \Lambda = B(p) R$$

$$B(p) = \frac{1}{\sqrt{2m(p_0 + m)}} \begin{bmatrix} p_0 + m & \vec{\sigma} \cdot \vec{p} \\ \vec{\sigma} \cdot \vec{p} & p_0 + m \end{bmatrix}$$

$\Lambda$  is an element of the Lorentz group,  $R$  is a pure spatial rotation generated by  $J_{12}, J_{23}, J_{31}$ , and  $m = \sqrt{p^2}$ .

- The action is given by

$$S = i m r_0^2 \int d\tau \operatorname{Tr} \left( \frac{\gamma_0}{r_0} g^{-1} \frac{dg}{d\tau} \right) + i \frac{n}{2} \operatorname{Tr}(J_{12} g^{-1} dg) - \mathcal{H}$$

Using  $B\gamma_0 B^{-1} = \gamma^\alpha p_\alpha / m$  and taking  $r_0 \rightarrow \infty$ , we find, for the Poincaré group,

$$S = - \int d\tau p_\mu \dot{x}^\mu + i \frac{n}{4} \int d\tau \operatorname{Tr}(\Sigma_3 \Lambda^{-1} \dot{\Lambda}) - \mathcal{H}, \quad \Sigma_3 = \begin{bmatrix} \sigma_3 & 0 \\ 0 & \sigma_3 \end{bmatrix}$$

- $\mathcal{H}$  generates  $\tau$ -evolution, so we should set it to zero as a constraint on quantum states. This leads to the wave equation.
- The addition of the term  $e A_\mu \dot{x}^\mu$  leads to relativistic charged point-particle dynamics, with magnetic moment ( $g = 2$ ) and spin-orbit coupling.

- Consider the point-particle *à la* WONG again. Take a collection of particles indexed by  $\lambda$ .

$$S = -in \int dt \text{Tr}(\sigma_3 g^{-1} \dot{g}) \rightarrow S = -i \int dt \sum_{\lambda} n_{\lambda} \text{Tr}(\sigma_3 g_{\lambda}^{-1} \dot{g}_{\lambda})$$

- We can take the continuum limit by  $\lambda \rightarrow \vec{x}$ ,  $\sum_{\lambda} \rightarrow \int d^3x/v$ ,  $n_{\lambda}/v \rightarrow \rho(x)$ .
- This leads to

$$S = -i \int d^4x \rho \text{Tr}(\sigma_3 g^{-1} \dot{g})$$

where  $g = g(\vec{x}, t)$ .

- This suggest the relativistic form

$$S = -i \int j^{\mu} \text{Tr}(\sigma_3 g^{-1} \partial_{\mu} g)$$

- The difficulty for Poincaré is about what replaces  $\dot{x}^{\mu}$ . Only 3 of the 4 components are independent; also positions of particles are not well defined in the fluid version.

- Ordinary fluid dynamics can be described by a Poisson bracket system

$$\begin{aligned} [\rho(x), \rho(y)] &= 0 \\ [v_i(x), \rho(y)] &= \partial_{xi} \delta^{(3)}(x - y) \\ [v_i(x), v_j(y)] &= -\frac{\omega_{ij}}{\rho} \delta^{(3)}(x - y) \end{aligned}$$

$$\omega_{ij} = (\partial_i v_j - \partial_j v_i).$$

$$H = \int d^3x \left[ \frac{1}{2} \rho v^2 + V(\rho) \right]$$

- We get the usual equations of fluid motion with pressure  $p = \rho \frac{\partial V}{\partial \rho} - V$ .
- The PBs can be summarized as

$$[F, G] = \int \left[ \frac{\delta F}{\delta \rho} \partial_i \left( \frac{\delta G}{\delta v_i} \right) - \frac{\delta G}{\delta \rho} \partial_i \left( \frac{\delta F}{\delta v_i} \right) - \frac{\omega_{ij}}{\rho} \frac{\delta F}{\delta v_i} \frac{\delta G}{\delta v_j} \right]$$

for any two functions  $F, G$ .



- The helicity  $C$  is given by

$$C = \frac{1}{8\pi} \int \epsilon^{ijk} v_i \partial_j v_k = \text{CS term for } v_i$$

- The helicity Poisson-commutes with all local observables,  $[F, C] = 0$  for all  $F$   
 $\implies C$  is superselected.
- Usually if  $[\xi^a, \xi^b] = K^{ab}$ , the Lagrangian is of the form  $\mathcal{A}_b \dot{\xi}^b$ , where  $\partial_a \mathcal{A}_b - \partial_b \mathcal{A}_a = K_{ab}^{-1}$ .  
 Here  $K$  is not invertible,  $\delta C / \delta v_i$  is a zero mode.  
 This is the difficulty in writing down a Lagrangian.
- The solution is also clear: **We must fix the value of  $C$  and seek a parametrization for the velocity which keeps the same value of  $C$ .**
- Such a parametrization exists. It is the so-called Clebsch parametrization,

$$v_i = \partial_i \theta + \alpha \partial_i \beta$$

$\theta, \alpha, \beta$  are arbitrary functions.

- For  $v_i$  parametrized in terms of well-defined  $\theta, \alpha, \beta$ ,

$$C = \int (\text{total derivative}) = 0$$

- A suitable action which gives the PBs is now (C.C. LIN)

$$S = \int \rho \dot{\theta} + \rho \alpha \dot{\beta} - \int \left[ \frac{1}{2} \rho v^2 - V \right]$$

- We can also write this as

$$S = \int J^\mu (\partial_\mu \theta + \alpha \partial_\mu \beta) - \int \left[ J^0 - \frac{J^i J^i}{2\rho} + V \right]$$

$J^0 = \rho$ ; elimination of the auxiliary  $J^i$  leads to the previous version.  $\int J^0$  is a constant.

- The relativistic generalization is

$$S = \int J^\mu (\partial_\mu \theta + \alpha \partial_\mu \beta) - \int f(n)$$

$$f(n) = n + V(n), \quad n^2 = J^2 = (J^0)^2 - J^i J^i.$$

- The lesson from this is to treat
  - Translational part of action  $\rightarrow$  Clebsch parametrization
  - Rest of the action in terms of the co-adjoint orbit version
- The general action is thus

$$S = \int d^4x \left[ j^\mu (\partial_\mu \theta + \alpha \partial_\mu \beta) - \frac{i}{4} j_{(s)}^\mu \text{Tr}(\Sigma_3 \Lambda^{-1} \partial_\mu \Lambda) + i \sum_a j_{(a)}^\mu \text{Tr}(q_a g^{-1} D_\mu g) - f(\{n\}) \right] + S(A)$$

- Generally, we must have different currents  $j^\mu, j_{(s)}^\mu, j_{(a)}^\mu$  for mass flow, spin flow and the transport of other quantum numbers.
- Coupling to gauge fields follow from covariant derivatives on the group elements

- $f(\{n\})$  depends on all invariant combinations of the currents and characterize the nature of the fluid,  $n = \sqrt{j^\mu j_\mu}$ ,  $n_a = \sqrt{j_{(a)}^\mu j_{\mu(a)}}$ , etc.
- The group-valued fields are related to flow velocities and currents and given by the equations of motion,

$$\begin{aligned} \frac{1}{n} \frac{\partial f}{\partial n} j_\mu &= \partial_\mu \theta + \alpha \partial_\mu \beta \\ \frac{1}{n_a} \frac{\partial f}{\partial n_a} j_{\mu(a)} &= i \text{Tr}(q_a g^{-1} D_\mu g), \quad \text{etc.} \end{aligned}$$

**Remark:** The Clebsch parametrization can also be written in a “group” form,

$$-i \text{Tr}(\sigma_3 g^{-1} dg) = d\theta + \alpha d\beta$$

where  $g \in SU(1, 1)$ ,

$$g = \frac{1}{\sqrt{1 - \bar{u}u}} \begin{pmatrix} 1 & u \\ \bar{u} & 1 \end{pmatrix} \begin{pmatrix} e^{i\theta/2} & 0 \\ 0 & e^{-i\theta/2} \end{pmatrix}, \quad \alpha = \frac{2\bar{u}u}{(1 - \bar{u}u)}, \quad \beta = -\frac{i}{2} \log(u/\bar{u})$$

- We will discuss 3 examples
  - $SU(2)$  internal symmetry (Nonabelian Magnetohydrodynamics)
  - Magnetohydrodynamics including spin, magnetic moment and spin-orbit effects
  - Spin and coupling to gravity
- We will also discuss generalization to include anomalies

- Consider the action (BISTROVIC, JACKIW, LI, NAIR, PI)

$$S = \int J^\mu (\partial_\mu \theta + \alpha \partial_\mu \beta) - i \int j^\mu \text{Tr}(\sigma_3 g^{-1} D_\mu g) - \int f(n) + S_{YM}$$

$$D_\mu g = \partial_\mu g + A_\mu g \qquad A_\mu = -i t^a A_\mu^a, \quad t^a = \frac{1}{2} \sigma^a$$

$$J^\mu = n_m U^\mu, \qquad U^2 = 1$$

$$j^\mu = n u^\mu, \qquad u^2 = 1$$

- We also include a background field which couples to the color charge.

- The current which couples to  $A_\mu^a$  is given by

$$J^{a\mu} = \text{Tr}(\sigma_3 g^{-1} t^a g) j^\mu = Q^a j^\mu, \quad Q^a = \text{Tr}(\sigma_3 g^{-1} t^a g)$$

This is the Eckart form for currents.

- The equations of motion are

$$\partial_\mu j^\mu = 0$$

$$(D_\mu J^\mu)^a = 0$$

$$n u^\mu \partial_\mu (u_\nu f') - n \partial_\nu f' = \text{Tr}(J^\mu F_{\mu\nu}) \quad (\text{"Euler equation"})$$

- The first two equations lead to the fluid generalization of the Wong equations

$$u^\mu (D_\mu Q)^a = (D_0 Q)^a + \vec{u} \cdot (\vec{D} Q)^a = 0$$

- We also have

$$\partial_\mu T^{\mu\nu} = \text{Tr}(J^\mu F_{\mu\nu})$$

$T^{\mu\nu}$  has the perfect fluid form.

- The nonabelian charge density  $\rho = \rho^a t^a$  (which is the time-component of  $J^{a\mu}$ ) transforms, under gauge transformations, as

$$\rho \rightarrow \rho' = h^{-1} \rho h, \quad h \in SU(2)$$

- We can diagonalize  $\rho$  at each point by an  $(\vec{x}, t)$ -dependent transformation,  $\rho_{diag} = \rho_0 \sigma_3$ . Then  $\rho = g \rho_{diag} g^{-1}$ , or

$$\rho^a = \rho_0 \text{Tr}(g \sigma_3 g^{-1} t^a) = j^0 \text{Tr}(g \sigma_3 g^{-1} t^a)$$

- $g$  is the transformation which diagonalizes the charge density at each point. The eigenvalues are gauge-invariant and are represented by  $n$ . Their flow is given by  $u^\mu$ . Under a gauge transformation,  $g \rightarrow h^{-1} g$ .



- There are two (related) charge densities,  $j^0$  and the nonabelian charge density  $\rho^a = J^{a0}$ .
- The basic (new) Poisson brackets are

$$\{j^0(\vec{x}), j^0(\vec{y})\} = 0$$

$$\{j^0(\vec{x}), g(\vec{y})\} = -i g(\vec{x}) \left(\frac{\sigma_3}{2}\right) \delta(x - y)$$

$$\{\rho^a(\vec{x}), \rho^b(\vec{y})\} = f^{abc} \rho^c(\vec{x}) \delta(x - y)$$

$$\{\rho^a(\vec{x}), g(\vec{y})\} = -i \left(\frac{\sigma_a}{2}\right) g(\vec{x}) \delta(x - y)$$

- Another interesting observation is that, since  $\Pi_3[SU(N)] = \mathbb{Z}$ , there are skyrmion-type solitons in any nonabelian magnetohydrodynamics. (DAI, NAIR)
- **Remark:** These equations of motion and charge algebra have some points of overlap with the work of GIBBONS, HOLM, KUPERSHMIDT

- Consider a special case where mass transport and charge transport are described by the same flow velocity.
- This applies when we have one species of particles with the same charge.
- Further, for dilute systems, if we neglect the possibility of spin-singlets forming (and moving independently), we can take spin flow velocity  $\approx$  charge flow velocity
- The action for this case is (KARABALI, NAIR)

$$S = S(A) + \int d^4x \left[ j^\mu (\partial_\mu \theta + \alpha \partial_\mu \beta + e A_\mu) - \frac{i}{4} j^\mu \text{Tr}(\Sigma_3 \Lambda^{-1} \partial_\mu \Lambda) - f(n, \sigma) \right]$$

$\Lambda = B R$  contains the same velocity  $u^\mu$  as in  $j^\mu = n u^\mu$ .

- $F$  depends on  $n$  and  $\sigma = S^{\mu\nu} F_{\mu\nu}$ , where  $S^{\mu\nu}$  is the spin density,

$$S^{\mu\nu} = \frac{1}{2} \text{Tr}(\Sigma_3 \Lambda^{-1} J^{\mu\nu} \Lambda), \quad J^{\mu\nu} = \frac{i}{4} [\gamma^\mu, \gamma^\nu]$$

- Having the same flow velocity can be satisfied by

$$\frac{2}{n} \frac{\partial f}{\partial n} \frac{\partial f}{\partial \sigma} = e$$

This is the fluid analog of the requirement of  $g = 2$  for point-particles.

- The equations of motion are the Maxwell equations +

$$\begin{aligned}
 u^\alpha \partial_\alpha (f' u_\nu) - \partial_\nu f' &= e \left[ u^\lambda F_{\lambda\nu} - \frac{4}{s^2 f'} \partial_\nu S^{\lambda\beta} (SFS - FSS)_{\lambda\beta} - \dots \right] \\
 u^\alpha \partial_\alpha S_{\mu\nu} &= \frac{1}{f'} \left[ S_\mu^\lambda (eF_{\lambda\nu} + G_{\lambda\nu}) - S_\nu^\lambda (eF_{\lambda\mu} + G_{\lambda\mu}) \right] \\
 &\quad - \frac{4e}{s^2 f'^2} (u_\mu S_\nu^\lambda - u_\nu S_\mu^\lambda) \partial_\lambda S^{\rho\beta} (SFS - FSS)_{\rho\beta} + \dots \\
 G_{\lambda\nu} &= u_\lambda \partial_\nu f' - u_\nu \partial_\lambda f' \\
 (SFS - FSS)_{\lambda\beta} &= S_\lambda^\rho F_{\rho\tau} S_\beta^\tau - F_\lambda^\rho S_{\rho\tau} S_\beta^\tau
 \end{aligned}$$

- Spin density is subject to precession effects due to pressure gradient terms as well as due to the external field.

- Consider spin transport in the fluid in a gravitational background.
- Specializing the general fluid action to the Lorentz group (to describe spin), with covariant derivatives appropriate to a general curved manifold,

$$S[e, \omega, j, \Lambda] = \int \det e \left[ -\frac{i}{2} j^\mu \text{Tr}(\Sigma_3 \Lambda^{-1} D_\mu(\bar{\omega}(e)) \Lambda) - f(n) \right] - \frac{1}{32 \pi G} \epsilon_{abcd} \int e^a \wedge e^b \wedge R^{cd}(\omega)$$

- Here we use  $\bar{\omega}(e)$  to avoid generating torsion via equations of motion,

$$\bar{\omega}_\mu^{ab} = (e^{-1})^{\nu a} \partial_{[\mu} e_{\nu]}^b - (e^{-1})^{\nu b} \partial_{[\mu} e_{\nu]}^a - (e^{-1})^{\rho a} (e^{-1})^{\sigma b} \partial_{[\rho} e_{\sigma]}^c e_{\mu c}.$$

- The energy-momentum tensor is now

$$T_{\mu\nu} = T_{\mu\nu}^{(f)} + 2 \nabla_\alpha (j_\mu Q_\nu^\alpha + j_\nu Q_\mu^\alpha).$$

where  $T_{\mu\nu}^{(f)}$  has the perfect fluid form,

$$T_{\mu\nu}^{(f)} = n f' u_\mu u_\nu - g_{\mu\nu} (n f' - f)$$

where  $Q^{\alpha\beta}$  is the spin density,

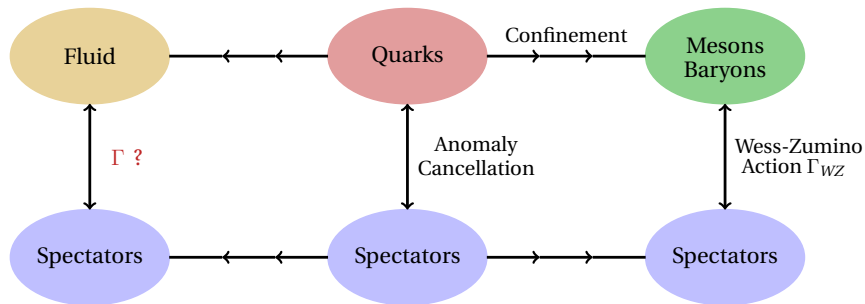
$$Q^{\alpha\beta} = (e^{-1})^\alpha_a (e^{-1})^\beta_b \frac{1}{2} \text{Tr} \left( \Sigma_3 \Lambda^{-1} J^{ab} \Lambda \right)$$

- The conservation law may be written as

$$\nabla_\mu T^{(f)\mu\nu} - 2 (R_{\alpha\beta})^\nu_\lambda j^\lambda Q^{\alpha\beta} = 0$$

- The second term is the fluid version of the spin-curvature coupling occurring in the **Mathisson-Papapetrou equation**.

- 't Hooft argument for the Wess-Zumino action for anomalies



- A similar argument for the fluid phase suggests an effective action for anomalies in terms of fluid variables. What is this action?

- Since we have formulated fluid dynamics using group variables, this is easy. We can use the same  $\Gamma_{WZ}$  but using fluid group element instead of meson fields.
- The suggestion is (NAIR, RAY, ROY)

$$\begin{aligned}
 S = & -i \int \left[ j_3^\mu \text{Tr} \left( \frac{\lambda_3}{2} g_L^{-1} D_\mu g_L \right) + j_8^\mu \text{Tr} \left( \frac{\lambda_8}{2} g_L^{-1} D_\mu g_L \right) + k_3^\mu \text{Tr} \left( \frac{\lambda_3}{2} g_R^{-1} D_\mu g_R \right) \right. \\
 & \left. + k_8^\mu \text{Tr} \left( \frac{\lambda_8}{2} g_R^{-1} D_\mu g_R \right) + j_0^\mu \text{Tr} \left( g_L^{-1} D_\mu g_L \right) + k_0^\mu \text{Tr} \left( g_R^{-1} D_\mu g_R \right) \right] \\
 & - f(n_3, n_8, n_0, m_3, m_8, m_0) + S_{YM}(A) + \Gamma_{WZ}(A_L, A_R, g_L g_R^\dagger)
 \end{aligned}$$

- $\Gamma_{WZ}(A_L, A_R, g_L g_R^\dagger)$  is the standard Wess-Zumino term  $\Gamma_{WZ}(A_L, A_R, U)$  with  $U \implies g_L g_R^\dagger$ .
- There are other ways to incorporate anomalies (SON & SUROWKA; SADOBYEV & ISACHENKOV; ABANOV *et al*; BASAR, DUNNE, KHARZEEV; + *many others*); an approach somewhat similar to ours is by SHU LIN.

- In full it is given by (WITTEN; KAYMAKÇALAN, RAJEEV, SCHECHTER; + ...)

$$\begin{aligned}
 \Gamma_{WZ} = & -\frac{iN}{240\pi^2} \int_D \text{Tr}(dU U^{-1})^5 - \frac{iN}{48\pi^2} \int_{\mathcal{M}} \text{Tr}[(A_L dA_L + dA_L A_L + A_L^3) dUU^{-1}] \\
 & - \frac{iN}{48\pi^2} \int_{\mathcal{M}} \text{Tr}[(A_R dA_R + dA_R A_R + A_R^3) U^{-1} dU] \\
 & + \frac{iN}{96\pi^2} \int_{\mathcal{M}} \text{Tr}[A_L dUU^{-1} A_L dUU^{-1} - A_R U^{-1} dU A_R U^{-1} dU] \\
 & + \frac{iN}{48\pi^2} \int_{\mathcal{M}} \text{Tr}[A_L (dUU^{-1})^3 + A_R (U^{-1} dU)^3 + dA_L dU A_R U^{-1} - dA_R d(U^{-1}) A_L U] \\
 & + \frac{iN}{48\pi^2} \int_{\mathcal{M}} \text{Tr}[A_R U^{-1} A_L U (U^{-1} dU)^2 - A_L U A_R U^{-1} (dUU^{-1})^2] \\
 & - \frac{iN}{48\pi^2} \int_{\mathcal{M}} \text{Tr}[(dA_R A_R + A_R dA_R) U^{-1} A_L U - (dA_L A_L + A_L dA_L) U A_R U^{-1}] \\
 & - \frac{iN}{48\pi^2} \int_{\mathcal{M}} \text{Tr}[A_L U A_R U^{-1} A_L dUU^{-1} + A_R U^{-1} A_L U A_R U^{-1} dU] \\
 & - \frac{iN}{48\pi^2} \int_{\mathcal{M}} \text{Tr}[A_R^3 U^{-1} A_L U - A_L^3 U A_R U^{-1} + \frac{1}{2} U A_R U^{-1} A_L U A_R U^{-1} A_L]
 \end{aligned}$$

with  $U \implies g_L g_R^\dagger$ .



- This action gives the chiral magnetic effect (& other anomaly related effects) for all flavor gauge fields and chemical potentials ( $A_0$  components become the chemical potentials  $\mu$ .)
- The electromagnetic current, for example, is given by (*previous refs, also CALLAN & WITTEN*)

$$\begin{aligned}
 J^\mu &= J_3^\mu + \frac{e}{16\pi^2} \epsilon^{\mu\nu\alpha\beta} \text{Tr} \left[ Q(\partial_\nu U U^{-1} \partial_\alpha U U^{-1} \partial_\beta U U^{-1} \right. \\
 &\quad \left. + U^{-1} \partial_\nu U U^{-1} \partial_\alpha U U^{-1} \partial_\beta U) \right] \\
 &\quad + i \frac{e^2}{4\pi^2} \epsilon^{\mu\nu\alpha\beta} \partial_\nu A_\alpha \text{Tr} \left[ Q^2 (\partial_\beta U U^{-1} + U^{-1} \partial_\beta U) + \frac{1}{2} (Q \partial_\beta U Q U^{-1} \right. \\
 &\quad \left. - Q U Q \partial_\beta U^{-1}) \right]
 \end{aligned}$$

- We can restrict to two flavors by choosing

$$U = e^{i\theta} \begin{bmatrix} V & 0 \\ 0 & 1 \end{bmatrix}$$

- The current is now

$$J^\mu = J_3^\mu + \frac{e}{48\pi^2} \epsilon^{\mu\nu\alpha\beta} \text{Tr}(\mathcal{I}_\nu \mathcal{I}_\alpha \mathcal{I}_\beta) + i \frac{e^2}{16\pi^2} \epsilon^{\mu\nu\alpha\beta} \partial_\nu A_\alpha \text{Tr}[(\Sigma_{3L} + \Sigma_{3R}) I_\beta] + J_\theta^\mu$$

$$J_\theta^\mu = -\frac{e^2}{4\pi^2} \epsilon^{\mu\nu\alpha\beta} \partial_\nu A_\alpha \partial_\beta \theta \left[ 2 + \frac{1}{4} \text{Tr}(\Sigma_{3L} \Sigma_{3R} - 1) \right]$$

$$\mathcal{I}_\beta = g_L^{-1} \partial_\beta g_L - g_R^{-1} \partial_\beta g_R, \quad \Sigma_{3L} = g_L^{-1} \sigma_3 g_L, \quad \Sigma_{3R} = g_R^{-1} \sigma_3 g_R.$$

- If we further restrict to  $g_L = g_R$ , we get

$$J_\theta^\mu = -\frac{e^2}{2\pi^2} \epsilon^{\mu\nu\alpha\beta} (\partial_\nu A_\alpha) \partial_\beta \theta$$

$$J_i = -\frac{e^2}{4\pi^2} (\mu_L - \mu_R) B_i$$

This reproduces the chiral magnetic effect which was originally calculated using Feynman diagrams (KHARZEEV, MCLERRAN, WARRINGA, FUKUSHIMA + ...).

- The full set of equations describe hydrodynamic transport of flavor charges.

- The anomaly term  $\Gamma_{WZ}$  also has terms proportional to  $Z_\mu$ , so there is also an induced isospin current (CAPASSO, NAIR, TEKEL).
- The relevant term is

$$\Gamma_{WZ} = -\frac{Ne^2}{6\pi^2} (\cot 2\theta_W) \int \epsilon^{\mu\nu\alpha\beta} Z_\mu \partial_\nu A_\alpha \partial_\beta \theta$$

- This leads to

$$J^{Z\mu} = -\frac{e}{8\pi^2} (\cos 2\theta_W) \epsilon^{\mu\nu\alpha\beta} F_{\nu\alpha} \partial_\beta \theta$$

$$J^{3\mu} = \frac{e}{8\pi^2} (\mu_L - \mu_R) B^i$$

- In terms of pion fields,  $J^{3\mu} \approx -\frac{1}{2} f_\pi \partial^\mu \Pi^0 + \dots$ . So we can interpret this as a pion field of gradient

$$\partial^i \Pi^0 = -\frac{e}{4\pi^2 f_\pi} (\mu_L - \mu_R) B^i$$

- This can manifest itself as an asymmetry in the neutral pion distribution.

- Generally, there is a contribution even when the background fields are zero.
- If we eliminate the group elements in favor of velocities, we get

$$\begin{aligned}
 J^\mu &= J_3^\mu + J_\theta^\mu + i \frac{e^2}{16\pi^2} \epsilon^{\mu\nu\alpha\beta} \partial_\nu A_\alpha \text{Tr} [(\Sigma_{3L} + \Sigma_{3R}) I_\beta] \\
 &\quad + \frac{1}{16\pi^2} \epsilon^{\mu\nu\alpha\beta} \partial_\nu \text{Tr}(\mathbf{g}_L^{-1} \partial_\alpha \mathbf{g}_L \mathbf{g}_R^{-1} \partial_\beta \mathbf{g}_R) \\
 &\quad + \frac{e}{12\pi^2} \epsilon^{\mu\nu\alpha\beta} \left[ \left( \frac{\partial f}{\partial n_3} \right)^2 u_{3L\nu} \partial_\alpha u_{3L\beta} - \left( \frac{\partial f}{\partial m_3} \right)^2 u_{3R\nu} \partial_\alpha u_{3R\beta} \right].
 \end{aligned}$$

- A left-right asymmetry with nonzero vorticity can generate an electromagnetic current
- Finally, we know that the standard model can have mixed gauge-gravity anomalies in some restricted cases, due to the 6-form index density,

$$I_6 = \frac{i}{384 \pi^3} (\text{Tr} F) \text{Tr} (R \wedge R)$$

- If **up**, **down** and **strange** are described by fluid variables, while **charm** is described as a Dirac field, the fluid part can show an anomaly.
- The hypercharge conservation law is changed to

$$\partial_\mu J^\mu = -i \frac{N_c}{768 \pi^2} \frac{\epsilon^{\mu\nu\alpha\beta}}{\sqrt{-g}} \text{Tr}(R_{\mu\nu} R_{\alpha\beta}).$$

- The corrected energy-momentum tensor is given by the previous result +

$$T^{\nu\sigma}]_{corr} = -i \frac{N_c}{192 \pi^2} \frac{1}{\sqrt{-g}} \nabla_\lambda \left[ \text{Tr}_Y(\partial_\mu \theta) (R_{\alpha\beta})^{\lambda\sigma} \epsilon^{\mu\nu\alpha\beta} + (\nu \leftrightarrow \sigma) \right].$$

The remaining trace is over the hypercharge values. If we replace  $\dot{\theta}$  by the chemical potentials, as can be done for the chiral magnetic effect,

$$\text{Tr}_Y(\dot{\theta}) \rightarrow \frac{1}{2} \left[ \frac{1}{3} (\mu_L^u + \mu_L^d + \mu_L^s) + \frac{2}{3} (\mu_R^d + \mu_R^s - 2\mu_R^u) \right].$$

- A number of new directions are suggested by this. Just to name a few:
  - Relating the Poincaré group version to the formulation of fluid dynamics in terms of the group of diffeomorphisms
  - The chiral vorticity effect and the mixed anomaly effects may be significant in certain gravitational backgrounds, such as cosmic strings.
  - Exploring soliton solutions. (Topological solitons in the fluid are possible in this formalism.)
  - Quantum ground states for fluids can be meaningful in some contexts, as in liquid Helium. How do they look in more general situations?

Thank you