

Daya Bay II: A multi-purpose reactor neutrino experiment

Liang Zhan

Institute of High Energy Physics

International Workshop on RENO-50 toward Neutrino Mass Hierarchy

June 13-14, 2013, Seoul National University

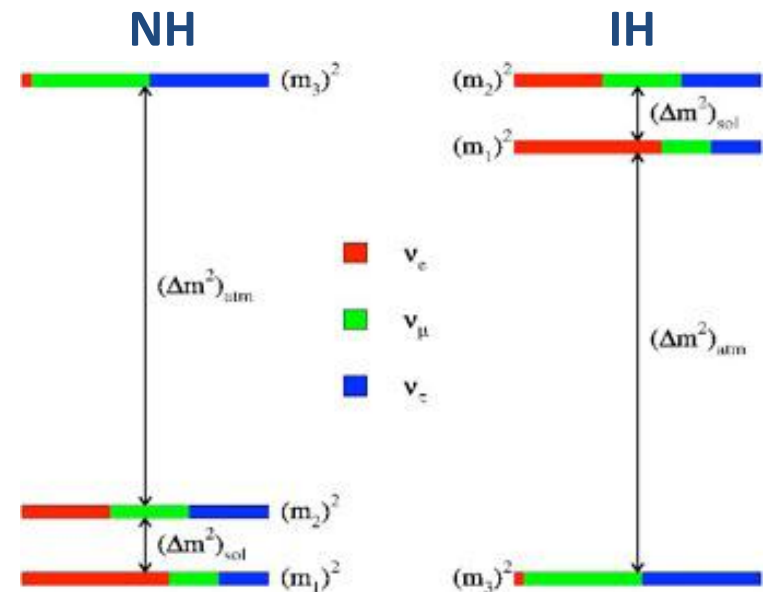
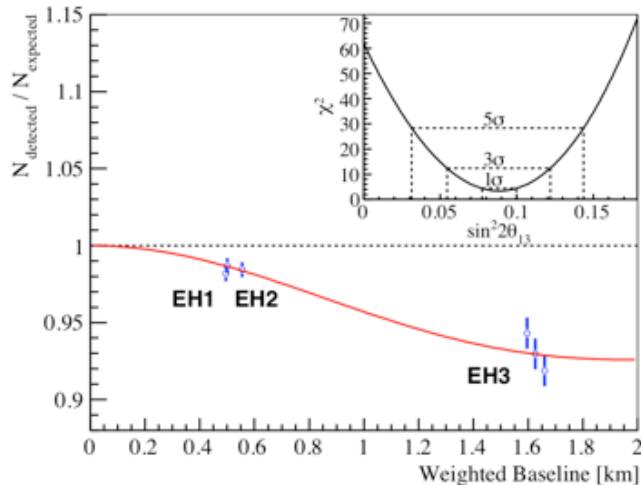
Outline

- Physics
- Challenges
- Detector concepts
- Site optimization and civil
- Schedule
- Summary

The large θ_{13} era

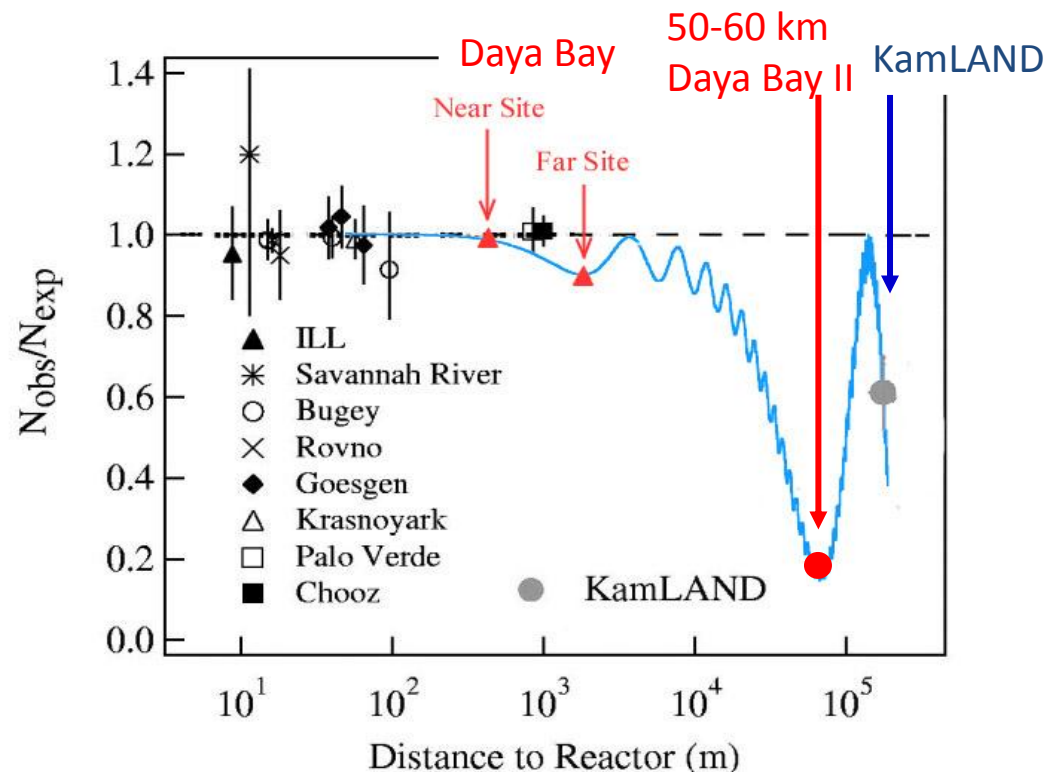
- The non-zero and large θ_{13} has been observed by Daya Bay, Double Chooz, Reno, and accelerator experiments
- Daya Bay will measure $\sin^2 2\theta_{13}$ to 4-5% precision in three years.
- Mass hierarchy and CP phase are the main focus of next generation neutrino experiments.
- A medium baseline reactor neutrino experiment can measure mass hierarchy independent of CP phase.

Daya Bay, Dec 24, 2011 - May 11, 2012
 $\sin^2 2\theta_{13} = 0.089 \pm 0.010(\text{stat}) \pm 0.005(\text{sys})$



Daya Bay II Experiment

- 20 kton LS detector
- 2-3 % energy resolution
- Rich physics possibilities
 - **Mass hierarchy**
 - Precision measurement of 3 mixing parameters
 - Supernova neutrino
 - Geoneutrino
 - Sterile neutrino
 - Atmospheric neutrinos
 - Exotic searches



Reactor antineutrino to determine MH

$$P_{ee}(L/E) = 1 - P_{21} - P_{31} - P_{32}$$

$$P_{21} = \cos^4(\theta_{13}) \sin^2(2\theta_{12}) \sin^2(\Delta_{21})$$

$$P_{31} = \cos^2(\theta_{12}) \sin^2(2\theta_{13}) \sin^2(\Delta_{31})$$

$$P_{32} = \sin^2(\theta_{12}) \sin^2(2\theta_{13}) \sin^2(\Delta_{32})$$

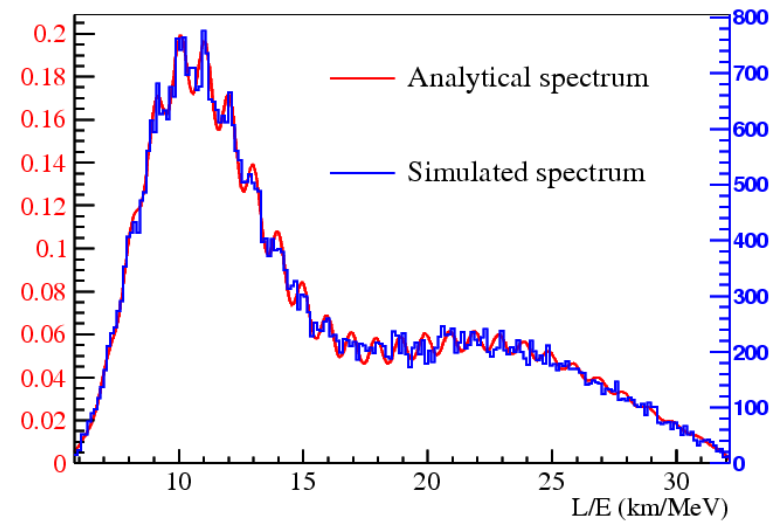
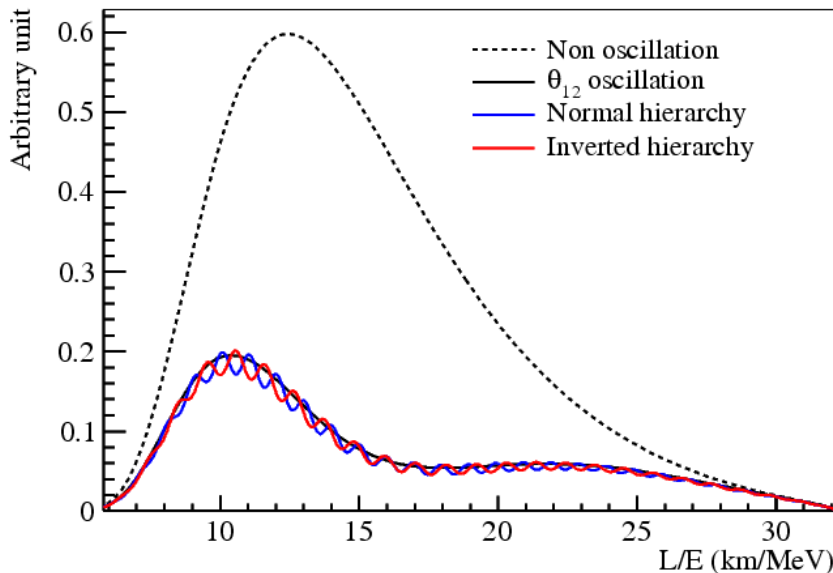
$$\Delta m_{31}^2 = \Delta m_{32}^2 + \Delta m_{21}^2$$

NH : $|\Delta m_{31}^2| = |\Delta m_{32}^2| + |\Delta m_{21}^2|$

IH : $|\Delta m_{31}^2| = |\Delta m_{32}^2| - |\Delta m_{21}^2|$

S.T. Petcov et al., PLB533(2002)94
 S.Choubey et al., PRD68(2003)113006
 J. Learned et al., hep-ex/0612022 L.

Zhan, Y. Wang, J. Cao, L. Wen,
 PRD78:111103, 2008
 PRD79:073007, 2009



50000 events

Fourier transform to L/E spectrum

- L/E spectrum \leftrightarrow δm^2 spectrum (oscillation frequency)

$$F(L/E) = \phi(E)\sigma(E)P_{ee}(L/E)$$

$$FST(\omega) = \int_{t_{min}}^{t_{max}} F(t) \sin(\omega t) dt$$

$$FCT(\omega) = \int_{t_{min}}^{t_{max}} F(t) \cos(\omega t) dt$$

- Clear distinctive features:

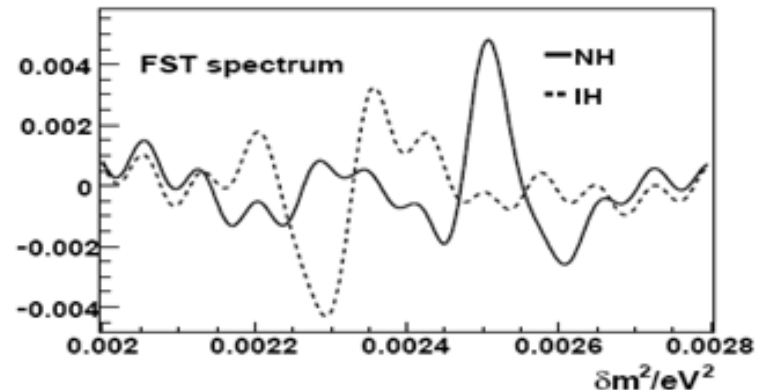
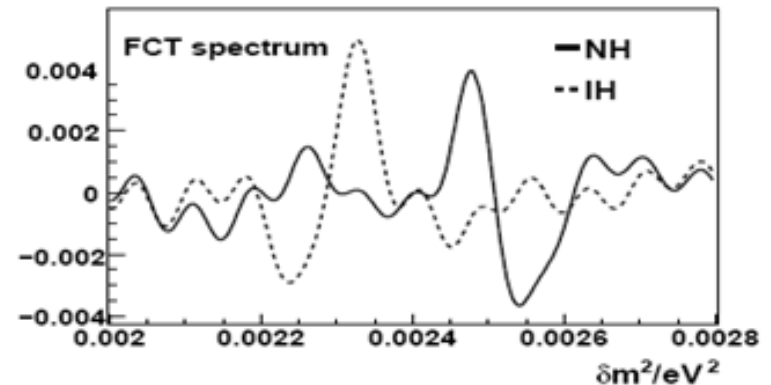
- FCT:

- NH: valley at the left side
- IH: valley at the right side

- FST:

- NH: prominent peak
- IH: prominent valley

- No pre-condition of Δm^2_{32} : features depends on shape but not absolute peak position.

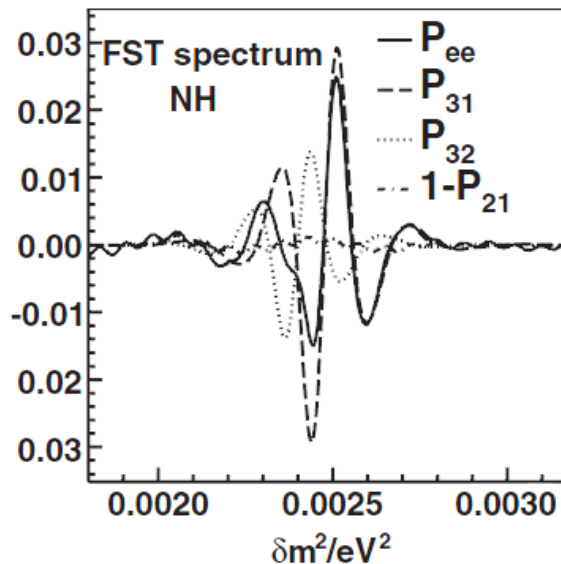
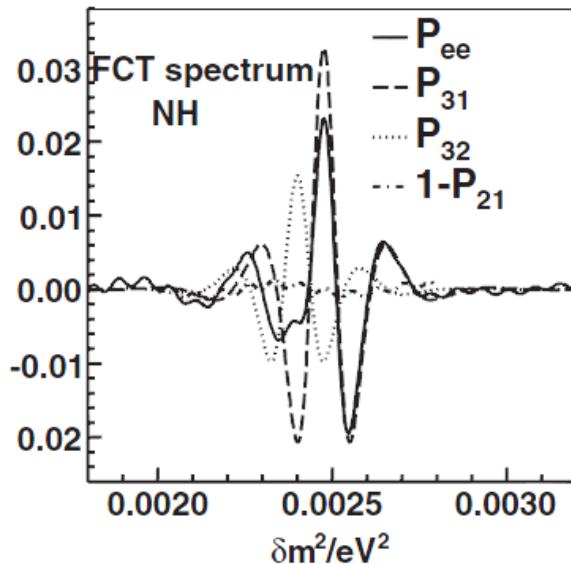
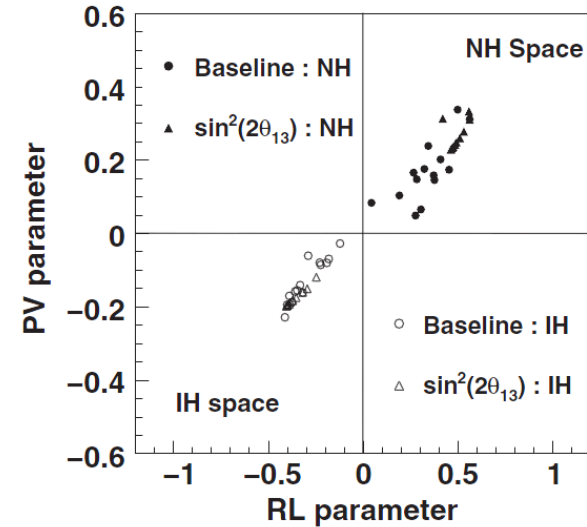


Quantify features of FCT and FST

- Define quantities

$$RL = \frac{RV - LV}{RV + LV}, \quad PV = \frac{P - V}{P + V}$$

- RV/LV: amplitude of the right/left valley in FCT
- P/V: amplitude of the peak/valley in FST
- NH: $PV > 0$ and $RL > 0$, IH: $PV < 0$ and $RL < 0$
- Combined to one quantity: $PV+RL$



Interference of two oscillation components of P_{31} and P_{32}

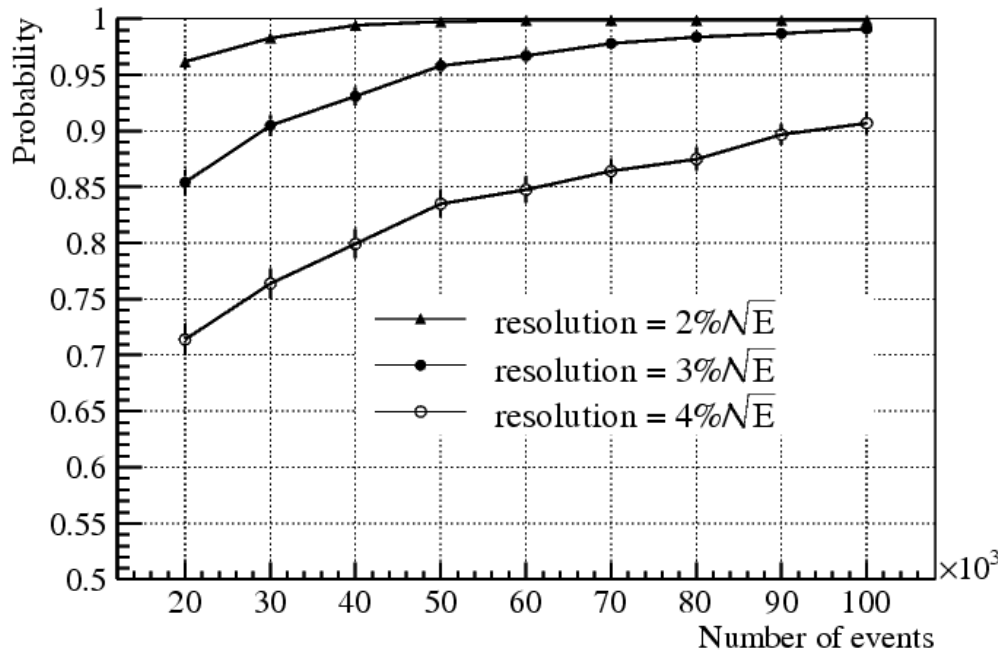
L. Zhan et al.,
PRD78:111103,2008

Experimental requirements

- Un-binned Fourier transform of N detected events

$$\text{FST}(\omega) = \sum_{i=1}^N \sin(\omega L/E_i^l), \quad \text{FCT}(\omega) = \sum_{i=1}^N \cos(\omega L/E_i^l),$$

- Energy resolution is very important for Δm^2_{32} and Δm^2_{31} oscillation measurement.



Energy resolution	3%/sqrt(E)
Baseline	58 km
Thermal Power	35 GW

20kt LS detector
 3 years \sim 2 sigma
 6 years \sim 3 sigma

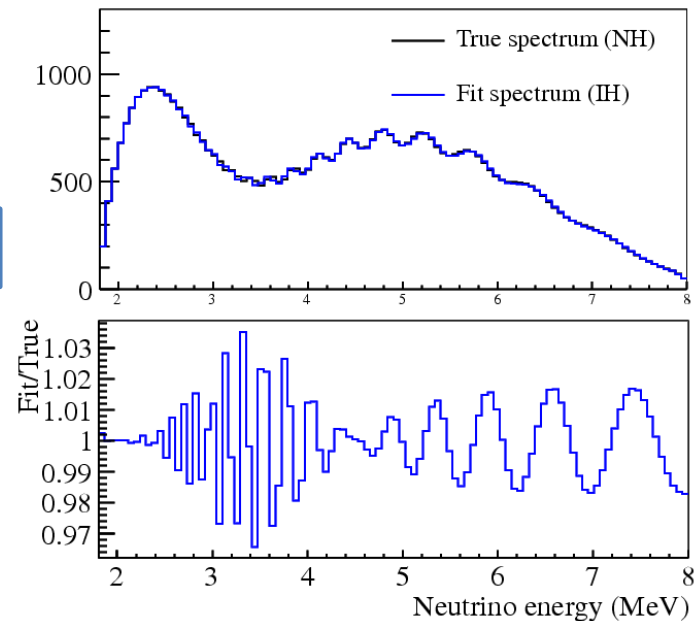
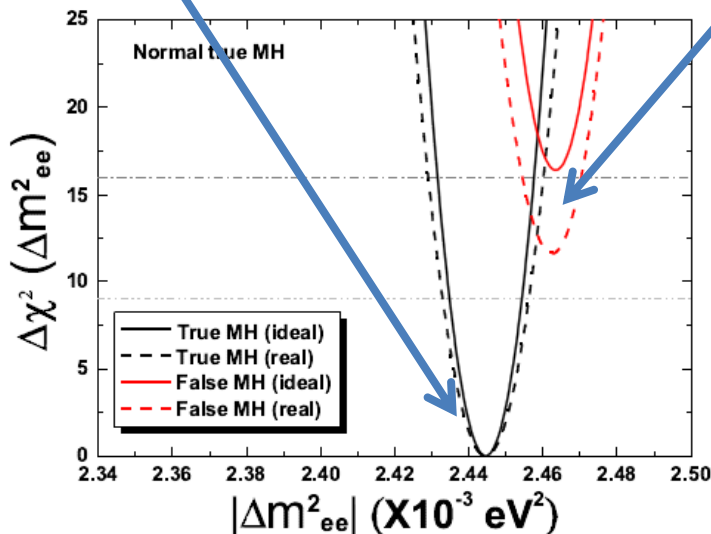
Alternative method: χ^2 fit

- Assume the truth is NH/IH, and calculate the truth spectrum.
- Calculate the spectra for NH and IH case and fit them to the truth spectrum respectively.
- Energy resolution is taken into account.

$$\chi^2 = \sum_i \frac{(F_i(\Delta m^2) - T_i(1 + \varepsilon + \varepsilon_i))^2}{T_i} + \left(\frac{\varepsilon}{\sigma}\right)^2 + \left(\frac{\varepsilon_i}{\sigma_i}\right)^2$$

NH spectrum fits to NH

IH spectrum fits to NH

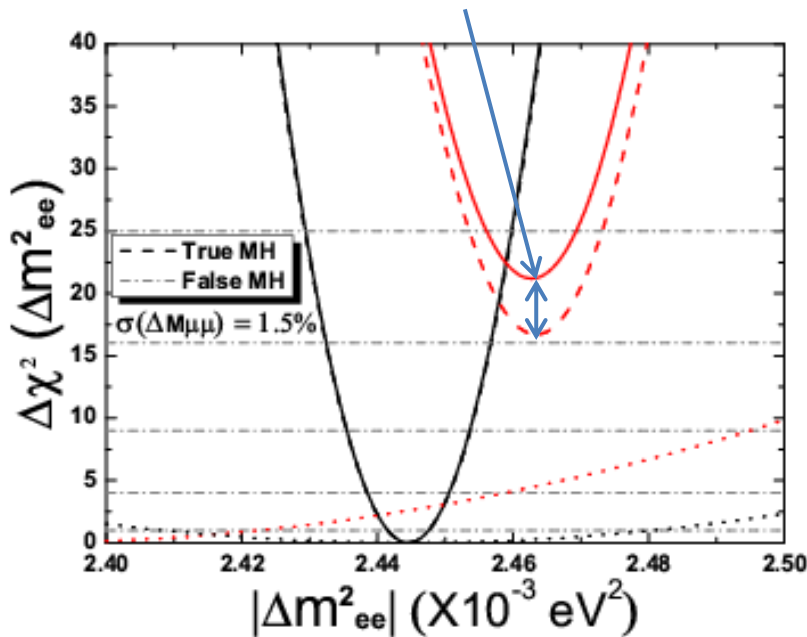


If truth is NH, NH spectrum may fit it better.
 Δm^2 is fitted without constraint.

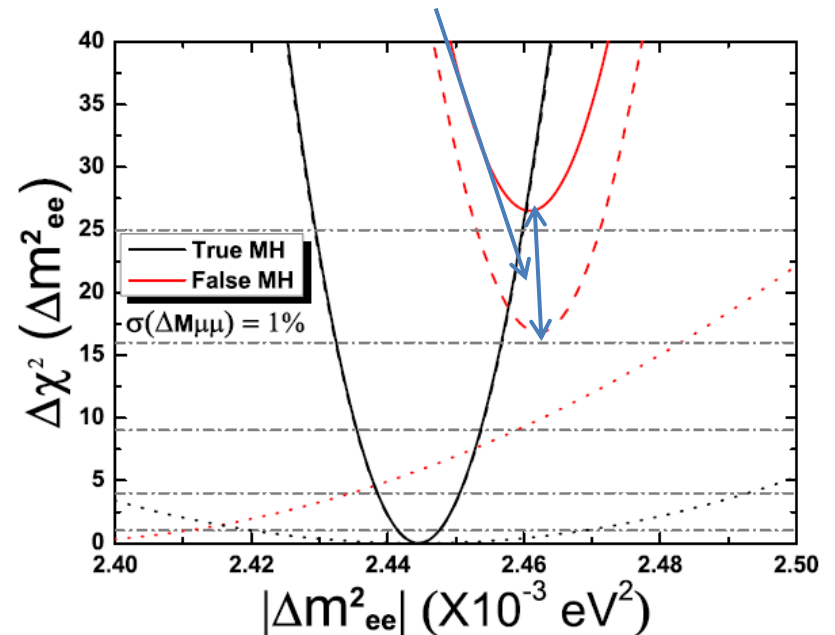
Taking into account Δm^2_{32}

- MH sensitivity improved by taking into account the Δm^2_{32} from T2K and Nova in the future

Improved by 1.5% precision



Improved by 1% precision



Other physics reach

1. Precision measurement of mixing parameters: θ_{12} , ΔM^2_{12} , ΔM^2_{31} → test the unitarity of the mixing matrix
2. Supernova neutrinos
3. Geo-neutrinos
4. Sterile neutrinos
5. Target for neutrino beams
6. Atmospheric neutrinos
7. Solar neutrinos
8. High energy cosmic-rays & neutrinos
 1. Point source: GRB, AGN, BH, ...
 2. Diffused neutrinos
 3. Dark matter

Precision measurement of mixing parameters

- Fundamental to the Standard Model and beyond
- Probing the unitarity of U_{PMNS} to $\sim 1\%$ level !

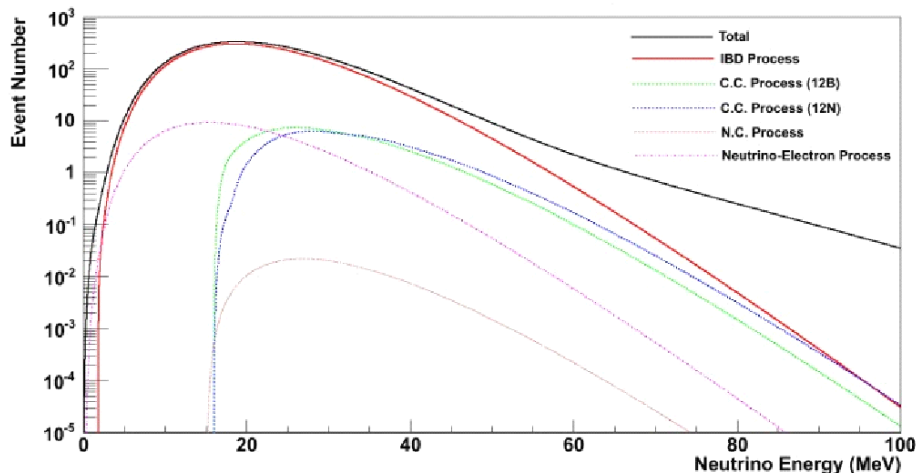
	Current	Daya Bay II
Δm^2_{21}	3%	0.6%
Δm^2_{31}	5%	0.6%
$\sin^2\theta_{12}$	6%	0.7%
$\sin^2\theta_{23}$	20%	-
$\sin^2\theta_{13}$	14% \rightarrow 5% (Daya Bay in 3 years)	15%

Will be more precise than CKM matrix elements.

Supernova neutrinos

- Less than 20 events observed so far
- Assumptions:
 - Distance: 10 kpc (our Galaxy center)
 - Luminosity: 3×10^{53} erg
 - Detector: 20 kt scintillator
- Many types of events:
 - $\bar{\nu}_e + p \rightarrow n + e^+$, ~ 3000 correlated events
 - $\bar{\nu}_e + {}^{12}\text{C} \rightarrow {}^{12}\text{B}^* + e^+$, ~ 10 -100 correlated events
 - $\nu_e + {}^{12}\text{C} \rightarrow {}^{12}\text{N}^* + e^-$, ~ 10 -100 correlated events
 - $\nu_x + {}^{12}\text{C} \rightarrow \nu_x + {}^{12}\text{C}^*$, ~ 600 correlated events
 - $\nu_x + p \rightarrow \nu_x + p$, single events
 - $\nu_e + e^- \rightarrow \nu_e + e^-$, single events
 - $\nu_x + e^- \rightarrow \nu_x + e^-$, single events

Water Cerenkov detectors can not see these correlated events



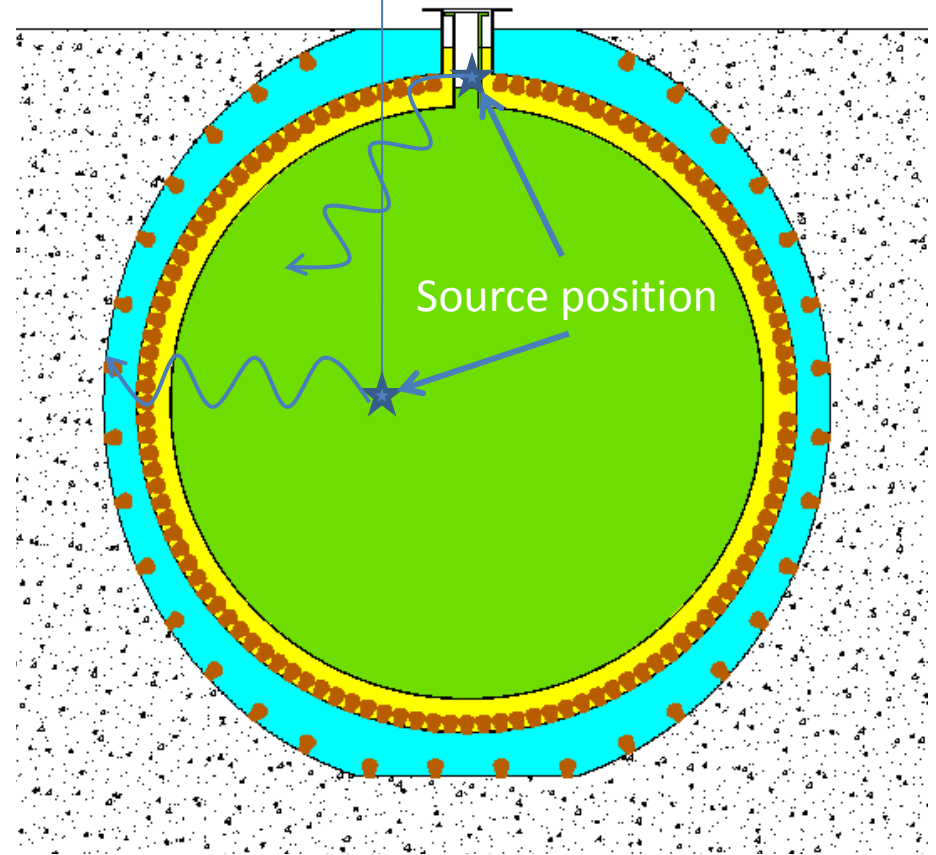
Energy spectra & fluxes of all types of neutrinos

Sterile neutrino

- Put radioactive source in detector center ($L = 0-17$ m, larger acceptance) or outside of detector ($L = 2-34$ m, smaller acceptance)
- Vertex reconstruction to determine baseline L .
- Measure the oscillation vs. L/E .

Isotopes produced by reactor with
 $E_\nu > 1.8$ MeV and $T_{1/2} > 10$ h

M	$T_{1/2}$	E_0/MeV	D	$T_{1/2}$	E_0/MeV
^{90}Sr	28.78a	0.546	Y	64.1h	2.282
^{91}Sr	9.63h	2.690	Y	58.51d	1.544
^{93}Y	10.18h	2.874	Zr	1.53e6a	0.091
^{97}Zr	16.9h	2.658	Nb	72.1m	1.934
^{106}Ru	373.6d	0.039	Rh	29.8a	3.541
^{112}Pd	21.03h	0.288	Ag	3.13h	3.956
^{125}Sn	9.64d	2.364	Sb	2.758a	0.767
$^{131\text{m}}\text{Te}$	30h	0.182	Te	25m	2.233
^{132}Te	3.204d	0.493	I	2.295h	3.577
^{159}Sm	9.4h	0.722	Eu	15.19d	2.451
^{140}Ba	12.75d	1.047	La	1.678d	3.762
^{144}Ce	284.9d	0.319	Pr	17.28m	2.997

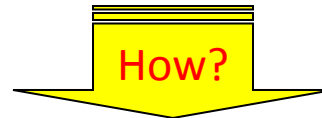


Technical challenges

- Requirements:
 - Large detector: 20 kt LS
 - Energy resolution: $3\%/\sqrt{E} \rightarrow 1200 \text{ p.e./MeV}$
- Ongoing R&D:
 - Low cost, high QE “PMT”
 - New type of PMT
 - Highly transparent LS: 15m \rightarrow >20m
 - Understand better the scintillation mechanism
 - Find out traces which absorb light, remove it from the production

Energy resolution

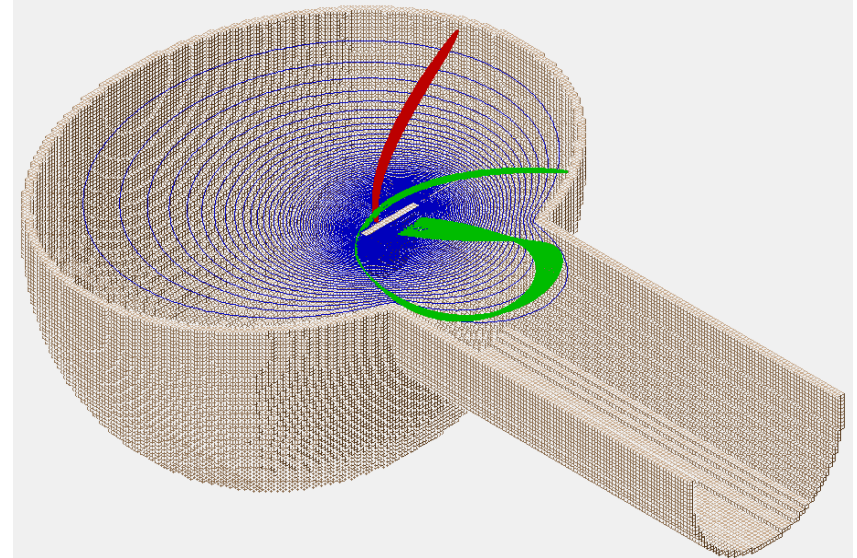
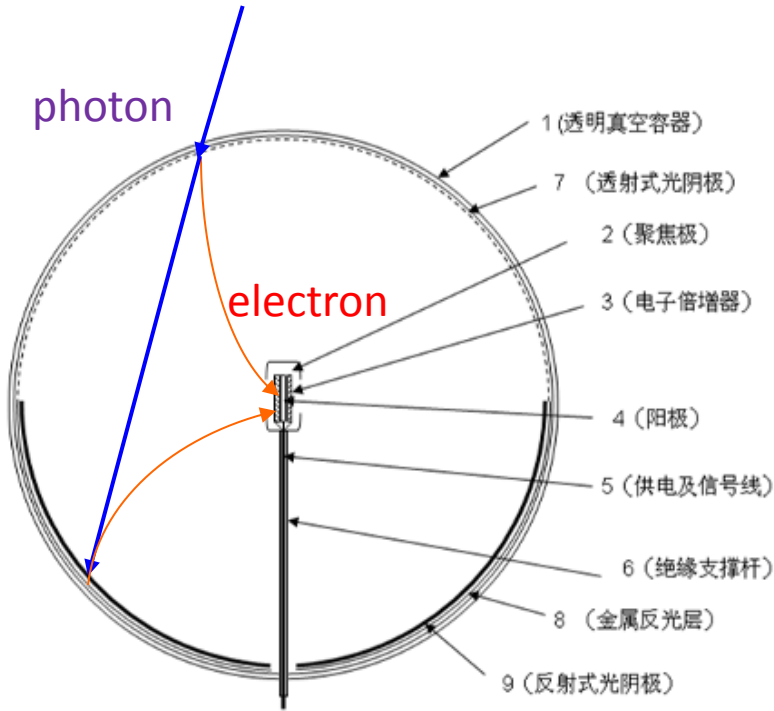
	KamLAND	Daya Bay II
Detector	~1 kt Liquid Scintillator	20 kt Liquid Scintillator
Energy Resolution	6%/√E	3%/√E
Light yield	250 p.e./MeV	1200 p.e./MeV



- Highly transparent LS
 - Attenuation length/D: 15m/16m → 30m/34m X 0.9
- High light yield LS:
 - KamLAND: 1.5g/l PPO → 5g/l PPO
 - Light Yield: 30% → 45% X 1.5
- Photocathode coverage:
 - KamLAND: 34% → ~80% X 2.3
- High QE “PMT”:
 - 20” SBA PMT QE: 25% → 35% X 1.4
 - or New PMT QE: 25% → 40% X 1.6
 - Both: 25% → 50% X 2.0

X 4.3-5.0 → (3.0-2.5)%/√E

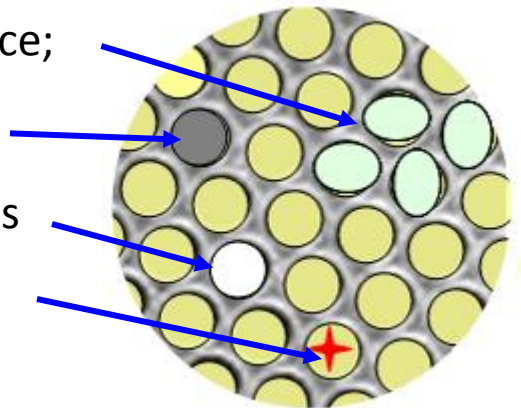
A new type of PMT: MCP-PMT



- Top: transmitted photocathode
- Bottom: reflective photocathode
additional QE: $\sim 80\% * 40\%$
- MCP (Microchannel Plate) to replace Dynodes → no blocking of photons

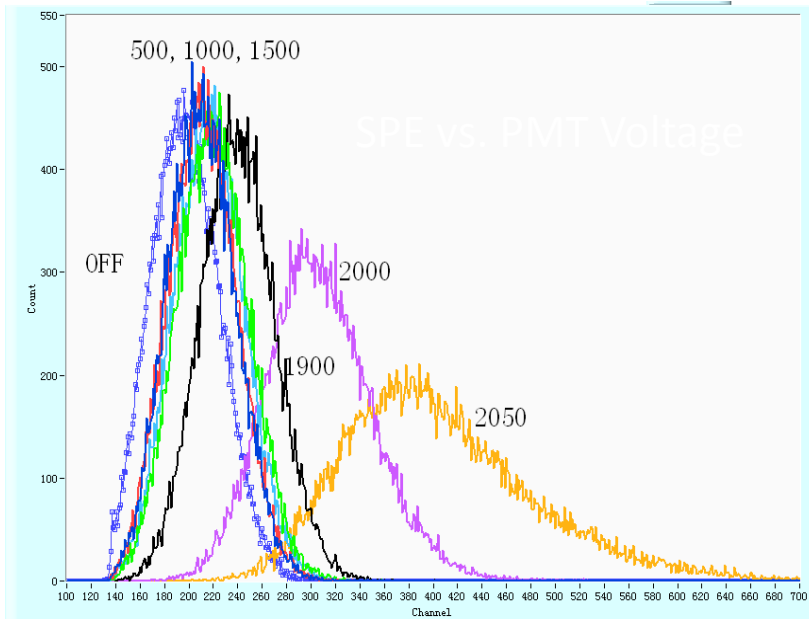
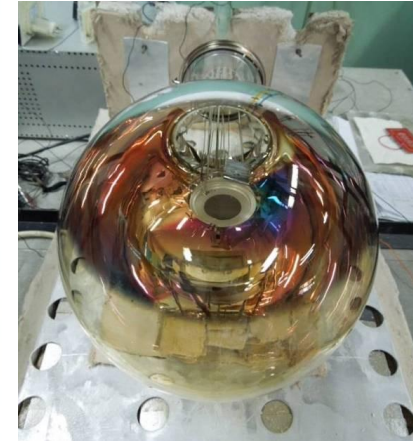
Low cost MCP by accepting the followings for SPE detection.

1. Asymmetric surface;
2. Blind channels;
3. Non-uniform gains
4. Flashing channels

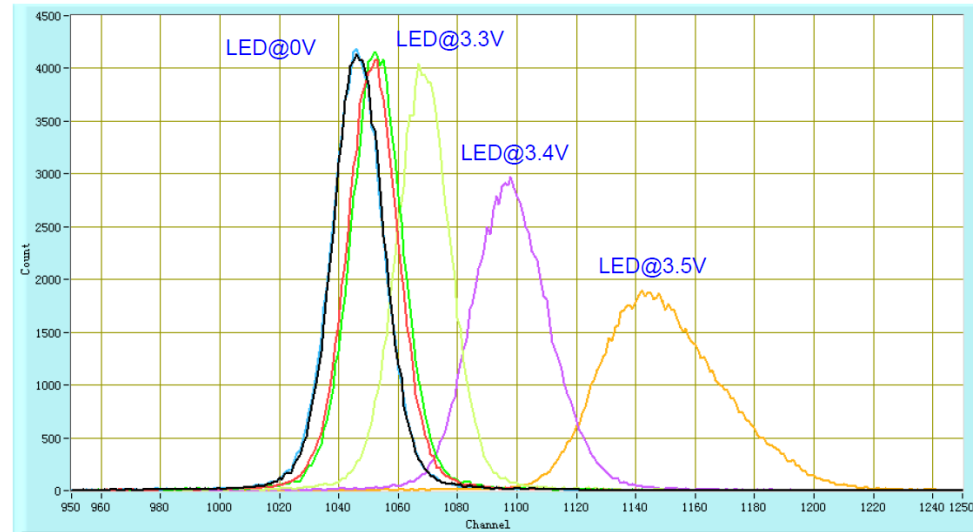


MCP surface

Prototypes



MPE vs the luminance of the LED light

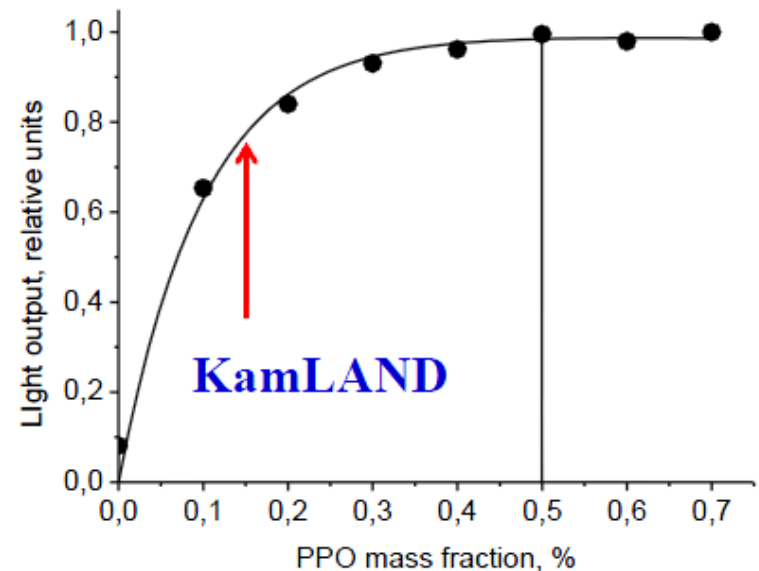
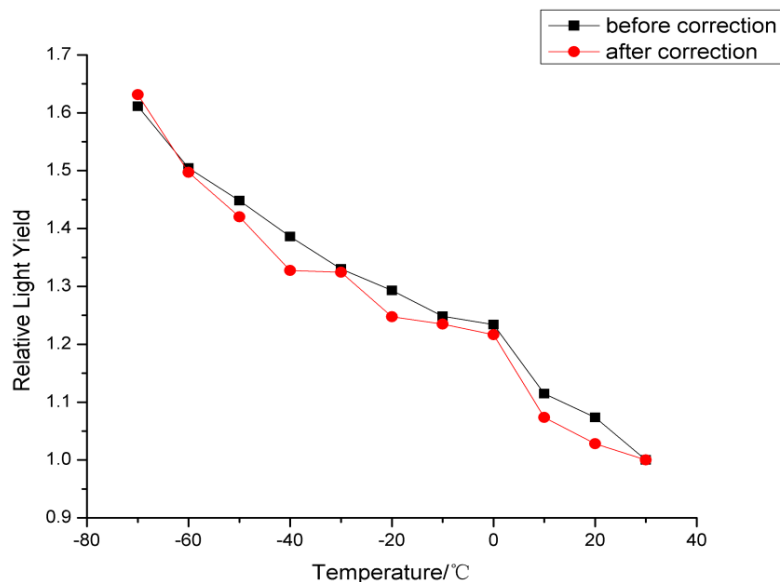


LAB based liquid scintillator

- To enhance the attenuation length
 - Improve raw materials (using Dodecane instead of MO)
 - Improve the production process for large volume
 - Purification
- High light yield
 - Lower temperature
 - Fluor concentration optimization

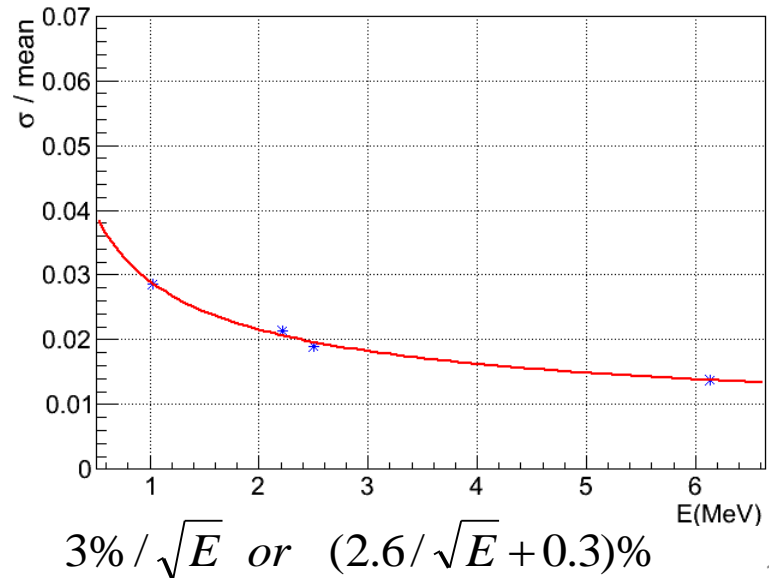
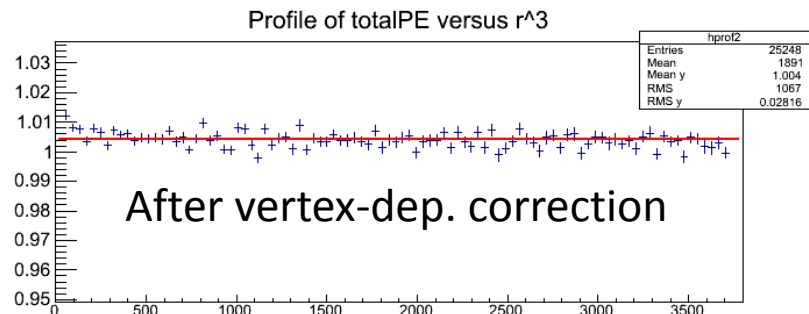
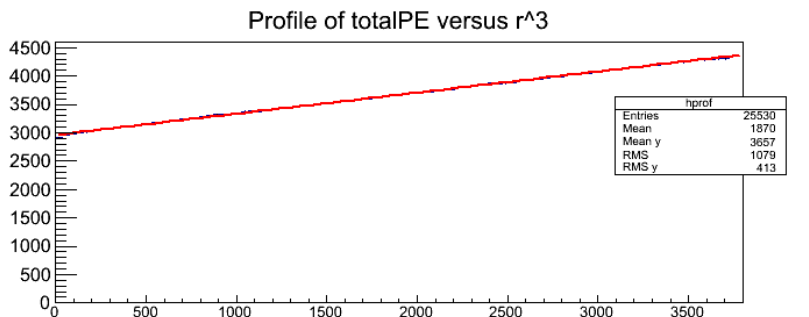
Test on purification

LAB	Atte. Length @ 430 nm
RAW	14.2 m
Vacuum distillation	19.5 m
SiO ₂ coloum	18.6 m
Al ₂ O ₃ coloum	22.3 m



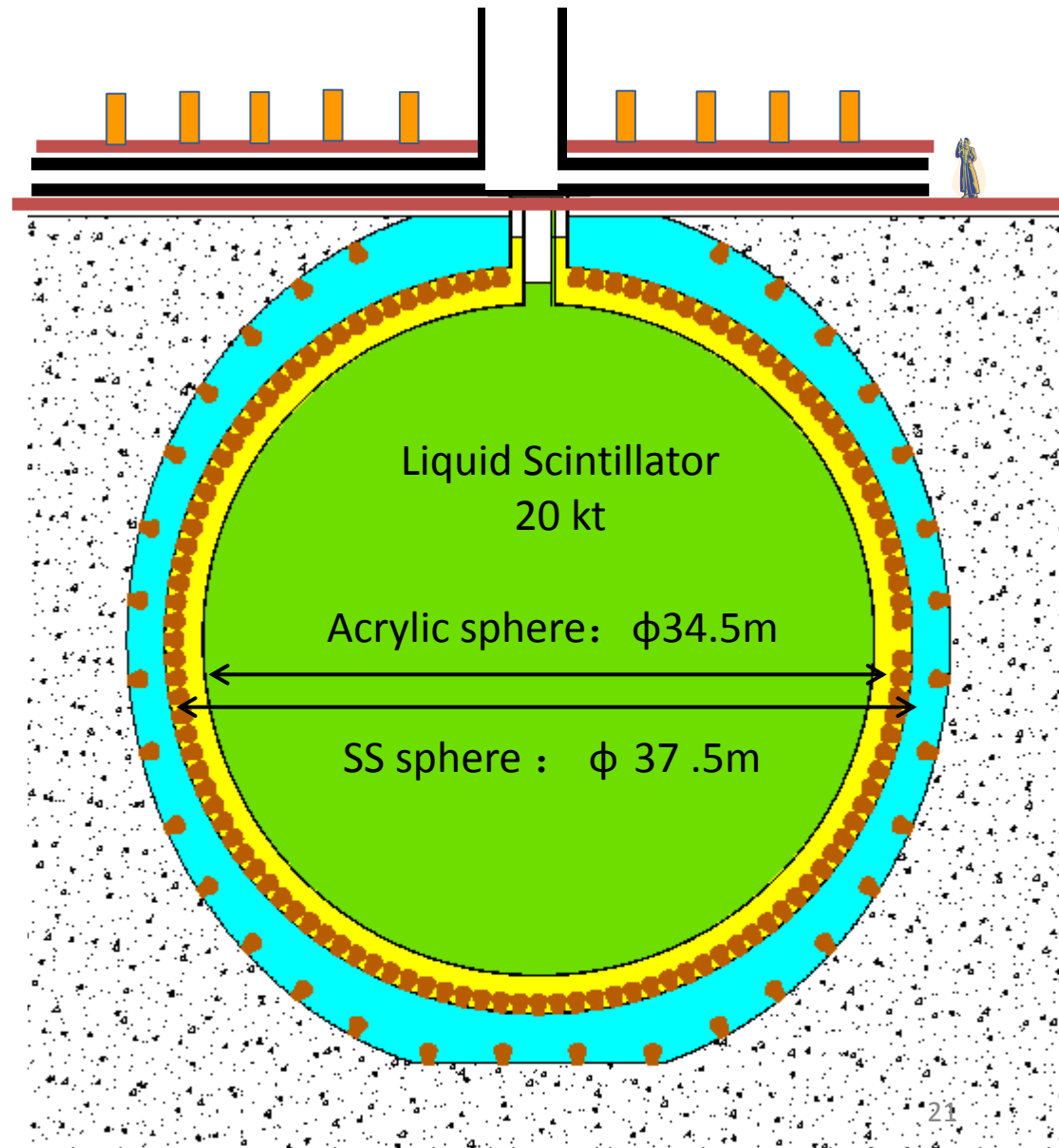
MC example

- DYBII MC, based on DYB MC (tuned to data), except
 - DYBII Geometry and 80% photocathode coverage
 - SAB PMT: maxQE from 25% → 35%
 - Lower detector temperature to 4 degree (+13% light)
 - LS attenuation length (1m-tube measurement@430nm)
 - From 15m = absorption 24 m + Raylay scattering 40 m
 - To 20 m = absorption 40 m + raylay scattering 40 m



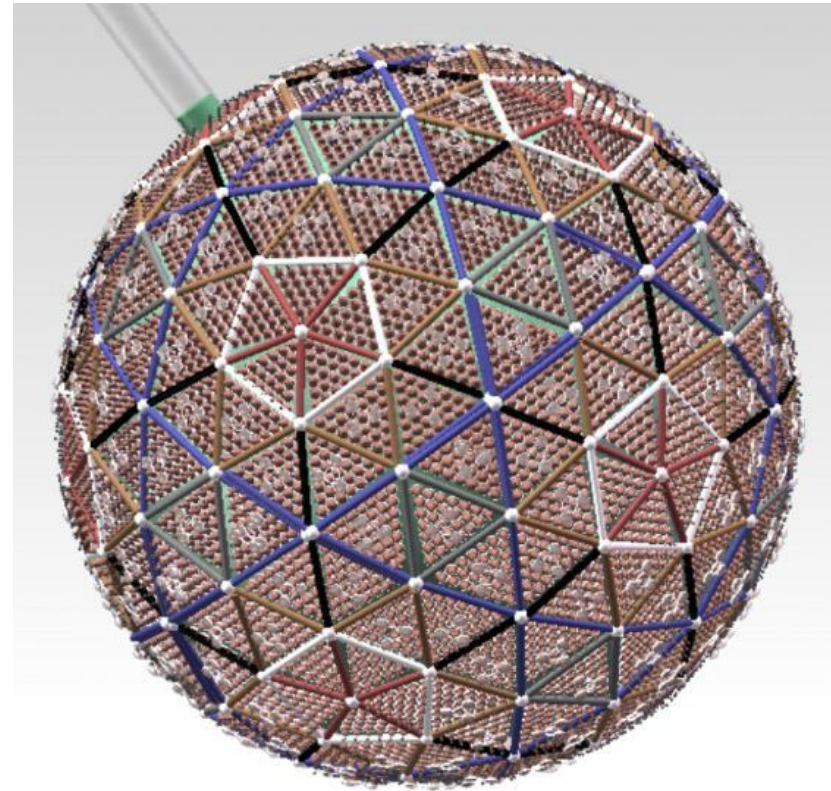
Detector Concept

- Extremely difficult to build both the stainless steel tank and the acrylic tank
- Options:
 - No steel tank, only acrylic tank
 - Steel tank +
 - Acrylic box/wall
 - Balloon
 - nothing



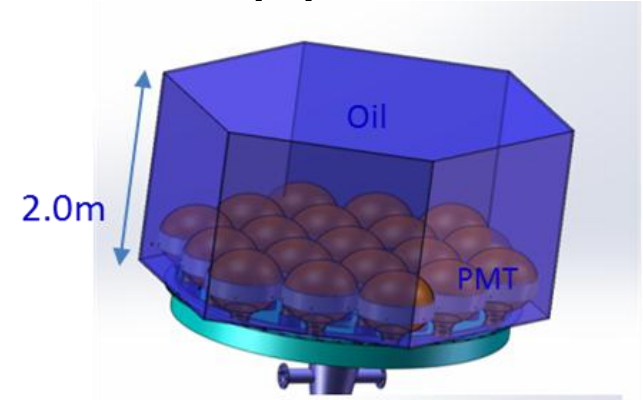
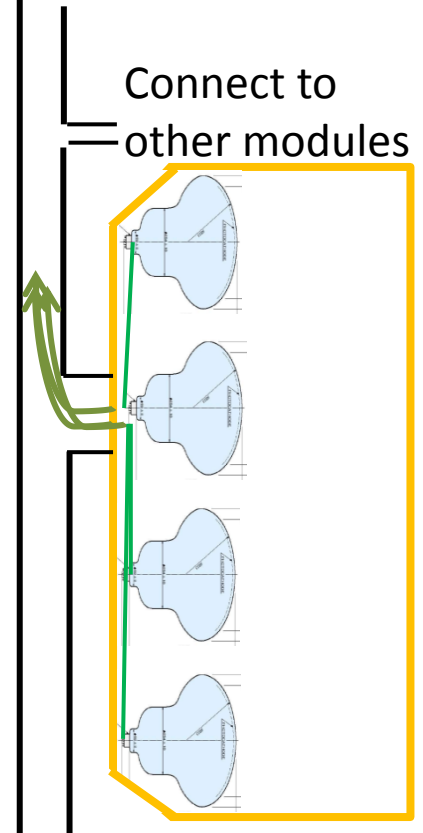
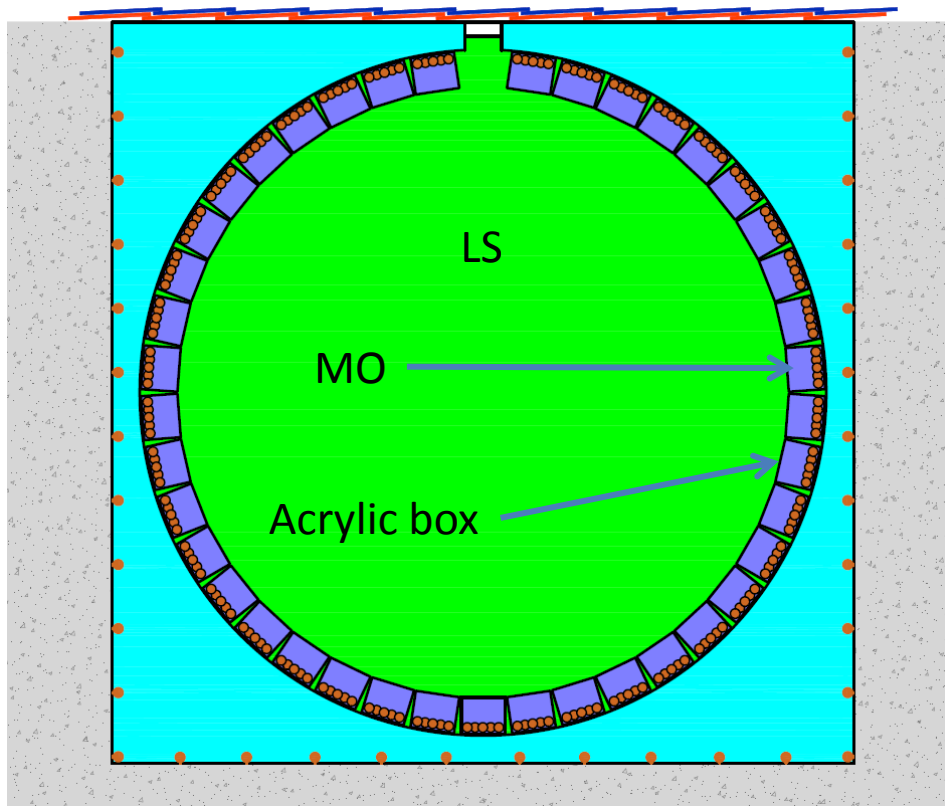
Option 1: no steel tank

- No more interference
- “Easy” for PMT holding
- Water replaces oil buffer
→ cheap
- Difficulties:
 - Larger pressure difference for the acrylic tank.



Option 2: acrylic box

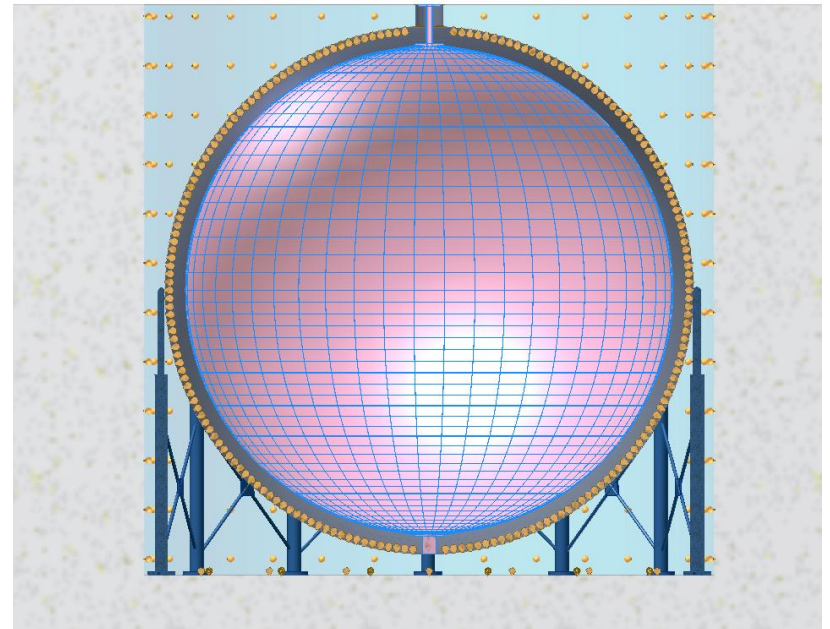
- Mineral oil in the optical modules
- Pipe for filling MO and cabling
- Concerns: leakage through cables



Option 3: balloon

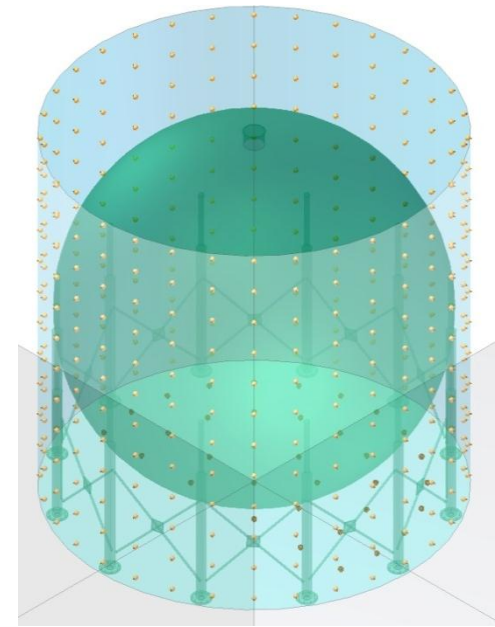
- “Cheap” for construction & quick for installation
- Experience from Borexino (0.5kt) & KamLAND (1kt)
- Need to consider film materials (mechanics, transparency, compatibility, welding technique, radon permeability, ...) , cleanness, leak check, deployment, backup plan if fails, ...

Not new to IHEP



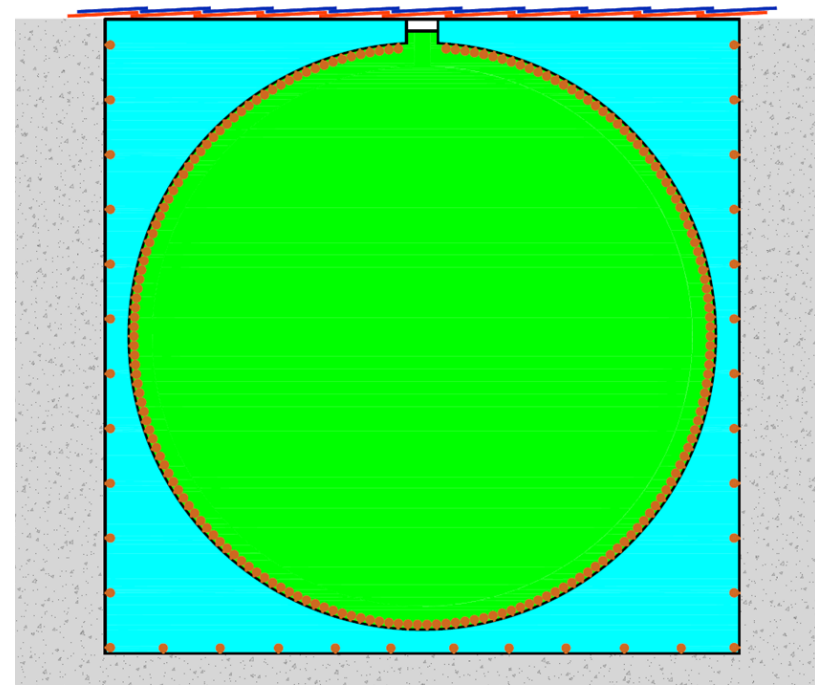
Option 4: Steel tank only

- No problem for construction
- A fall back plan of the balloon option
- But
 - PMT protection
 - Trigger rate by backgrounds
 - Resolution affected by backgrounds



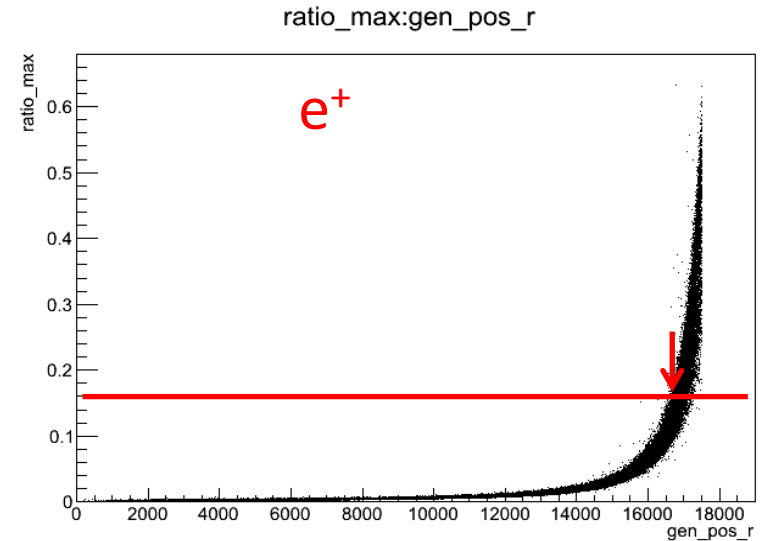
If the PMT glass is the same as Daya Bay, radioactivity will be 44 Bq/PMT, or 3.3 MHz in total

If better glass is used, it may be reduced to 1 MHz



Online background suppression

- Divide PMTs to 1476 regions
- Look at the charge ratio q_i/q_{total} (i: the region ID)
 - Cut charge ratio < 0.16
 - Cut also $N_{\text{p.e.}} < 500$ (~ 0.4 MeV)
- Event rates is reduced to 0.6kHz



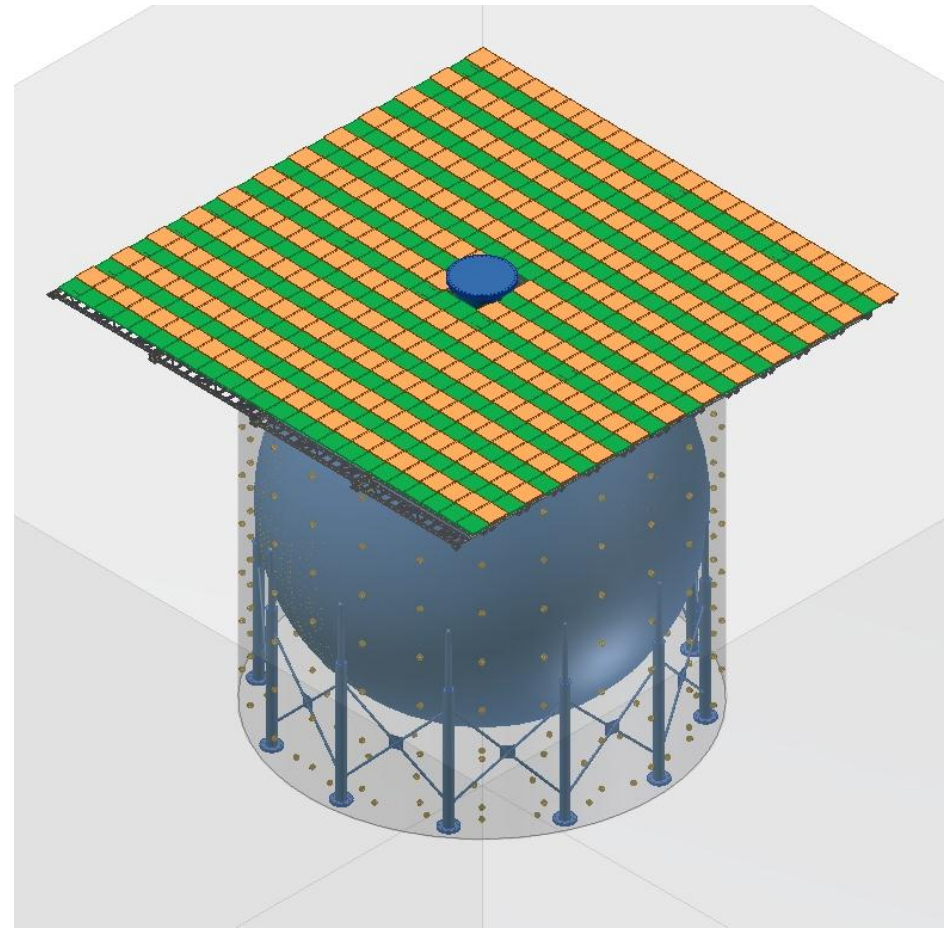
Equivalent to fiducial volume cut.

Resolution is affected:

Energy(MeV)	No Background (vertex corrected)		Mix Background(1MHz, 500ns) (vertex corrected)	
	sigma	mean	sigma	mean
2*0.511	0.030	1	0.035	0.94
2.22	0.024	1	0.027	0.97
1.173+1.333	0.021	1	0.024	0.97
6.13	0.016	1	0.017	0.99

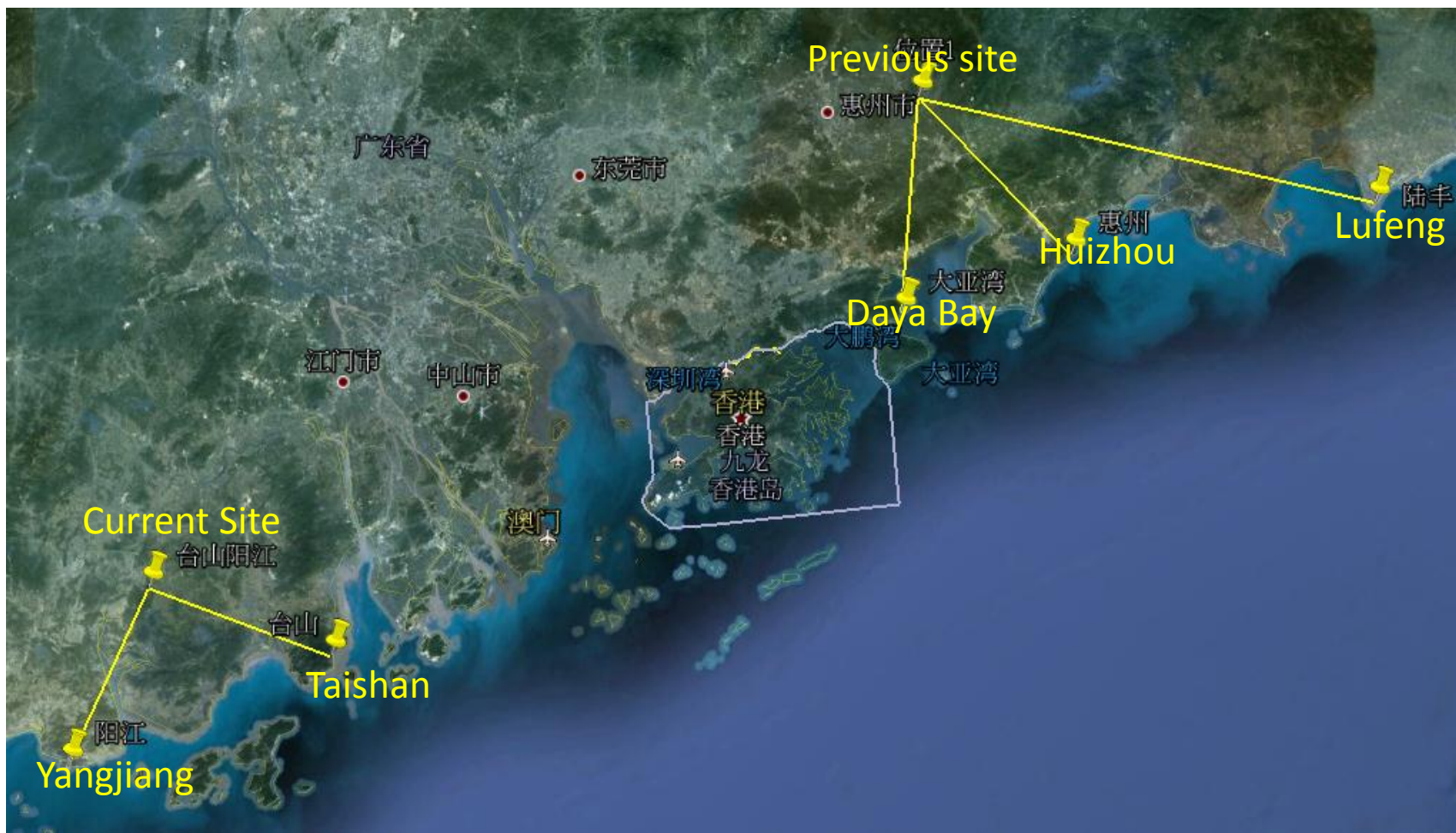
VETO

- Water
 - A MC simulation show that $\sim 2\text{m}$ water, 1500 20" PMT is good enough
- Top VETO Options:
 - RPC
 - Plastic scintillator
 - Liquid scintillator
 - Two layers?
 - precise muon tracking



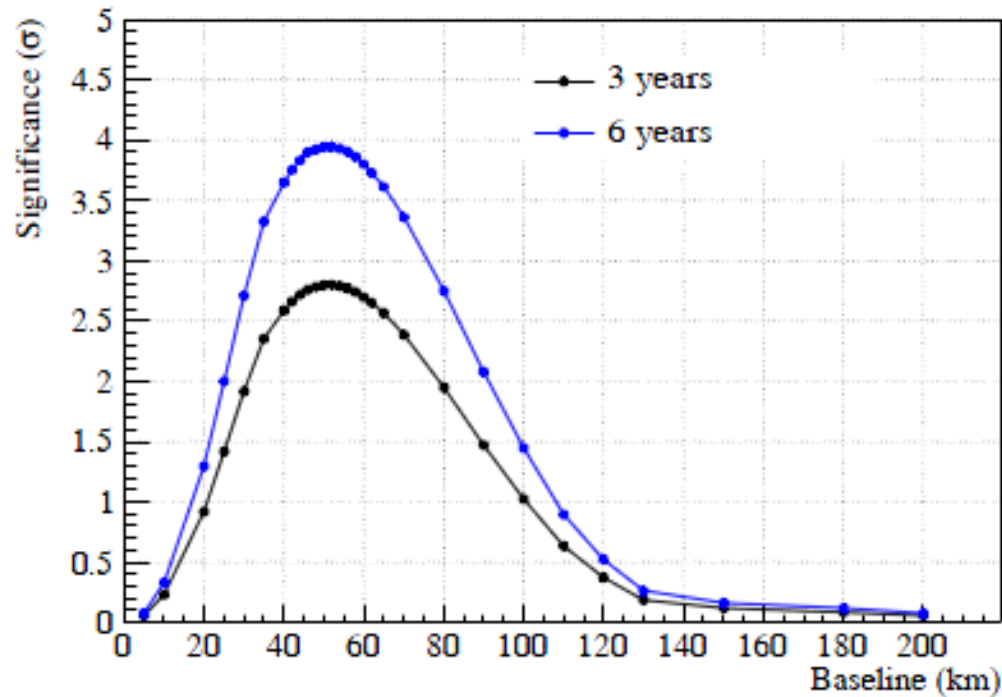
The site: Kaiping county, Jiangmen City

	Daya Bay	Huizhou	Lufeng	Yangjiang	Taishan
Status	Operational	Planned	Planned	Under construction	Under construction
Power	17.4 GW	17.4 GW	17.4 GW	17.4 GW	18.4 GW



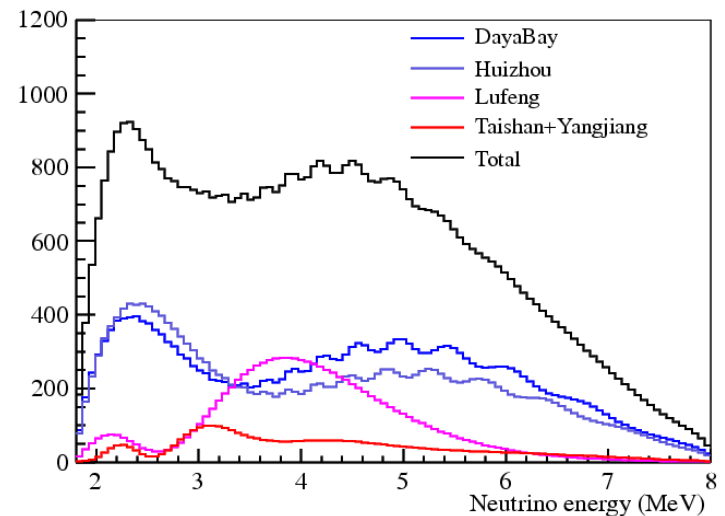
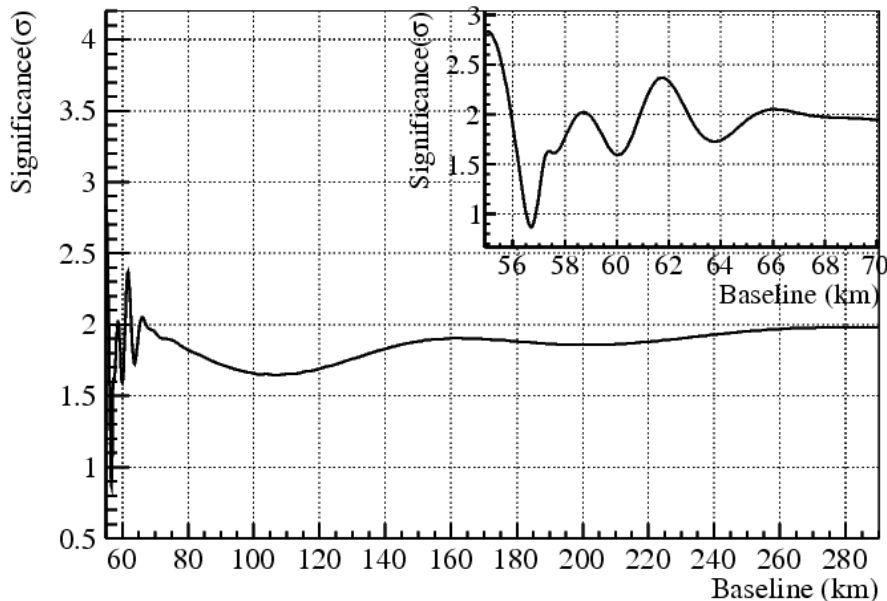
Optimal baseline

- Proper baseline: ~ 50 km, around θ_{12} oscillation maximum.



Complex interference between reactors

- Adding one reactor (more statistics) is not always helpful.
- Example:
 - One reactor (6X2.9 GW) at 55 km, the significance is 2σ .
 - Add another reactor
 - Statistics doubles with equal baseline.
 - Helpful, if the baseline difference < 1 km.
 - Harmful as background, if the baseline difference > 1 km.
 - The worst baseline difference is 2 km due to θ_{13} oscillation cancellation.



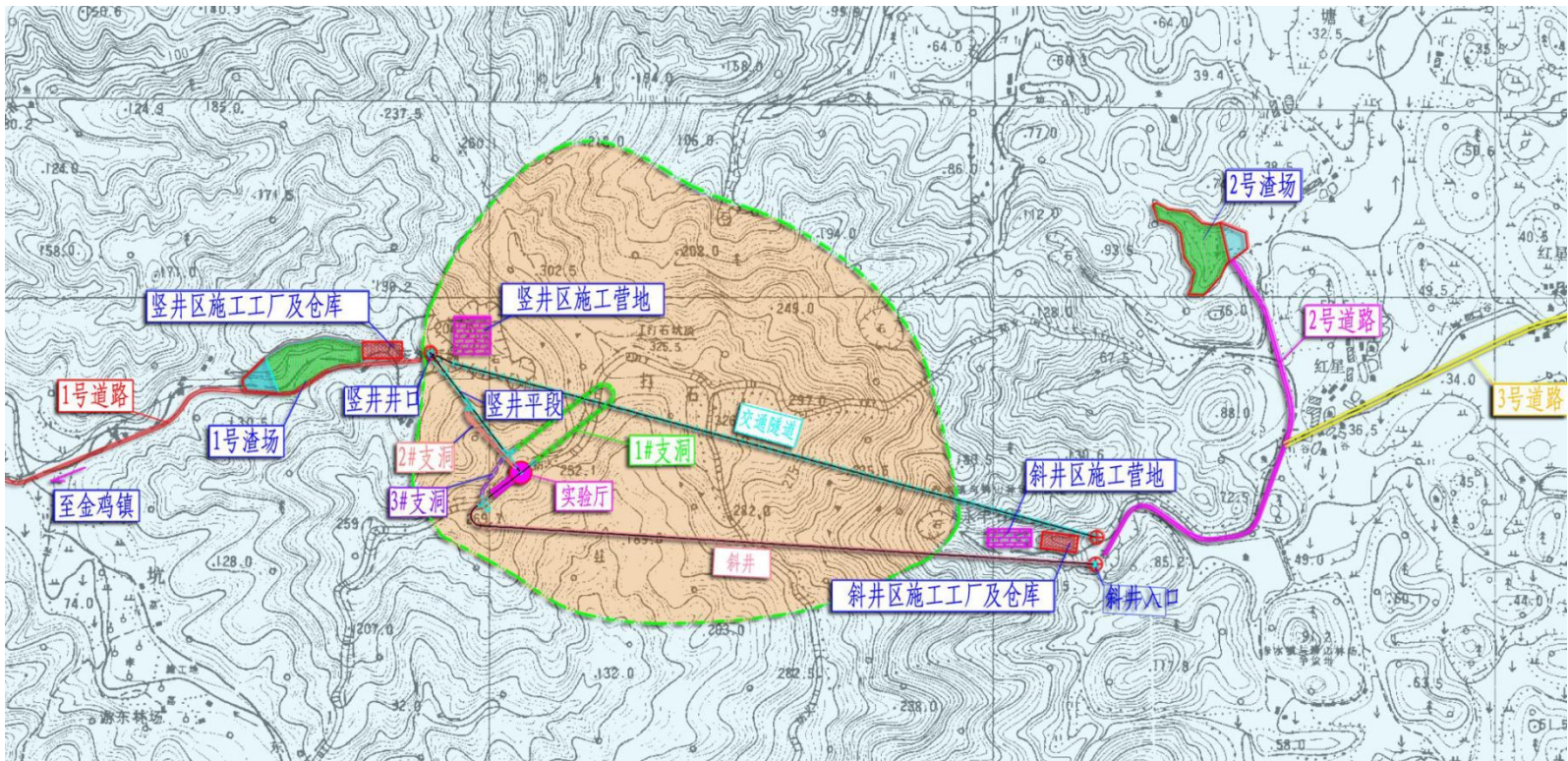
Example of peak and valley cancellation

Kaiping county, Jiangmen City
in Guangdong province

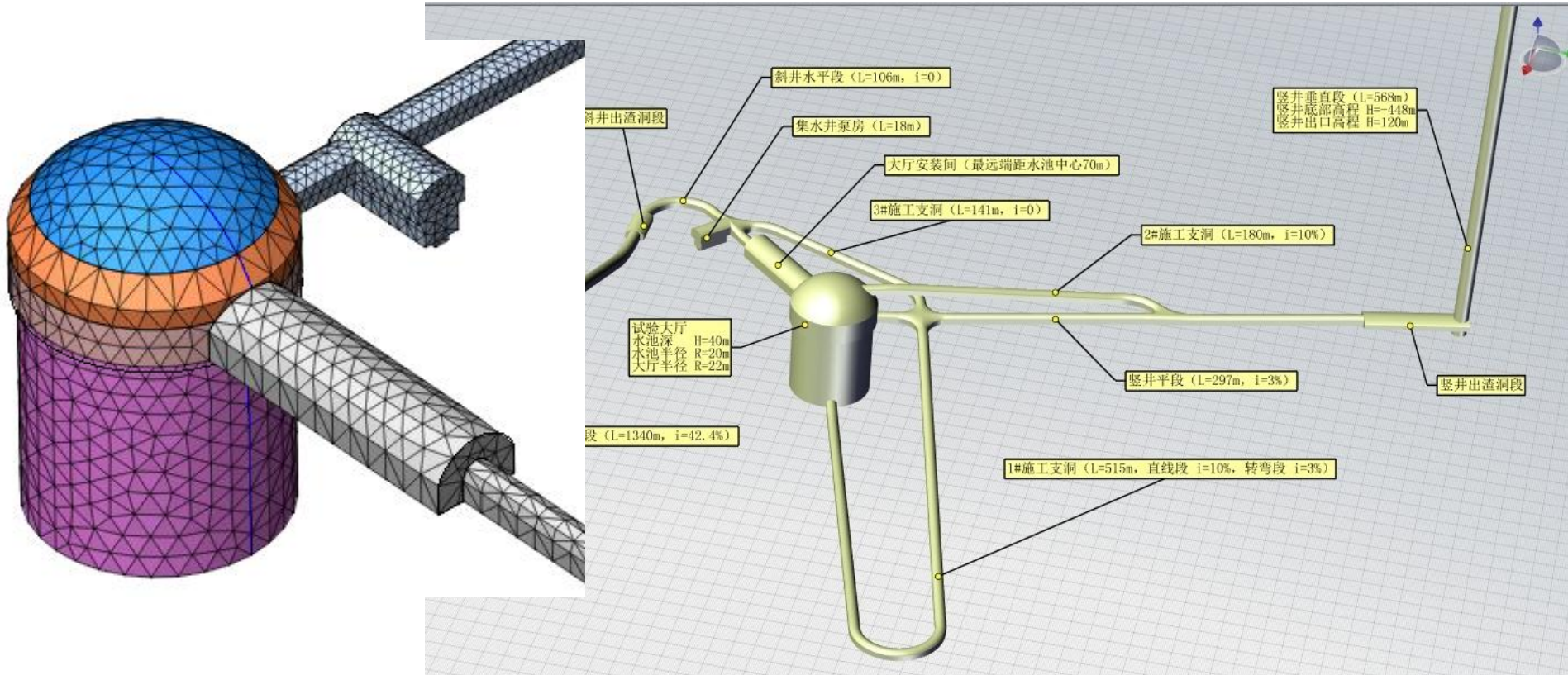


Construction plan

- Two options considered
 - Rails(40%, 1100m) + vertical shaft (600m)
 - Rails(40%, 1100m) + horizontal tunnel (6600m)
- Conceptual design completed. Review held on Dec.17, 2012.
 - Rails + vertical shaft is chosen for cost and schedule reasons
 - No show-stoppers



Experimental hall



- Preliminary study shows that:
 - Stability of the hall is not a problem
 - Total time needed for construction is 3 years

Brief schedule

- Civil preparation: 2013-2014
- Civil construction: 2014-2017
- Detector R&D: 2013-2016
- Detector component production: 2016-2017
- PMT production: 2016-2019
- Detector assembly & installation: 2018-2019
- Filling & data taking: 2020

After a number of reviews, we are approved by the CAS(~ CD1)

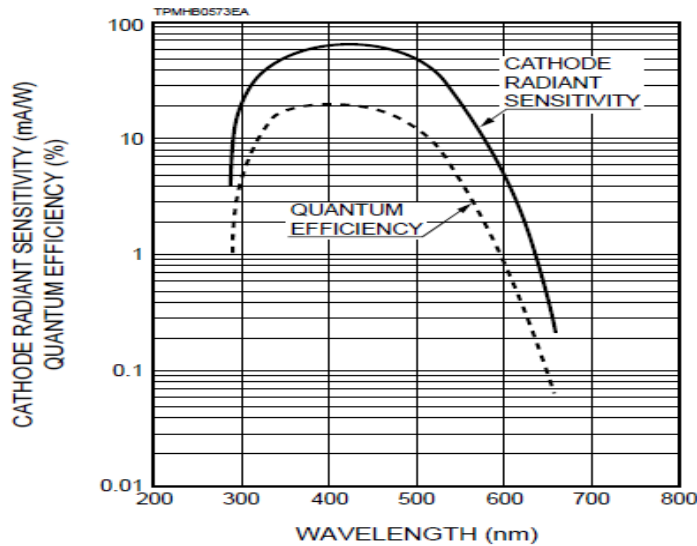
Summary

- Daya Bay II proposed in 2008-2009, now boosted by the large θ_{13}
- Rich physics potential
- Although challenging, preliminary study shows it is not impossible.
- A few R&D efforts has been started, more will come.
- Detector design and civil design have been started.
- Good support from the local government and the Chinese funding agencies.

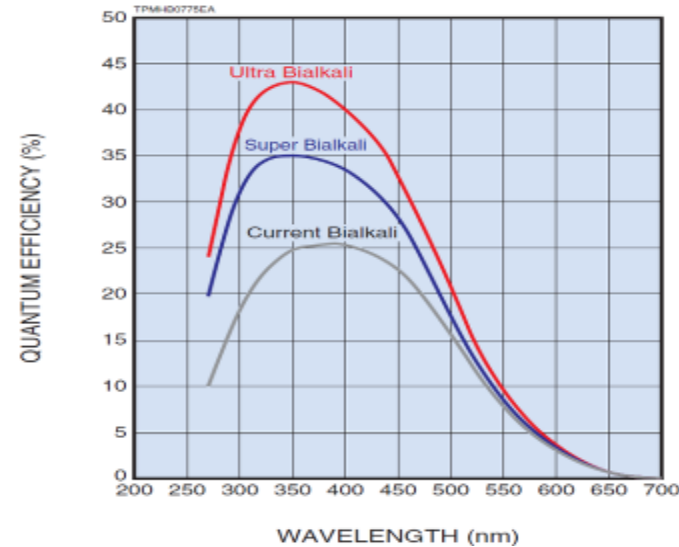
backup

The Quantum Efficiency of PMT

The QE of 20" PMT-R3600



The QE of SBA/UBA



High QE PMTs: SBA (35%) and UBA (43%)

are only available in small format (< 5" diameter ?)

➤ Can we improve the Quantum Efficiency of Photocathode or Photon Detection Efficiency for the large area 20" PMT ?

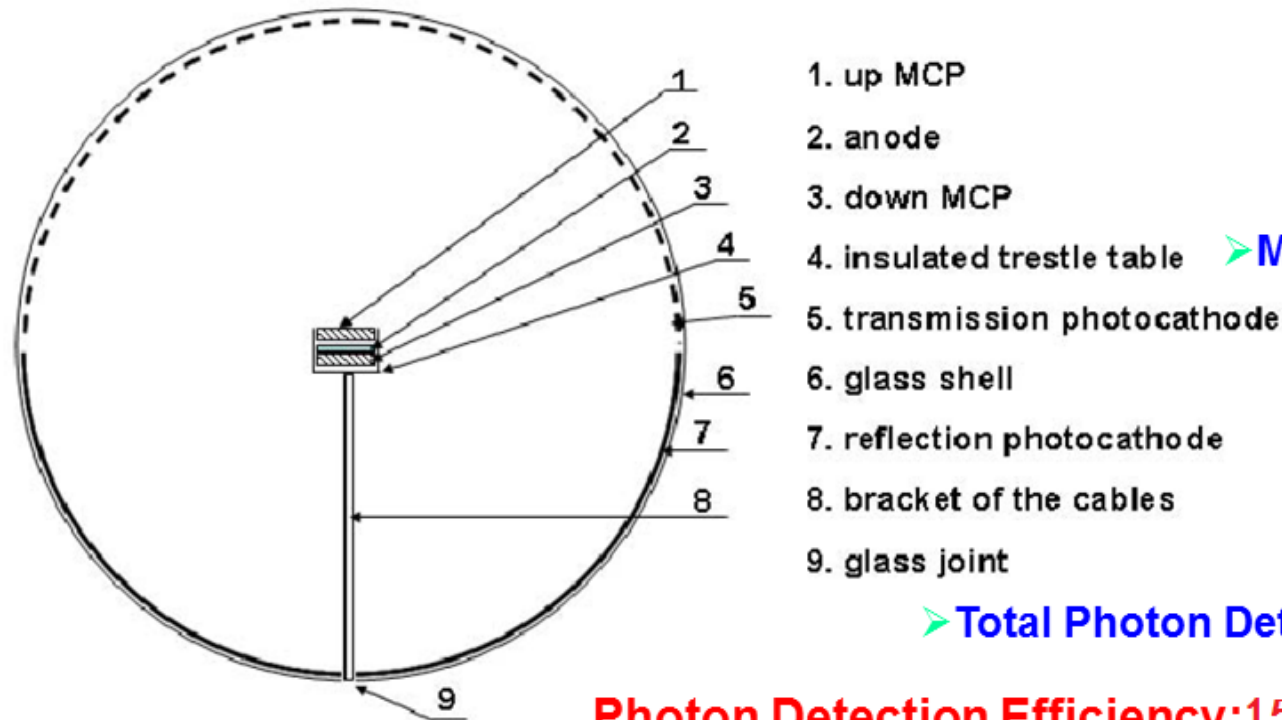
QE: 20% → 30%

The new design of a large area PMT

High photon detection efficiency + Single photoelectron Detection + Low cost

1) Using transmission photocathode (front hemisphere) and reflective photocathode (back hemisphere) $\Rightarrow \sim 4\pi$ viewing angle!!

2) Using two sets of Microchannel plates (MCPs) to replace the dynode chain



1. up MCP
2. anode
3. down MCP
4. insulated trestle table
5. transmission photocathode
6. glass shell
7. reflection photocathode
8. bracket of the cables
9. glass joint

➤ Quantum Efficiency:

Transmission photocathode: 30%

Reflection photocathode: 35%

➤ MCP Collection Efficiency: 70%

Photon detection efficiency:

$$30\% * 70\% = 21\%$$

$$40\% * 35\% * 70\% = 9.8\%$$

➤ Total Photon Detection Efficiency: $\sim 30.8\%$

Photon Detection Efficiency: 15.4% \rightarrow 30.8% $\times \sim 2$ at least